

Particle Physics in China

Yuanning Gao
Tsinghua University
June 30, 2016

BEPC & BES

The Starting Point of Particle Physics in China

对流层风场特征及应用：

这种风场产生于 1990-9
-1991 年间 10-12 级强寒潮。
西风带被切断。首先切断了西
风带的风向，而且已经切断
了北半球中高纬度的风。

现在讨论这种风场特征
今年 9 月以后：这种风场才
开始形成并持续到今春。

西风带被切断后，
有关方面一再强调这种风

场，是造成这次寒潮的主要
原因之一。

这种风场可能与过去
许多次寒潮和强寒潮有
一定的相似之处，但也有很
多不同之处。这种风场的特
点是：在冬季风带切断后，
在冬季风带切断后，即形成
一个巨大的低气压系统，这
时风场的特征与过去完全不同。

周卫星
- 1991 年 10 月 1 日



方毅同志听取李政道教授对
中国第一台高能加速器建设方案的意见



1984年10月7日，北京正负电子对撞机工程破土动工，邓小平等党和国家领导人亲自为工程奠基

A bit of history

1984

Construction start

1989

Physics with BESI

1993

CAS approved the 1st upgrade plan: higher luminosity for BEPC, BESI to BESII.

1998

Physics with BESII

2003

State council approved 2nd upgrade plan:
BEPC to BEPCII, BESII to BESIII.

2004

BEPC shutdown & dismantling for BEPCII.

2008

Physics with BESIII

BEPC and BESI

- Beijing Electron-Positron Collider (BEPC) and Beijing Spectrometer (BES) is the 1st large facility for high energy physics ever built in China.



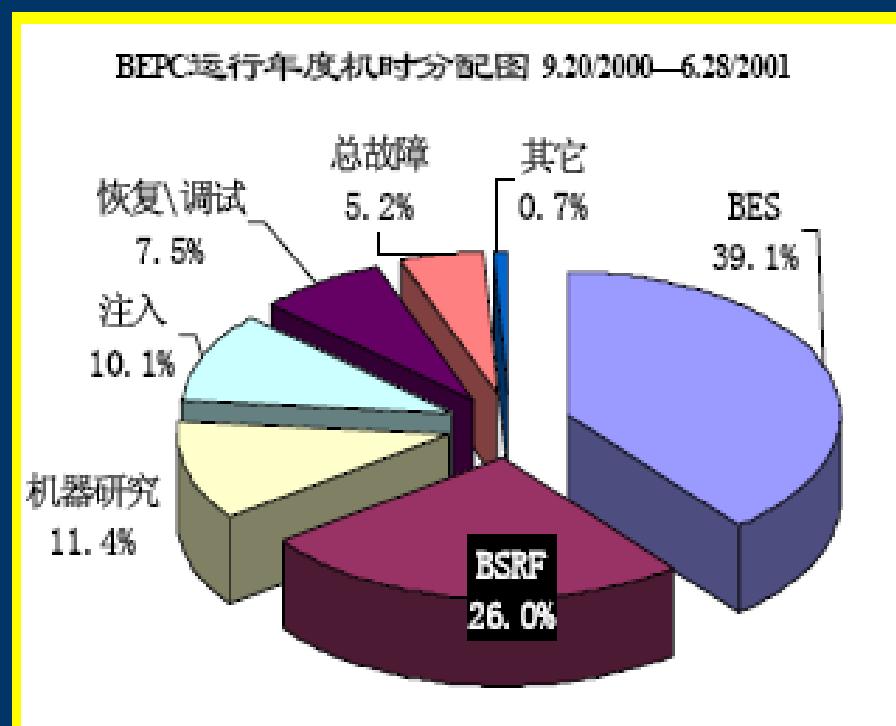
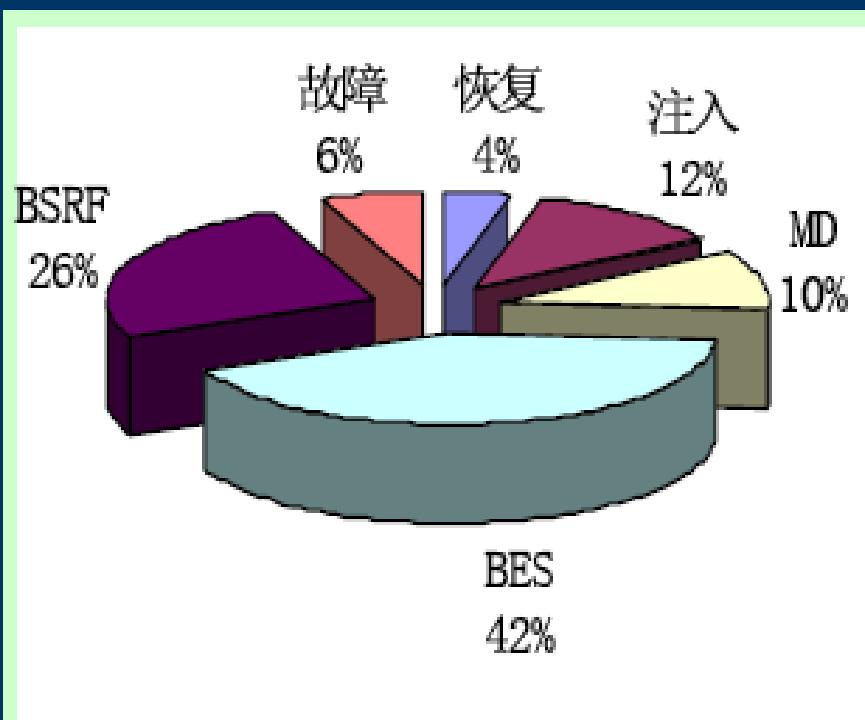
E_{cm} : 2-5 GeV

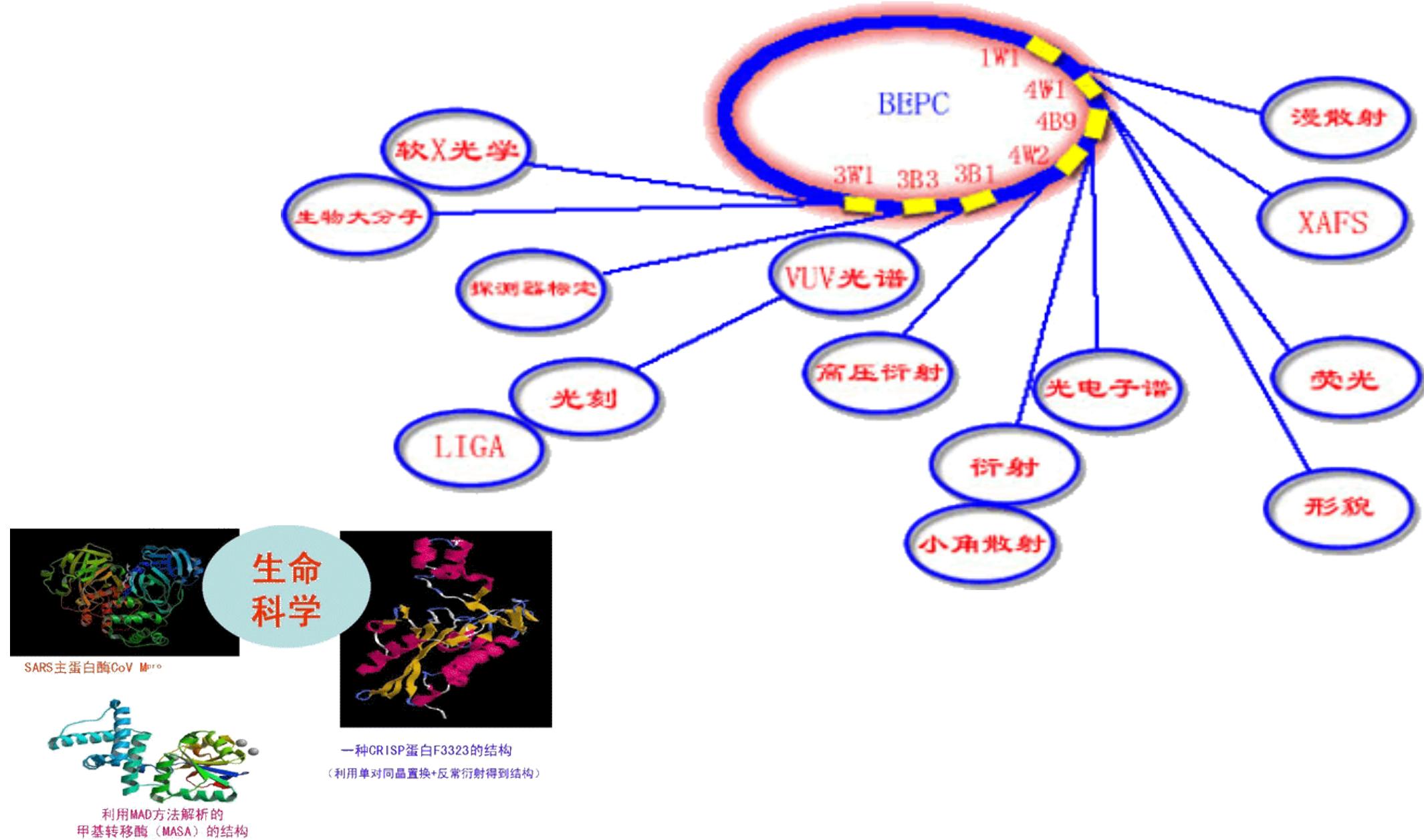
Luminosity: $\sim 10^{30}-10^{31} \text{ cm}^{-2}\text{s}^{-1}$



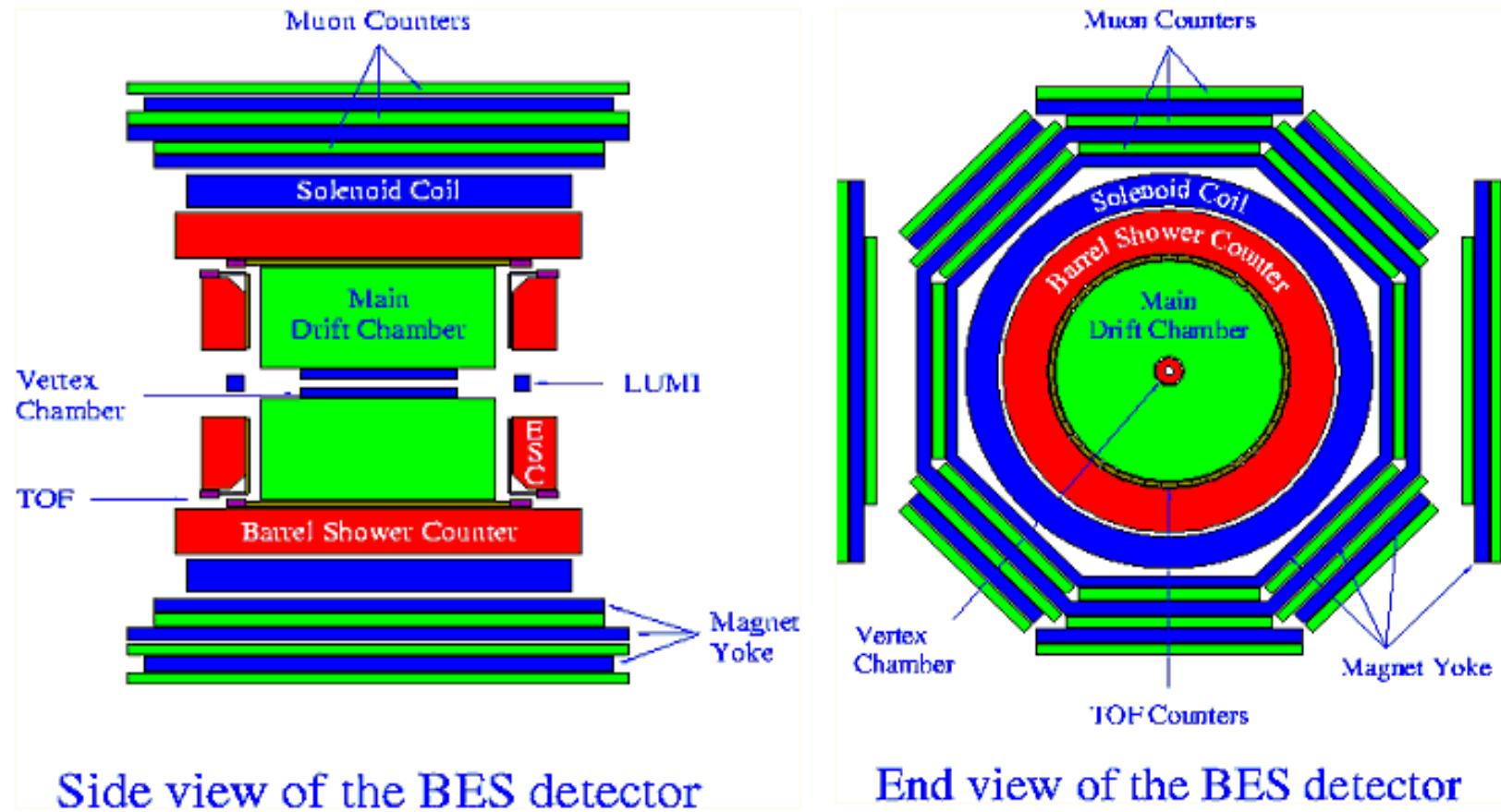
年度机时分配

年度		运行	恢复	注入	调试	MD	BES	BSRF	其它	故障
00— 01	计划	6648	266	798	-	694	2791	1700	0	399
	实际	6581	490.4	664.7	-	750.2	2574	1708	47.8	345.2





BESII Detector (1995-1997 upgraded)



$$\text{VC: } \sigma_{xy} = 100 \text{ } \mu\text{m}$$

$$\text{MDC: } \sigma_{xy} = 250 \text{ } \mu\text{m}$$

$$\sigma_{dE/dx} = 8.4 \text{ \%}$$

$$\Delta p/p = 1.8\sqrt{(1+p^2)}$$

$$\text{TOF: } \sigma_T = 180 \text{ ps}$$

$$\text{BSC: } \Delta E/\sqrt{E} = 22 \text{ \%}$$

$$\sigma_\phi = 7.9 \text{ mr}$$

$$\sigma_z = 2.3 \text{ cm}$$

$$\mu \text{ counter: } \sigma_{r\phi} = 3 \text{ cm}$$

$$\sigma_z = 5.5 \text{ cm}$$

$$\text{B field: } 0.4 \text{ T}$$

$$\text{Dead time/event: } \langle 10 \text{ ms}$$

Institutions (ever) in BES Collaboration

Institute of High Energy Physics
California Institute of Technology

CCAST

Chonhuk University

Colorado State University

Gyeongsang National University

Henan Normal University

Huazhong Normal University

Hunan University

KEK

Korea University

Liaoning University

Miyazaki University

Nankai University

Nihon University

Queen Mary University of London

Peking University
Shandong University
Shanghai Jiaotong University

Sichuan University
SLAC

Seoul National University
University of Hawaii

University of California at Irvine
University of Science and Technology of China
University of Texas at Dallas

Tokyo Institute of Technology
University of Tokyo
Tsinghua University
Wuhan University
Zhejiang University

Highlights of BESI/BESII Physics Results

- Precise measurement of the mass of tau lepton
- Precise measurement of R value in 2-5 GeV
- $\psi(2S)$ decays: Study many decay modes, the 12% rule
- J/ψ decays: Light hadron spectroscopy, search for
 - multi-quark candidates, glueballs
- $\psi(3770)$ decays and D physics

Selected results from BESI/BESII

- τ mass measurement (1991)

PDG2007

VALUE (MeV)
 1776.90 ± 0.20 OUR AVERAGE

$1776.81^{+0.25}_{-0.23} \pm 0.15$

81

τ MASS

DOCUMENT ID

TECN

COMMENT

$1775.1 \pm 1.6 \pm 1.0$

13.3k

¹ ABBIENDI

00A

O

1995 LEP runs
 $E_{cm}^{ee} = 10.6$ GeV

$1778.2 \pm 0.8 \pm 1.2$

ANASTASSOV 97

CLE

$E_{cm}^{ee} = 10.6$ GeV

$1776.96^{+0.18+0.25}_{-0.21-0.17}$

65

² BAI

96

BES

$E_{cm}^{ee} = 3.54-3.57$ GeV

$1776.3 \pm 2.4 \pm 1.4$

11k

³ ALBRECHT

92M

ARG

$E_{cm}^{ee} = 9.4-10.6$ GeV

1783^{+3}_{-4}

692

⁴ BACINO

78B

DLC

$E_{cm}^{ee} = 3.1-7.4$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1777.8 \pm 0.7 \pm 1.7$

35k

⁵ BALEST

93

CLEO

Repl. by ANAS-

TASSOV 97

$1776.9^{+0.4}_{-0.5} \pm 0.2$

14

⁶ BAI

92

BES

Repl. by BAI 96

¹ ABBIENDI 00A fit τ pseudomass spectrum in $\tau \rightarrow \pi^\pm \leq 2\pi^0\nu_\tau$ and $\tau \rightarrow \pi^\pm \pi^+ \pi^- \leq 1\pi^0\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$.

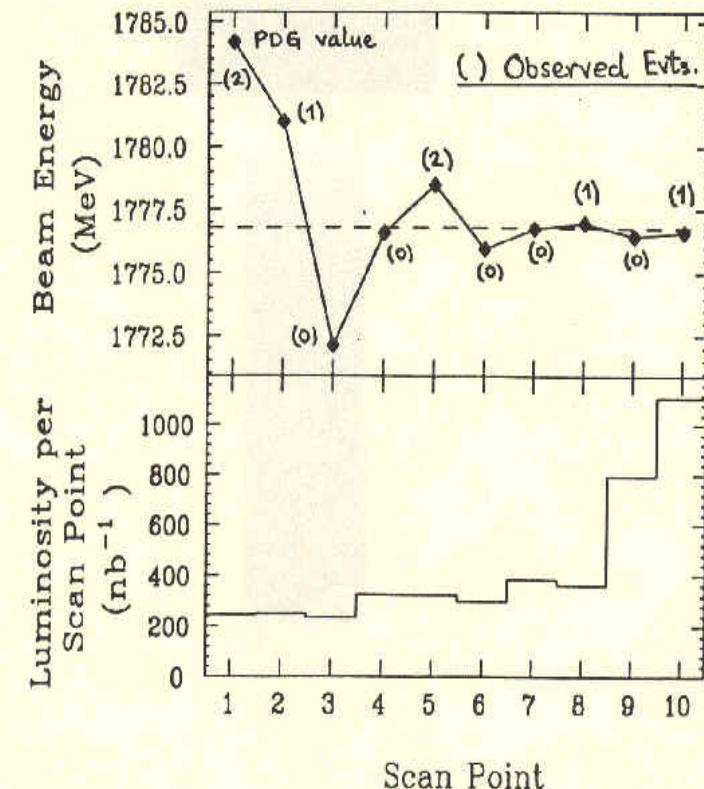
² BAI 96 fit $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$ at different energies near threshold.

³ ALBRECHT 92M fit τ pseudomass spectrum in $\tau^- \rightarrow 2\pi^- \pi^+ \nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$.

⁴ BACINO 78B value comes from $e^\pm X\bar{X}$ threshold. Published mass 1782 MeV increased by 1 MeV using the high precision $\psi(2S)$ mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.

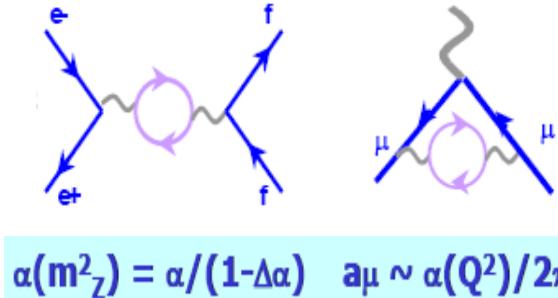
⁵ BALEST 93 fit to τ pseudomass spectrum in $\tau^- \rightarrow 2\pi^- \pi^+ \nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$.

History of the Energy Scan



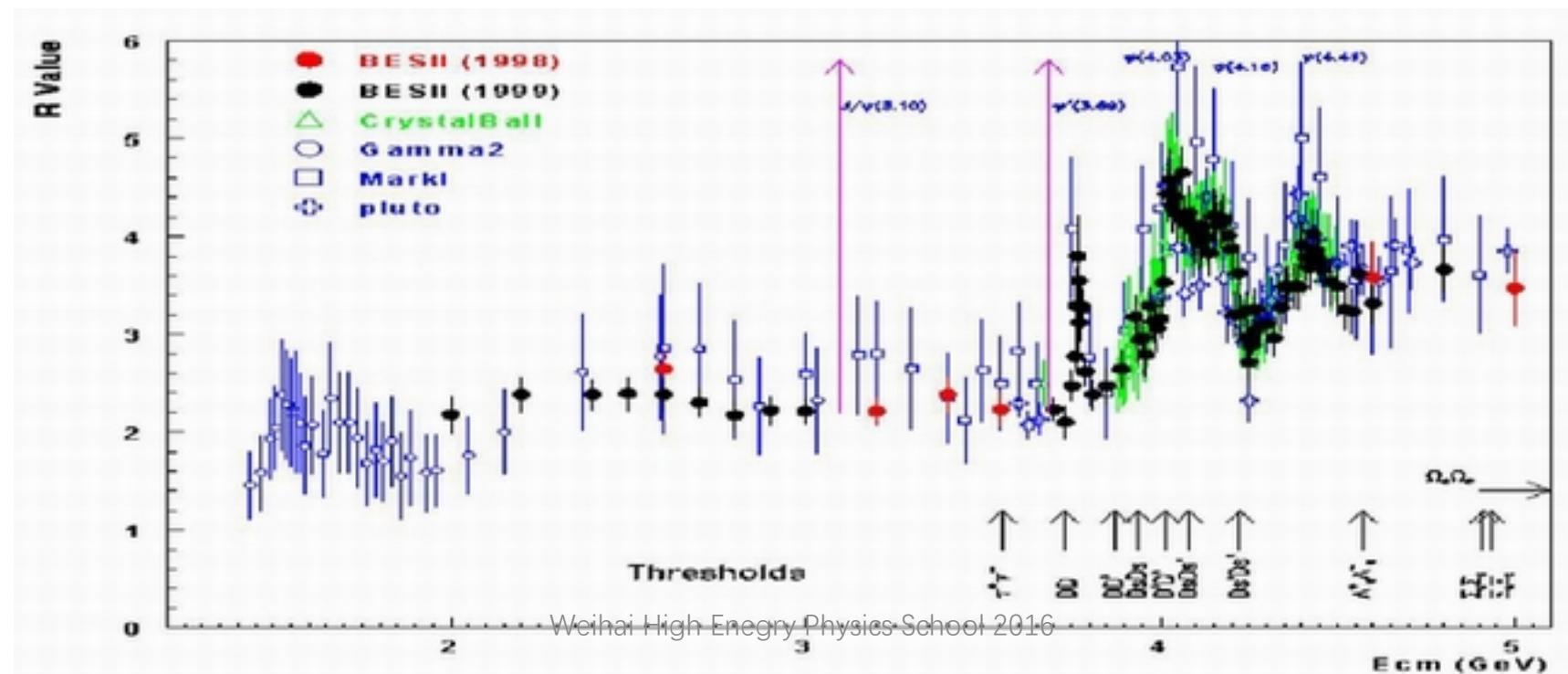
Selected results from BESI/BESII

- R measurement (1998-1999)



Optical Theorem

$$\Delta\alpha_{hadron} = -\frac{\alpha}{3\pi} \int_{4m_\pi^2}^{\infty} \frac{m_Z^2 ds'}{s' [s' - m_Z^2]} \frac{\sigma(e^+e^- \rightarrow \gamma^* \rightarrow q\bar{q})}{\sigma(e^+e^- \rightarrow \gamma^* \rightarrow \mu^+\mu^-)}$$



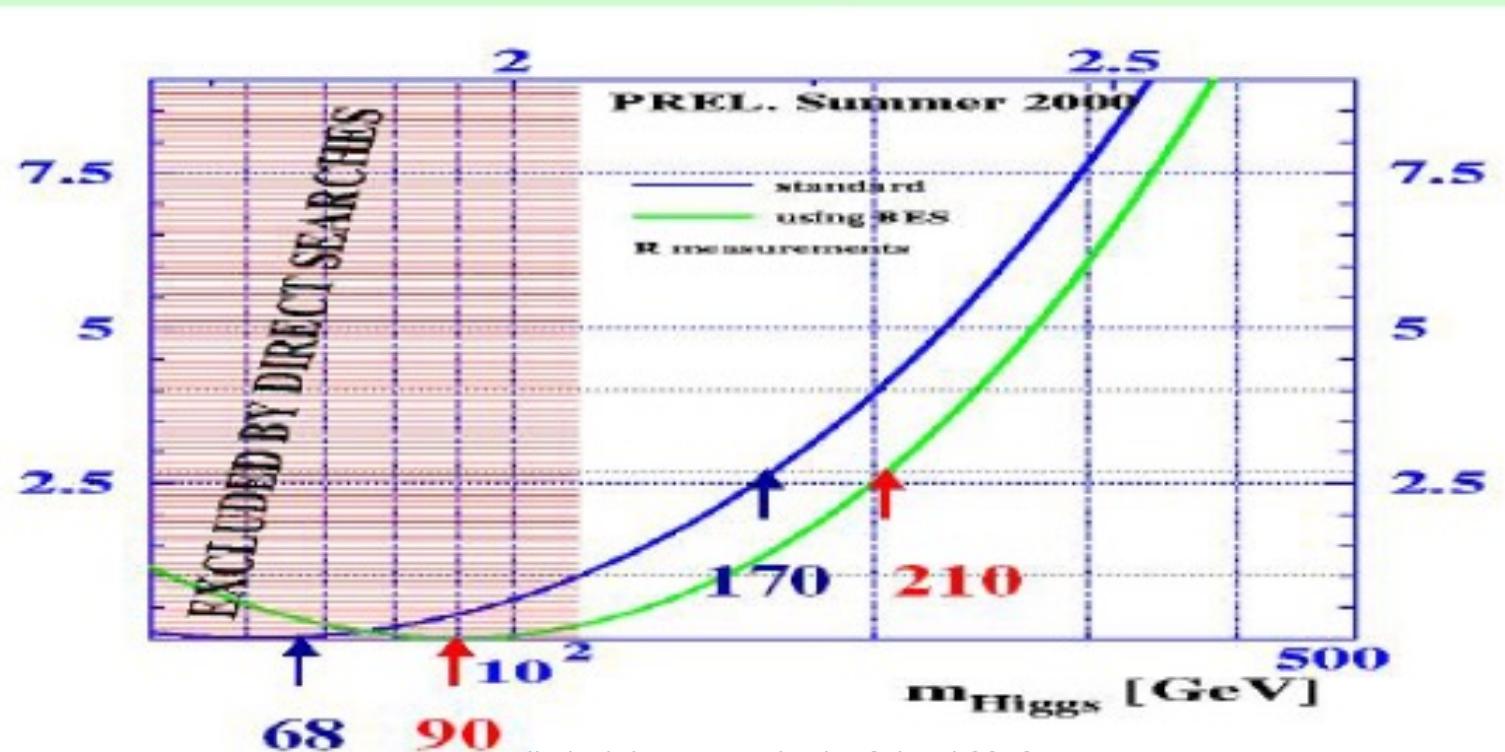
Impact of BES's New R Values on the SM Fit for $\alpha(M_Z^2)$ and Higgs mass

1995 before BES R data

$$\alpha(M_Z^2)^{-1} = 128.890 \pm 0.090$$

2001 after BES R data

$$\alpha(M_Z^2)^{-1} = 128.936 \pm 0.046$$



Selected results from BESI/BESII

- Test the 12% rule

$$Q_h = \frac{B_{\psi'} \times}{B_{J/\psi} \times} = \frac{B_{\psi'} e^+e^-}{B_{J/\psi} e^+e^-} = 12\%$$

M. Appelquist and H. D. Politzer, PRL34, 43 (1975)

$$\begin{aligned} \Gamma_h &= |M_h|^2 |\Psi(0)|^2 \\ &= (2/9\pi)(\pi^2 - 9) \frac{5}{18} \alpha_s^3 (\frac{4}{3} \alpha_s)^3 m_\phi'. \end{aligned} \quad (3)$$

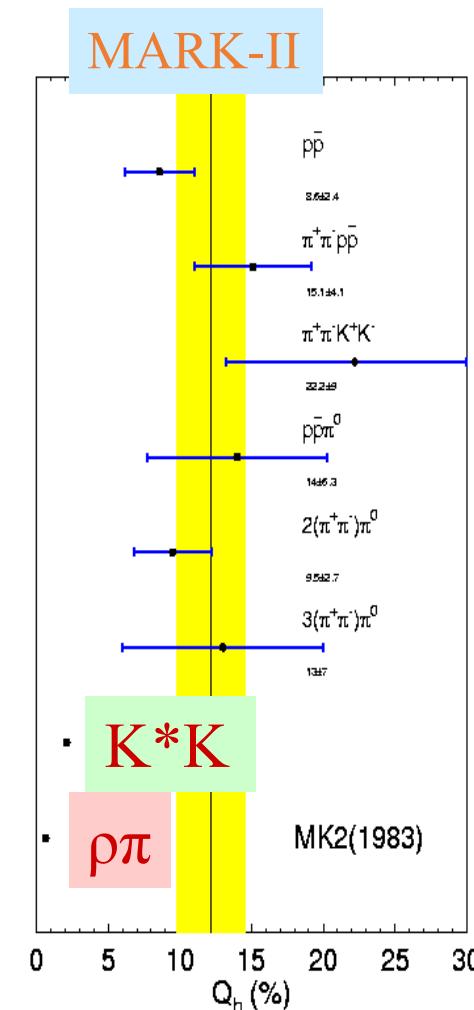
The leptonic width via one photon into $\bar{l}l$ is

$$\Gamma_l = |M_l|^2 |\Psi(0)|^2 = \frac{1}{2} (\frac{2}{3} \alpha)^2 (\frac{4}{3} \alpha_s)^3 m_\phi', \quad (4)$$

where $\alpha \approx \frac{1}{137}$. Although separately these calculations are not trustworthy, the ratio

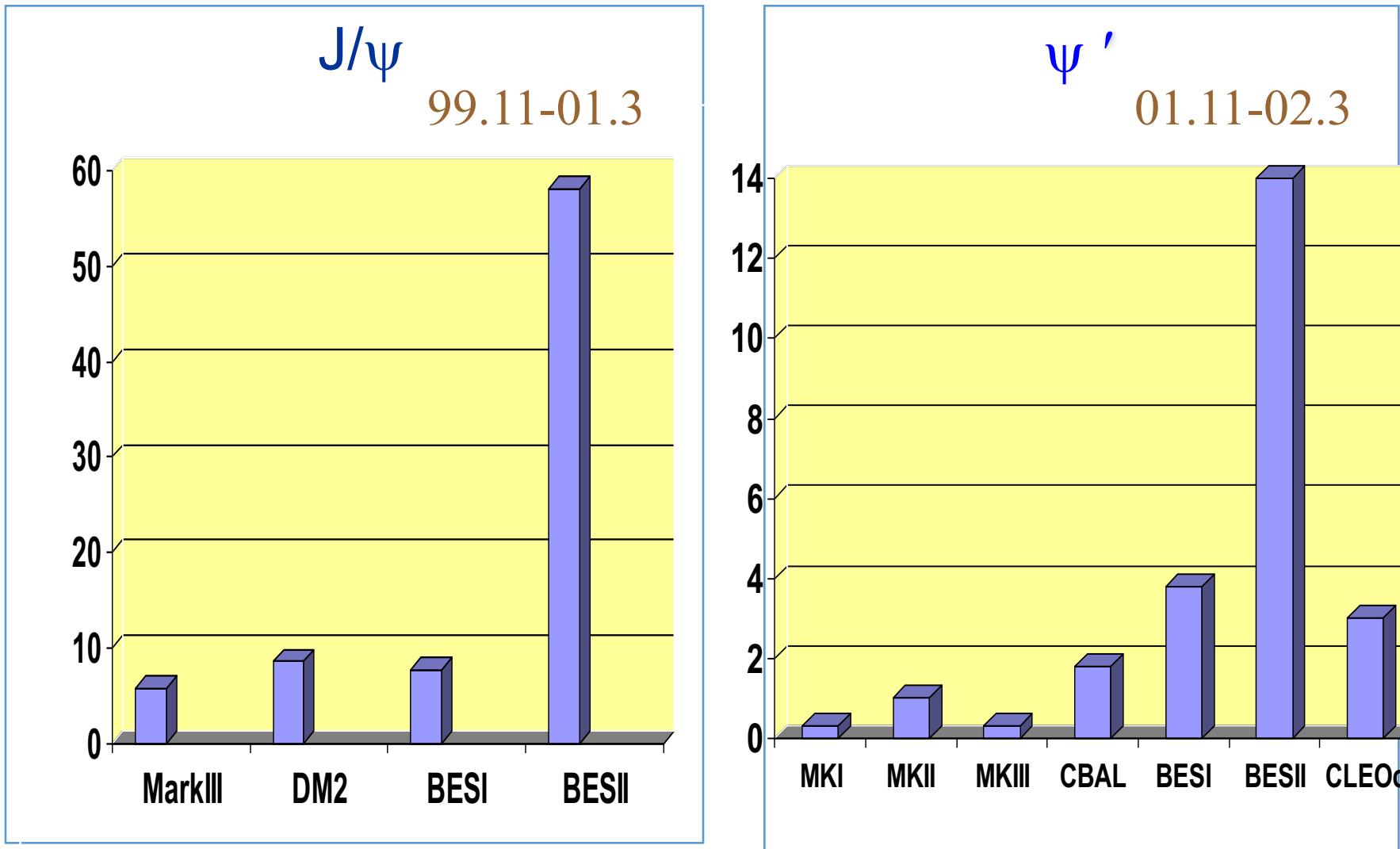
$$\frac{\Gamma_l}{\Gamma_h} = \frac{\frac{2}{9} \alpha^2}{(2/9\pi)(\pi^2 - 9) 5/\alpha_s^3} \quad (5)$$

is independent of wave-function effects.



