



Revisiting the scalar leptoquark solution to the $R(D^{(*)})$ anomalies and its implications

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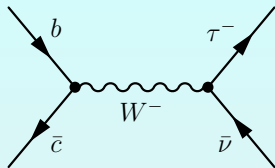
X. Q. Li, Y. D. Yang and X. Zhang, JHEP **08** (2016) 054



- 1 Motivation
- 2 Framework
- 3 Our work
- 4 Conclusion

Motivation

$$B \rightarrow D^{(*)} \tau \nu$$



$$\mathcal{R}(D^{(*)}) \equiv \mathcal{B}(B \rightarrow D^{(*)} \tau \nu) / \mathcal{B}(B \rightarrow D^{(*)} \ell \nu)$$

$$\mathcal{R}_{\text{exp}}(D) = 0.397 \pm 0.049$$

BaBar+Belle

$$\mathcal{R}_{\text{SM}}(D) = 0.300 \pm 0.008$$

2.0 σ \rightarrow 1.9 σ

$$\mathcal{R}_{\text{exp}}(D^*) = 0.316 \pm 0.019$$

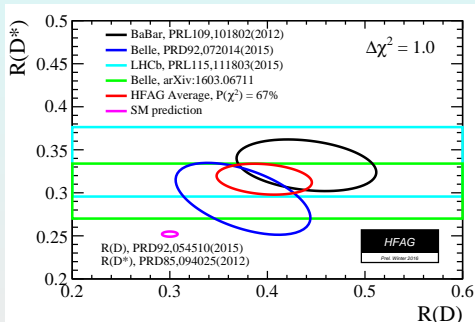
BaBar+Belle+LHCb

$$\mathcal{R}_{\text{SM}}(D^*) = 0.252 \pm 0.003$$

2.7 σ \rightarrow 3.3 σ

- combined excess: 3.4 σ \rightarrow 4.0 σ
- the: V_{cb} , hadronic uncertainties canceled
- BSM: scalar leptoquark scenario
- exp: Belle II @ 50ab $^{-1}$
 $\sigma \approx 0.010$ for $R(D)$
 $\sigma \approx 0.005$ for $R(D^*)$

B. Golob et al., Report No. BELLE2- NOTE-0021



The scalar leptoquark scenario

The scalar ϕ transforms as $(3, 1, -\frac{1}{3})$ with couplings to fermions described by

$$\mathcal{L}_{\text{int}}^{\phi} = \bar{Q}_L^c \lambda^L i\tau_2 L \phi^* + \bar{u}_R^c \lambda^R \ell_R \phi^* + \text{h.c.},$$

M. Bauer and M. Neubert, *Phys. Rev. Lett.* 116 (2016), 141802

The ϕ mediates the $b \rightarrow c \tau^- \bar{\nu}_\tau$ at tree level, and gives rise to \mathcal{H}_{eff} including SM

$$\mathcal{H}_{\text{eff}} = \frac{4G_F V_{cb}}{\sqrt{2}} \left[C_V(M_\phi) \bar{c} \gamma_\mu P_L b \bar{\tau} \gamma^\mu P_L \nu_\tau + C_S(M_\phi) \bar{c} P_L b \bar{\tau} P_L \nu_\tau - \frac{1}{4} C_T(M_\phi) \bar{c} \sigma_{\mu\nu} P_L b \bar{\tau} \sigma^{\mu\nu} P_L \nu_\tau \right],$$

$$C_V(M_\phi) = 1 + \frac{\lambda_{b\nu_\tau}^L \lambda_{c\tau}^{L*}}{4\sqrt{2} G_F V_{cb} M_\phi^2},$$

$$C_{S,T}(M_\phi) = - \frac{\lambda_{b\nu_\tau}^L \lambda_{c\tau}^{R*}}{4\sqrt{2} G_F V_{cb} M_\phi^2}. \quad \text{X.Q. Li et.al, JHEP 08 (2016) 054}$$

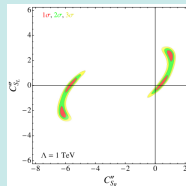
In order to re-sum potentially large logarithmic effects, the $C_{V,S,T}$ should be run down to $\mu_b \sim m_b$ scale.

