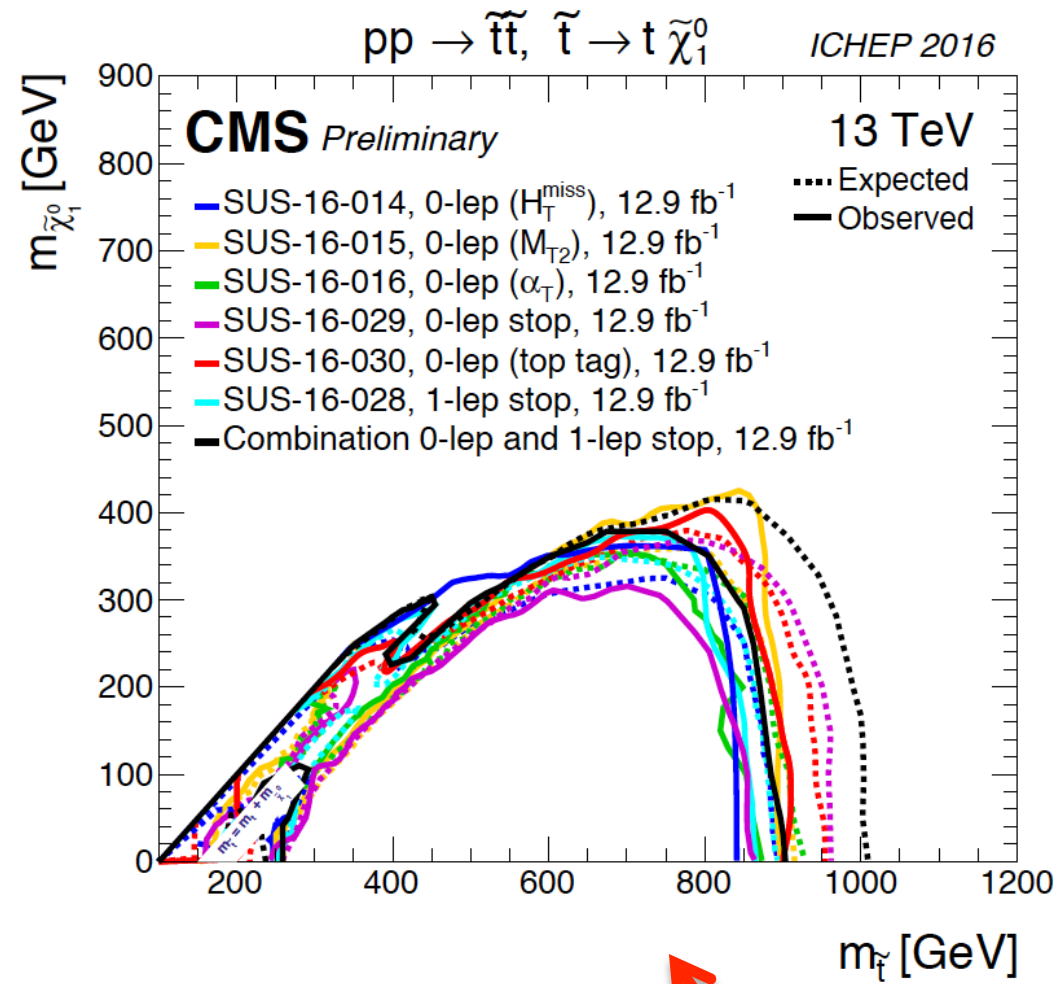


Prospect of LHC searches

LianTao Wang
University of Chicago

CLHCP, Dec 19, 2016

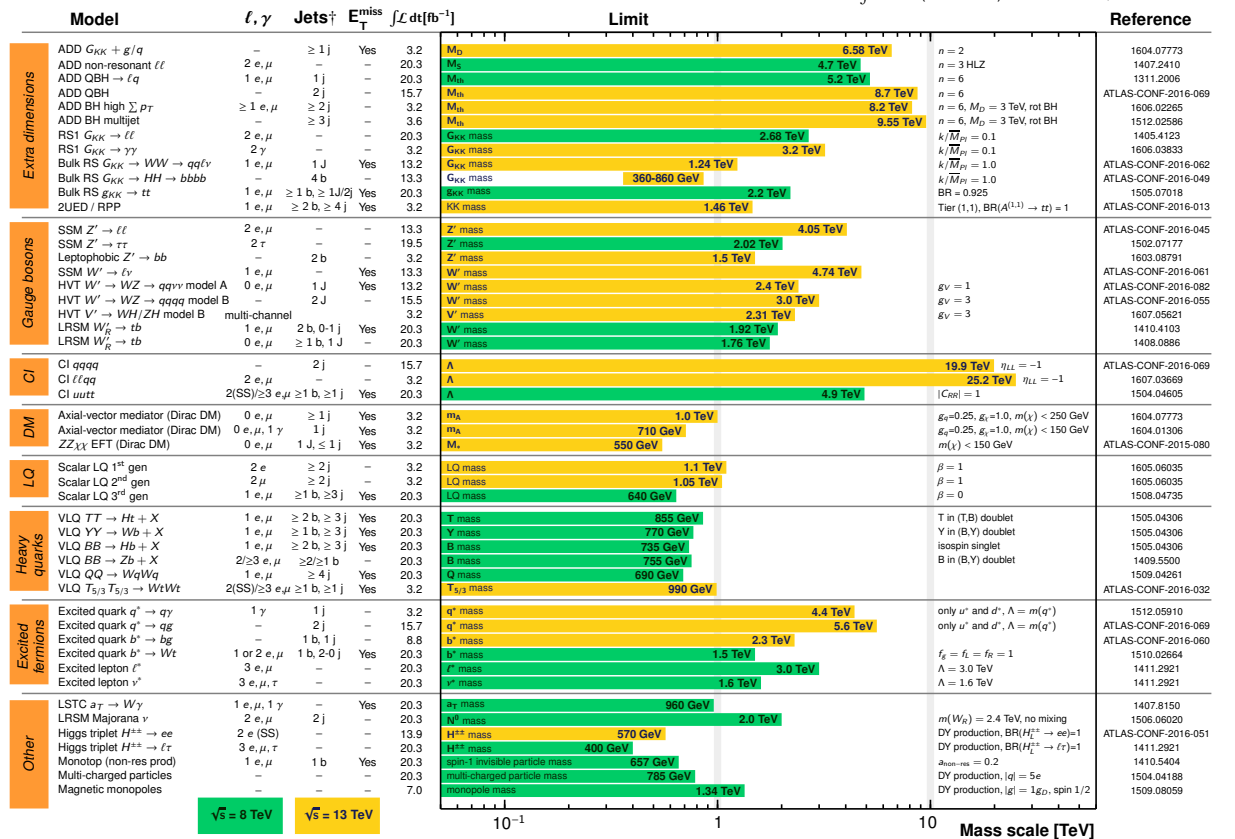
Present



ATLAS Exotics Searches* - 95% CL Exclusion

Status: August 2016

ATLAS Preliminary
 $\int \mathcal{L} dt = (3.2 - 20.3) \text{ fb}^{-1}$
 $\sqrt{s} = 8, 13 \text{ TeV}$