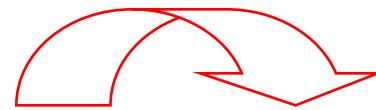
" $\rho\pi$ puzzle"

World largest J/ψ and ψ' data samples (10^6)



Extensively studied by BESII

- VP mode: $\rho \pi$, $K^{*+}K^- + c.c.$, $K^{*0}K^0 + c.c.$, $\omega\pi^0, \dots$
- PP mode: $K_S K_L$, $K^+ K^-$, $\pi^+ \pi^-$
- VT mode: $K^* K_{2^+}$, $\phi f_2^{'+}$, $\rho a2$, ωf_2
- 3-body: $p\bar{p}\pi^0$, $p\bar{p}\eta$, $p\bar{n}\pi^-$, $p\bar{n}\pi^+$, \dots
- Multi-body: $\pi^+ \pi^- \pi^0$, $K^+ K^-$, $3(\pi^+ \pi^-)$, \dots
- Radiative mode: $\gamma K \bar{K} \pi$, $\gamma \eta \pi \pi$

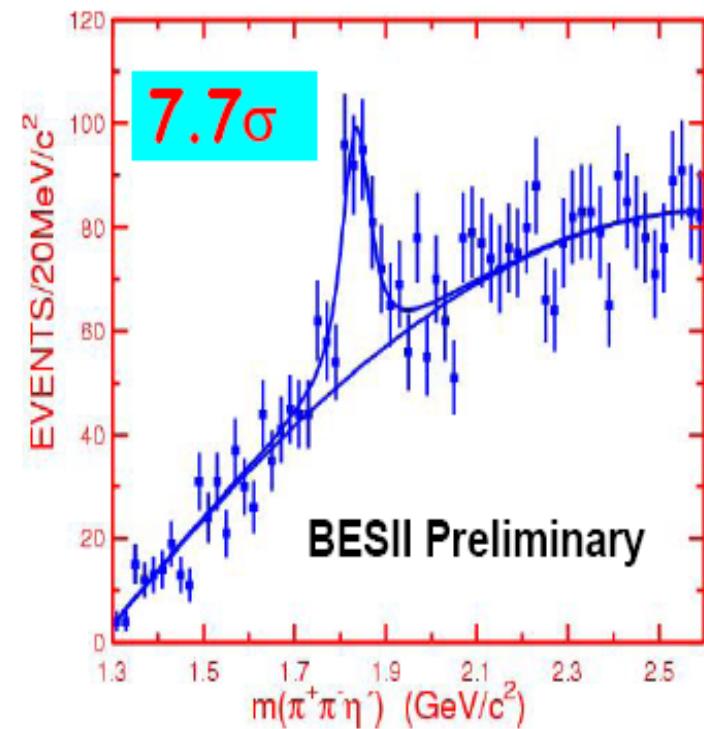
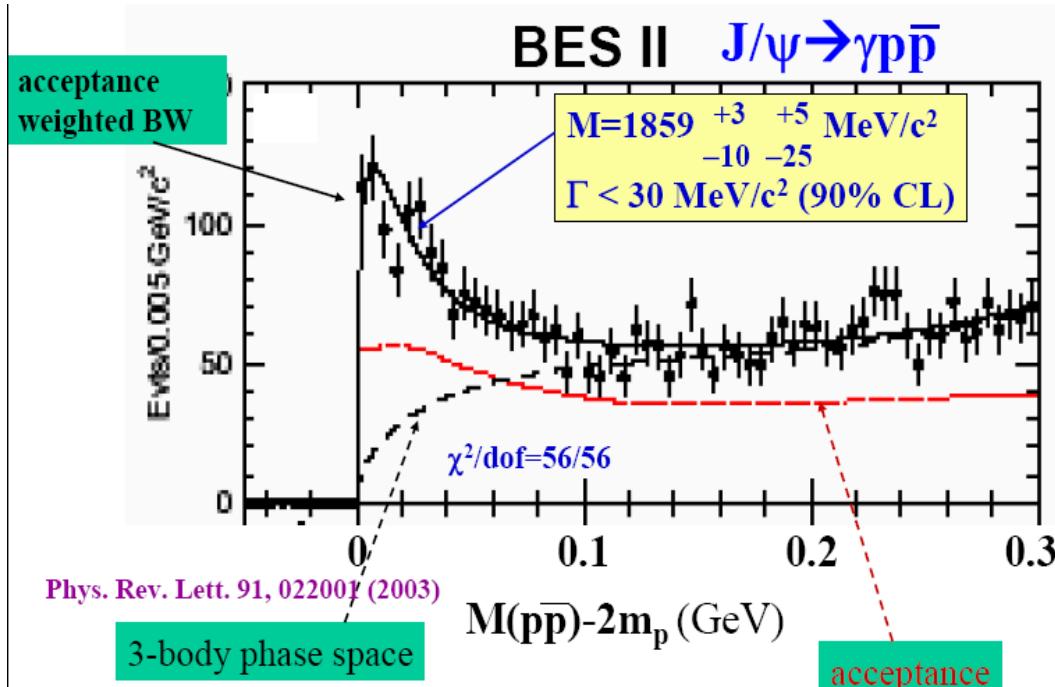


Is there a rule here?

ψ' → VP : suppressed
 ψ' → PP : enhanced/suppressed
 ψ' → VT : suppressed
Multi-body : obey/sup
Radiative : enhanced/suppressed

Selected results from BESI/BESII

- X(1835)



The $\pi^+\pi^-\eta'$ mass spectrum
for η' decaying into
 $\eta' \rightarrow \pi^+\pi^-\eta$ and $\eta' \rightarrow \gamma p$

$$N_{obs} = 264 \pm 54$$

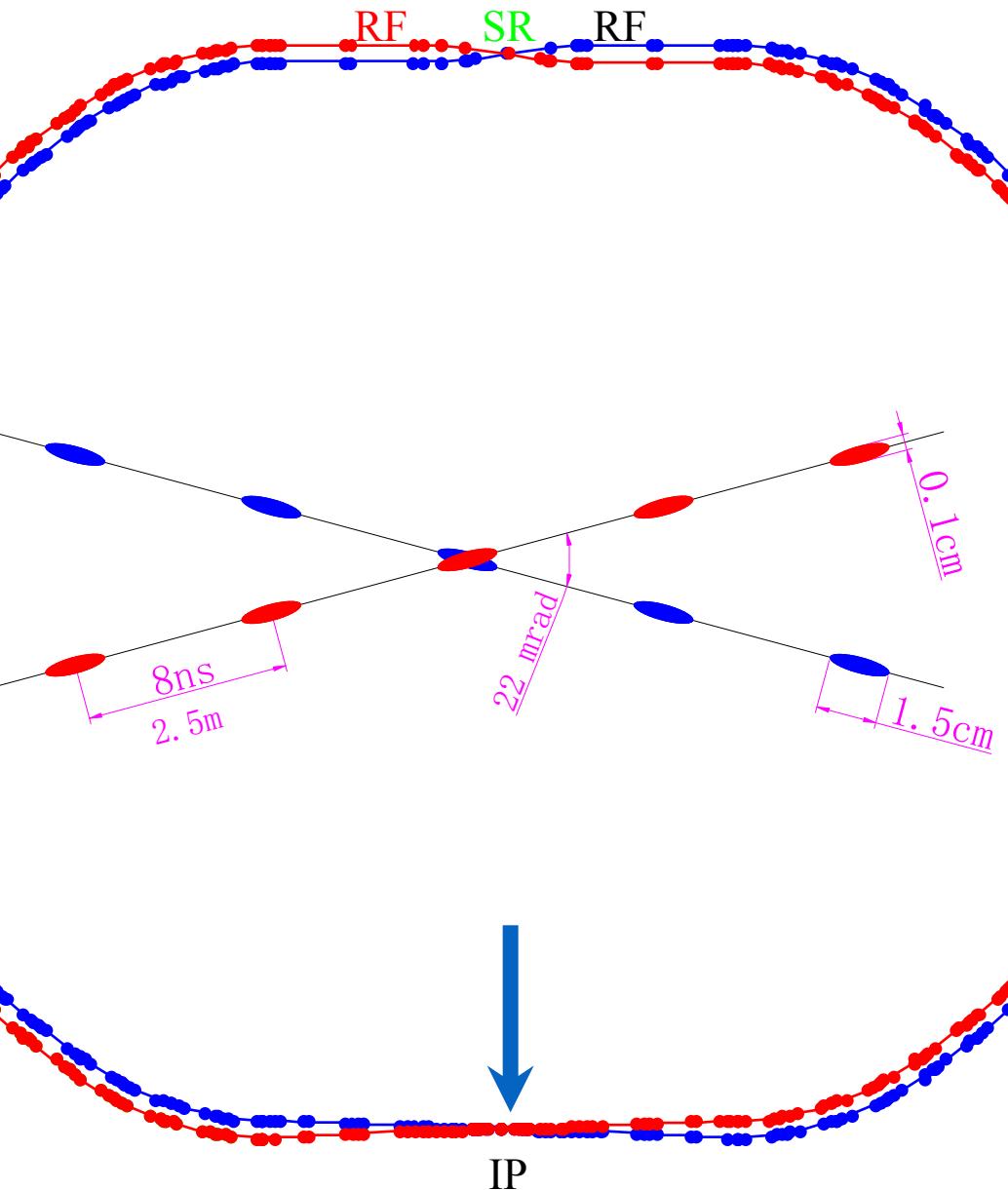
$$m = 1833.7 \pm 6.1 \pm 2.7 \text{ MeV}$$

$$\Gamma = 67.7 \pm 20.3 \pm 7.7 \text{ MeV}$$

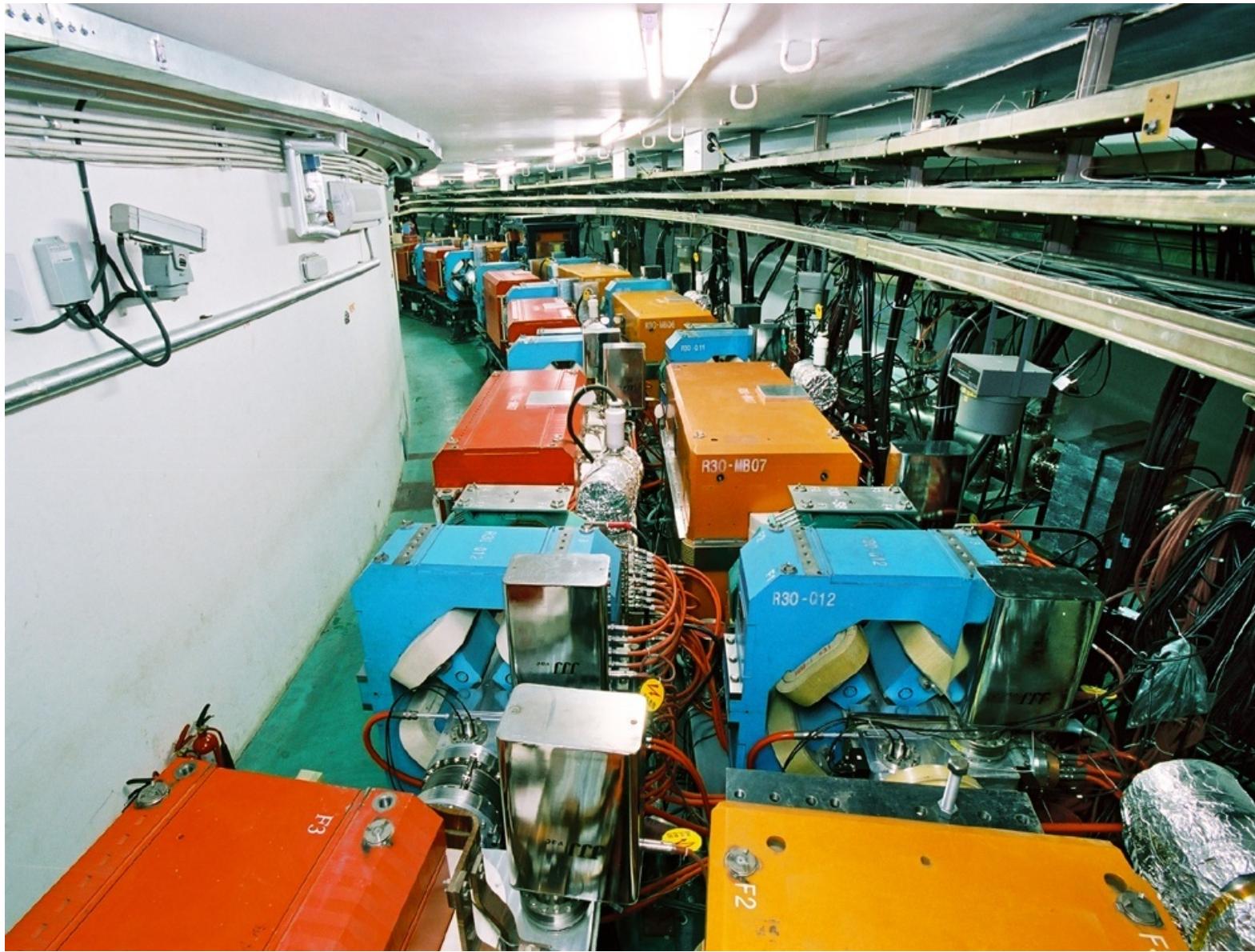
BEPCII and BESIII upgrade

- The state council approved the upgrade plan in 2003, total budget 650M RMB
- Upgrade BEPC to a tau-charm factory: BEPCII
- Build a state of the art detector: BESIII
- First physics run in 2008

BEPC II Storage ring: Large angle, double-ring



Beam energy:
1.0-2.3 GeV
Luminosity:
 $1 \times 10^{33} \text{ cm}^{-2}\text{s}^{-1}$
Optimum energy:
1.89 GeV
Energy spread:
 5.16×10^{-4}
No. of bunches:
93
Bunch length:
1.5 cm
Total current:
0.91 A
SR mode:
0.25A @ 2.5 GeV

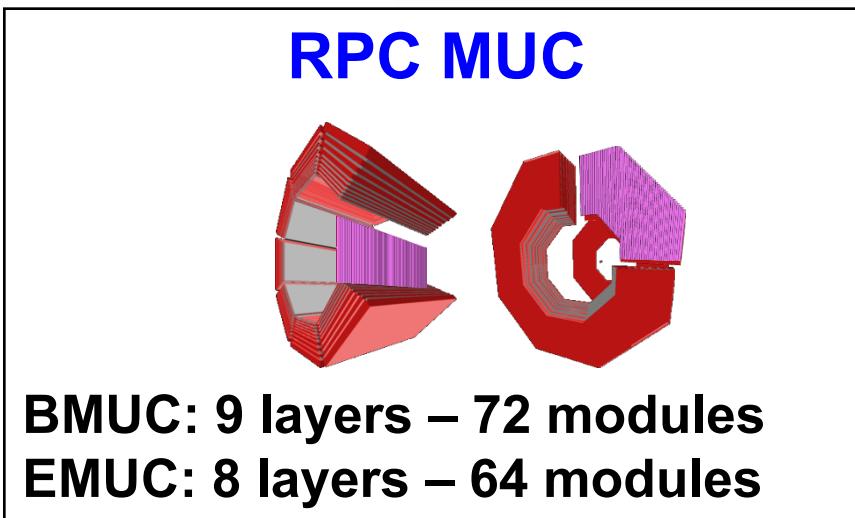
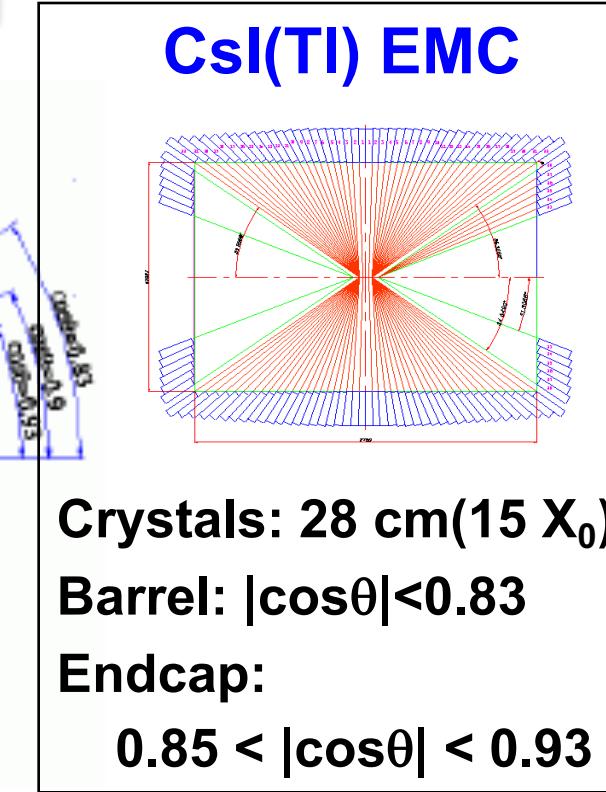
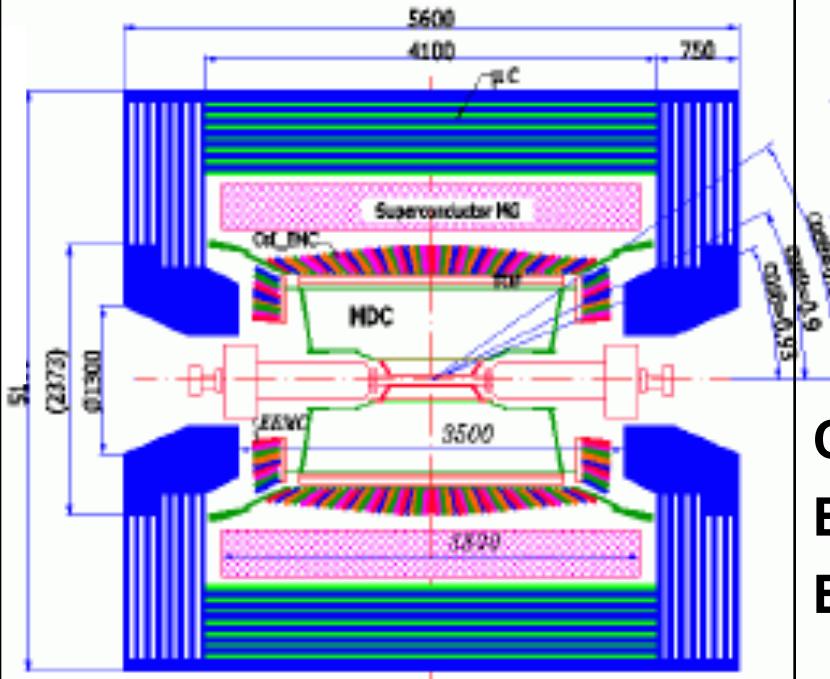
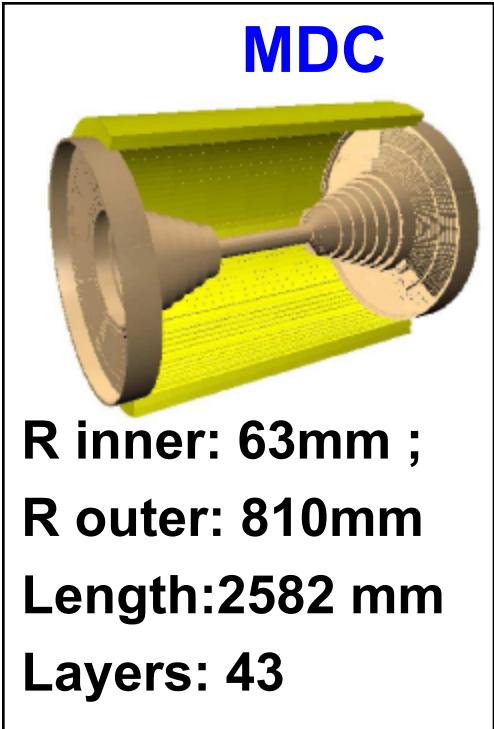


2016/7/30

Weihai High Energy Physics School 2016

22

BESIII Detector



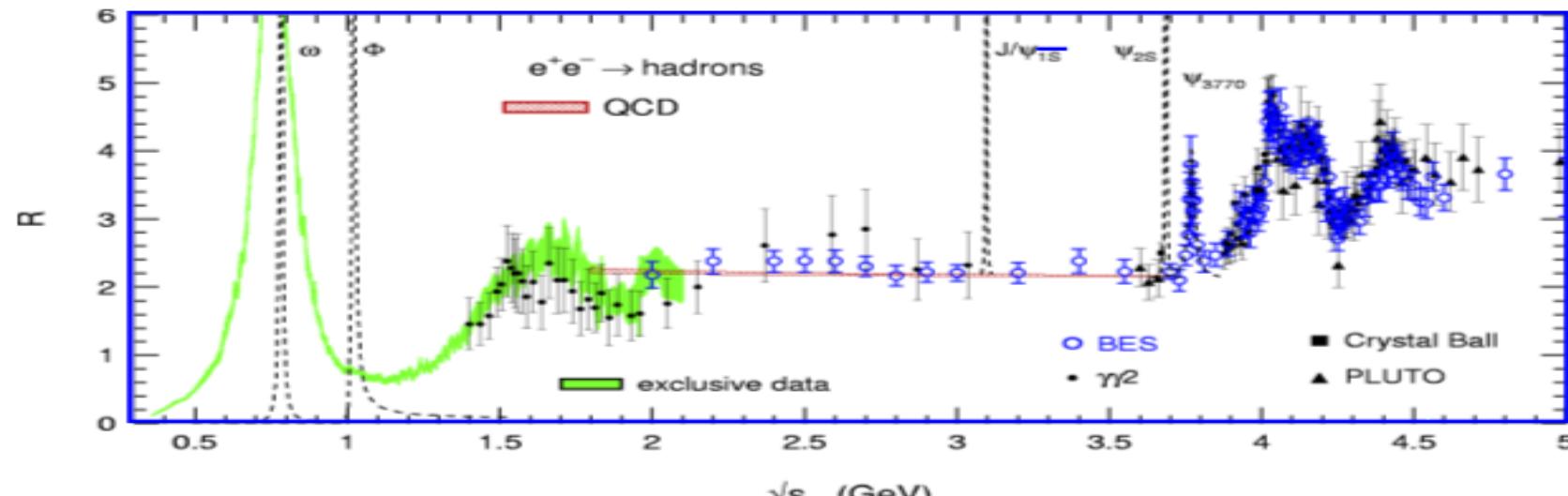
- Drift chamber and its electronics (IHEP, Sichuan, Tsinghua)
- CsI(Tl) calorimeter and its electronics (IHEP, Tsinghua)
- TOF (IHEP, USTC, Tokyo, Hawaii)
- TOF electronics (USTC)
- RPC (IHEP, Uni. of Washington)
- RPC electronics (USTC)
- Trigger (IHEP, USTC)
- DAQ & online software (IHEP, Tsinghua)
- Offline software (IHEP, Peking, Shangdong, Nanjing)
- Superconducting magnet (IHEP, Wang NMR)
- Mechanics (IHEP)
- Technical support (IHEP, Tsinghua)

BESIII International Collaboration



BESIII Data Taking Status & Plan

	Previous Data set	BESIII now	Goal
J/psi	BESII: 58M	1.2 B	10B
Psi'	CLEO: 28 M	0.5 B	3B
Psi''	CLEO: 0.8 /fb	3.0/fb	20 /fb
$\psi(4040)/\psi(4160)$ /X(4260) etc.	CLEO: 0.6/fb @ $\psi(4160)$	0.5/fb $\psi(4040)$; 2.3/fb @~4260, 0.5/fb@~4360; 1/fb@~4420; 0.5/fb@~4600	~ 10 /fb
R scan & Tau	BESII @10K/pnts	105 pnts@3.8-4.6 GeV	100K/pnts

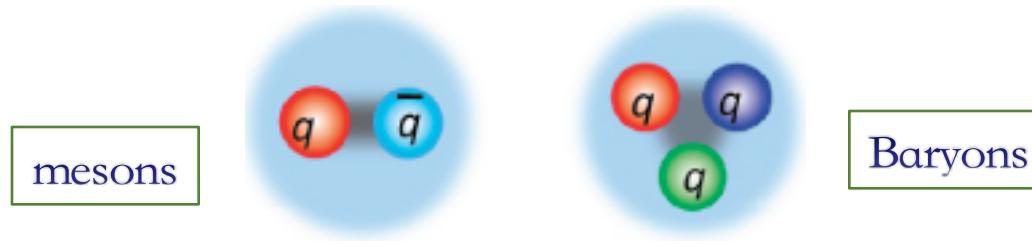


BESIII will continue for the next 8-10 years: Unique in the world

New Type of Hadrons

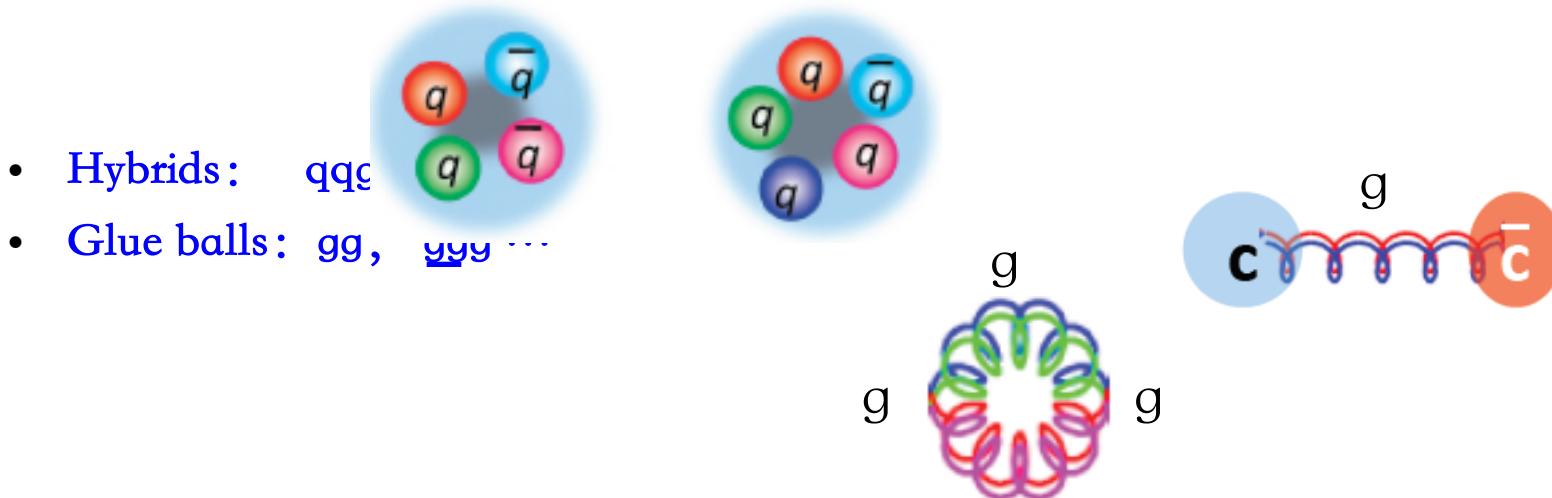
- Normal hadrons are made of 2 or 3 quarks :

Quark Model:



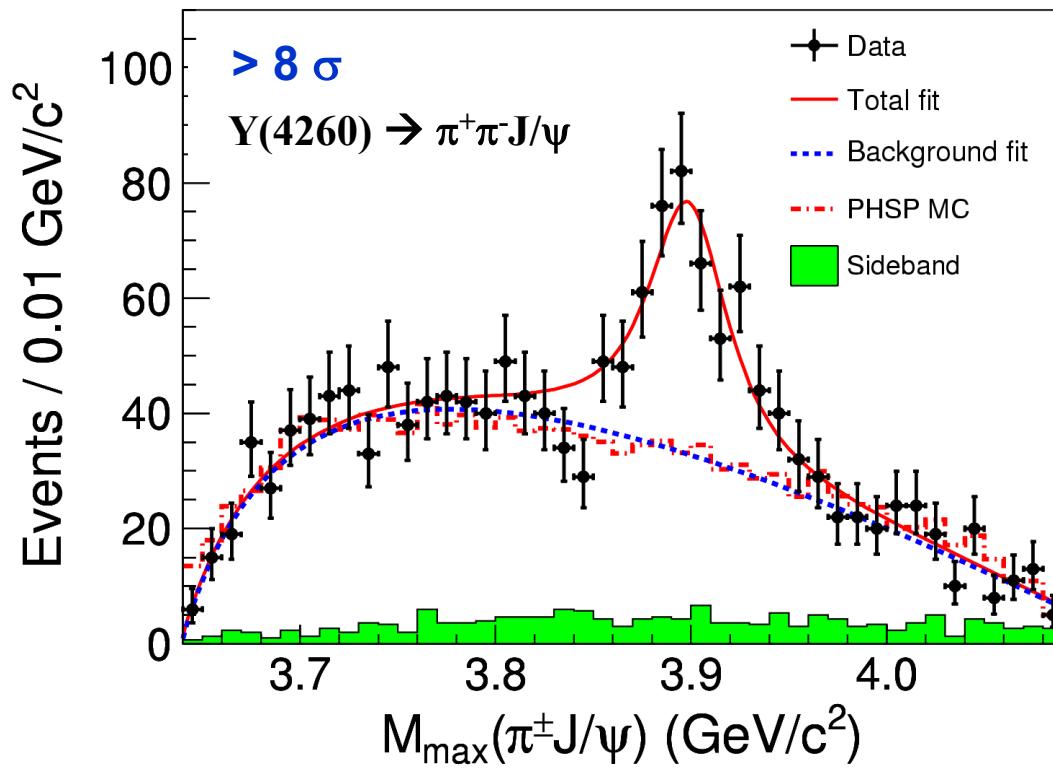
- QCD predict new type of hadrons:

- Multi-quark states: No. of quarks ≥ 4



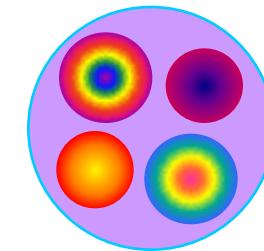
No experimental proof of new types of hadrons

BESIII: Observation of Zc(3900)



PRL110, 252001 (2013)

- Close to M(DD*)
- Couples to $\bar{c}c$
- Has electric charge
- At least 4-quarks
- What is its nature?

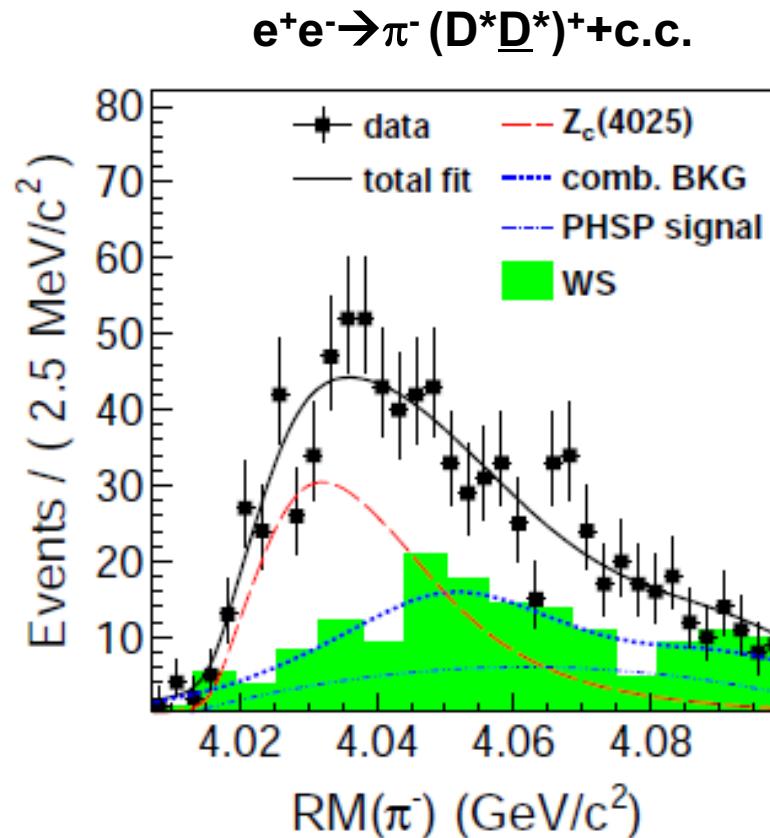


- S-wave Breit-Wigner with efficiency correction
- Mass = (3899.0 ± 3.6 ± 4.9) MeV
- Width = (46 ± 10 ± 20) MeV
- Fraction = (21.5 ± 3.3 ± 7.5)%

By collecting a lot of data, we may understand the nature of Y(4260), Zc and probably, many XYZ particles with the help of LQCD

$Z_c(4025)/Z_c(4020)$: Excited State of $Z_c(3900)$?