I. Doršner et.al, *JHEP* 11 (2013) 084

Parameters

M. Freytsis *et al.*, *Phys. Rev. D*92 (2015), 054018

$$(\lambda_{b\nu\tau}^L, \lambda_{c\tau}^{L*}, \lambda_{b\nu\tau}^L, \lambda_{c\tau}^{R*}) = \begin{cases} (0.35, -0.03), & P_A \\ (0.96, 2.41), & P_B \\ (-5.74, 0.03), & P_C \\ (-6.34, -2.39), & P_D \end{cases}$$



Only P_A is adopted by Bauer and Neubert.

M. Bauer and M. Neubert, *Phys. Rev. Lett.* 116 (2016), 141802

It would be worth investigating whether the four best-fit solutions could be discriminated from each other.

Purely leptonic decay $B_c^- \rightarrow \tau^- \bar{\nu}_\tau$

$$\Gamma(B_c^- \rightarrow \tau^- \bar{\nu}_\tau) = \frac{G_F^2}{8\pi} |V_{cb}|^2 f_{B_c}^2 m_{B_c}^3 \frac{m_\tau^2}{m_{B_c}^2} \left(1 - \frac{m_\tau^2}{m_{B_c}^2}\right)^2 \left| C_V - C_S \frac{m_{B_c}^2}{m_\tau [m_b(\mu_b) + m_c(\mu_b)]} \right|^2$$

	P_A	P_B	P_C	P_D
C_V	1.129	1.354	-1.117	-1.338
C_S	0.018	-1.465	-0.018	1.453

$$\Gamma(B_c^- \rightarrow \tau^- \bar{\nu}_\tau) = \begin{cases} 2.22 \times 10^{-2} \Gamma_{B_c}, & \text{SM} \\ 2.45 \times 10^{-2} \Gamma_{B_c}, & P_A \quad \text{😊} \\ 1.33 \Gamma_{B_c}, & \cancel{P_B} \\ 2.39 \times 10^{-2} \Gamma_{B_c}, & P_C \quad \text{😊} \\ 1.31 \Gamma_{B_c}, & \cancel{P_D} \end{cases}$$

Comparison between P_A and P_C :

$$\mathcal{H}_{\text{fit}} = \frac{4G_F}{\sqrt{2}} V_{cb} \left\{ \underbrace{\left[1 + \begin{pmatrix} 0.129 & P_A \\ -2.117 & P_C \end{pmatrix} \right]}_{C_V^{\text{fit}}} \bar{c} \gamma_\mu P_L b \bar{\tau} \gamma^\mu P_L \nu_\tau + \underbrace{\begin{pmatrix} 0.018 & P_A \\ -0.018 & P_C \end{pmatrix}}_{C_S^{\text{fit}}} \bar{c} P_L b \bar{\tau} P_L \nu_\tau \right. \\ \left. + \underbrace{\begin{pmatrix} -0.002 & P_A \\ 0.002 & P_C \end{pmatrix}}_{C_T^{\text{fit}}} \bar{c} \sigma_{\mu\nu} P_L b \bar{\tau} \sigma^{\mu\nu} P_L \nu_\tau \right\}.$$

$$B_c^- \rightarrow \gamma \tau^- \bar{\nu}_\tau$$

Figure: The leading order Feynman diagrams resulting from the \mathcal{H}_{eff} .

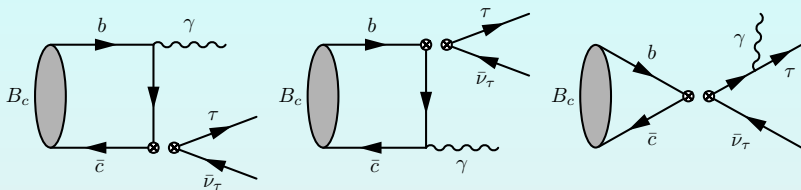


Table: Branching ratios within the SM and beyond with $E_\gamma \geq 1$ GeV.

