The International Linear Collider a physics project at an electron-positron collider in an energy range from Z to TeV

Introduction



For a detailed study of the electroweak breaking

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What is ILC ?

A brief historical survey

The previous frontier electron-positron collider was the LEP at CERN, a circular collider from Z to W pairs. First beams summer 89, end in 2000 The first linear collider was at the same time running at SLAC (SLC), with polarised beams.

The first meetings I know about what is now ILC were in 1991 for physics and detector, 1992 for accelerator.

The lepton community was already considering the next step to come in parallel with LHC. Introduction

Under the leadership of DESY a project of superconducting accelerator (S band) to be built at DESY Hamburg in the energy range Z to TeV (TDR in 2000) in parallel with a synchrotron light source, the EXFEL. At the same time two warm projects were considered the Next Linear Collider in the US (Xband) the Global Linear Collider in Japan.

Germany accepted EXFEL and turned down TESLA.

The world community joined then efforts in a unique project when the superconducting technology was chosen (2004). That was ILC which under the control of ICFA developped a TDR in 2012. Today the only site under consideration is in Japan (Kitakami).

ICFA created by IUPAP in 1976, in regions : ECFA, ACFA, APS DPB

At the time of letters of intent for ILC experiments there were 4 proposals, the committee turned down the 4th concept,

two of them, GLD and LDC, merged into ILD SiD remained.

Today remain SiD, silicon detector, and ILD international large detector



Another approach to the same questions but at higher energies : the CLIC

developped at CERN on the double accelerator scheme to reach 3 TeV. A lot of synergies with ILC, in particular for the detector



Why electrons?

What else?

stable or quasi stable particles: electrons, protons, ions with their antiparticles photons (an option) muons?

A well defined initial state

parton monokinetic beam CM = laboratory interaction point spin state

Angular distributions → detector but!

An initial state properly coupled to what we intend to study

The proton-proton cross section grows with s

Introduction

The e⁺e⁻ cross section for a given final state far from the threshold goes like 1/s (why ?) low backgrounds → no pile-up, no trigger, no bias

except $\gamma\gamma$ (including beamstrahlung) which is at low p_t hence low angle

A clean collider





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Y, Z, ..

e



What accelerator?

Do we know how to make an electrons-positrons collider going to the adequate energy (~ TeV) with an adequate luminosity ?

For cross-sections $\sigma \sim 100$ fb we will need integrated luminosities above few 100 fb⁻¹ per year (10000 events detected) counting on a year of 10⁷ s to take into account efficiency this means intantaneous luminosities of 10³⁴ cm⁻² s⁻¹.



What physics?

All what may appear in the range of energy we can dream of today, from Z to TeV.

is there a physics to invest the price of the collider (8 G€) with 2 (?!) detector(s)(10 %)?



What detector?

Do we know how to make a detector to collect fully, in a reasonable time, the physics present at these energies?

> Measurement precision, efficiency are equivalent to luminosity

Example: in the study of

Control the systematics at the detector level as well as accelerator

 $e^+ e^- \rightarrow ZH, Z \rightarrow q\bar{q}, H \rightarrow WW^x$

improve the jet resolution by a factor 2

~ gain of a factor 1.5 on \mathscr{L}

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Remark

The way an e^+e^- collects physics is strongly different from a proton machine.

The proton machine covers a wide band of parton-parton collisions energy and the machine will always run at its maximal energy.

Introduction

The electron collider has optimal energies for given studies, thresholds (many), peak of cross sections.

As a consequence, the time needed to fully explore an energy domain for example GigaZ, W threshold, top threshold, Higgs, high energy, may be considerable.

The choice of a running programme is strategic



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Socio-scientific remark :

A scientific result has to be reproducible but should it be reproduced ?

Should we then draw one, two or three experiments?

A circular collider offers space for few experiments easily LEP had 4, a slight overkill with an almost equivalent luminosity a linear one shares at best its luminosity between the experiments

with rather ugly constraints.

The funny game of push-pull



as well as what is beyond what we already know

A detector to collect it



Miscellaneous



Unit prefixes

milli	m	- 3	latin	kilo	k	3	greek
micro	μ	- 6	greek	mega	Μ	6	greek
nano	n	- 9	greek	giga	G	9	greek
pico	р	- 12	italian	tera	Т	12	greek
femto	f	- 15	danish	peta	Ρ	15	greek
atto	а	- 18	danish	exa	Е	18	greek



A question of units

Even though we may use SI units for some numerical applications we favour, as it is customary in HEP, using one unit only to be chosen as « eV ». this is reached by defining c=1, $\hbar=1$, $\epsilon_0=1$ $\mu_0=1$

Introduction

mass, momentum and energy are in eV, or MeV, or GeV

time and length are in the same unit E^{-1} $E = \hbar \omega$ the metre is indeed defined from the second energy and wavelength

Exercise : knowing that the earth-moon distance is about 30000km or 1s and that $\hbar c = 200 MeV fm$ give this distance in eV. or what is the energy of a photon with a wavelength equal to the earth-moon distance ?

writing the field energy, what is the dimension of E field, B field, charge ?

To remember : eV and μm^{-1} are of the same order of magnitude





 $\hbar c = 200 M eV f m$

 $1=0,2eV\mu m \qquad 1\mu m=5eV^{-1}$ $30000km=310^{14}\mu m=3\times510^{14}eV^{-1}=1,5(feV)^{-1}$



END