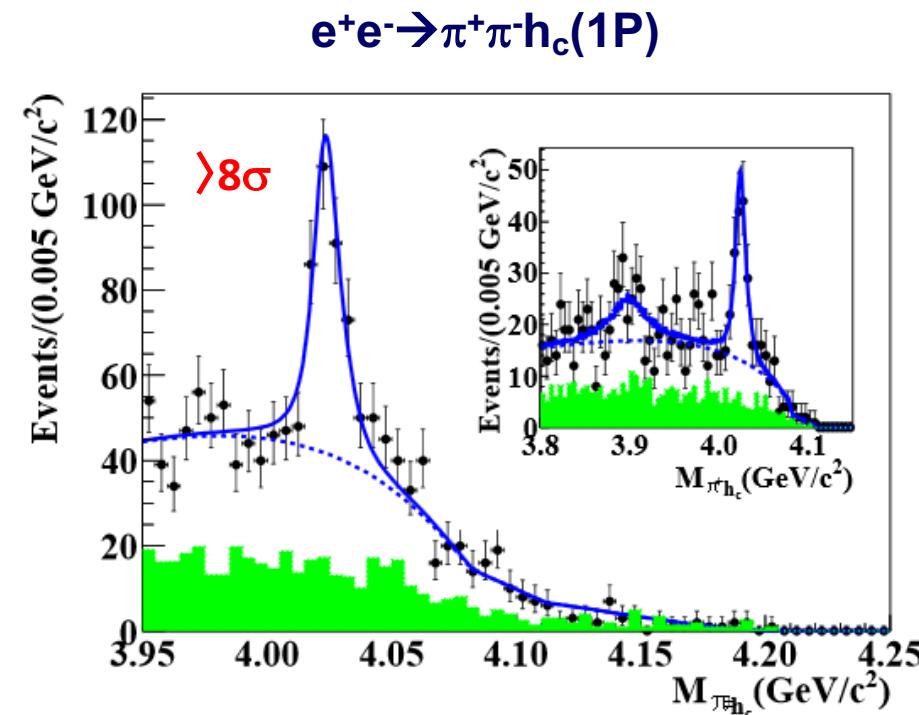
PRL 112, 132001 (2014)



$$M(Z_c(4025)) = 4026.3 \pm 2.6 \pm 3.7 \text{ MeV};$$

$$\Gamma(Z_c(4025)) = 24.8 \pm 5.7 \pm 7.7 \text{ MeV}$$

PRL 111, 242001 (2013)



$$M(Z_c(4020)) = 4022.9 \pm 0.8 \pm 2.7 \text{ MeV}$$

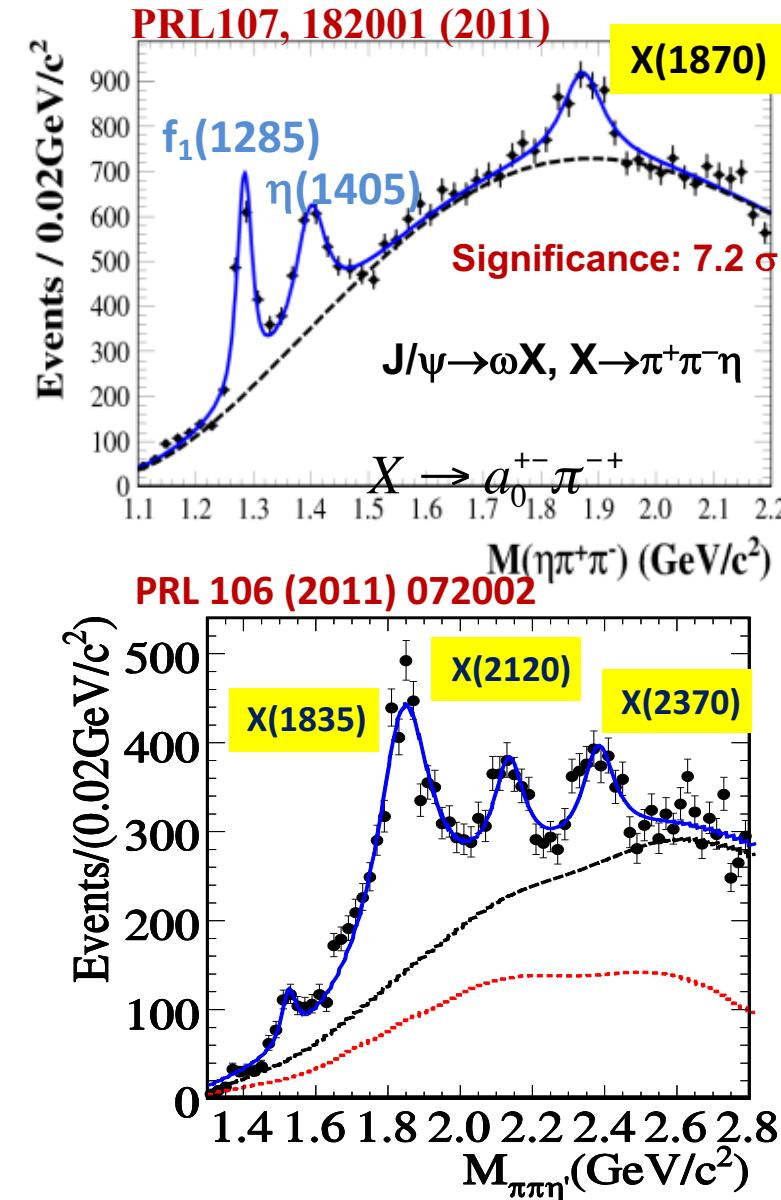
$$\Gamma(Z_c(4020)) = 7.9 \pm 2.7 \pm 2.6 \text{ MeV}$$

2016/04/05 02:35:20

Luminosity	9.98	E32/cm^2/s
e+		e-
Energy [GeV]	1.8832	1.8831
Current [mA]	843.12	857.67
Lifetime [hr]	1.61	2.30
Inj. Rate [mA/min]	0.00	0.00

Summary of physics at BEPCII/BESIII

- BEPCII/BESIII is the best facility for light hadrons spectroscopy and Charmonium physics: glueballs, m_τ , f_D , f_{D_s} , R , etc.
- ~ 20 papers/year, ~ 100 papers in total so far.
- BESIII will continue to operate for another ~ 8 years.
- After 2020, a possible scenario:
 - a circular e^+e^- collider at 90GeV and 250 GeV (Z & Higgs factory)
 - a pp collider in the same tunnel afterwards



#13

25-AUG-1986 04:11:24

NEWMAIL

From: VXCRNA::SHUQIN
To: STEINBERGER
Subj: link

dear jack,i am very glad to send this letter to you via computing link which i think is first sucessful link test between cern and china.i would like to thank you again for your visit which leads this valuable test to be success. now i think each collaborators amoung aleph collaboration have computing link which

is very important.ofcause we still have problems to use this link effectively for analizing dst of aleph in being. and need to find budget in addition, but most

important thing is to get start.at the moment,we use the ibm-pc in 710 institute to connect to you, later we will try to use the microwave communicated equipment which we have used for linking m160h before,to link to you directly from our institute.

Please send my best regards to all of our colleagues and best wishes to you. cynt hia
and your family.

by the way, how about the carpet you bought in shanghai?
weimin

吴为民向瑞士的斯坦伯格教授发出的中国第一封电子邮件原件
现代物理知识 2009 Vol. 21 (3): 57-61

Particle Physics & Internet in China

1986 1st email to CERN

1991 Network link IHEP-SLAC

1993 64K BPS internet connection to the world. Email service for institutions and universities.

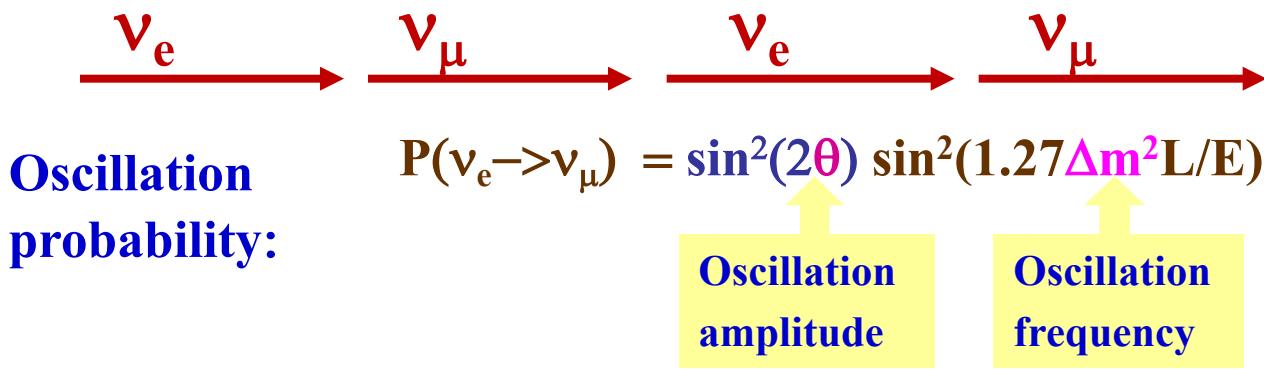
1994 Connected to WWW

2008 LCG

DayaBay & JUNO

Neutrino Oscillation

- If the neutrino mass eigenstate is different from that of the weak interaction, neutrinos can oscillate: from one type to another during the flight:



Oscillation matrix for 3 generations:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} V_{e1} & V_{e2} & V_{e3} \\ V_{\mu 1} & V_{\mu 2} & V_{\mu 3} \\ V_{\tau 1} & V_{\tau 2} & V_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- Known parameters: θ_{23} , θ_{12} , $|\Delta M^2_{23}|$, ΔM^2_{12} ,
- Recent progress: θ_{13}
- Unknown parameters: mass hierarchy(ΔM^2_{23}), CP phase δ



Бруно Понтекорво

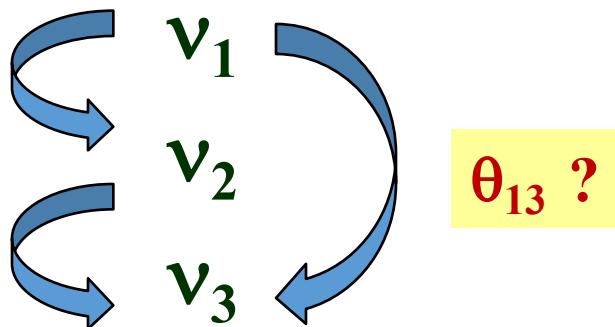
Bruno
Pontecorvo

Next Question in 2003: θ_{13} ?

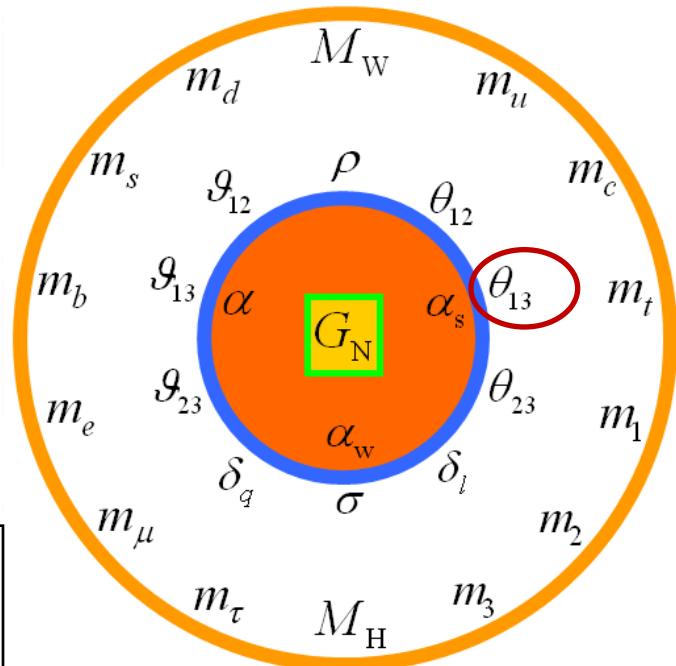
- Fundamental principles
- Fundamental parameter
- Direction of future neutrino physics:
 - If θ_{13} is too small, CPV cannot be figured out in the near future

Solar ν
Oscillation
 $\sin^2 2\theta_{12} \sim 0.9$

Atm. ν
Oscillation
 $\sin^2 2\theta_{23} \sim 1$

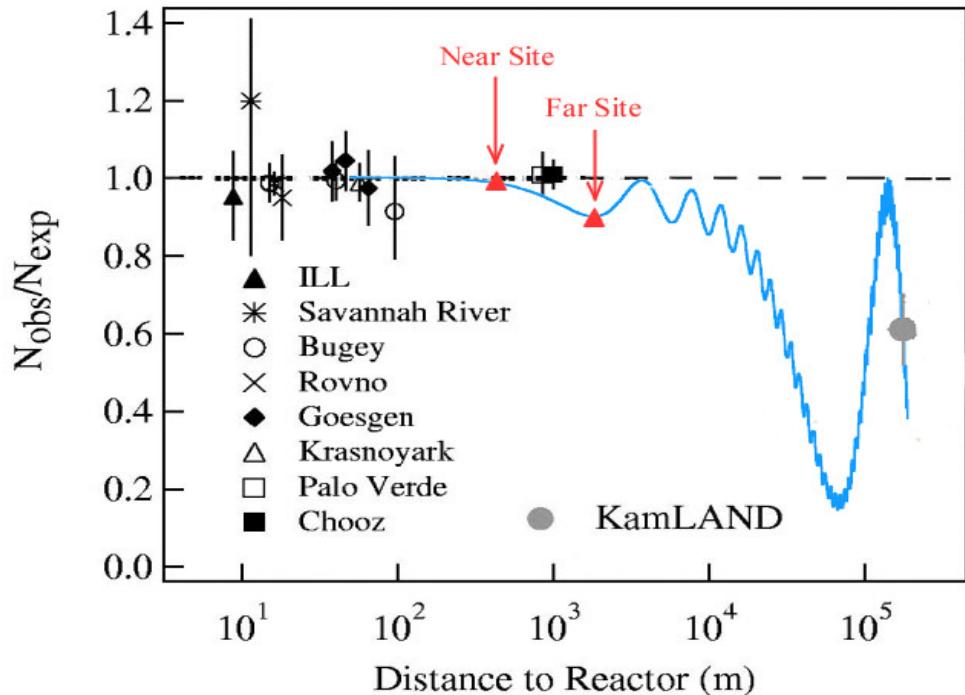


Two ways to measure θ_{13} :
At reactor
At accelerators

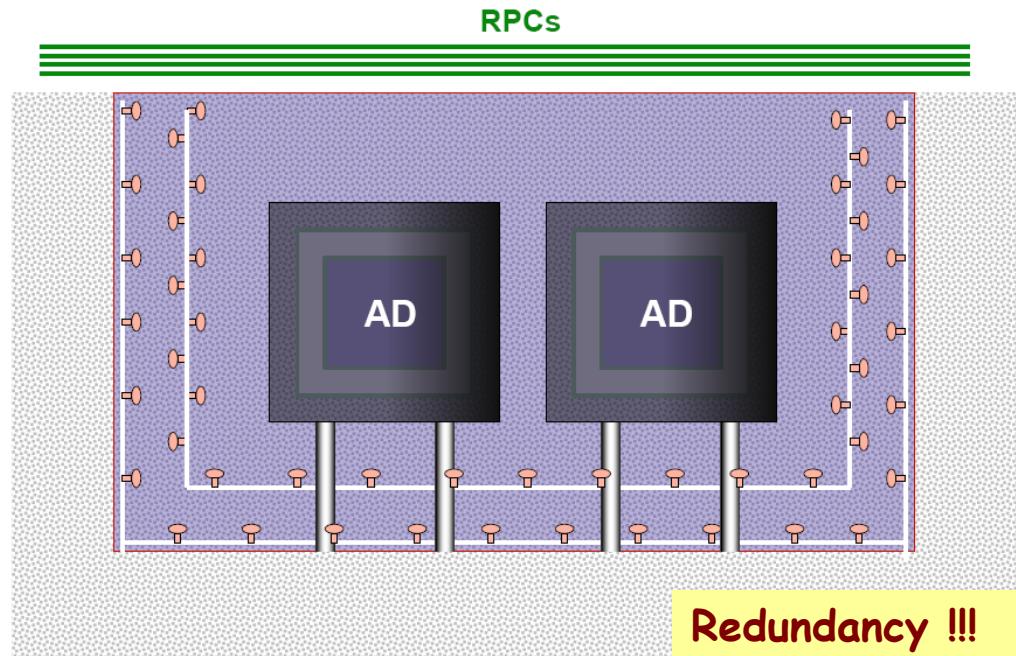


How to Measure θ_{13} at Reactors ?

$$P_{e \rightarrow e} \approx 1 - \sin^2 2\theta_{13} \sin^2 (1.27 \Delta m_{13}^2 L/E) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 (1.27 \Delta m_{12}^2 L/E)$$

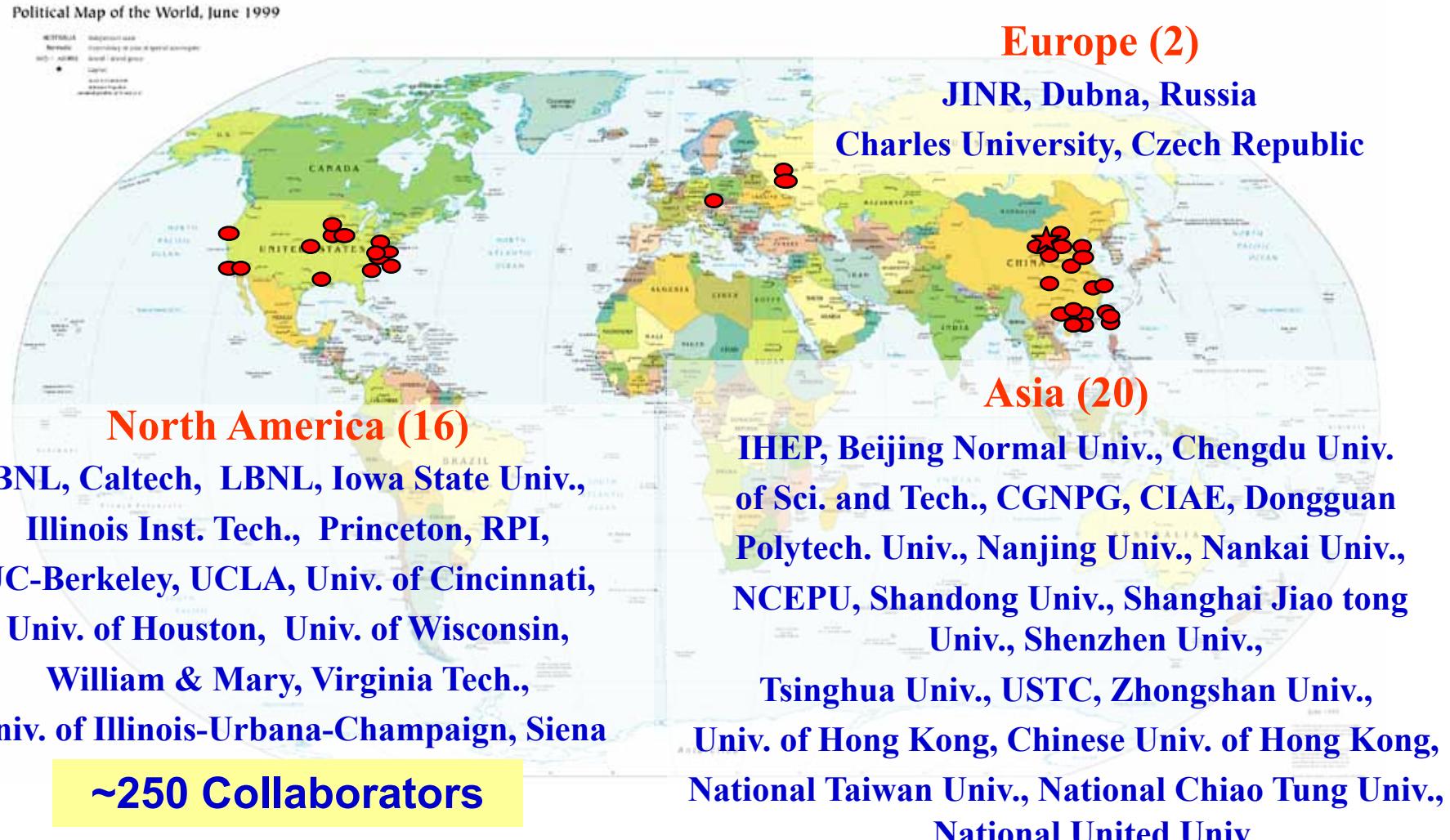


Daya Bay Experiment: Layout

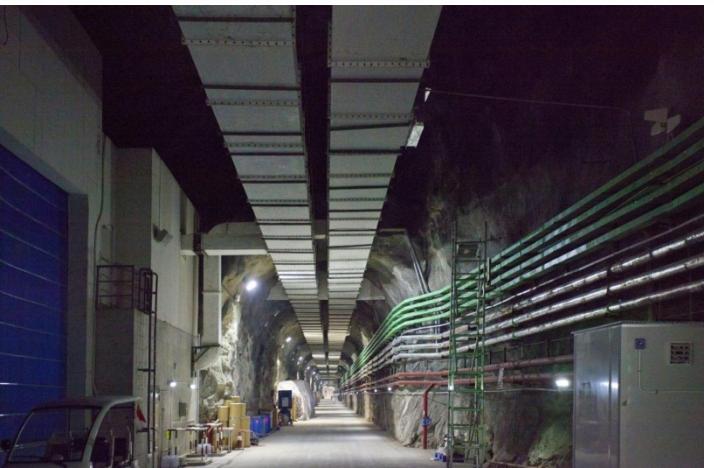
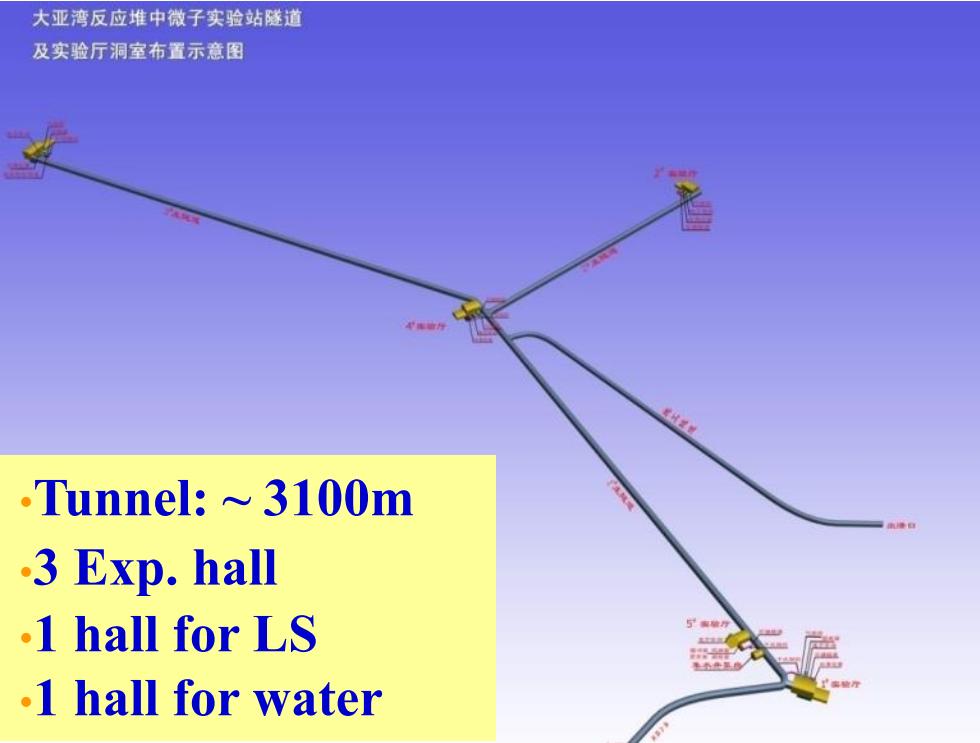


- Relative measurement to cancel Corr. Syst. Err.
 - **2 near sites, 1 far site**
- Multiple AD modules at each site to reduce Uncorr. Syst. Err.
 - **Far: 4 modules, near: 2 modules** Cross check; Reduce errors by $1/\sqrt{N}$
- Multiple muon detectors to reduce veto eff. uncertainties
 - **Water Cherenkov: 2 layers**
 - **RPC: 4 layers at the top + telescopes**

The Daya Bay Collaboration

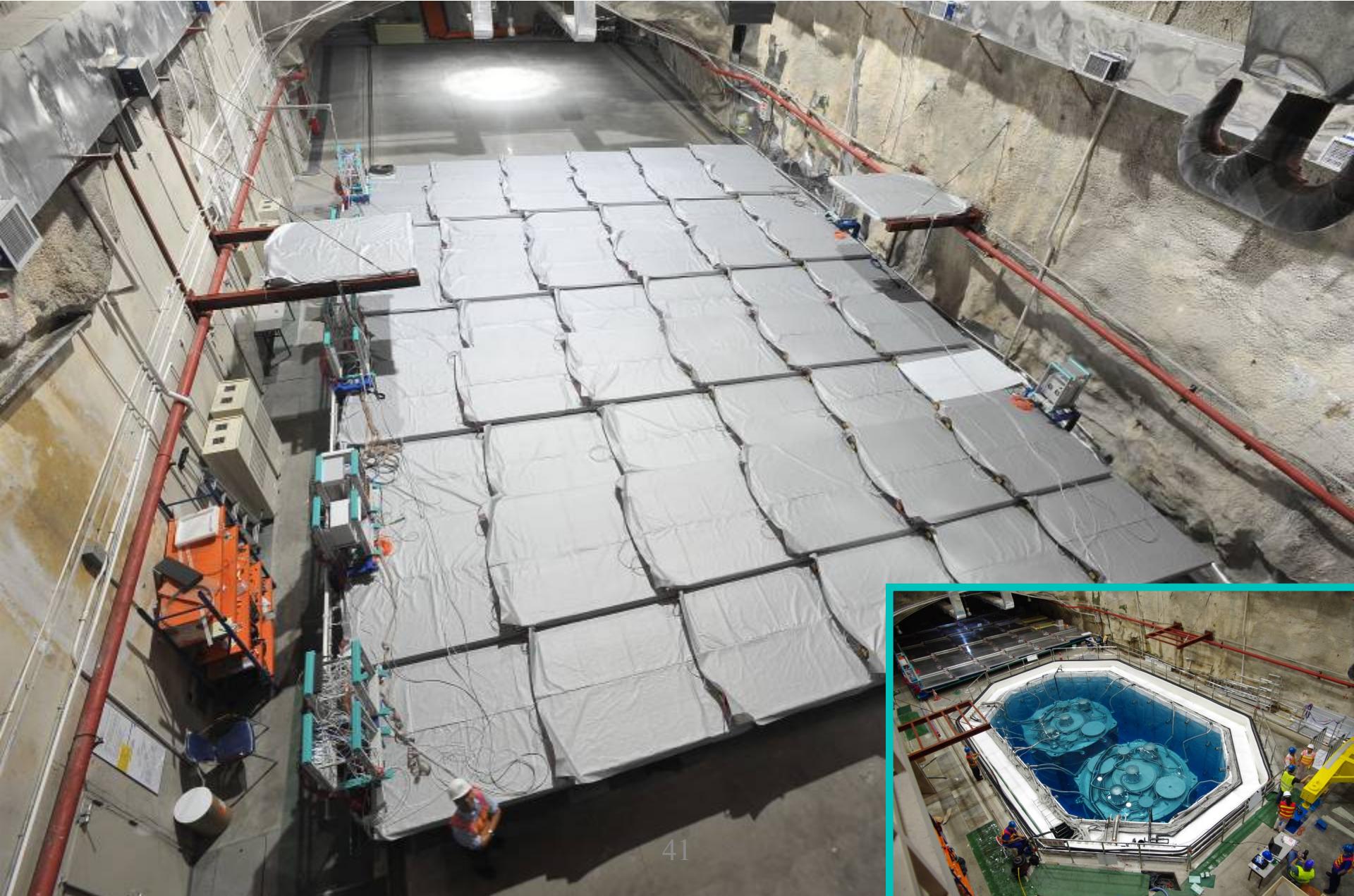


Tunnel and Underground Lab



A total of ~3000 blasting right next reactors. No one exceeds safety limit set by National Nuclear Safety Agency (0.007g)

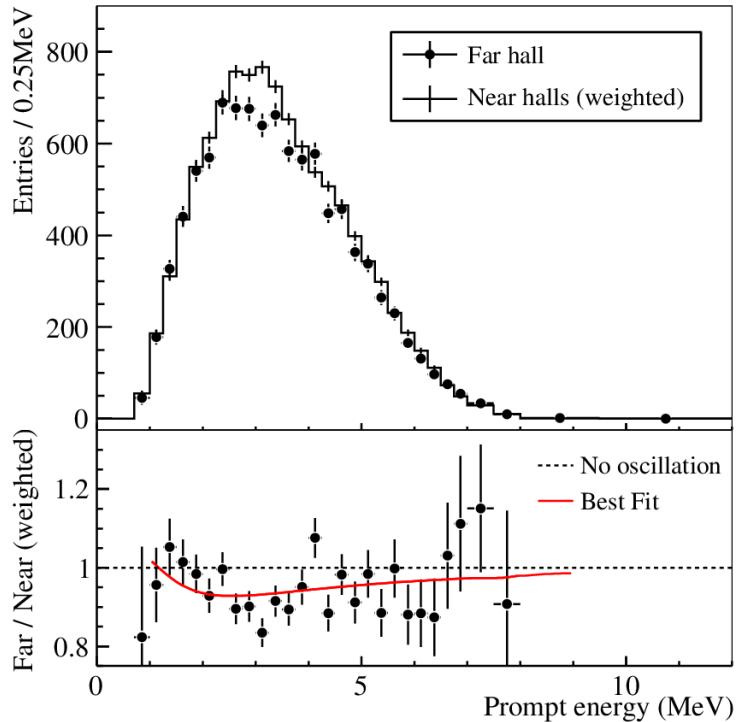
Experimental Hall in Operation



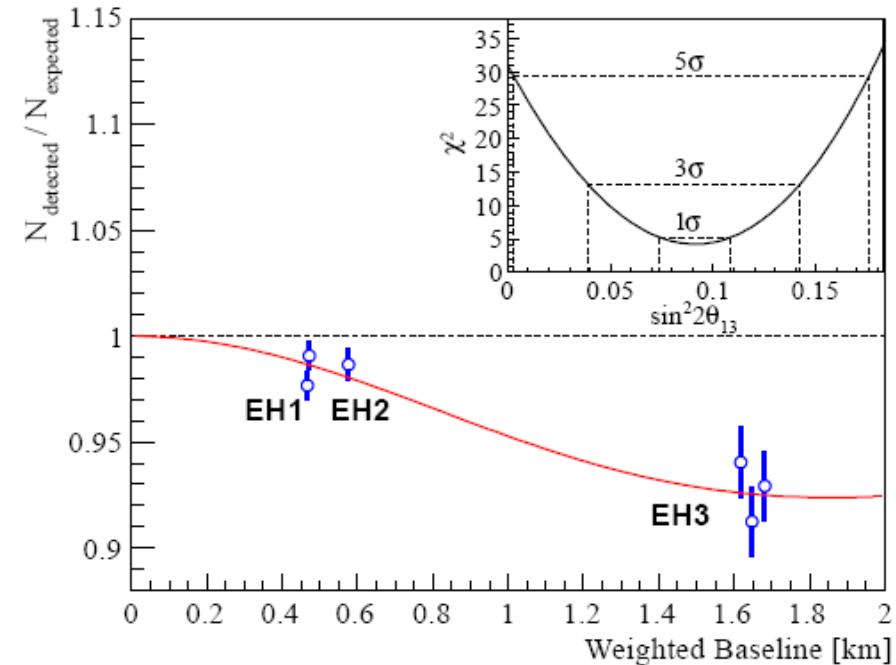
- Observation of electron anti-neutrino disappearance:

$$R = 0.940 \pm 0.011 \text{ (stat)} \pm 0.004 \text{ (syst)}$$

announced on
Mar. 8, 2012



$\sin^2 2\theta_{13} = 0.092 \pm 0.016 \text{ (stat)} \pm 0.005 \text{ (syst)}$
 $\chi^2/\text{NDF} = 4.26/4, \text{ 5.2 } \sigma \text{ for non-zero } \theta_{13}$