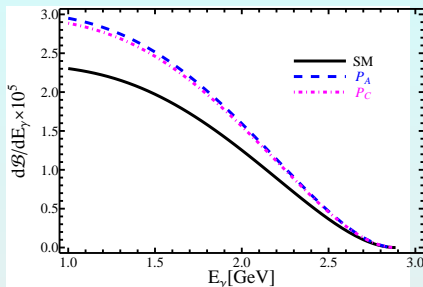
Observable	SM	P_A	P_C
$\mathcal{B}(B_c^- \rightarrow \gamma \tau^- \bar{\nu}_\tau) \times 10^5$	2.36	3.01	2.94

$\mathcal{B}(B_c^- \rightarrow \gamma \tau^- \bar{\nu}_\tau)$ are enhanced by $\sim 27\%$ in both P_A and P_C cases.

$$B_c^- \rightarrow \gamma \tau^- \bar{\nu}_\tau$$



Figure: The dependence of the differential branching ratios on E_γ .



The enhancement is significant when $1 \text{ GeV} < E_\gamma < 1.5 \text{ GeV}$, but tiny near the end. P_A and P_C are almost indistinguishable.

$$\bar{B} \rightarrow D^{(*)} \tau \bar{\nu}_\tau$$

- $R(D^{(*)})$ and $\mathcal{B}(\bar{B} \rightarrow D^{(*)} \tau \nu_\tau)$ (10^{-2})

Observable	SM	P_A	P_C	Exp
$R(D)$	0.298 ± 0.009	0.388 ± 0.012	0.380 ± 0.011	0.397 ± 0.049
$R(D^*)$	0.253 ± 0.002	0.325 ± 0.002	0.318 ± 0.002	0.316 ± 0.019
$\mathcal{B}(\bar{B} \rightarrow D \tau \nu)$	0.72 ± 0.06	0.94 ± 0.07	0.92 ± 0.07	1.07 ± 0.18
$\mathcal{B}(\bar{B} \rightarrow D^* \tau \nu)$	1.30 ± 0.04	1.67 ± 0.05	1.64 ± 0.04	1.64 ± 0.15

$$\begin{aligned}
 V_1(w) &= V_1(1)[1 - 8\rho_D^2 z + (51\rho_D^2 - 10)z^2 - (252\rho_D^2 - 84)z^3], & \rho_D^2 &= 1.186 \pm 0.054, \\
 h_{A_1}(w) &= h_{A_1}(1)[1 - 8\rho_{D^*}^2 z + (53\rho_{D^*}^2 - 15)z^2 - (231\rho_{D^*}^2 - 91)z^3], & \rho_{D^*}^2 &= 1.207 \pm 0.026, \\
 R_1(w) &= R_1(1) - 0.12(w - 1) + 0.05(w - 1)^2, & R_1(1) &= 1.403 \pm 0.033, \\
 R_2(w) &= R_2(1) + 0.11(w - 1) - 0.06(w - 1)^2, & R_2(1) &= 0.854 \pm 0.020, \\
 R_3(w) &= 1.22 - 0.052(w - 1) + 0.026(w - 1)^2. & V_1(1) &= 1.074 \pm 0.024, \\
 & & h_{A_1}(1) &= 0.908 \pm 0.017.
 \end{aligned}$$

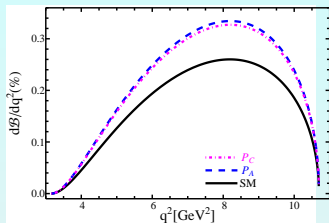
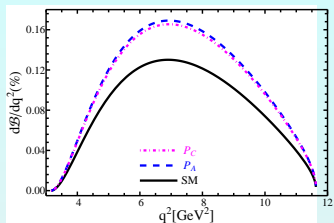
I. Caprini *et al.*, Nucl.Phys. B530 (1998) 153,

Lattice, Nucl.Phys.Proc.Supp. 140 (2005) 461

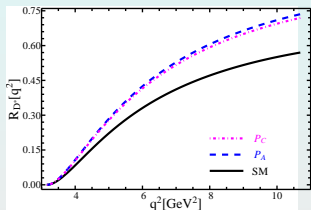
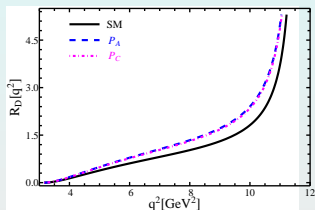
HFAG arXiv:1207.1158

$$\bar{B} \rightarrow D^{(*)} \tau \bar{\nu}_\tau$$

- The q^2 distributions of $d\mathcal{B}(\bar{B} \rightarrow D^{(*)} \tau \nu_\tau)/dq^2$