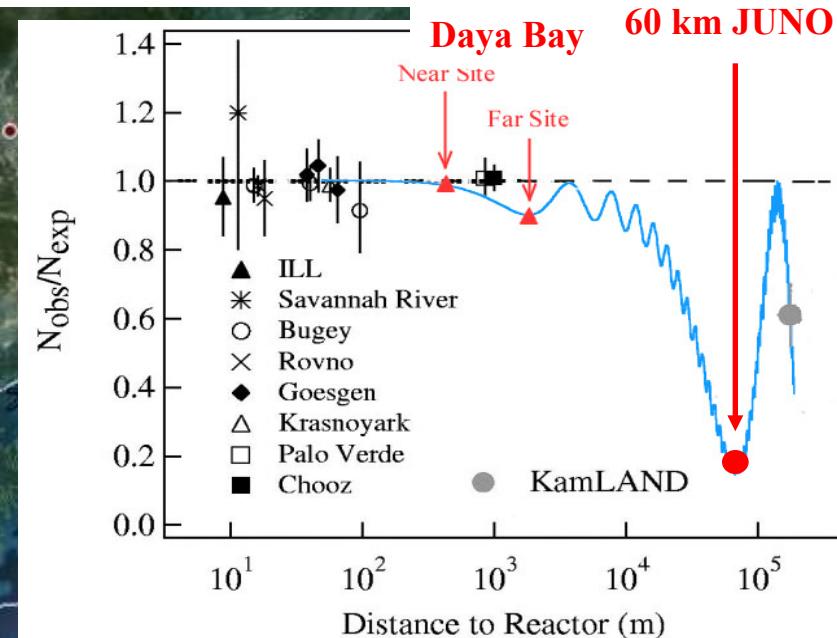
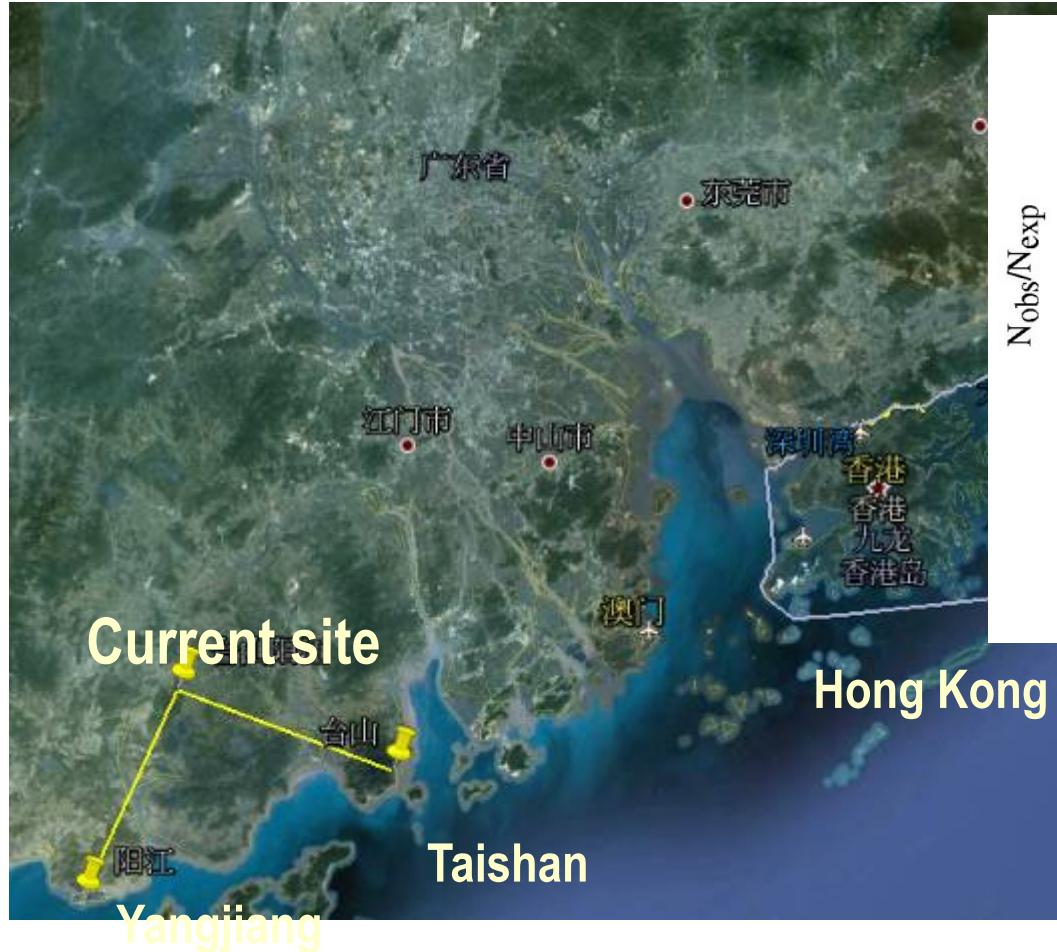
F.P. An et al., NIM A 685(2012)78
F.P. An et al., Phys. Rev. Lett. 108,
(2012) 171803

Still a Lot of Unknowns

- Neutrino oscillation:
 - **Neutrino mass hierarchy ?**
 - **Unitarity of neutrino mixing matrix ?**
 - Θ_{23} is maximized ?
 - CP violation in the neutrino mixing matrix as in the case of quarks ? Large enough for the matter-antimatter asymmetry in the Universe ?
- What is the absolute neutrino mass ?
- Neutrinos are Dirac or Majorana ?
- Are there sterile neutrinos ?
- Do neutrinos have magnetic moments ?
- Can we detect relic neutrinos ?
-

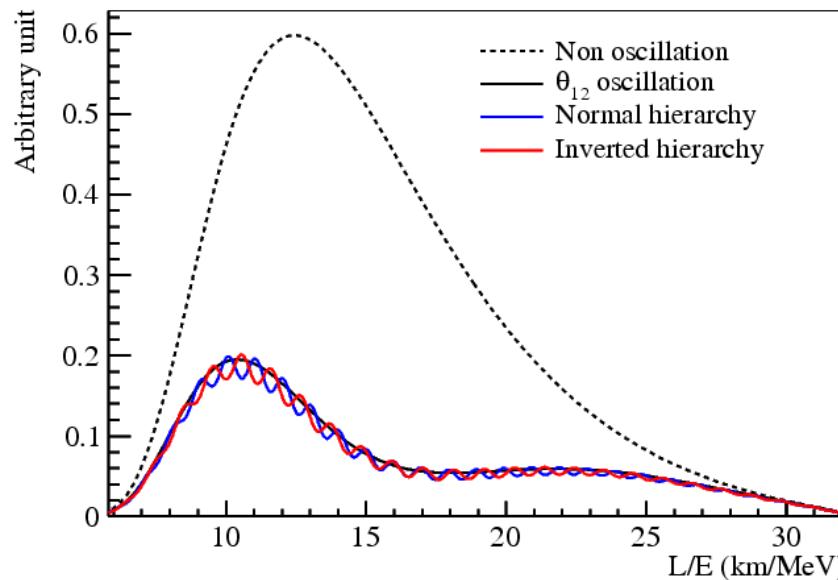
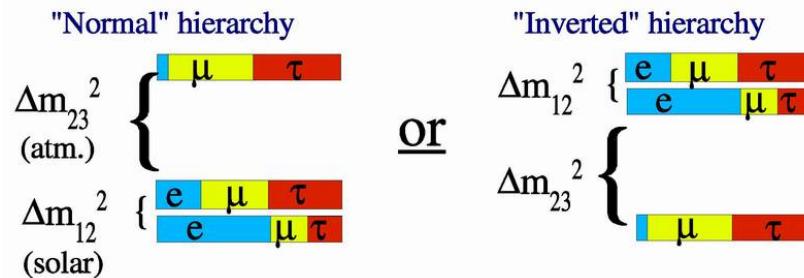
Next Step: Mass Hierarchy

	Daya Bay	Huizhou	Lufeng	Yangjiang	Taishan
Status	running	planned	approved	Construction	construction
power/GW	17.4	17.4	17.4	17.4	18.4



Talk by YFW at ICFA seminar 2008,
Neutel 2011; by J. Cao at NuTurn 2012 ;
Paper by L. Zhan, YFW, J. Cao, L.J. Wen,
PRD78:111103,2008; PRD79:073007,2009

Mass Hierarchy at Reactors



$$P_{ee}(L/E) = 1 - P_{21} - P_{31} - P_{32}$$

$$P_{21} = \cos^4(\theta_{13}) \sin^2(2\theta_{12}) \sin^2(\Delta_{21})$$

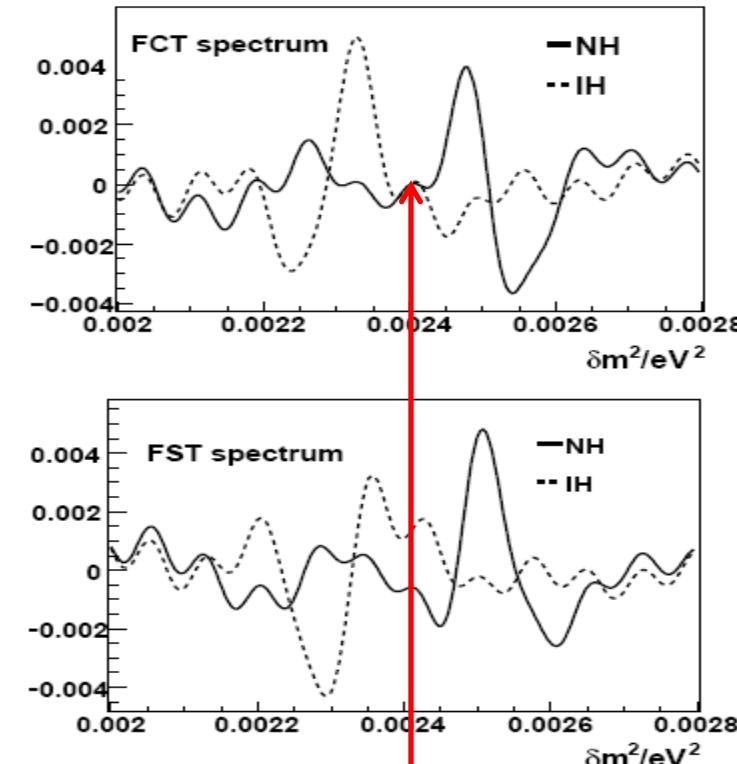
$$P_{31} = \cos^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{31})$$

$$P_{32} = \sin^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{32})$$

$$\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$$

NH : $|\Delta m_{31}^2| = |\Delta m_{32}^2| + |\Delta m_{21}^2|$

IH : $|\Delta m_{31}^2| = |\Delta m_{32}^2| - |\Delta m_{21}^2|$

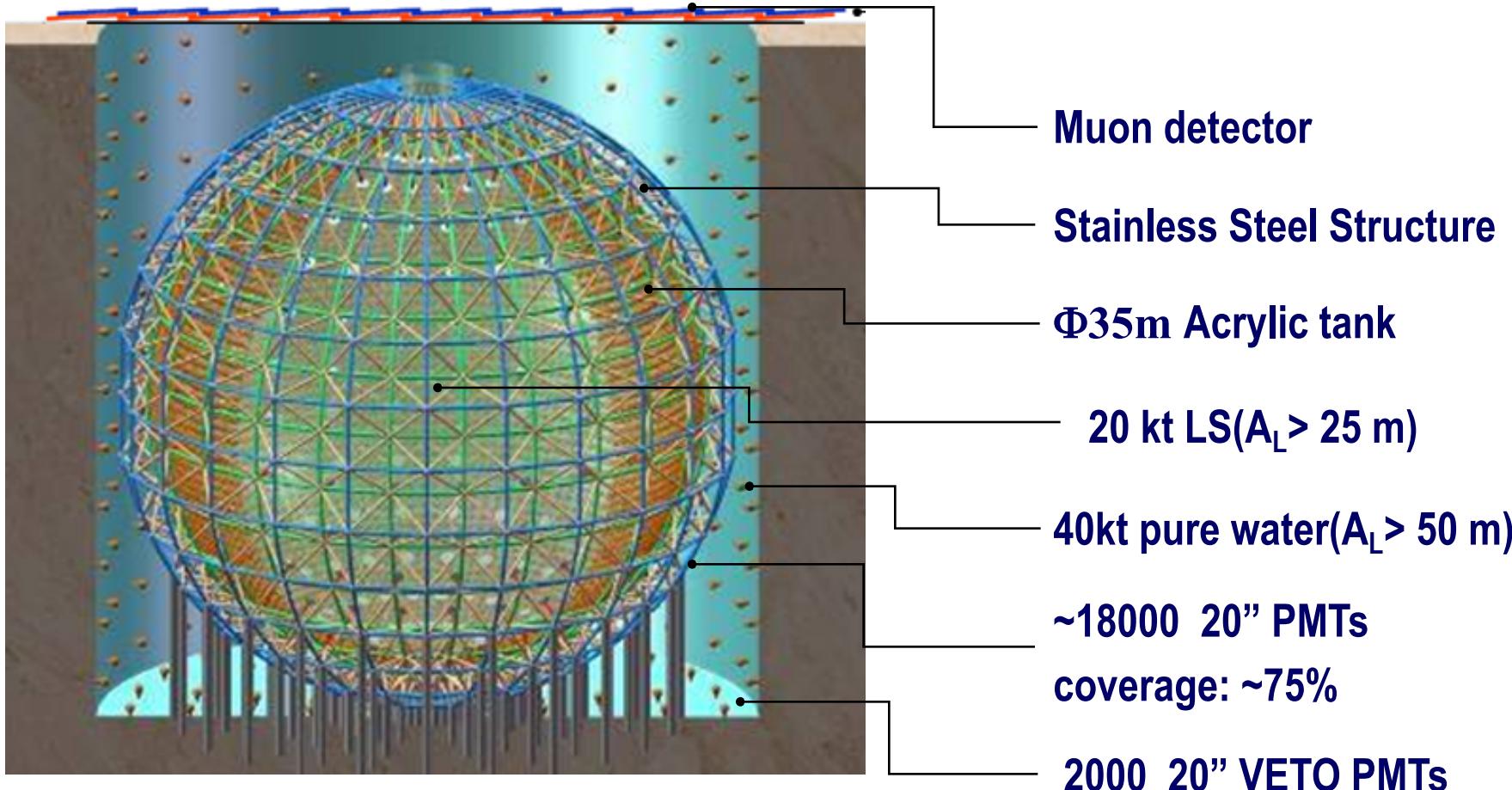


ΔM^2_{23}

L. Zhan et al., PRD78:111103,2008;
PRD79:073007,2009

The Plan: a Large LS Detector

- LS volume: $\times 20 \rightarrow$ for more statistics (40 events/day)
- light(PE) $\times 5 \rightarrow$ for better resolution ($\Delta M^2_{12}/\Delta M^2_{23} \sim 3\%$)



Physics Reach

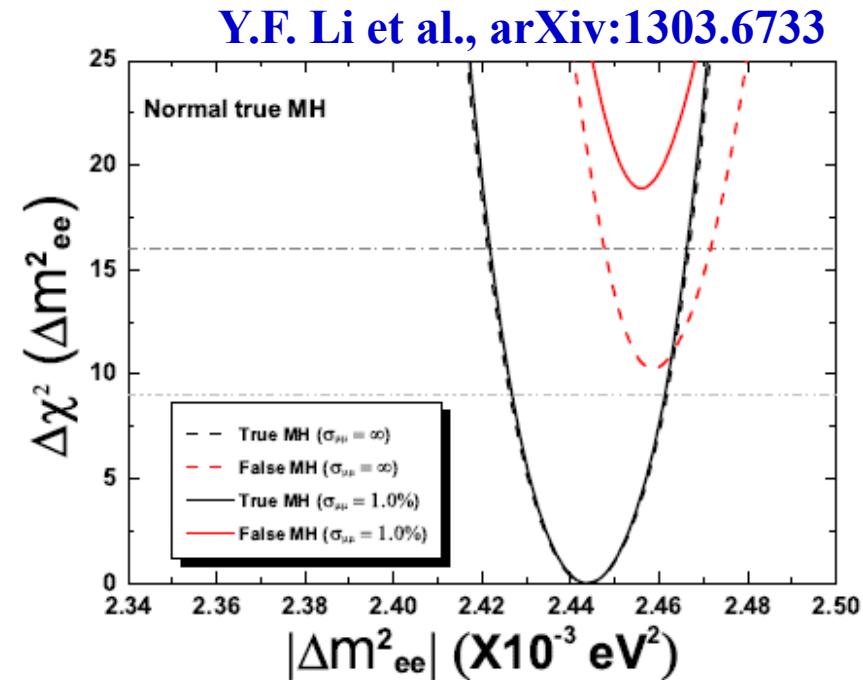
Thanks to a large θ_{13}

- Mass hierarchy
- Precision measurement of mixing parameters
- Supernova neutrinos
- Geoneutrinos
- Sterile neutrinos
-

	Current	Daya Bay II
Δm^2_{12}	4%	0.6%
Δm^2_{23}	4%	0.6%
$\sin^2 \theta_{12}$	6%	0.7%
$\sin^2 \theta_{23}$	10%	N/A
$\sin^2 \theta_{13}$	6% \rightarrow 4%	$\sim 15\%$

47

Detector size: 20kt
Energy resolution: 3%/ \sqrt{E}
Thermal power: 36 GW



For 6 years, mass hierarchy can be determined at 4σ level, if $\Delta m^2_{\mu\mu}$ can be determined at 1% level

Schedule & Current Status

Schedule:

Civil preparation: 2013-2014

Civil construction: 2014-2017

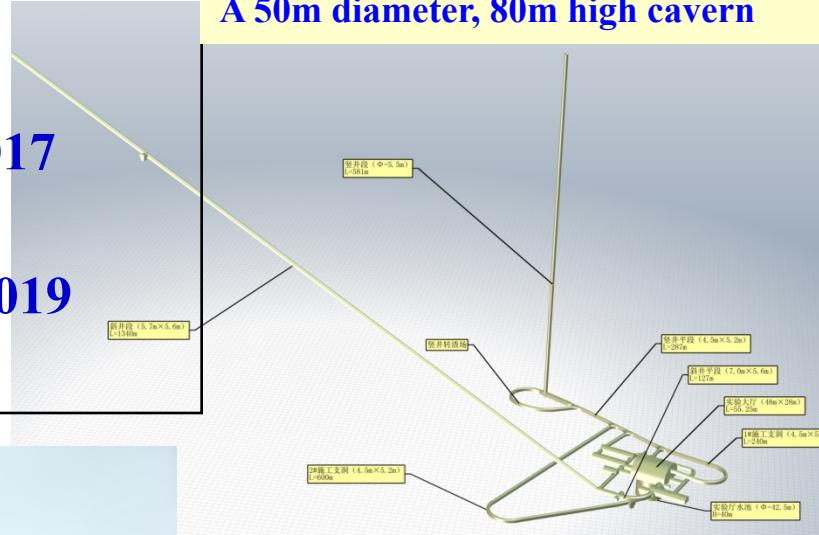
Detector component production: 2016-2017

PMT production: 2016-2019

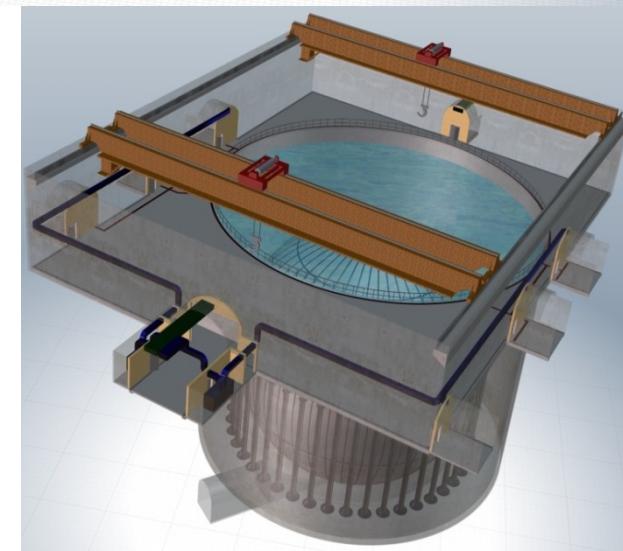
Detector assembly & installation: 2018-2019

Filling & data taking: 2020

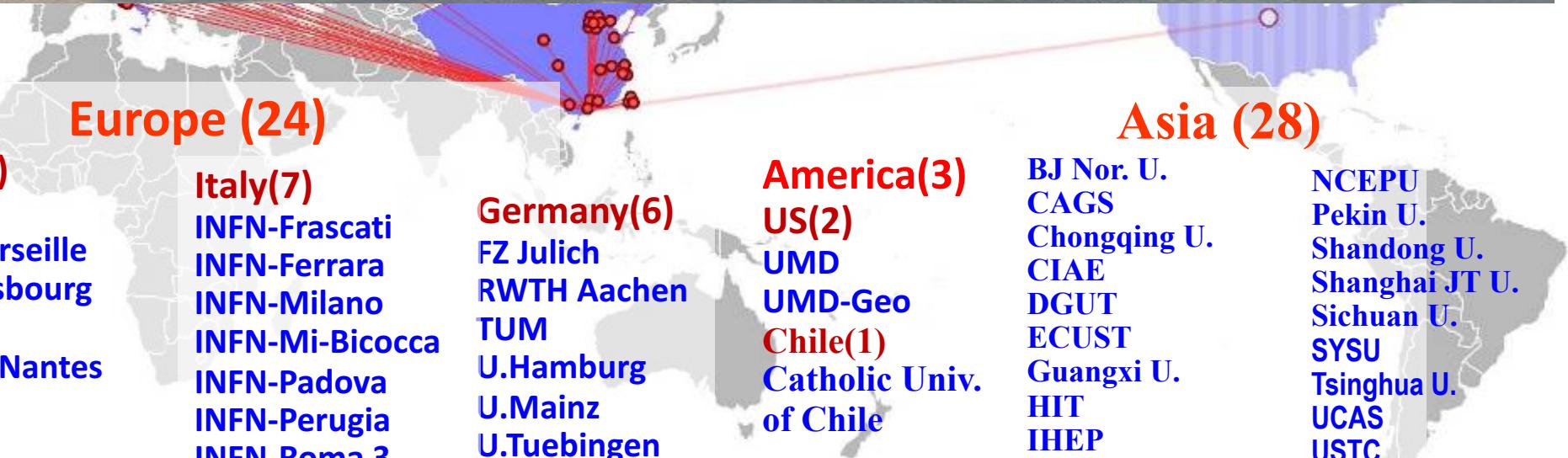
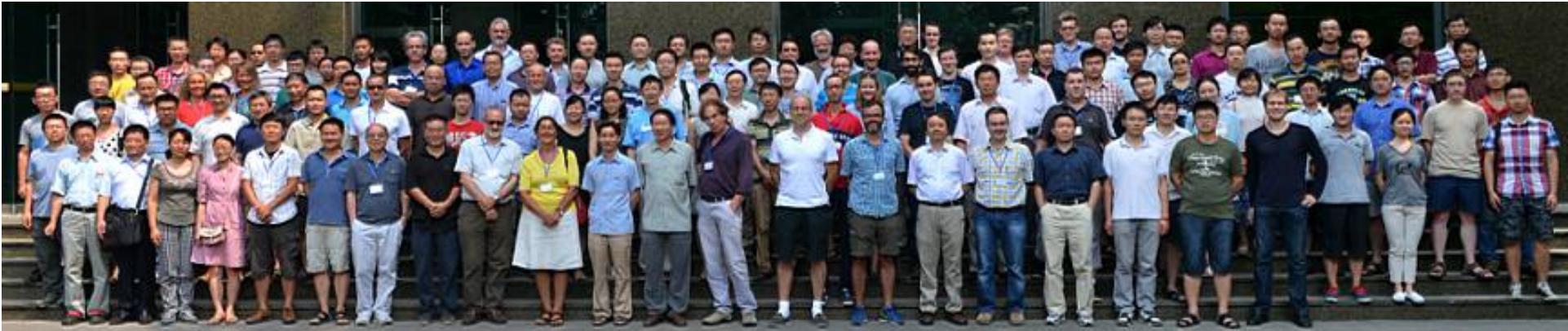
A 600m vertical shaft
A 1300m long tunnel(40% slope)
A 50m diameter, 80m high cavern



Grounding breaking on Jan. 10, 2015



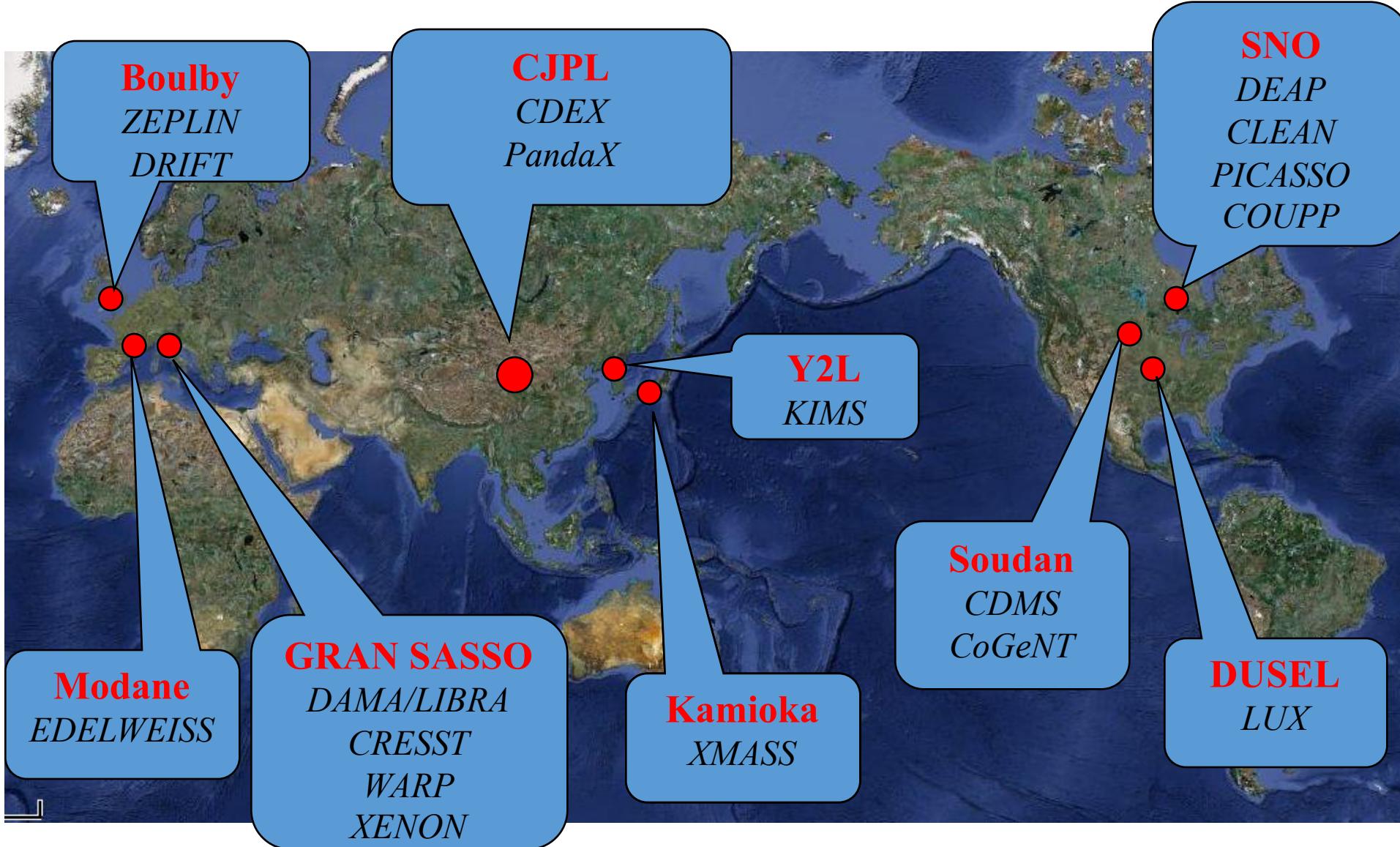
JUNO collaboration



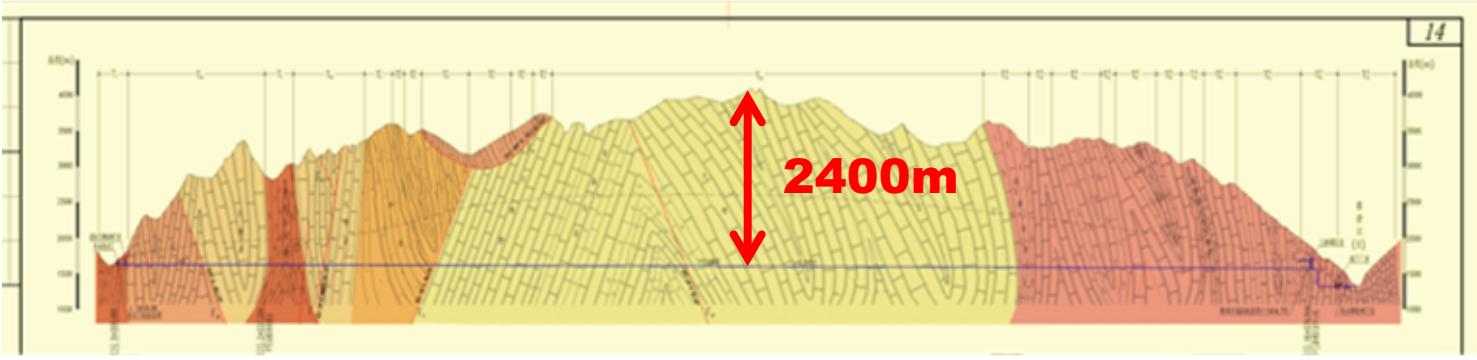
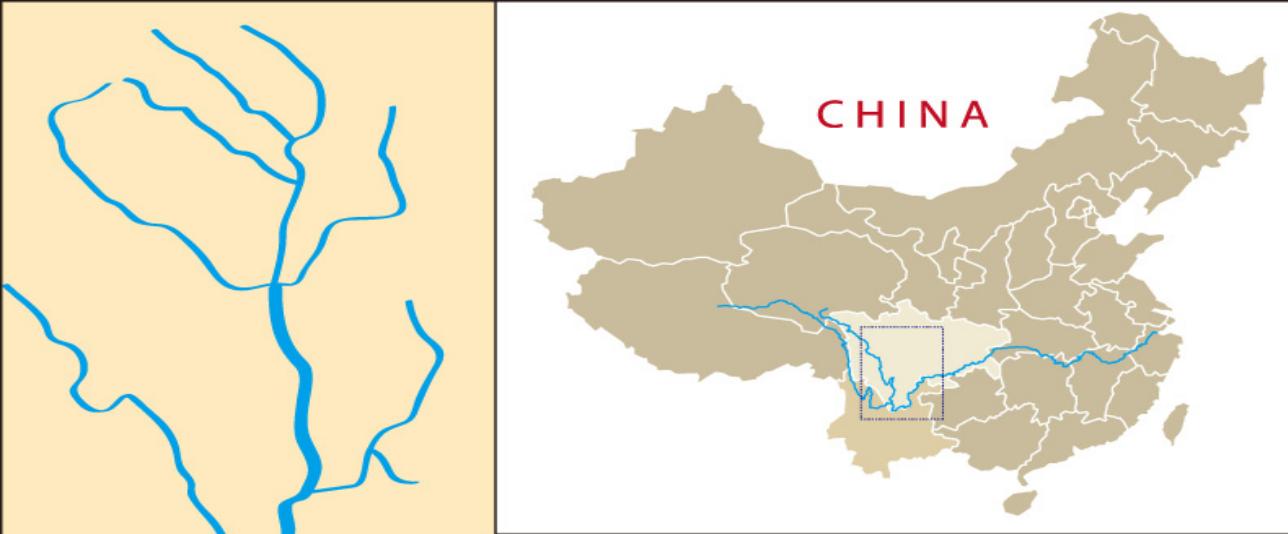
U.S. (2)
University of Michigan
University of Wisconsin
University of Texas at Austin
University of Illinois Urbana-Champaign
University of Minnesota
University of Colorado Boulder
University of Florida
University of California Berkeley
University of Pennsylvania
University of Rochester
University of Washington
University of Wisconsin-Madison
University of Texas at Dallas
University of Wisconsin-Milwaukee
University of Texas at San Antonio
University of Missouri
University of Tennessee Knoxville
University of Louisville
University of Wisconsin-Stevens Point
University of Wisconsin-Whitewater
University of Wisconsin-Platteville
University of Wisconsin-Eau Claire
University of Wisconsin-Oshkosh
University of Wisconsin-Green Bay
University of Wisconsin-La Crosse
University of Wisconsin-Milwaukee
University of Wisconsin-Whitewater
University of Wisconsin-Platteville
University of Wisconsin-Eau Claire
University of Wisconsin-Green Bay
University of Wisconsin-La Crosse

JinPing Underground Lab

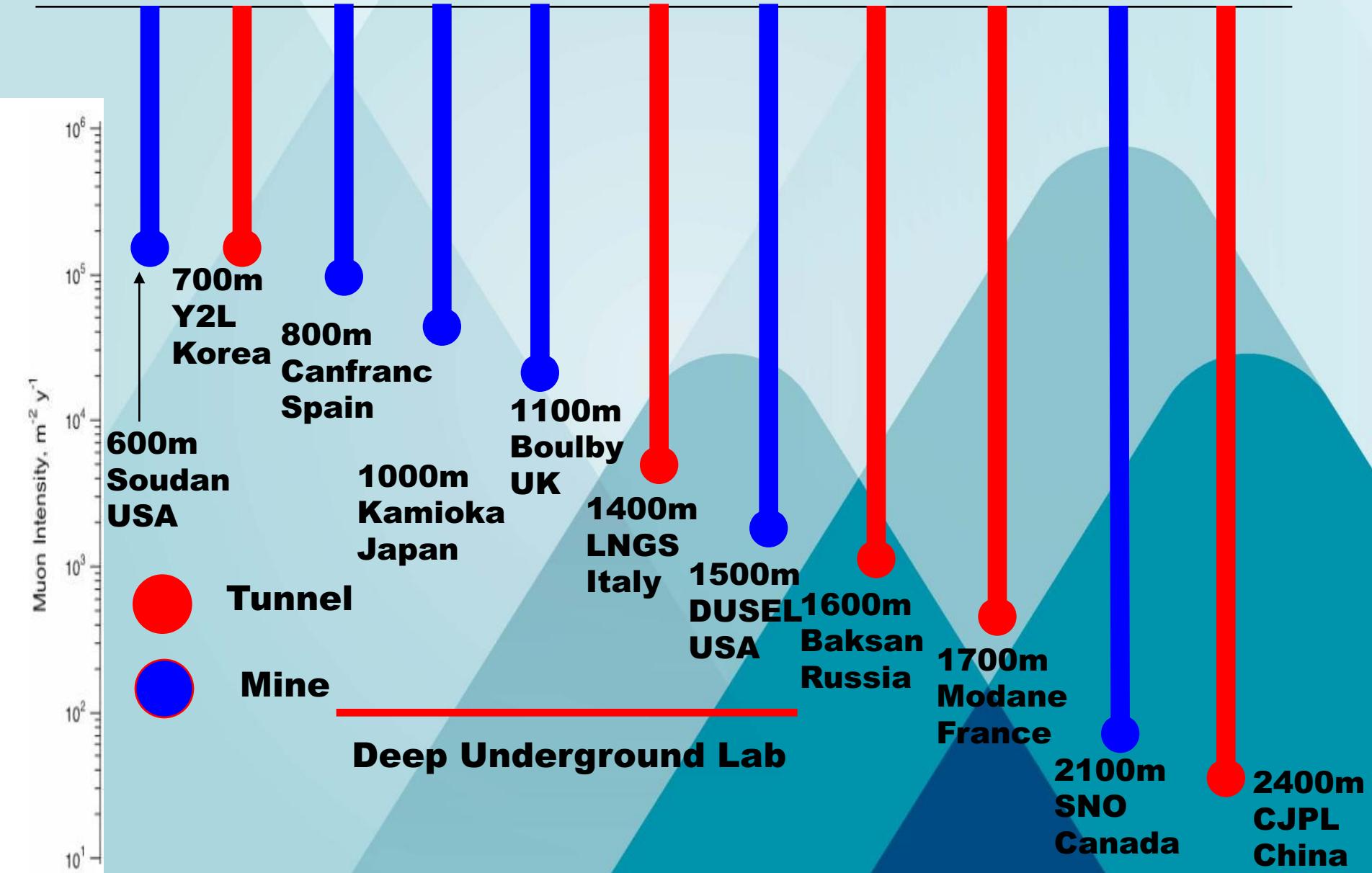
Underground abs in the world



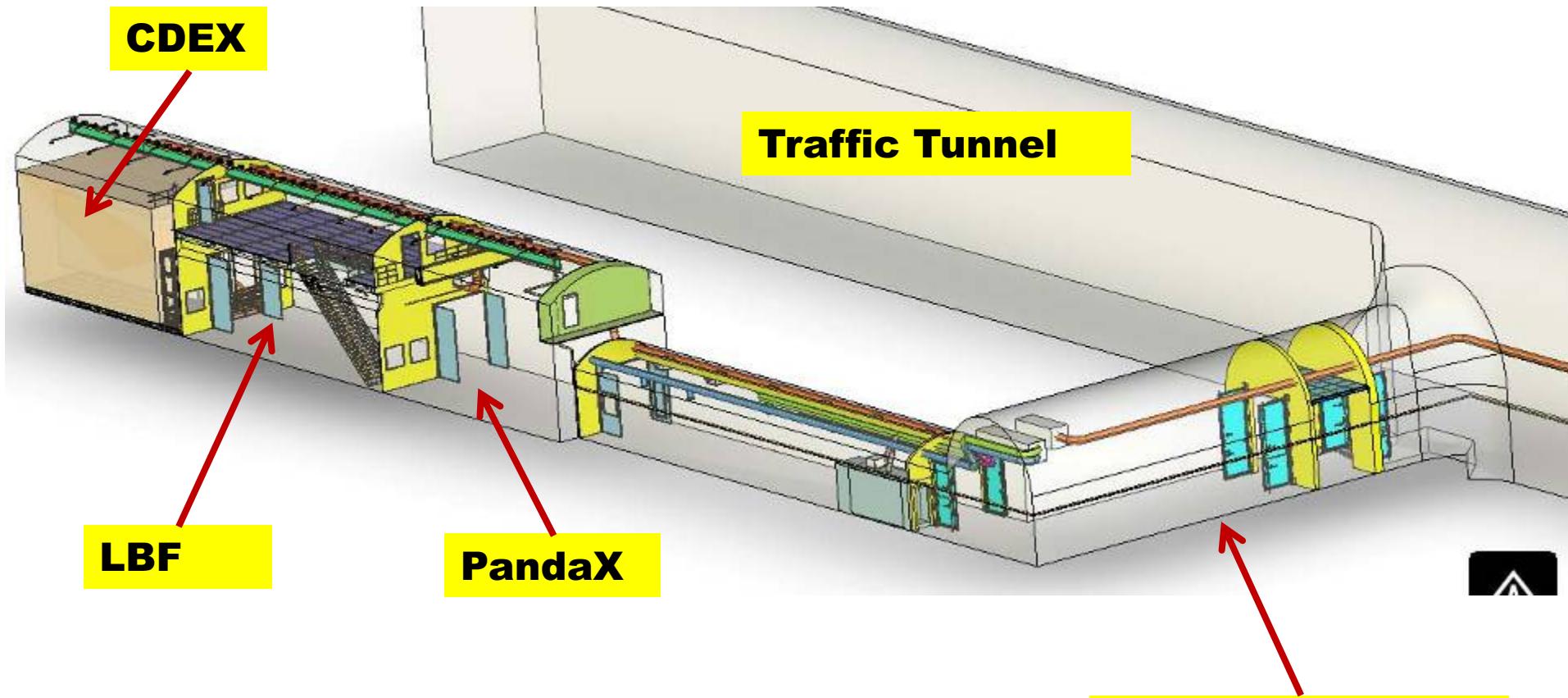
CJPL site



UL in the world(rock overburden)

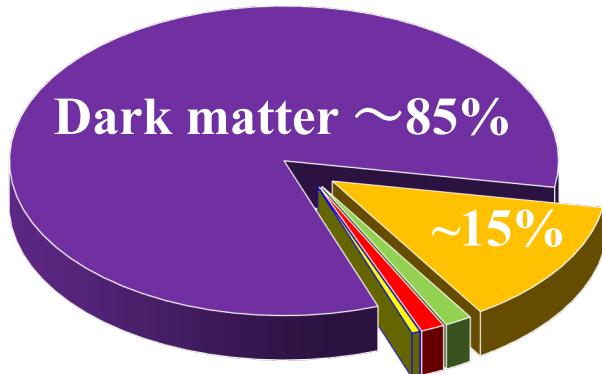


Layout of CJPL-I

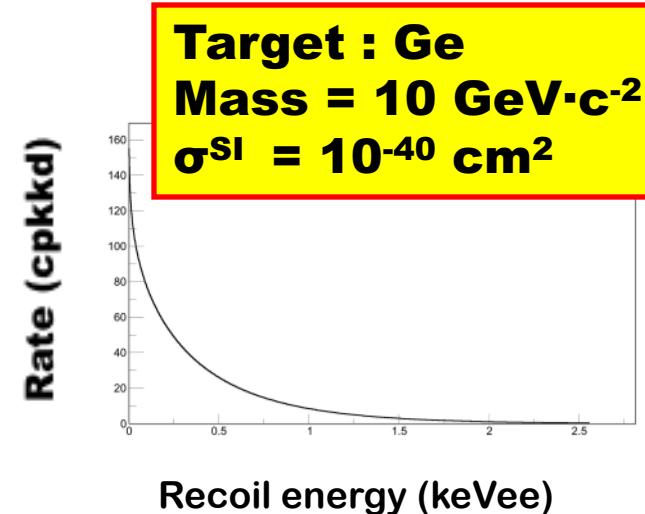
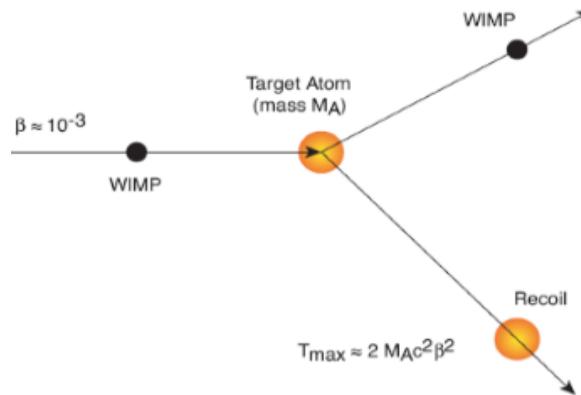


- Total space: 4000 m³
- Main Lab Space: 6.5(W) x 6.5(H) x 42(L)

China Darkmatter Experiment (CDEX)



- Nature of dark matter unknown.
- WIMPs is one kind of well motivated candidate.



Point-contact HPGe detector (PCGe) :
✓ Low energy threshold (~ 100eVee)
✓ Very good energy resolution
✓ Easy to scale up

CDEX target:
Direct detection of low mass dark matter with tonne-scale PCGe array!

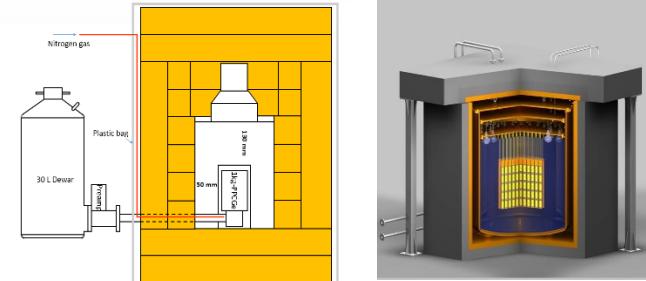
CDEX: China Dark matter EXperiment



Munich Ge workshop, April 25-27

CDEX development stages

- **CDEX-1:** Development of large-mass prototype PCGe detector, data analysis methods, and its background understanding and suppression;
- **CDEX-10:** Performances of HPGe array detector system and its passive/active shielding systems;
- **CDEX-10X:** Fabrication of HPGe, Ge crystal growth and ULB-Cu;
-
- **CDEX-1T:** Multi-purpose experiment for dark matter and double beta decay.



CDEX-1 experiment

1. HPGe detector

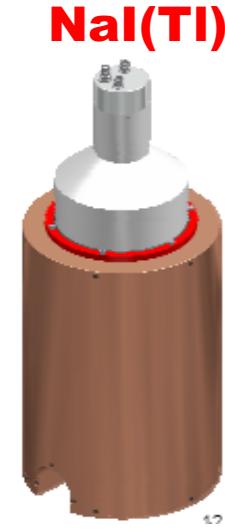
- ✓ Designed and studied the first 1kg-scale p-type point-contact Ge detector (1kg-PPCGe) based on our simulation and experience.
C1A: readout from both P+ point and N+ Li drift layer.
- ✓ Improved the second 1kg-PPCGe. **C1B with lower Eth**

2. NaI(Tl) used as anti-Compton detector

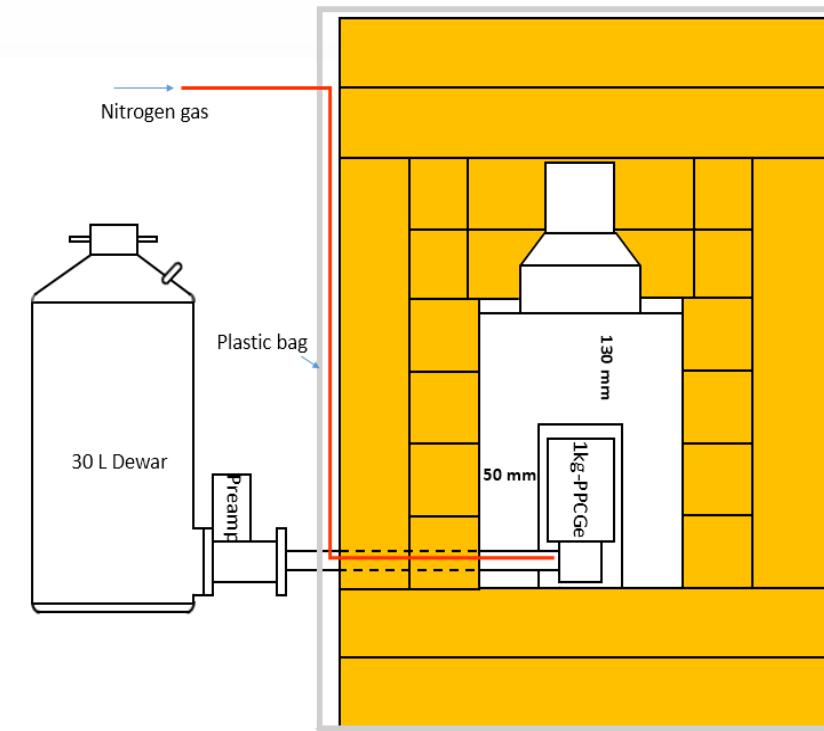
- ✓ C1A without AC
- ✓ C1A with NaI(Tl) AC
- ✓ C1B with NaI(Tl) AC
- ✓ C1 20g Ge + NaI(Tl)



1kg-PPCGe



NaI(Tl)



CDEX in CJPL-I

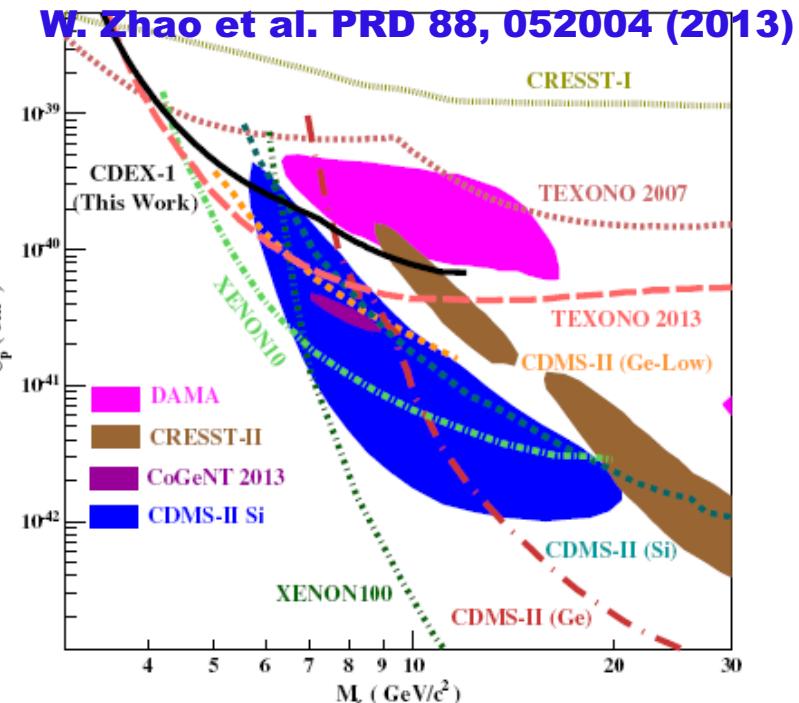
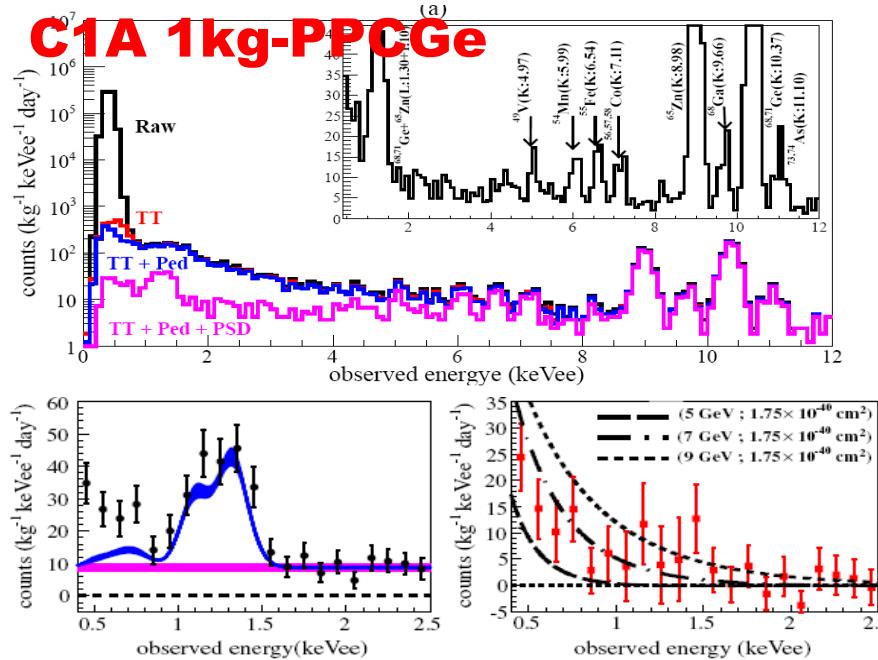


CDEX-1 PE Room

CDEX-1

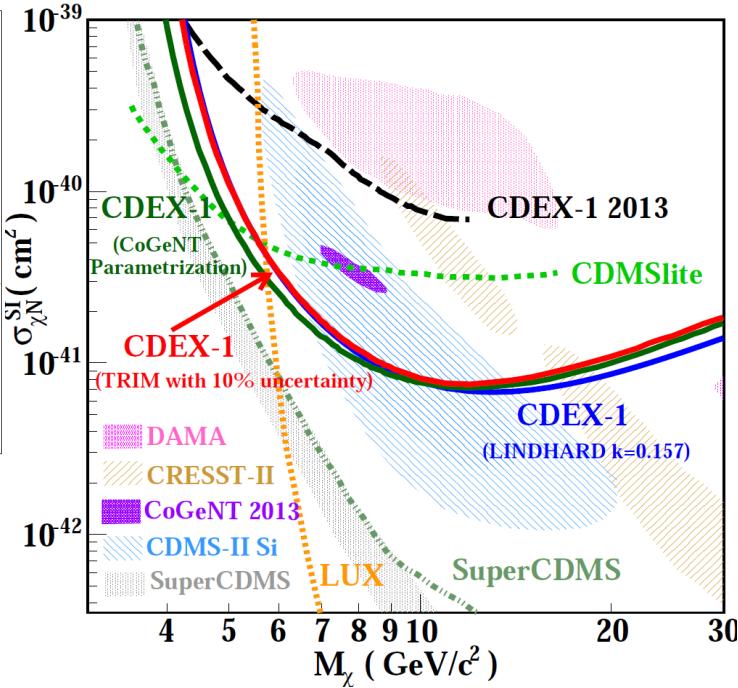
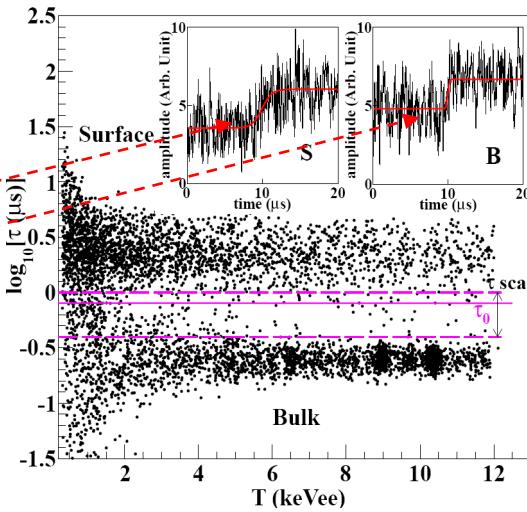
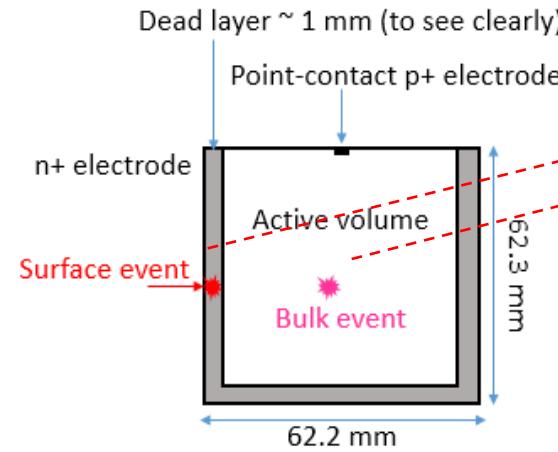


CDEX-1A experiment



- ✓ Ge crystal mass: 994g
- ✓ Energy threshold ~400eV
- ✓ Background level @ 3-5keV: 10cpkdd
- ✓ K/L X-ray peaks from cosmogenic nuclei identified.
- ✓ The first large pPCGe design by CDEX successful;
- ✓ The first dark matter physical result from China.

CDEX-1A upgrade



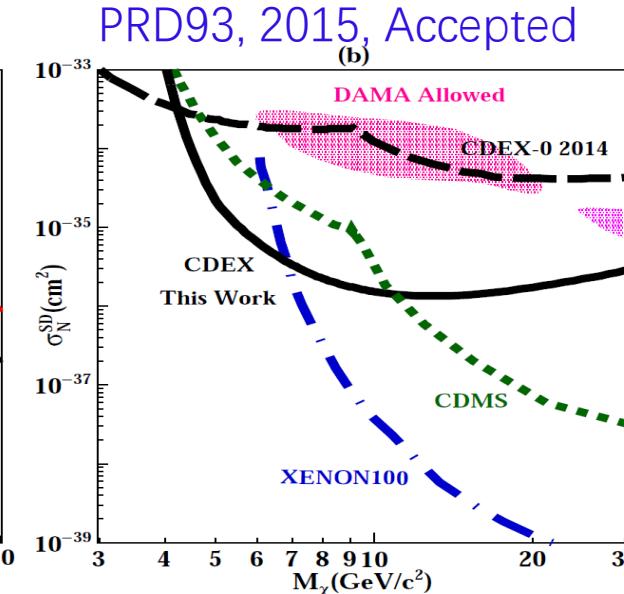
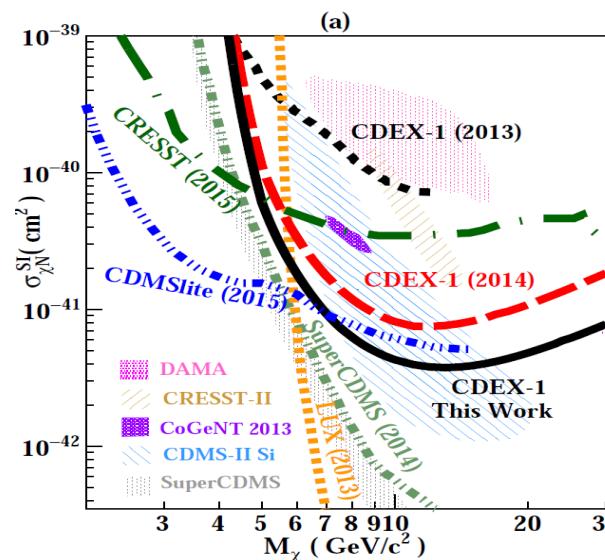
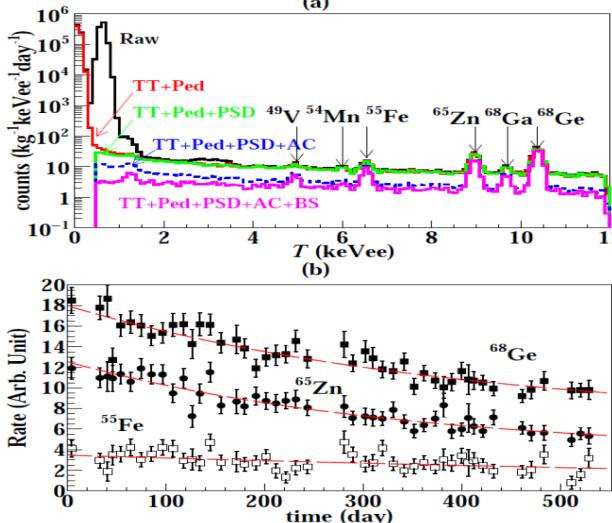
Q.Yue et al., PRD 90(RC) 091701

- **Bulk/Surf Discrimination**
- **C1A 1kg-PPCGe + NaI(Tl)**

- ✓ B/S disc. to get rid of the events with slow rise-time and partial charge collection;
- ✓ 1/3 background lower with/without AC detector;
- ✓ The best sensitivity by PCGe and Exclude the regions favored by CoGeNT.

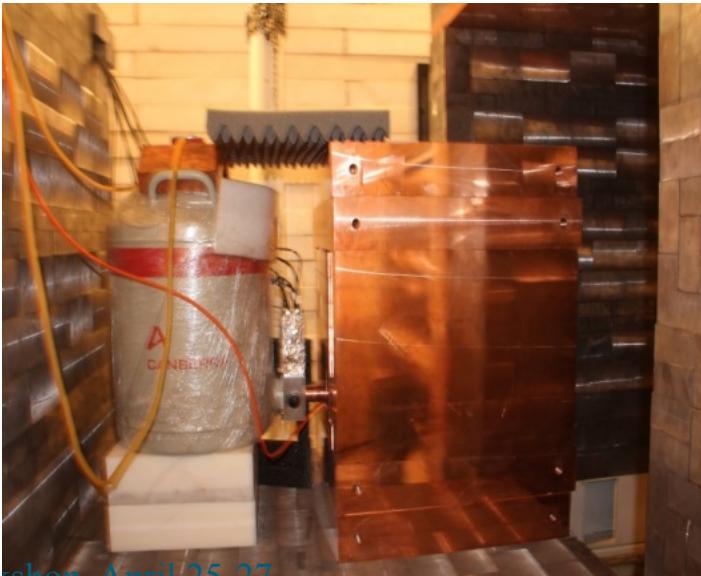
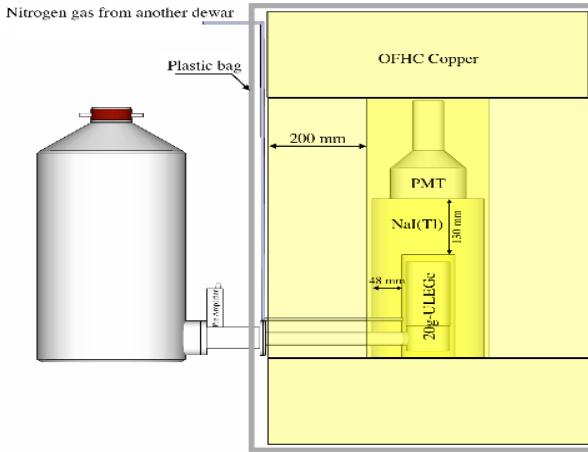
CDEX-1A new results

- CDEX-1A run > 500day, totally $\sim 336 \text{ d}\cdot\text{kg}$ dataset;
- Flat background level decay from ~ 10 to $\sim 4 \text{ cpkkd}$;
- Based on the decay of K/L x-ray peaks from the cosmogenic nuclei, The crystal history may be traced;
- ~ 2 times more sensitive than 2014 result;
- AM analysis with >1 year data going on.

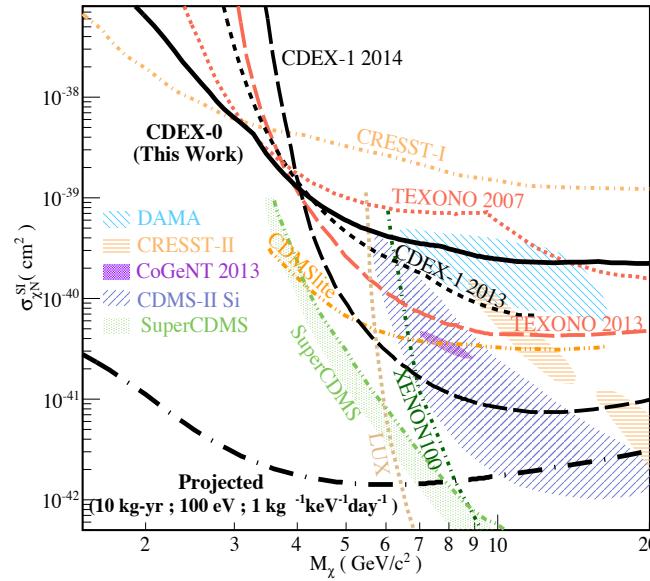


CDEX-0 result from 20g Array detector

CDEX-1 5g*4 + AC detector



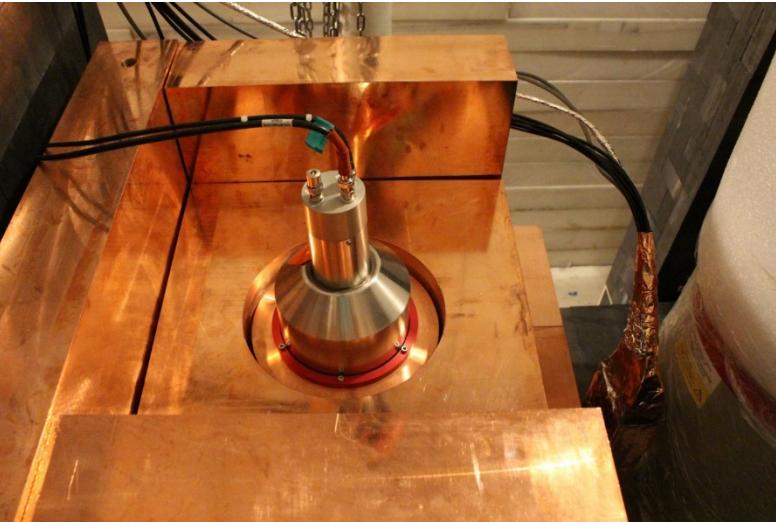
PRD90,032003,2014



Highlight:

- ✓ **177eV lowest ionization energy threshold**
- ✓ **Good direction to lower energy threshold further**

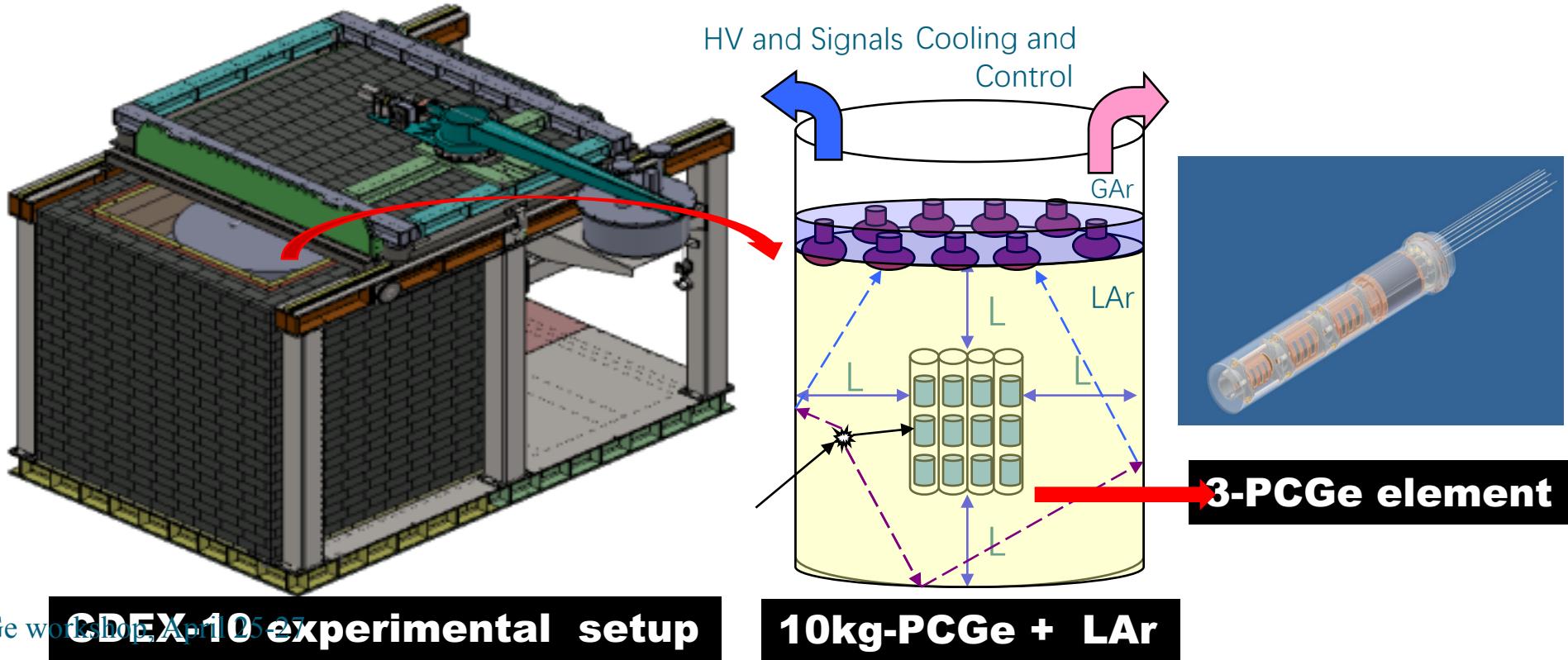
CDEX-1B experiment



- ✓ Ge crystal mass: 1008g
- ✓ Energy threshold <300eV
- ✓ Background level @ 3-5keV: 3cpkkd
- ✓ K/L X-ray peaks from cosmogenic nuclei identified
- ✓ The first physical result under preparation

CDEX-10 experiment

- The important stage towards tonne-scale Ge experiment;
- PCGe array with lower energy threshold: <300eV;
- Performance study of the Ge array detector;
- Feasibility of LAr anti-Compton detector for CDEX.