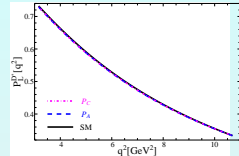
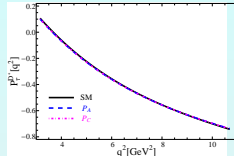
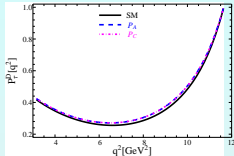


- The q^2 distributions of the ratios $R_{D^{(*)}}(q^2)$

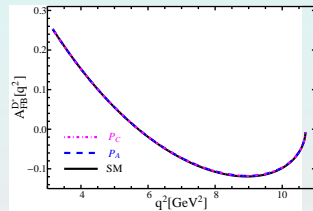
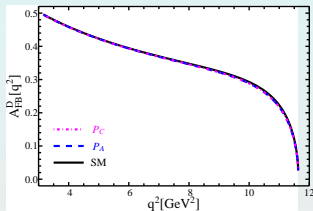


$$\bar{B} \rightarrow D^{(*)} \tau \bar{\nu}_\tau$$

$$\bullet P_\tau(q^2) = \frac{d\Gamma^{\lambda_\tau=1/2}/dq^2 - d\Gamma^{\lambda_\tau=-1/2}/dq^2}{d\Gamma/dq^2}, \quad P_L^{D^*}(q^2) = \frac{d\Gamma^{\lambda_{D^*}=0}/dq^2}{d\Gamma/dq^2}$$



$$\bullet A_{FB}^D(q^2) = \frac{\int_0^1 d \cos \theta (d^2\Gamma/dq^2 d \cos \theta) - \int_{-1}^0 d \cos \theta (d^2\Gamma/dq^2 d \cos \theta)}{d\Gamma/dq^2}$$



$$B \rightarrow X_c \tau^- \bar{\nu}_\tau$$

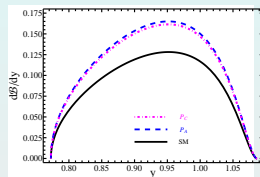
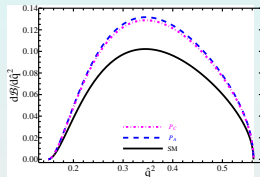
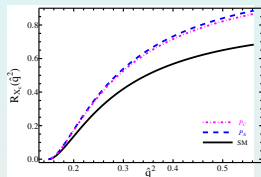
$R(X_c) = \frac{\mathcal{B}(\bar{B} \rightarrow X_c \tau \bar{\nu}_\tau)}{\mathcal{B}(\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell)}$, calculated with OPE, with $\mathcal{B}(B^- \rightarrow X_c e \bar{\nu}_e) = (10.92 \pm 0.16)\%$

F. U. Bernlochner *et al.* Phys. Rev. D85 (2012) 094033

Observable	SM	P_A	P_C
$R(X_c)$	0.230	0.297	0.290
$\mathcal{B}(B^- \rightarrow X_c \tau^- \bar{\nu}_\tau)$	2.51%	3.24%	3.17%

Our SM value of $R(X_c)$ is roughly consistent with $R(X_c) = 0.223 \pm 0.004$.

Z. Ligeti *et al.*, Phys. Rev. D90 (2014) 034021





- The scalar leptoquark can solve the $R(D^{(*)}), R_K$ and $(g-2)_\mu$ anomalies simultaneously with four best-fit solutions obtained.
- The two solutions $P_{B,D}$ are excluded by the decay $B_C^- \rightarrow \tau^- \bar{\nu}_\tau$.
- $P_{A,C}$ lead to different sign of \mathcal{H}_{fit} with just 1.2% difference in $|C_V^{\text{fit}}|$, which is much larger than $|C_S^{\text{fit}}|$ and $|C_T^{\text{fit}}|$.
- The scalar ϕ could give sizeable enhancements to almost all the observables except for $P_L^\tau, P_L^{D^*}$ and A_{FB} .
- With future measurement at LHCb or at Belle-II, one could further decipher various NPs.



Thanks for your attention !!