CDEX-10 Array detectors



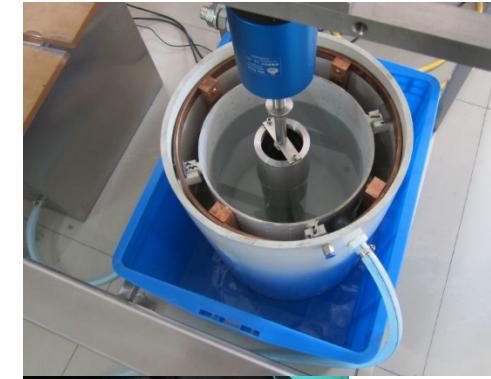
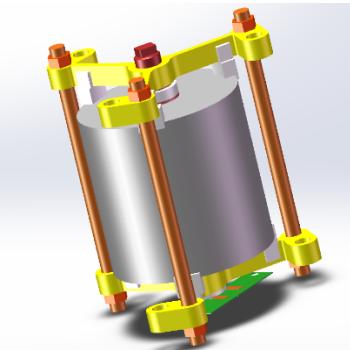
CDEX-10 Array detectors

- 3×3 kg Array detectors;
- $E_{\text{th}} < 300 \text{ eV}$;
- Better resolution in LN₂ than with cooling finger.
- Under testing in CJPL



CDEX-10X

- ✓ Two 0.5kg PCGe with <350eV under preparation by CDEX;
- ✓ Totally new design by CDEX and background control with pure cable, PreAMP substrate and structure materials;
- ✓ Two PreAMP types: JFET and ASIC;
- ✓ ULB-Cu production by CDEX in CJPL this year.



PandaX collaboration

~50 people



Started in 2009

- Shanghai Jiao Tong University (2009-)
- Peking University (2009-)
- Shandong University (2009-)
- Shanghai Institute of Applied Physics, CAS (2009-)
- University of Science & Technology of China (2015-)
- China Institute of Atomic Energy (2015-)
- Sun Yat-Sen University (2015-)
- Yalong Hydropower Company (2009-)
- University of Maryland (2009-)
- Alternative Energies & Atomic Energy Commission(2015-)
- University of Zaragoza(2015-)
- Suranaree University of Technology(2015-)

PandaX experiment

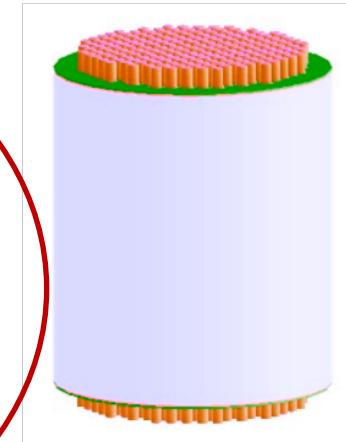
PandaX = Particle and Astrophysical Xenon Experiments



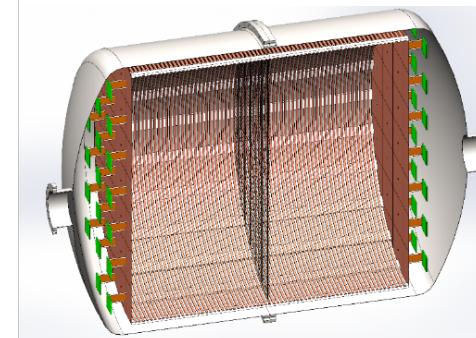
Phase I:
120 kg DM
2009-2014



Phase II:
500 kg DM
2014-2017

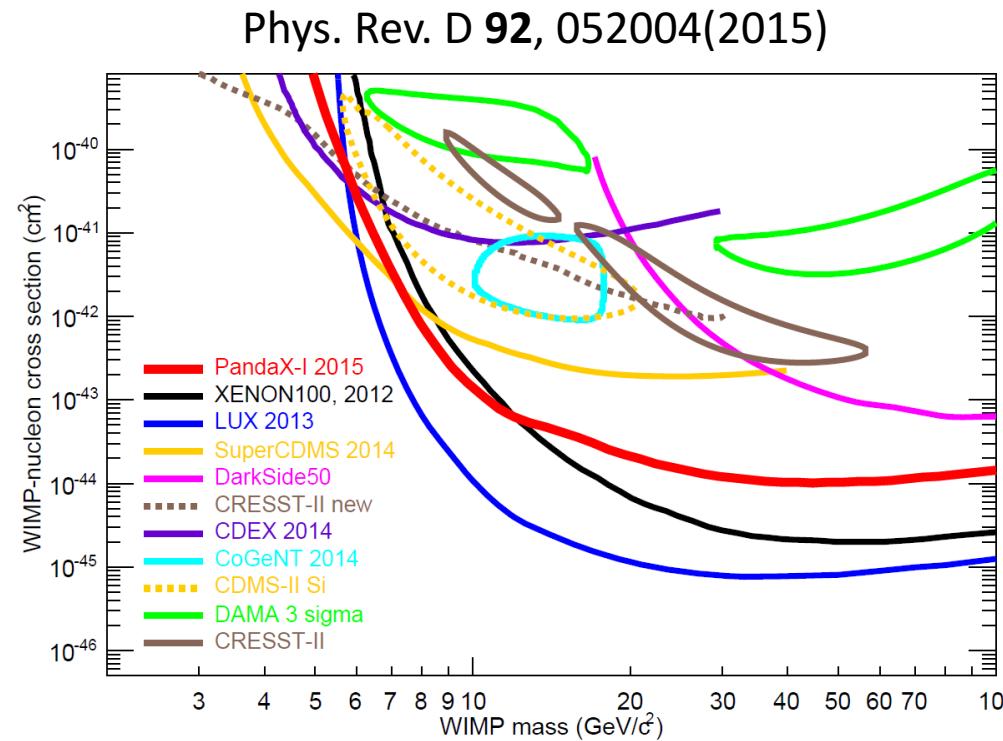
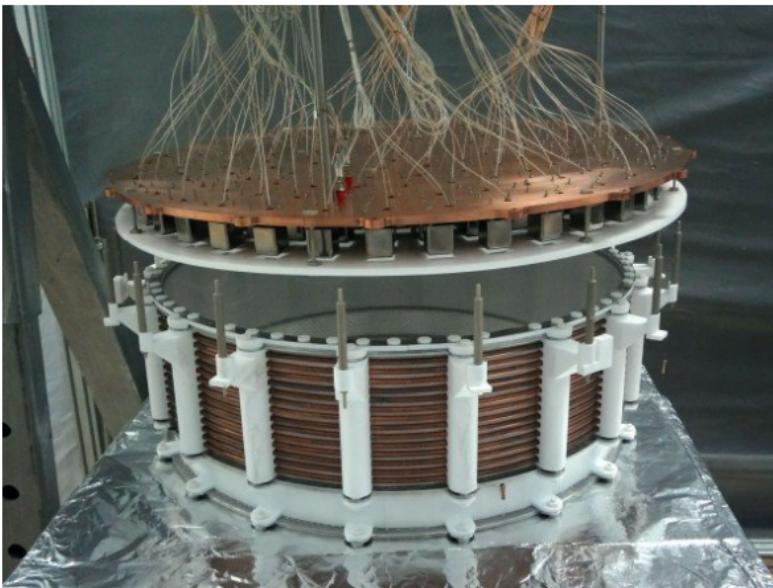


PandaX-xT:
multi-ton DM
future



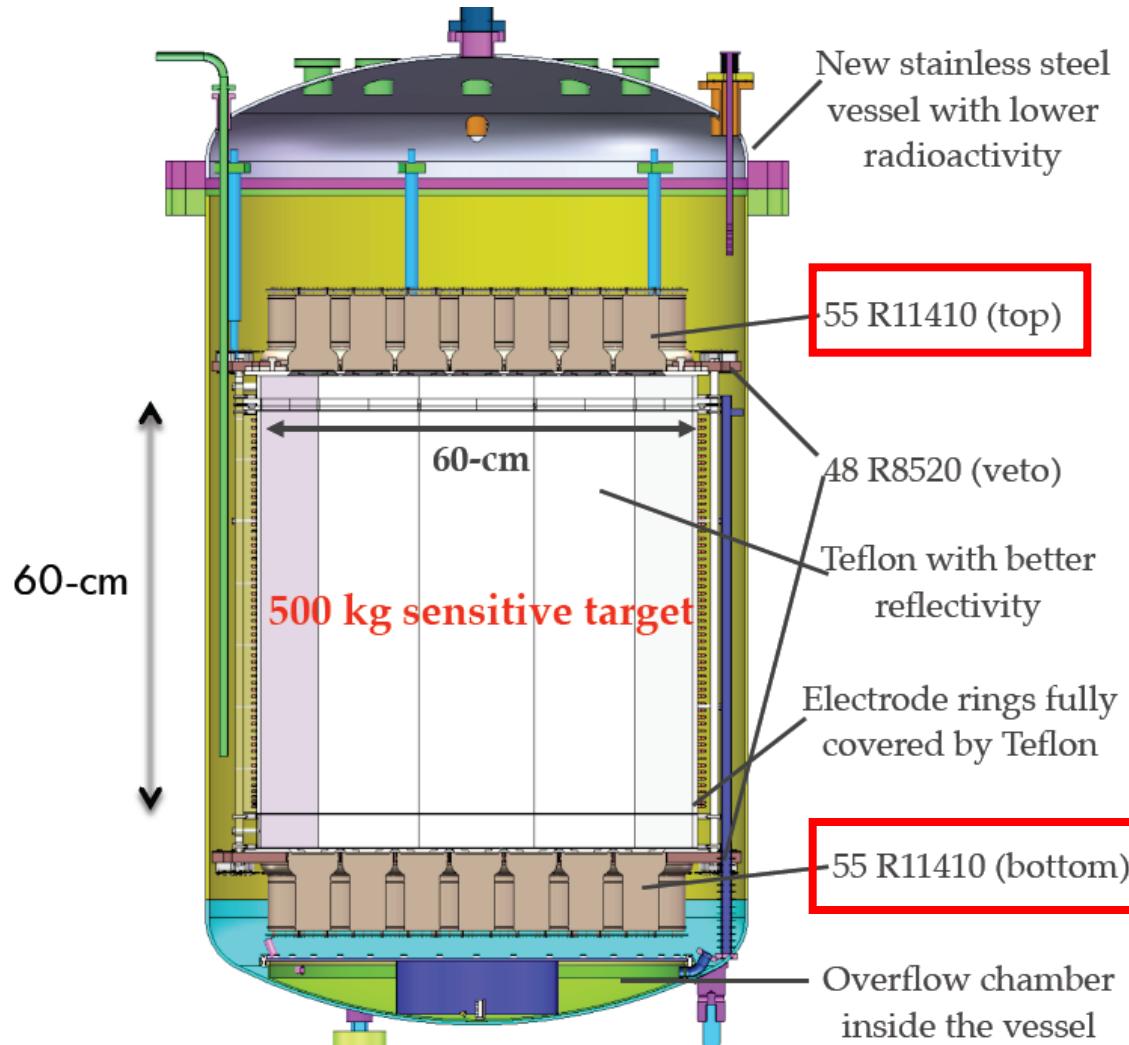
PandaX-III:
200 kg to 1 ton
 ^{136}Xe 0vDBD
future

Final Results from PandaX-I



- Completed in **Oct. 2014**, with $54.0 \times 80.1 \text{ kg-day}$ exposure
- Data strongly disfavor **all** previously reported claims
- Competitive upper limits for low mass WIMP in xenon experiments

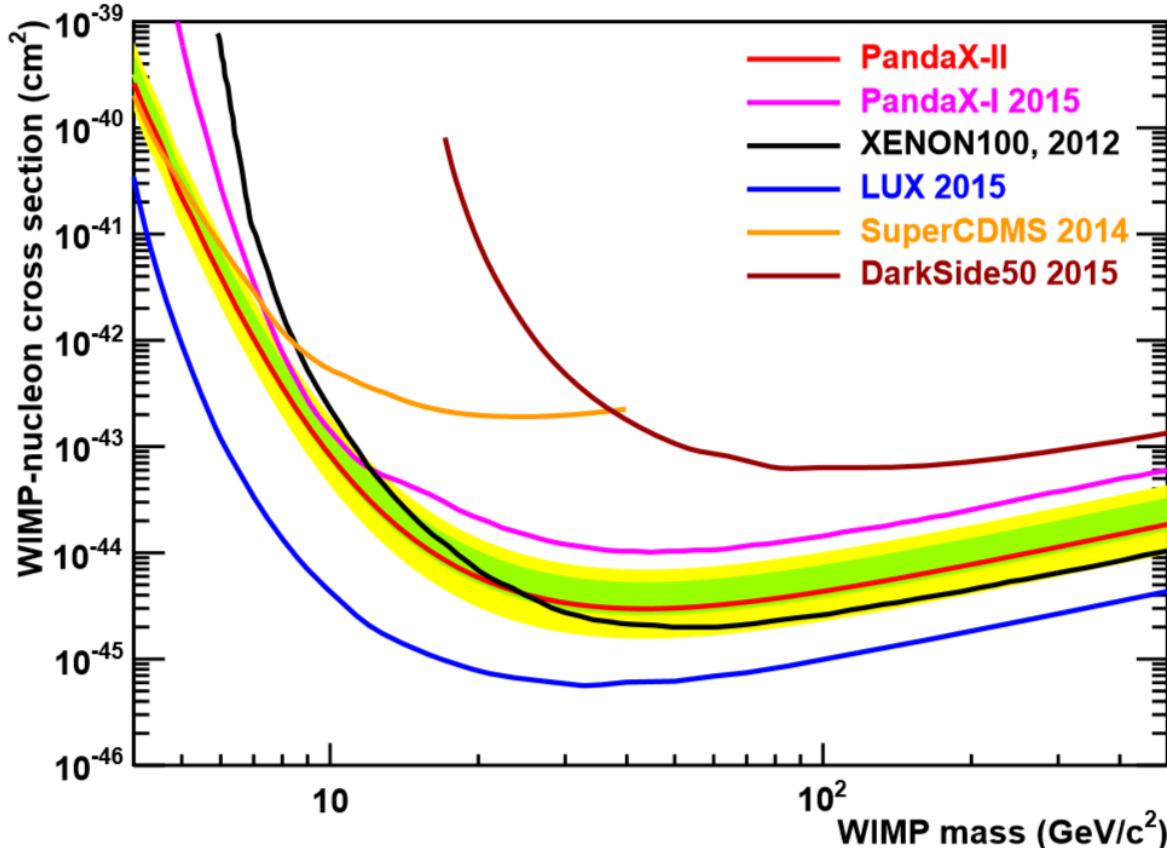
PandaX-II



- New inner vessel with clean SS
- New and taller TPC with brand-new electrodes
- More 3" PMTs and improved base design
- New separate skin veto region

Results from PandaX-II Commissioning

Phys. Rev. D. 39, 122009 (2016), Run 8

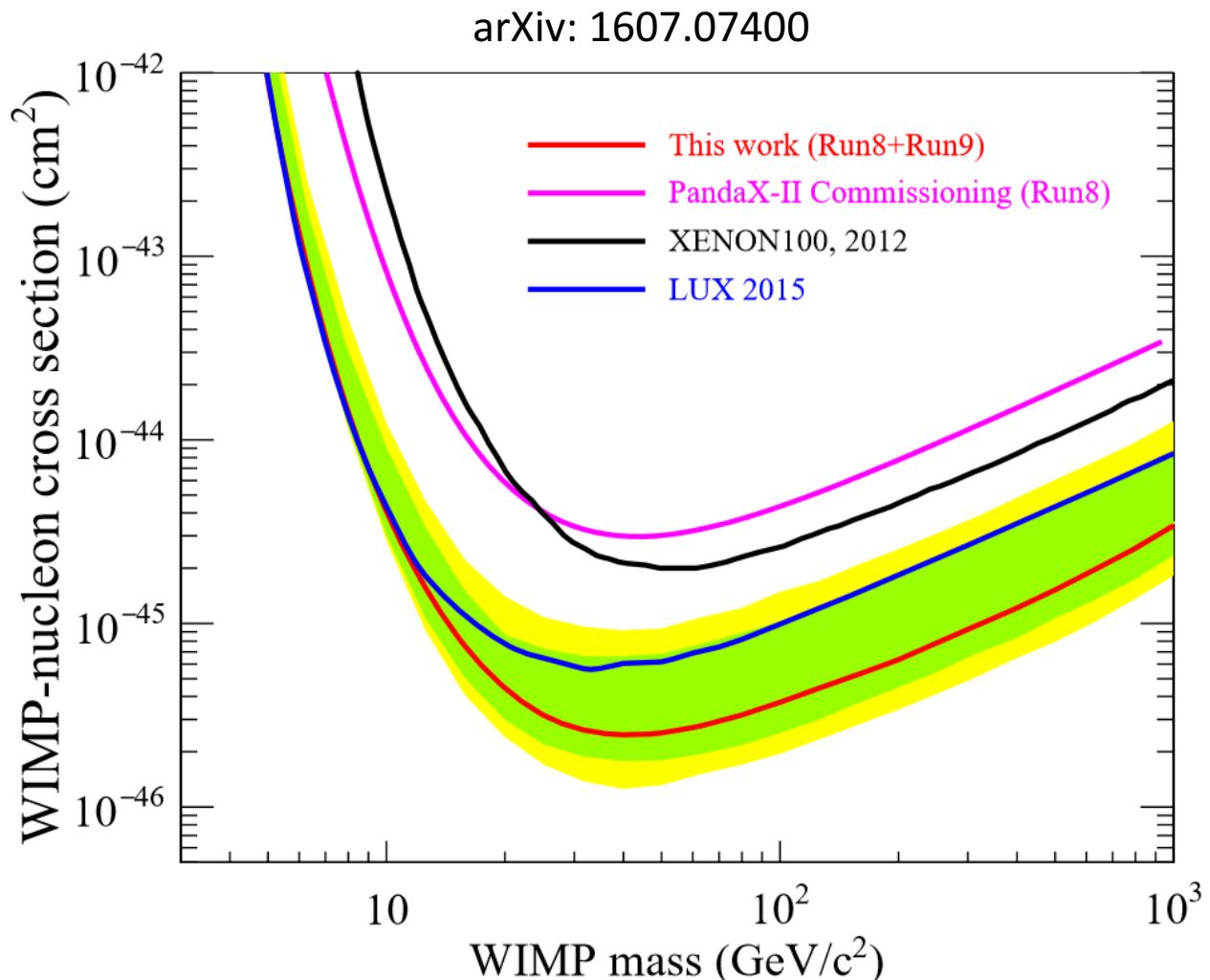


- ❑ Simple counting analysis based on an expected background of 3.2(0.7) evts and 2 observed evts
- ❑ Sizable (x2) difference of using original NEST or tuned NEST to predict DM distribution due to DM acceptance, but within 1σ band
- ❑ Low mass: competitive with SuperCDMS; high mass: similar exclusion limit as XENON100 225-day

Major upgrades in Run 9

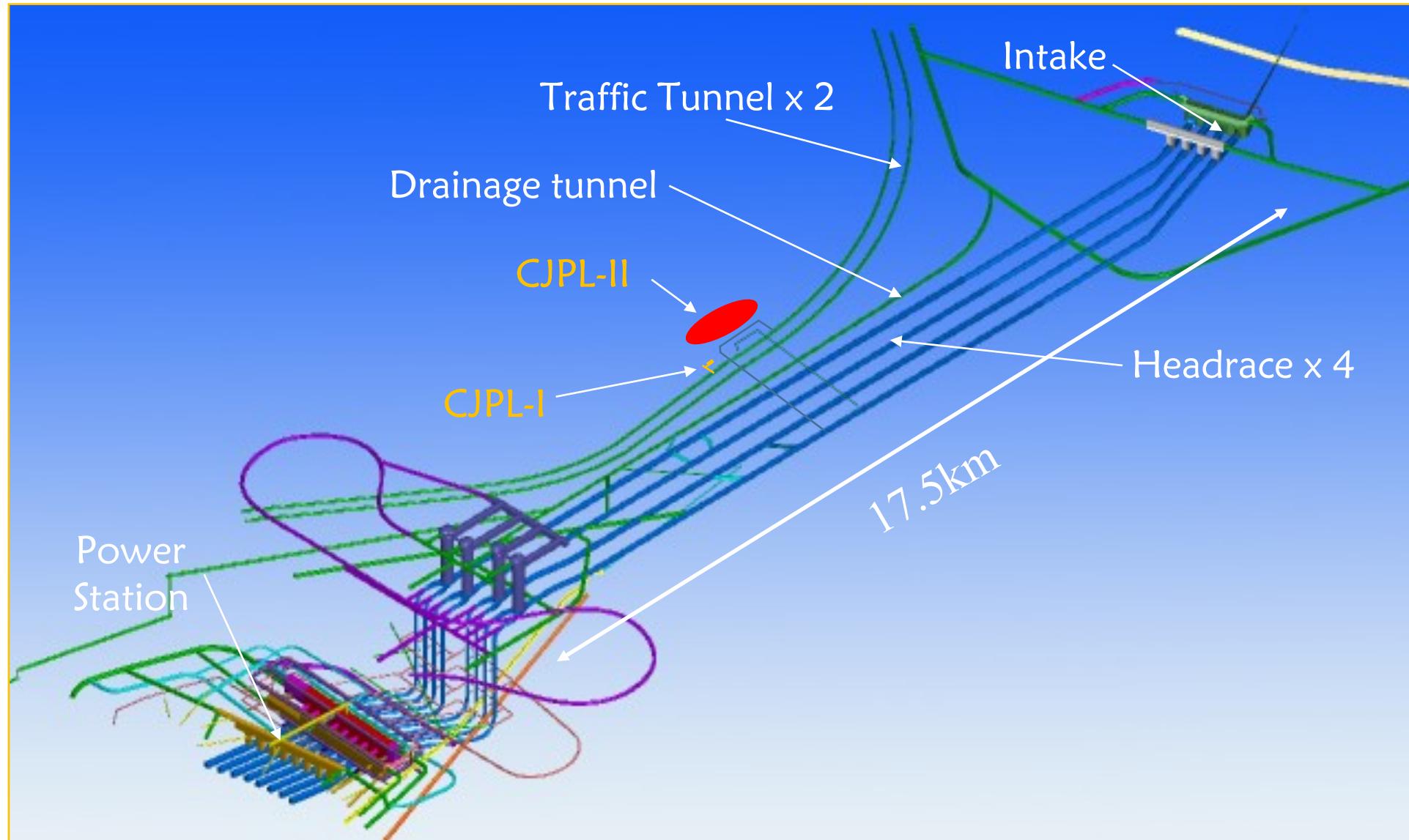
Items	Status in Run 9
Krypton level	Reduced by x10
Exposure	Increased x4 (79.6 vs 19.1 day)
ER calibration	Now have tritium calibration
NR calibration	Statistics x6
Analysis	Improved position reconstruction
Background	Accidental background suppressed more than x3 using BDT

Combined results

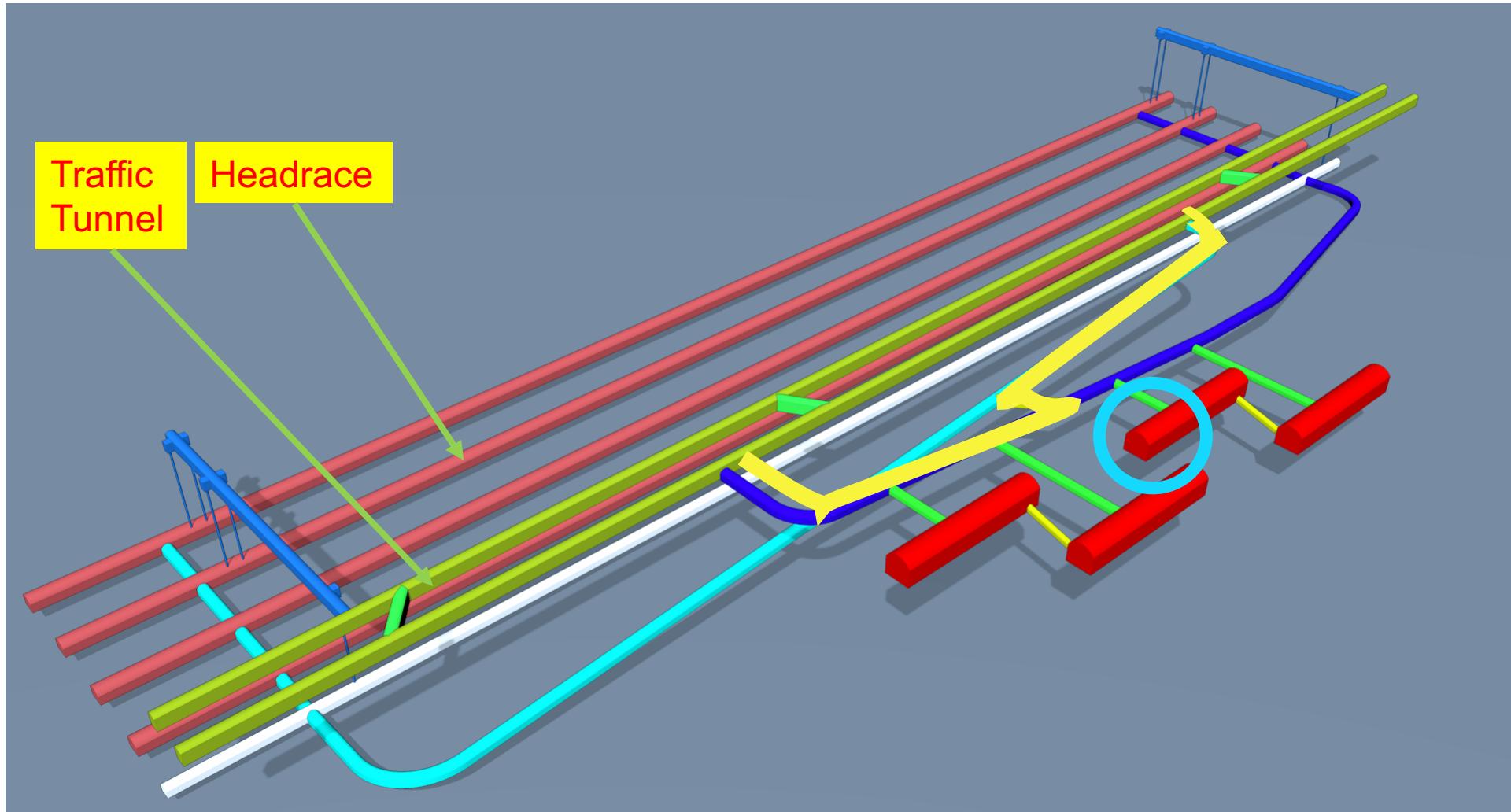


- Combination of Run 8 and Run 9 with an exposure of $3.3 \times 10^4 \text{ kg-day}$.
- Minimum upper limit for isoscalar SI elastic cross section at $2.5 \times 10^{-46} \text{ cm}^2$, more than a factor of 2 improvement compared to the LUX 2015 results

CJPL in Jinping tunnels

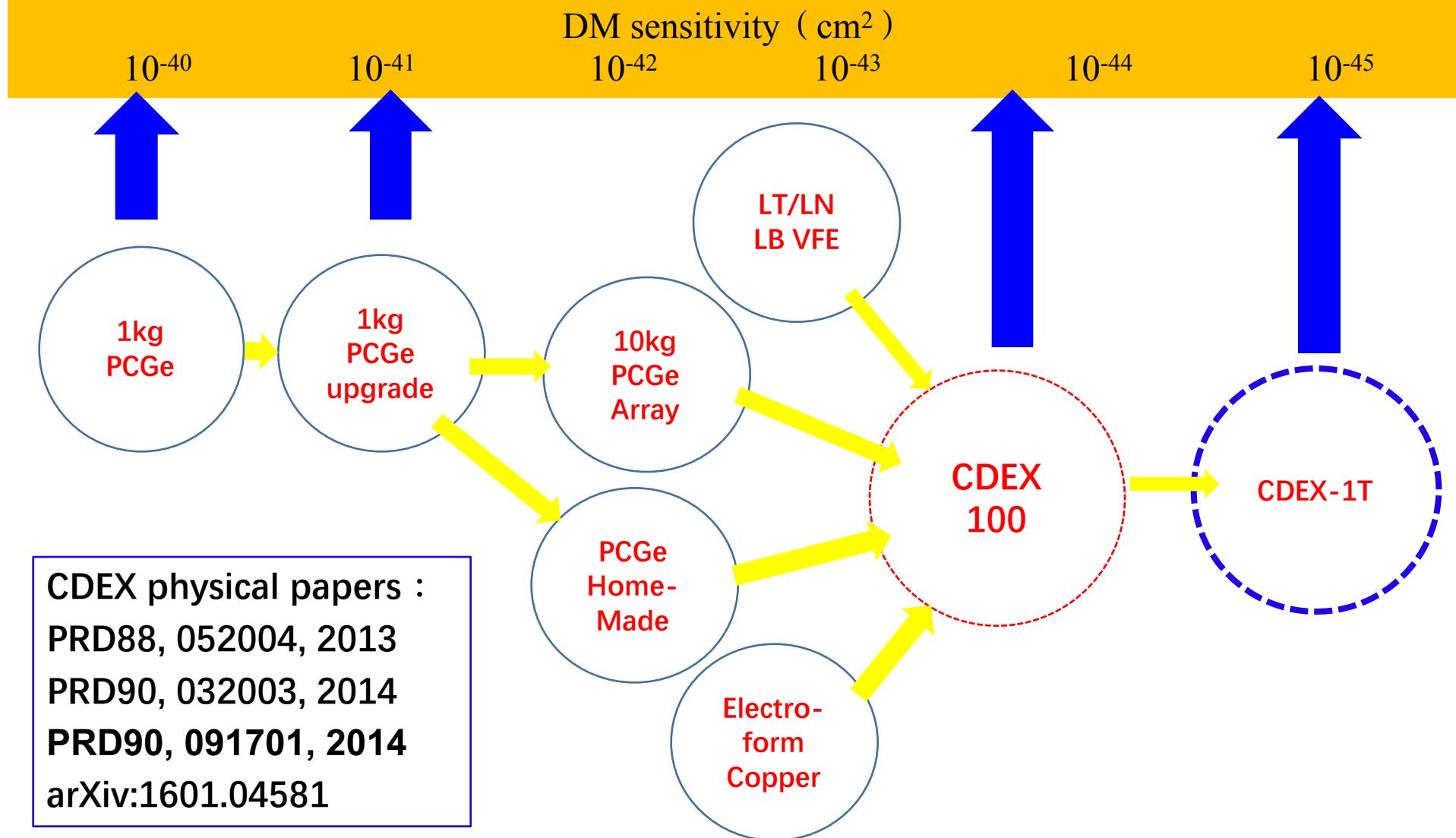


CDEX in CJPL-II



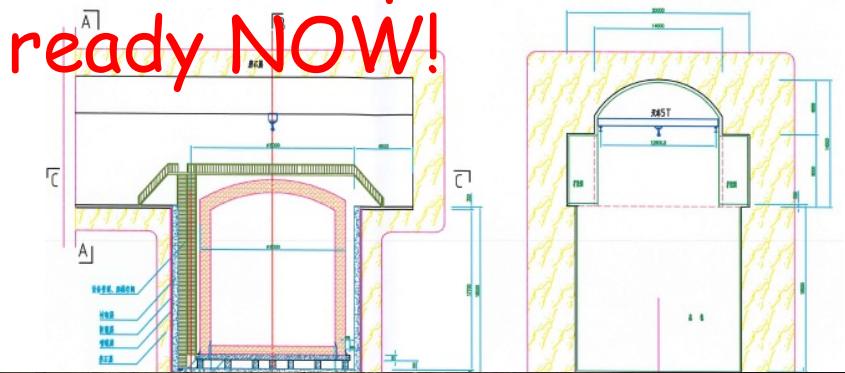
- Four 14m*14m*130m main halls
- Total Volume : ~300K m³

CDEX plan

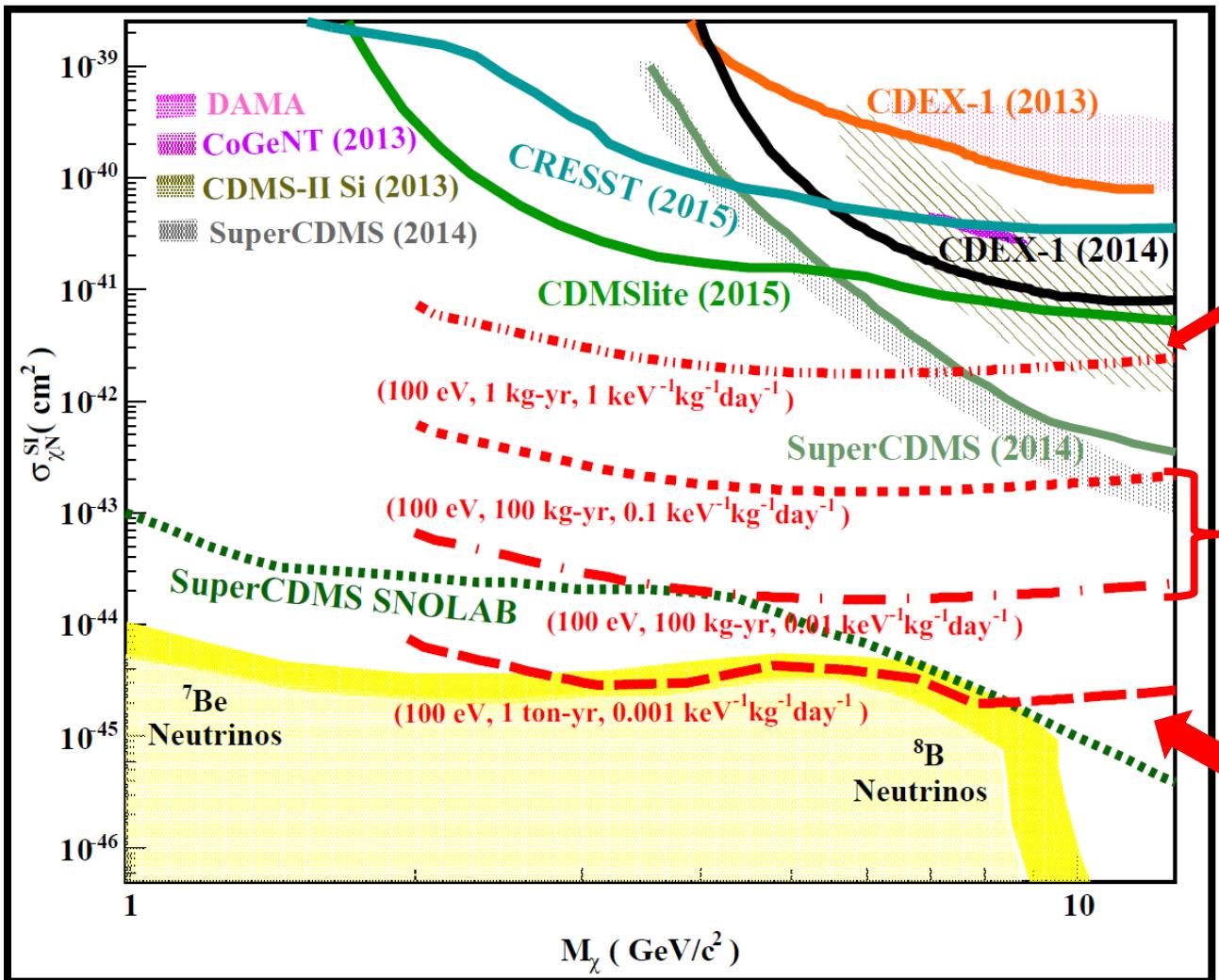


CDEX in CJPL-II

CDEX-1T space
ready NOW!



CDEX-1T Multi-purpose experiment



CDEX-1 : Realistic Reach

CDEX-10/100 :

- Reach @ 0νββ Grade
- Bkg Control
- Spread -- H3 Bkg & Subtraction

CDEX-1ton :

- Reach @ 0νββ Grade
- Bkg & Underground
- Ge-Growth, detector fabrication and ULB-Cu production

PandaX experiment

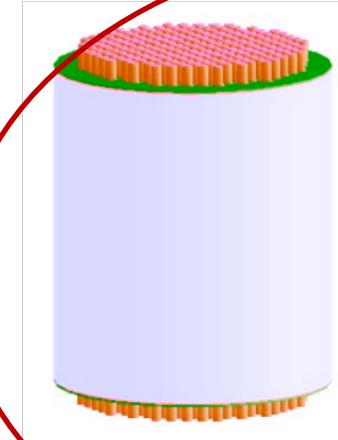
PandaX = Particle and Astrophysical Xenon Experiments



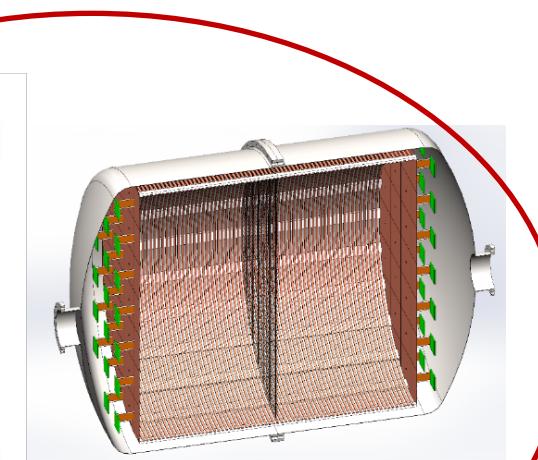
Phase I:
120 kg DM
2009-2014



Phase II:
500 kg DM
2014-2017



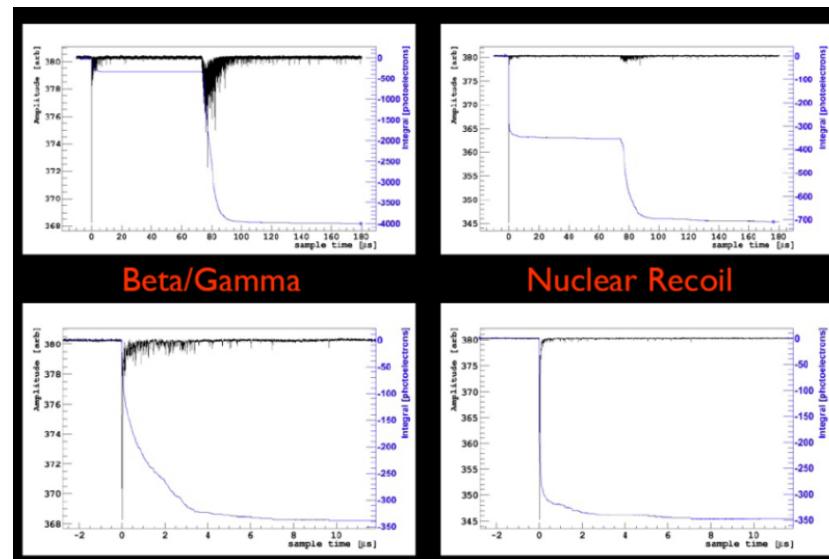
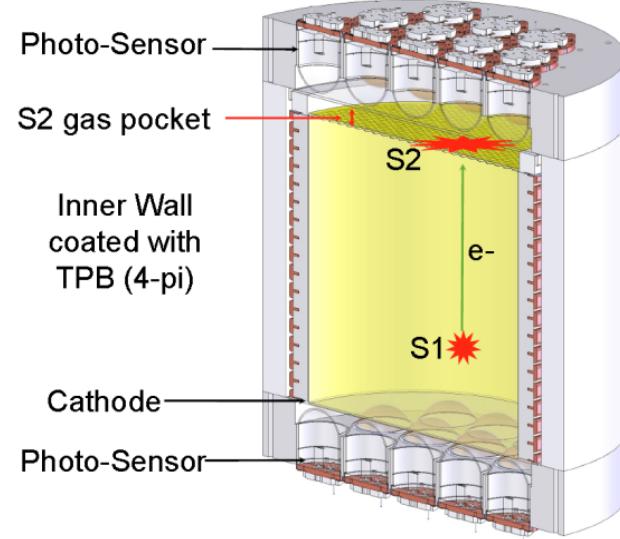
PandaX-xt:
multi-ton DM
future



PandaX-III:
200 kg to 1 ton
 ^{136}Xe 0vDBD
future

液氩探测器优势

- 液氩是优秀的闪烁材料，在两相测量中利用S2/S1可去除本底100–1000倍，这一点和液氩探测器是相同的；
- 液氩发光衰减时间为7纳秒快成分(单态发光)和1.6微秒慢成分(三态发光)，对于中子信号和光子信号，两种发光成分的比例是不同的，因而从波形上可以识别是哪一种事例。对于多于60光电子的光强，实验上证明了液氩具备 10^8 的伽玛-中子分辨能力；



液氩探测器能够测出更清楚的暗物质“脸谱”

液Ar与液氙探测器确定WIMP特性的互补

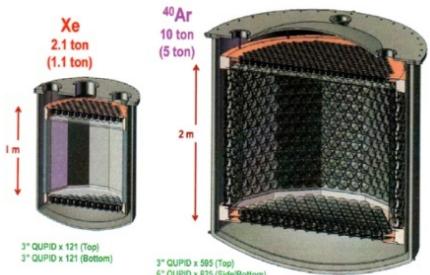


Fig. 1. Main parameters of the 1 ton/5 ton (fiducial) G2 system.

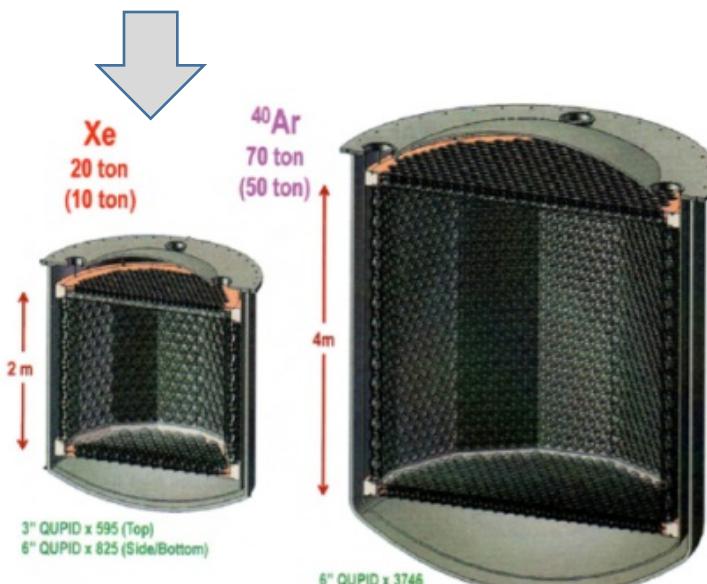


Fig. 2. Main parameters of the 10 ton/50 ton (fiducial) G3 system.

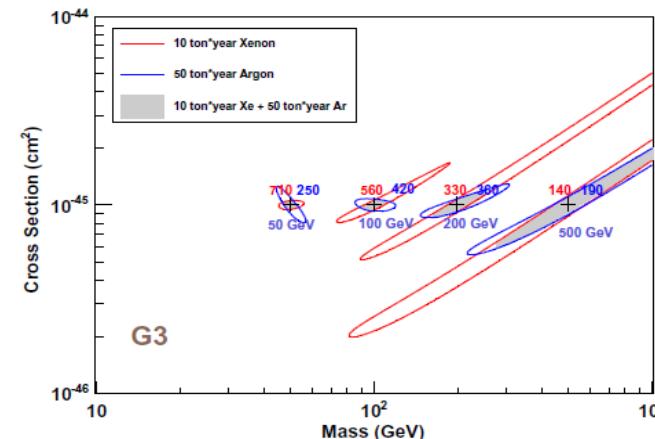
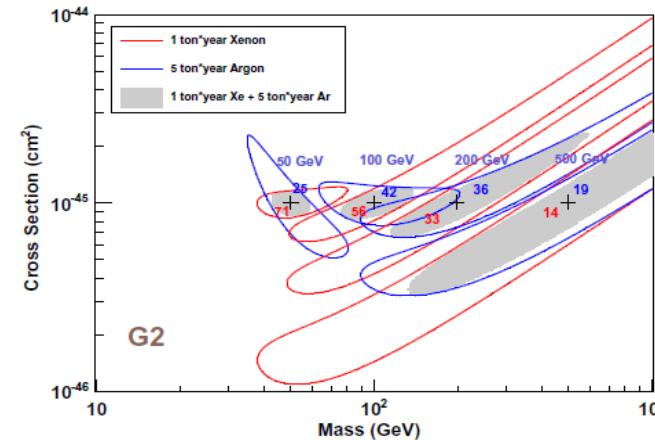
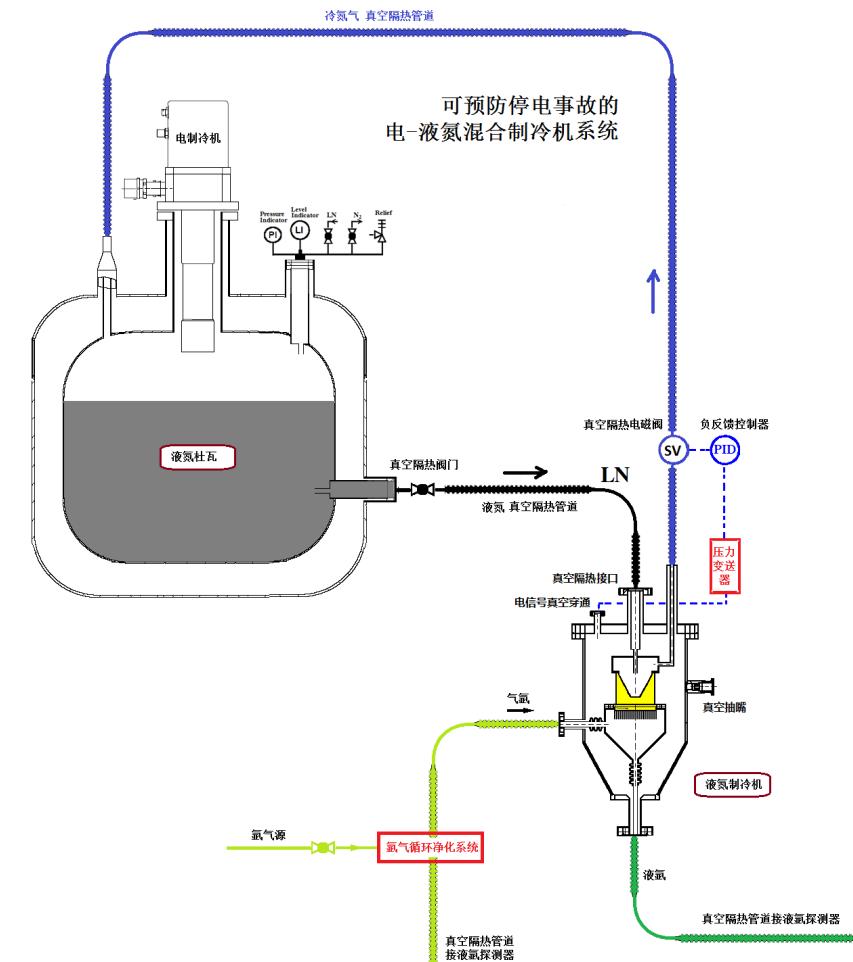
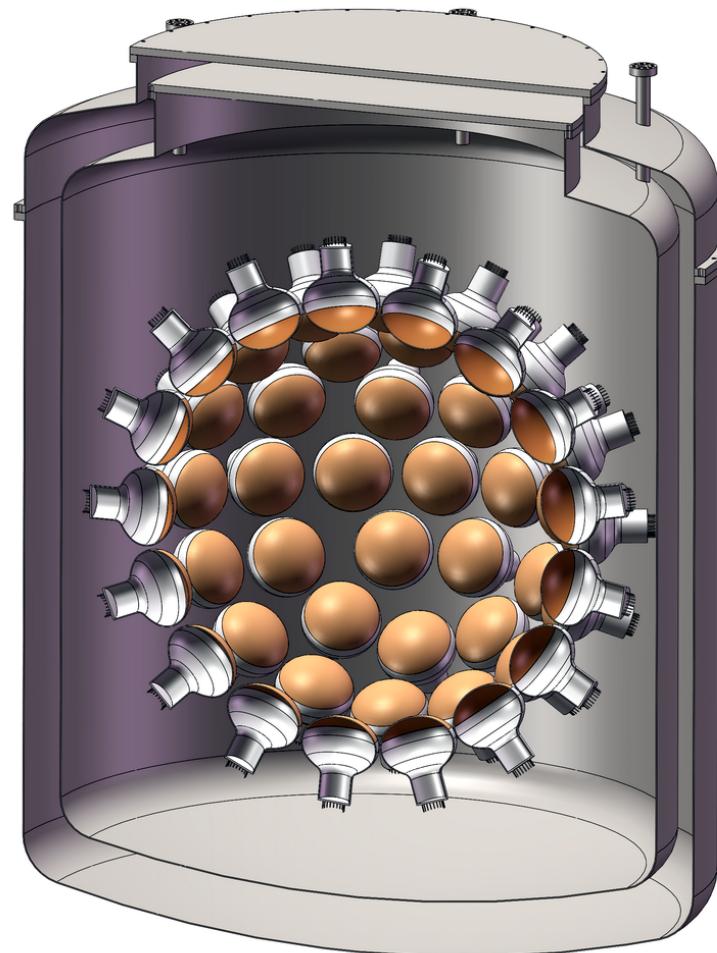


Fig. 12. Examples of overlap of 1σ Xe and Ar contours for WIMP masses 50–500 GeV (upper) cross-section 10^{-45} cm^2 with G2 system: 1 ton-y Xe & 5 ton-y Ar or cross-section 10^{-46} cm^2 with G3 system: 10 ton-y Xe & 50 ton-y Ar or cross-section 10^{-47} cm^2 with G4 system: 100 ton-y Xe & 500 ton-y Ar. (lower) cross-section 10^{-45} cm^2 with G3 system: 10 ton-y Xe & 50 ton-y Ar or cross-section 10^{-46} cm^2 with G4 system: 100 ton-y Xe & 500 ton-y Ar.

高能所液氩探测器实验室

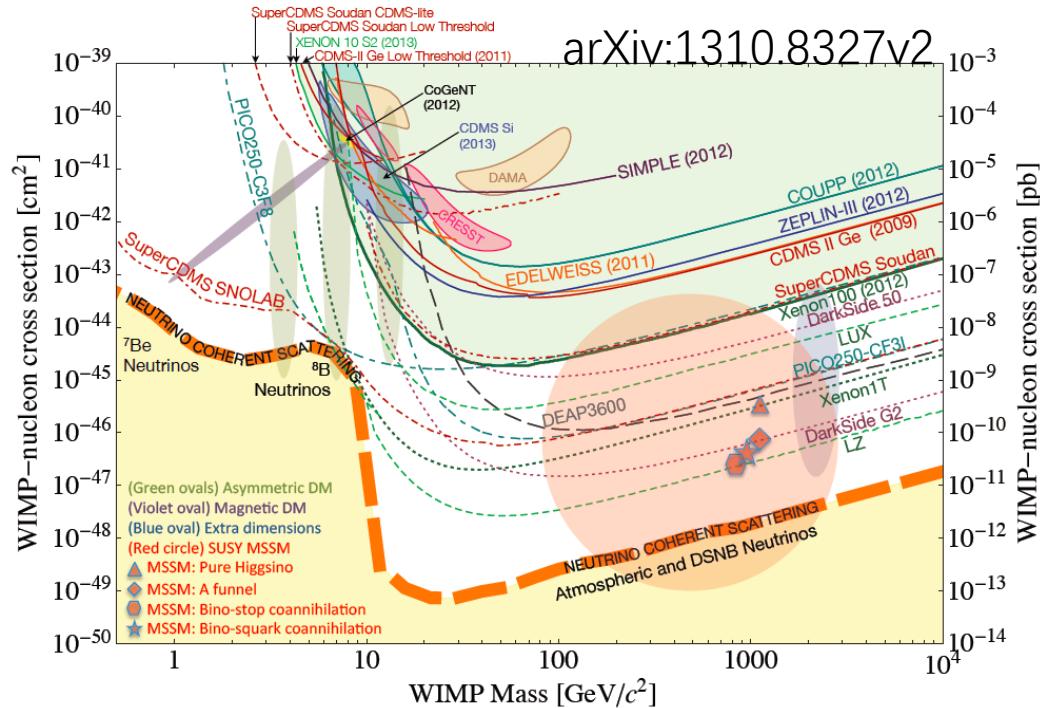


目前R&D: 1吨液氩探测器示意图



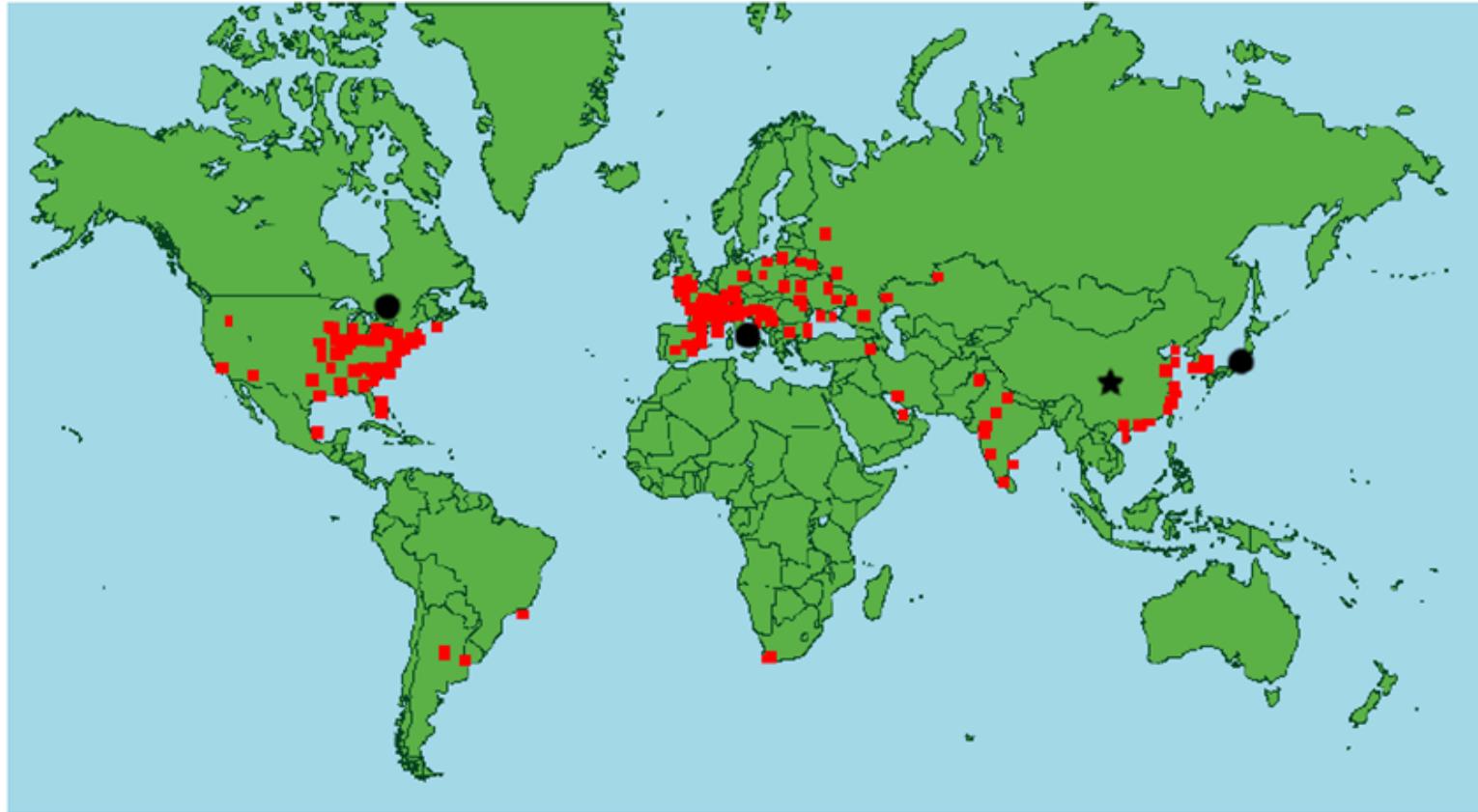
- 左图：吨级单相液氩探测器示意图，
- 右图：电-液氮混合制冷机概念设计图

目标: 锦屏液氩暗物质探测



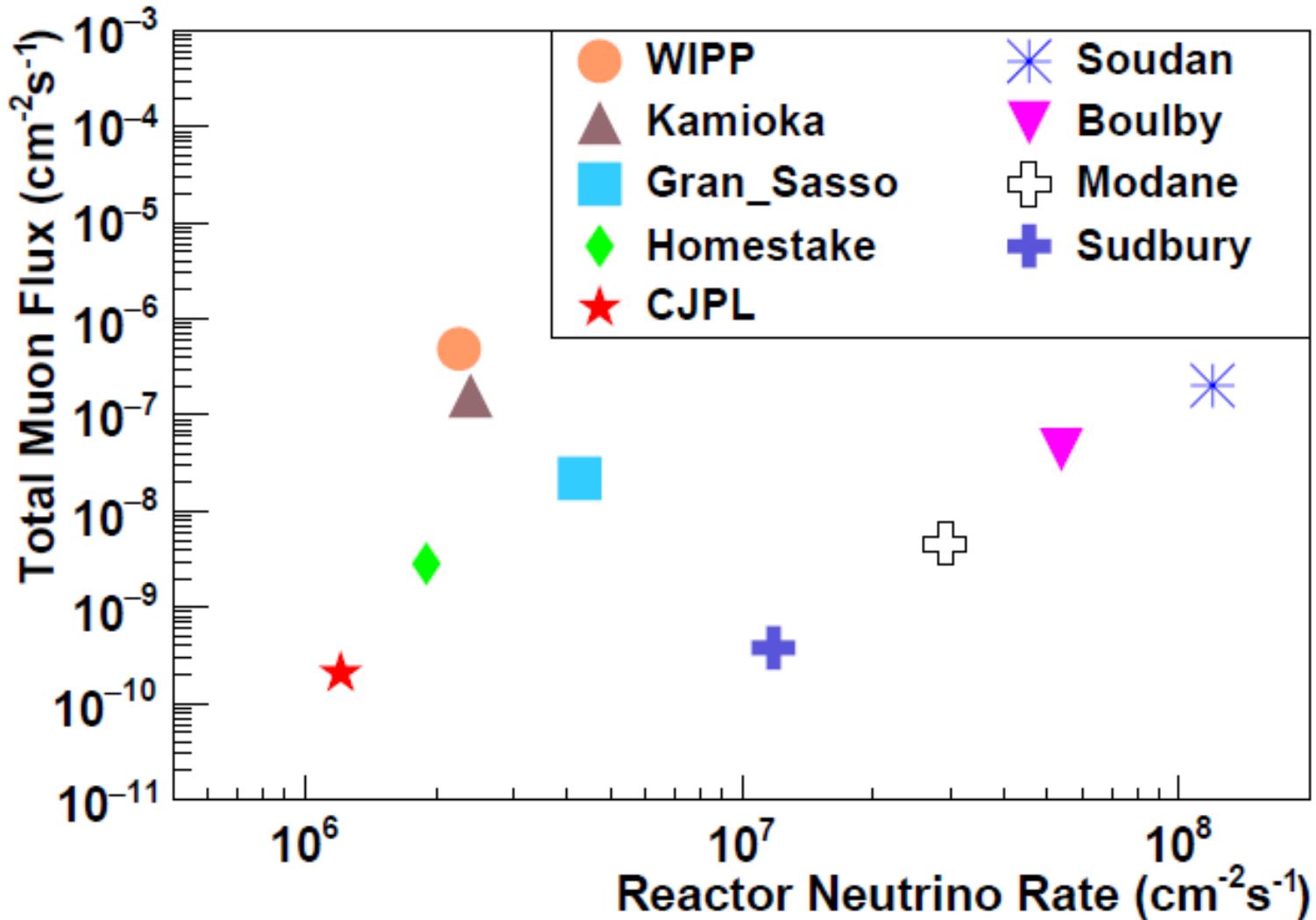
- 目前暗物质探测的灵敏度仍在提高，最终目标是实现暗物质的“终极探测”，达到中微子本底极限，这需要几百吨·年的曝光量。
- 氩是惰性气体中价格最便宜，最易做成液体探测器的一种材料，已运用于加速器实验的量能器中，也是中微子探测和质子衰变探测的重要媒介。

反应堆本底



距最近的反应堆1200 km，低反应堆本底

优良的地下实验室



太阳标准模型(恒星演化模型)的四个要素: John Bahcall

1. 太阳供能方式:

质子-质子聚变链 pp

碳氮循环 CNO

2. 太阳内能量的传输机制:

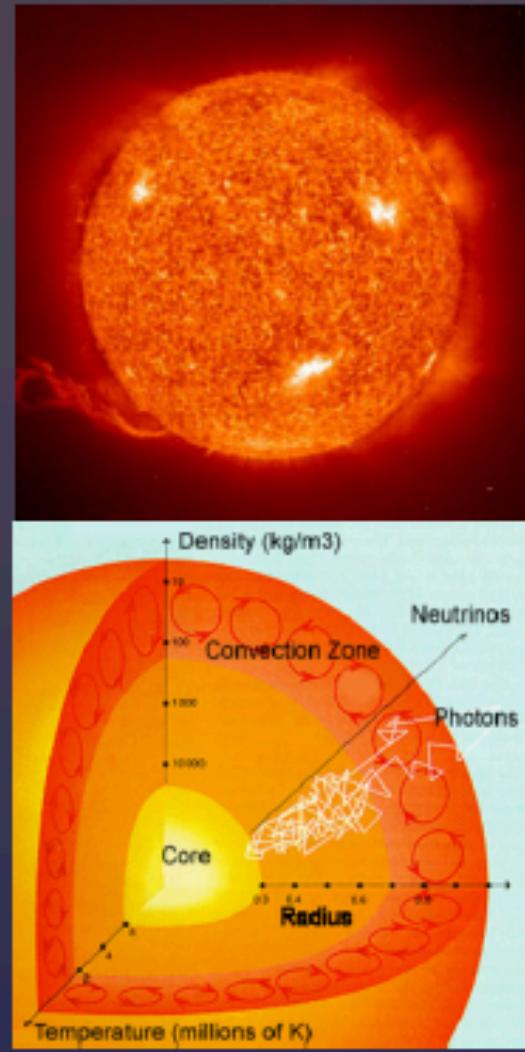
辐射 (内部), 依赖于辐射透明度
对流 (外部)

3. 太阳各处处于一个液态局域平衡状态:

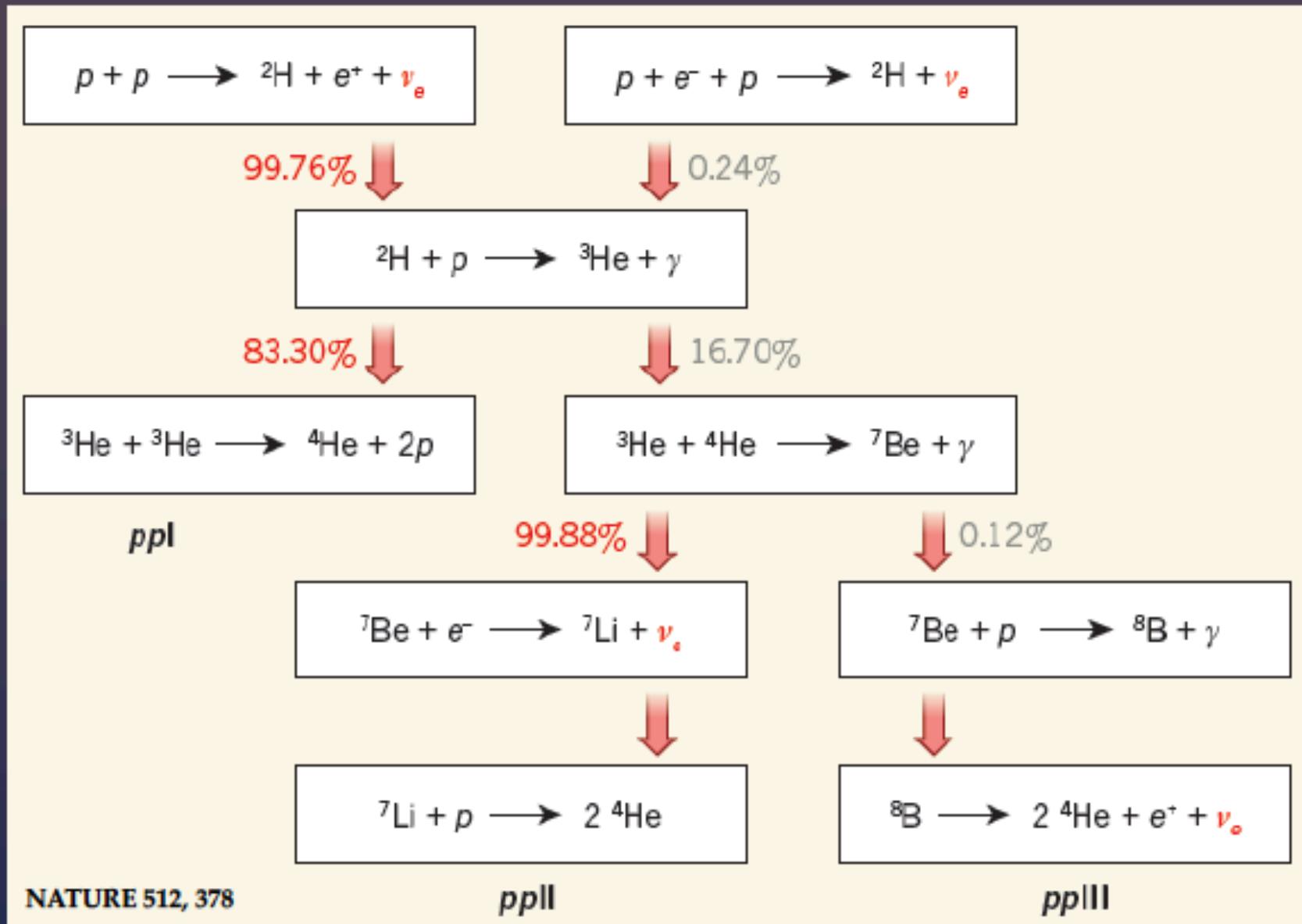
引力与辐射压力

4. 边界条件:

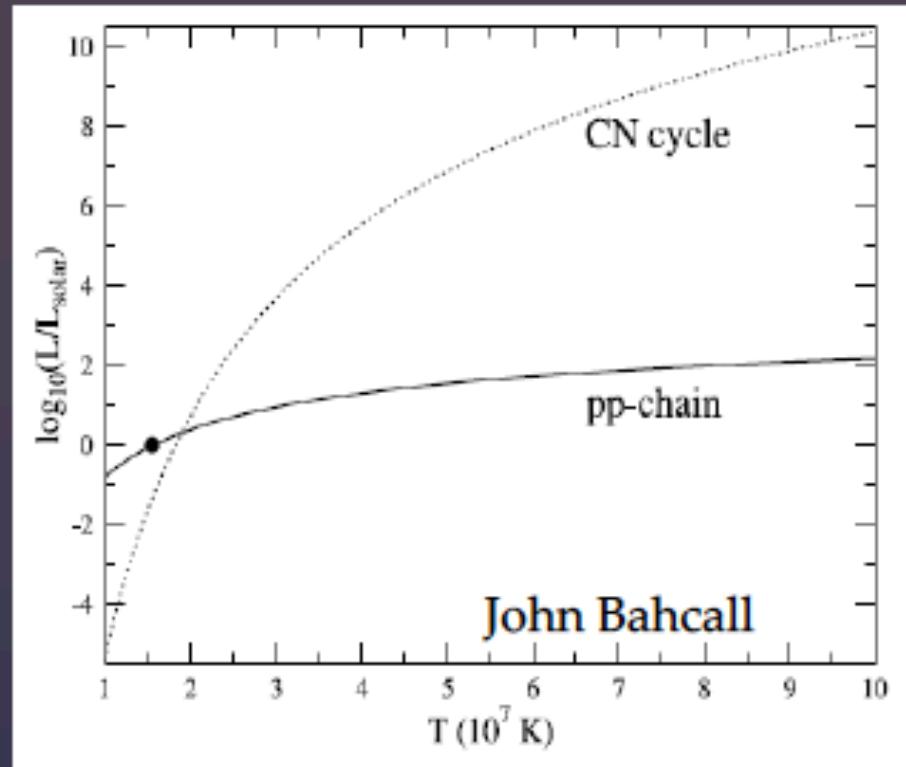
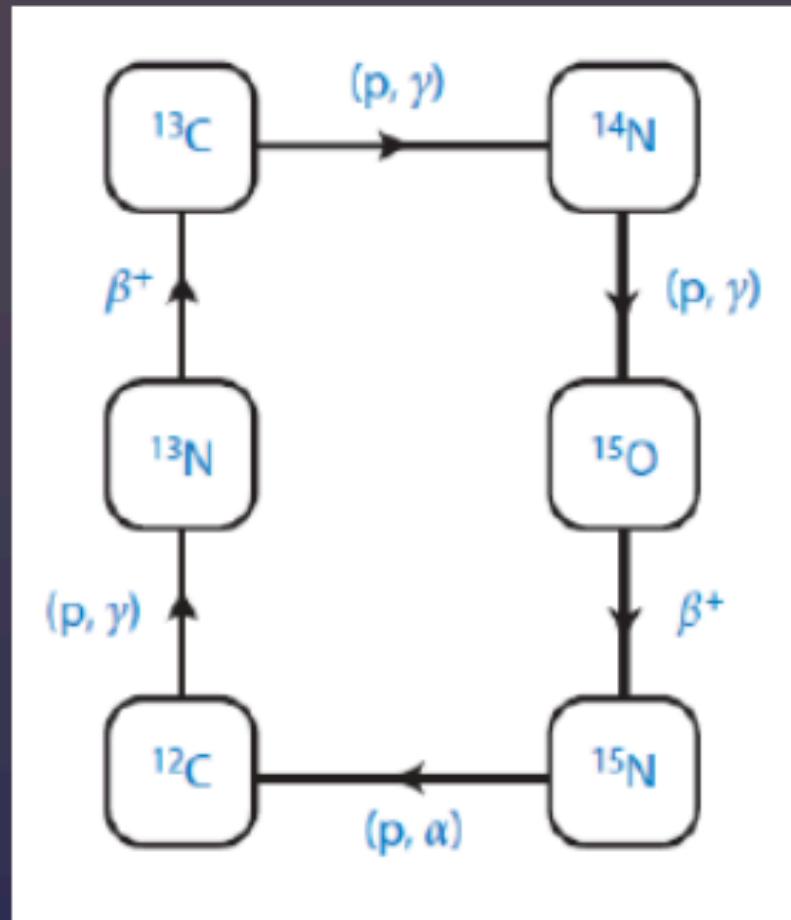
原初星的氢、氦、金属含量,
目前太阳半径, 质量, 亮度约束、年龄。



中微子产生-1：质子-质子聚变链 pp



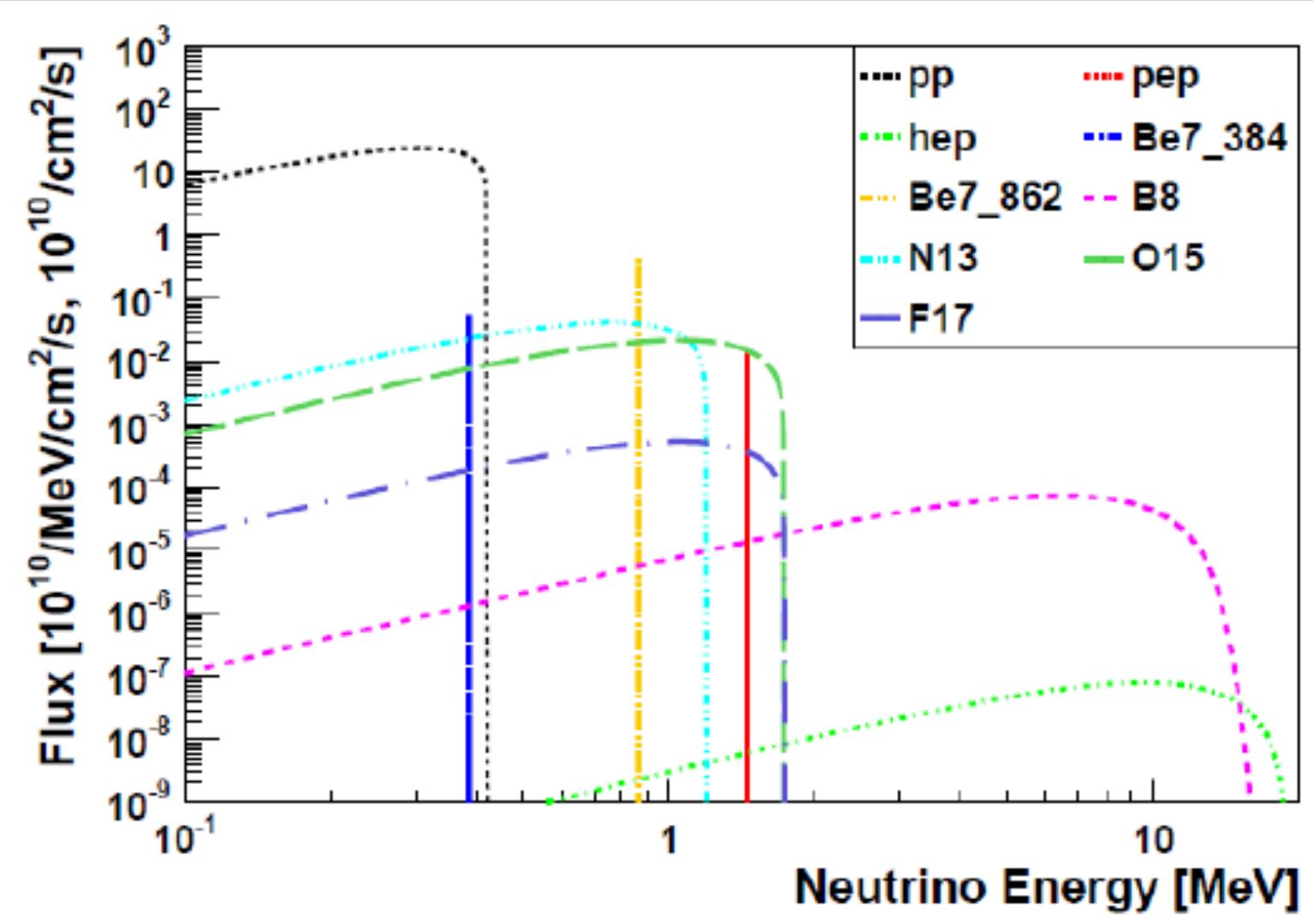
中微子产生-2：碳氮循环 CNO



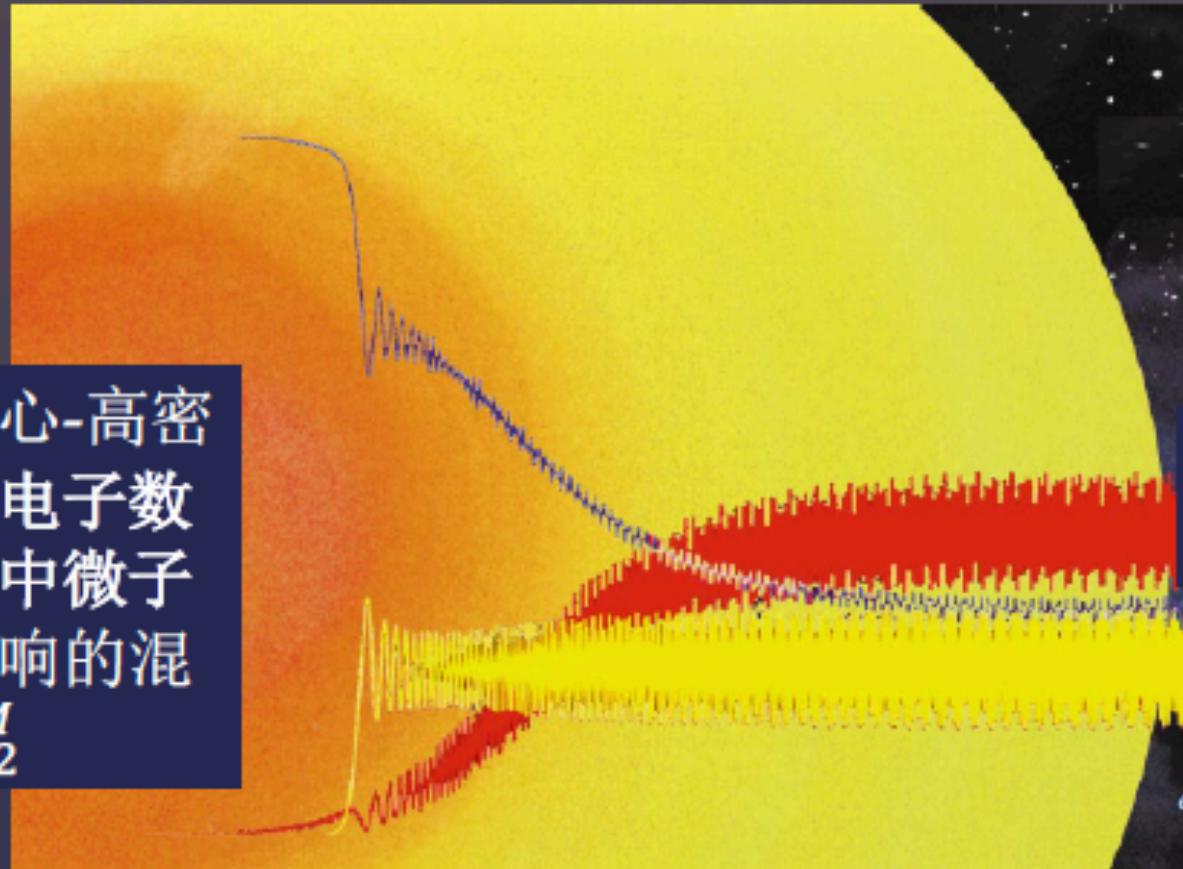
质子-质子聚变链：
(太阳温度下的主要过程)

碳氮循环：
(高温主序星的主要模式)

太阳中微子能谱



太阳中微子振荡与传播



太阳中心-高密度：
受电子数密度和中微子能量影响的混合角 θ_{12}^M

外围：真空
混合角 θ_{12}

$$P_{ee}^\odot = \cos^4 \theta_{13} \left(\frac{1}{2} + \frac{1}{2} \cos 2\theta_{12}^M \cos 2\theta_{12} \right)$$

*如果中微子穿过地球，其通量将
继续受到调制，幅度变化~3%

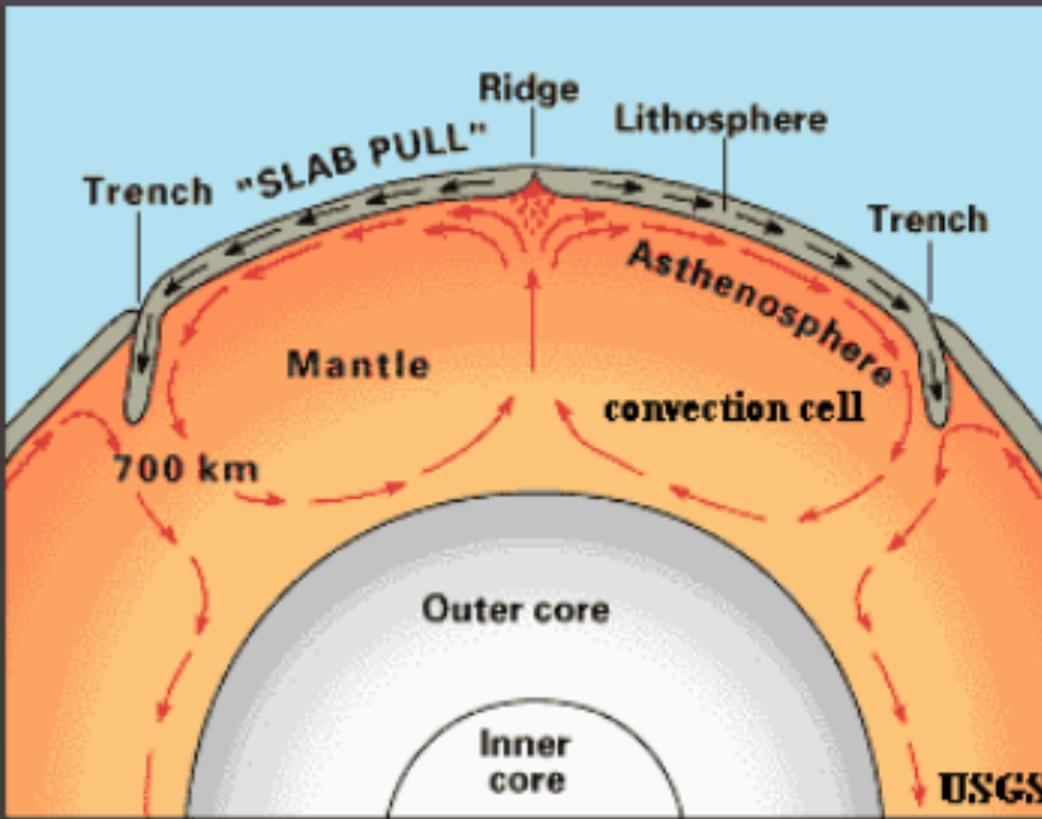
板块运动的原动力？

- 地球引力塌缩势能？
- 地球内部的核衰变，核裂变？



目前的知识：

- 全球地热测量 47 ± 3 TW
- 对核衰变热的预期：
 - Cosmochemical模型：10 TW
 - Geodynamical模型：15-30 TW
 - Geochemical模型：20 TW
- 地球中微子实验测量10-30 TW之间

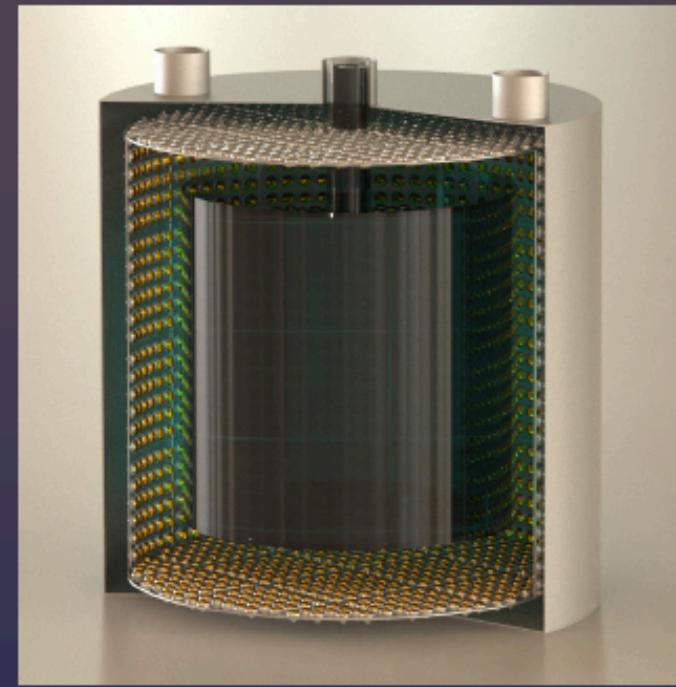
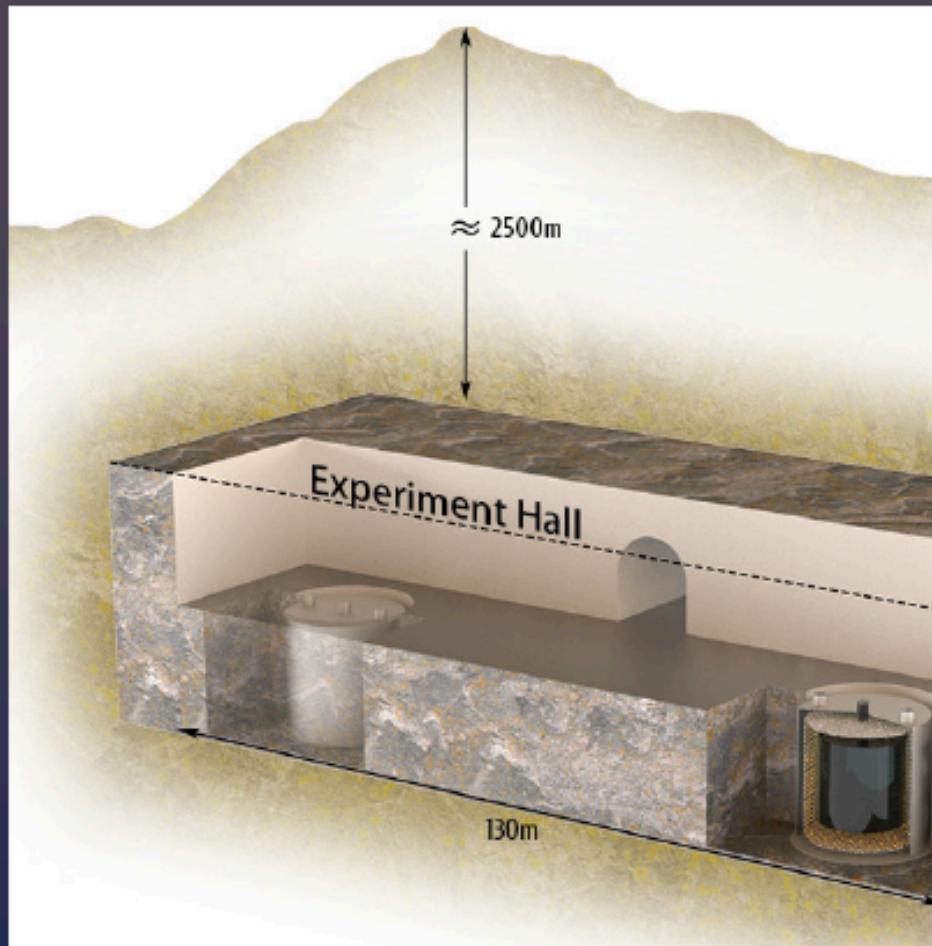


答案：

- 我们还在消耗地球形成之初的引力势能
- 核衰变情形基本未知
- 需要测量地球中微子

中微子探测器

锦屏中微子实验



太阳中微子：有效靶体积 $\approx V(H-2 \times 4\text{米}, D-2 \times 4\text{米}) \approx 1000\text{吨}$
地球中微子：有效靶体积 $\approx V(H-2 \times 3\text{米}, D-2 \times 3\text{米}) \approx 1500\text{吨}$

2016年6月2日

- 共两个中微子探测器
- 每个的太阳中微子有效靶体积为1000吨，共2000吨
- 三层主要结构：钢桶，PMT层，亚克力桶
- 钢桶和亚克力间为纯水
- 亚克力容器内为液闪或水基液闪
- 总高H、外直径D约20米
- 能量分辨 $> 500 \text{ PE/MeV}$
- 各方向外部光子缓冲 4米

锦屏中微子实验调研-统计相对误差

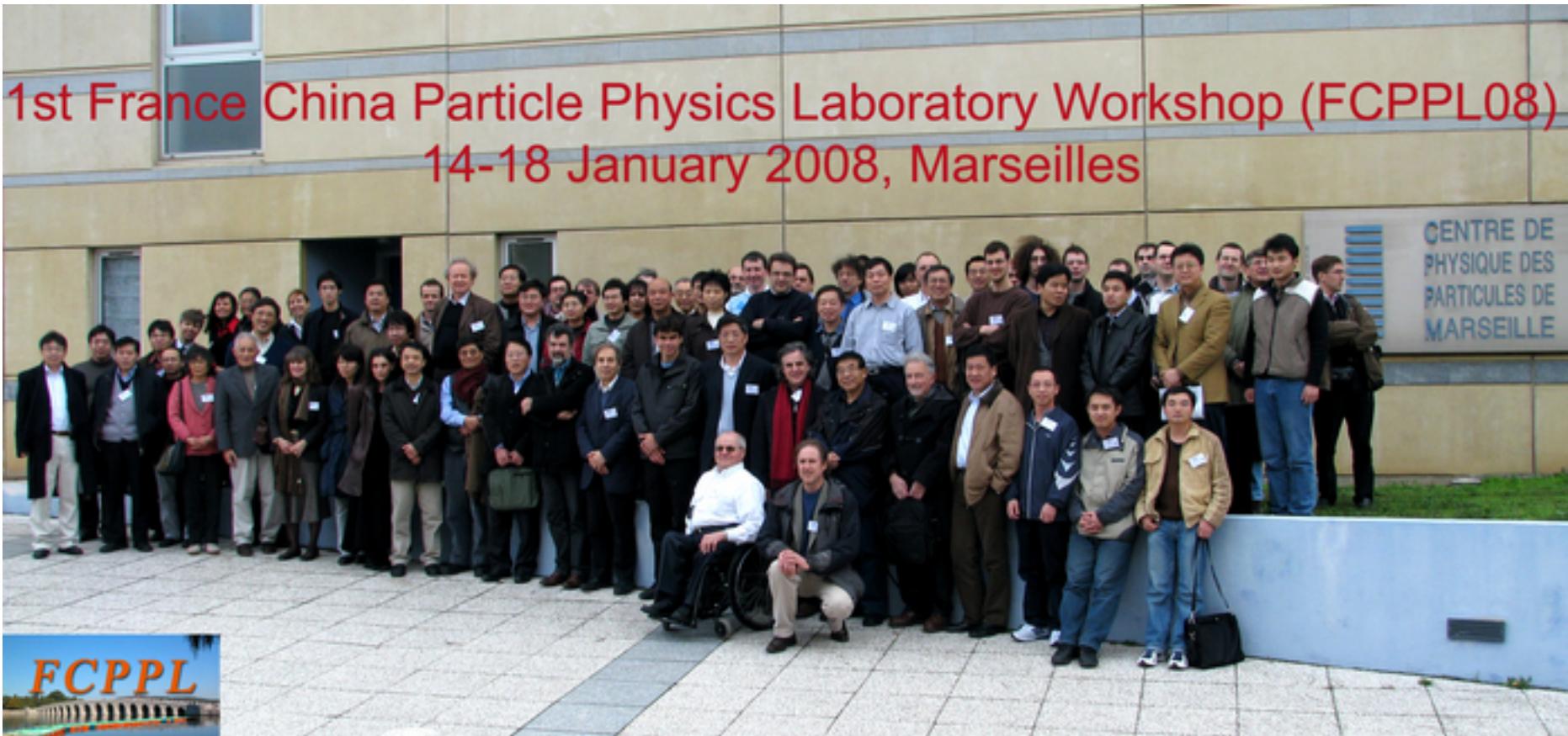
	Neutrino component	Energy resolution		
		200 PE/MeV	500 PE/MeV	1000 PE/MeV
Fiducial mass 1000 ton	pp	0.02	0.007	0.005
	⁷ Be	0.007	0.006	0.005
	pep	0.07	0.05	0.04
	¹³ N	NA	0.5 (NA)	0.3 (0.4)
	¹⁵ O	0.3	0.2 (0.4)	0.1 (0.2)
	⁸ B	0.02	0.02	0.02
Fiducial mass 2000 ton	pp	0.01	0.005	0.004
	⁷ Be	0.005	0.004	0.004
	pep	0.06	0.03	0.03
	¹³ N	0.4	0.3	0.2 (0.3)
	¹⁵ O	0.2	0.1	0.08 (0.1)
	⁸ B	0.02	0.02	0.02
Fiducial mass 4000 ton	pp	0.01	0.004	0.003
	⁷ Be	0.004	0.003	0.003
	pep	0.04	0.03	0.02
	¹³ N	0.3	0.2 (0.3)	0.2 (0.3)
	¹⁵ O	0.1 (0.2)	0.07 (0.1)	0.06 (0.09)
	⁸ B	0.01	0.01	0.01

相对不确定性，例如0.3
约对应 3σ

International Collaborations



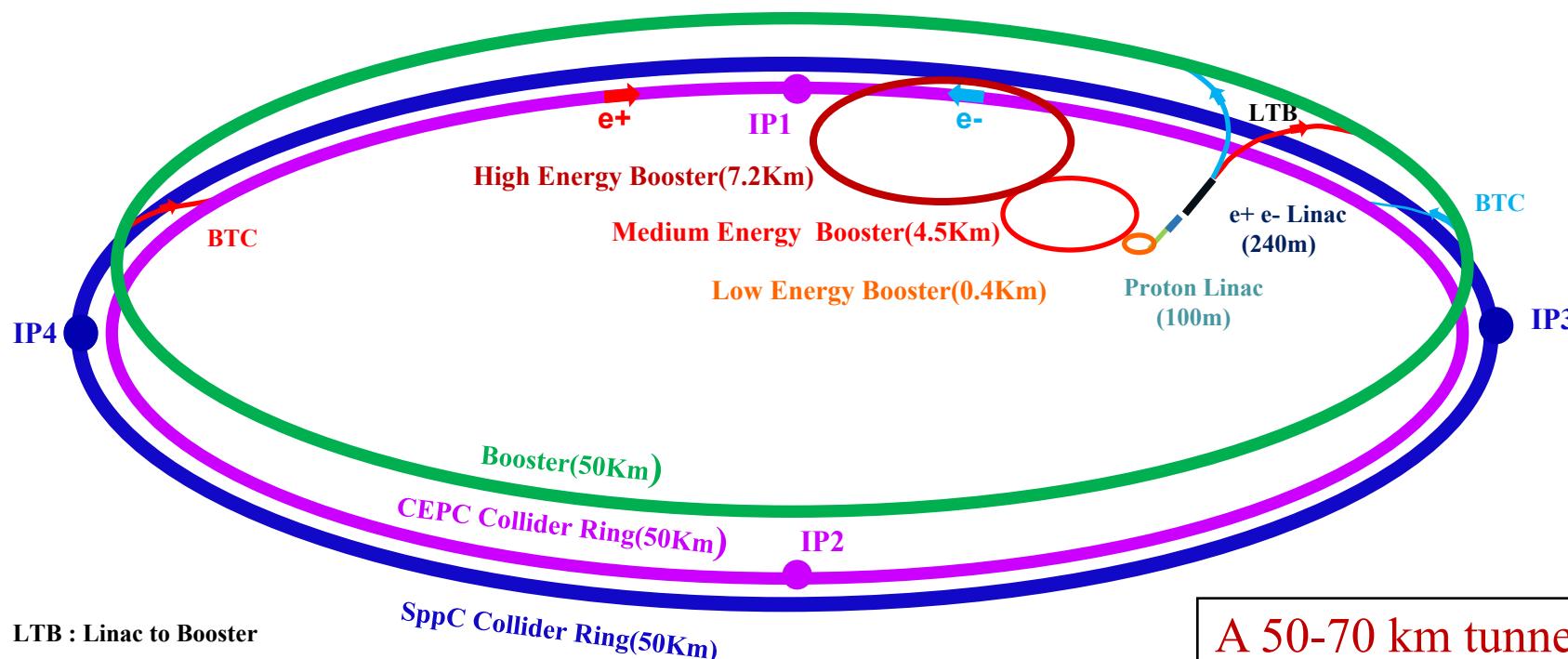
http://www.ihep.cas.cn/gjjl/hzxm/huitan/201111/t20111108_3392423.html



http://www.ihep.cas.cn/gjjl/hzxm/201404/t20140416_4091841.html

The Future: CEPC+SppC

- For about 8 years, we have been talking about “What can be done after BEPCII in China”
- Thanks to the discovery of the low mass Higgs boson, and stimulated by ideas of Circular Higgs Factories in the world, CEPC+SppC configuration was proposed in Sep. 2012



A 50-70 km tunnel is
relatively easier NOW
in China

Scientific Goals

- CEPC (e+e-: 90-250 GeV)
 - Higgs Factory: Precision study of Higgs(m_H , J^{PC} , couplings)
 - Same as SM prediction ? Other Higgs ? Composite ? New properties ? CP effect ?
 - Z & W factory: precision test of SM
 - New phenomena ? Rare decays ?
 - Flavor factory: b, c, τ and QCD studies
- SppC (pp: 50-100 TeV)
 - Directly search for new physics beyond SM
 - Precision test of SM
 - e.g., h^3 & h^4 couplings

Precision measurement & searches: Complementary with each other

Bob Wilson (first Director of Fermilab), when asked by a Congressional Committee "*What will your lab contribute to the defense of the US?*", replied "*Nothing, but it will make it worth defending*".