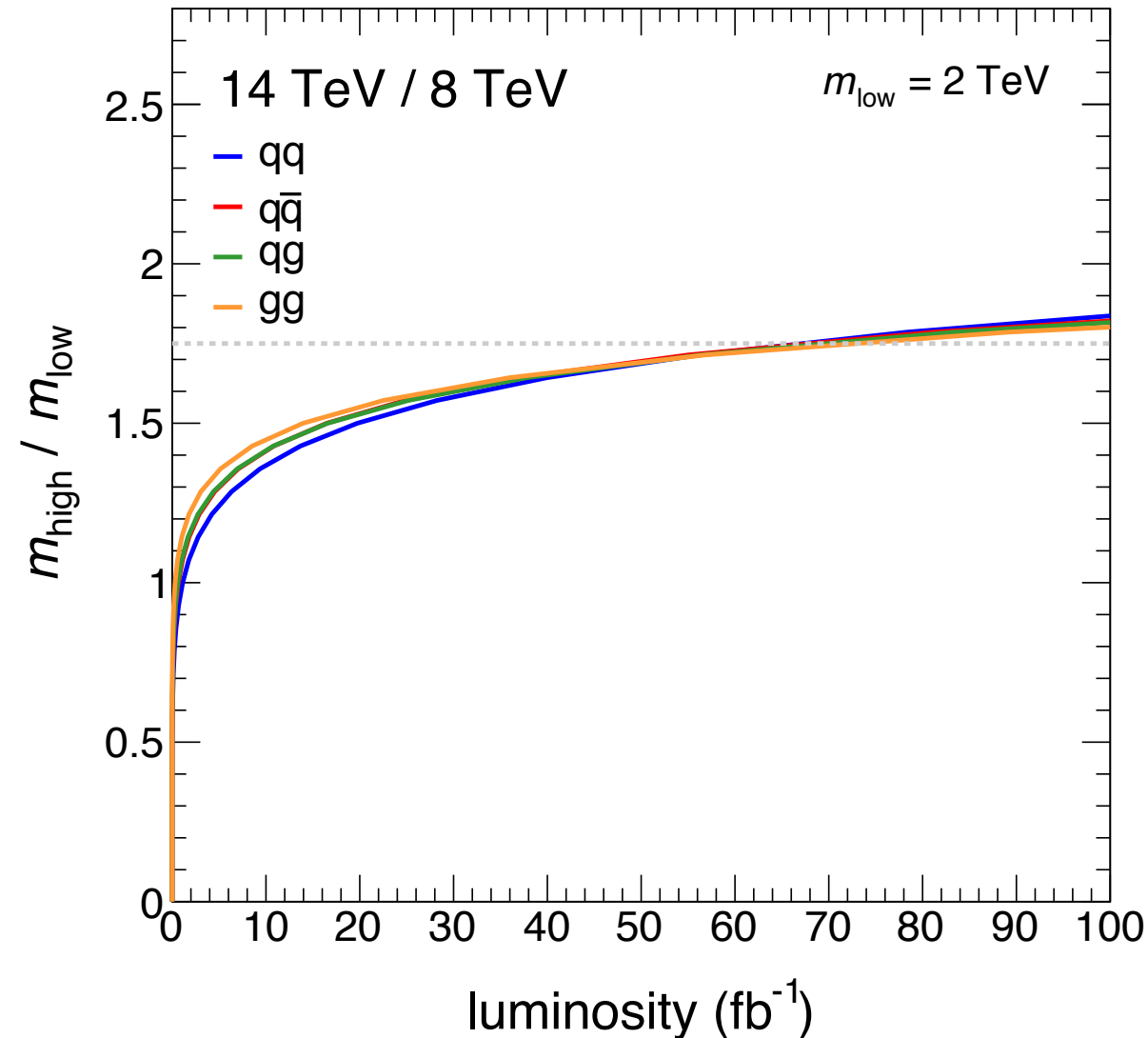


— No “early” discovery.

— Disappointed? Yes. Surprised? Not much.

As data accumulates

Run I limit 2 TeV, e.g. pair of 1 TeV gluino.



Rapid gain initial 10s fb^{-1} , slow improvements afterwards.

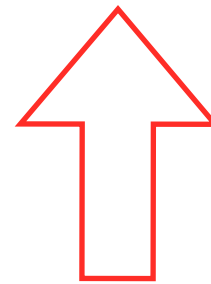
Reaching the “slow” phase after Moriond 2017

Road ahead:

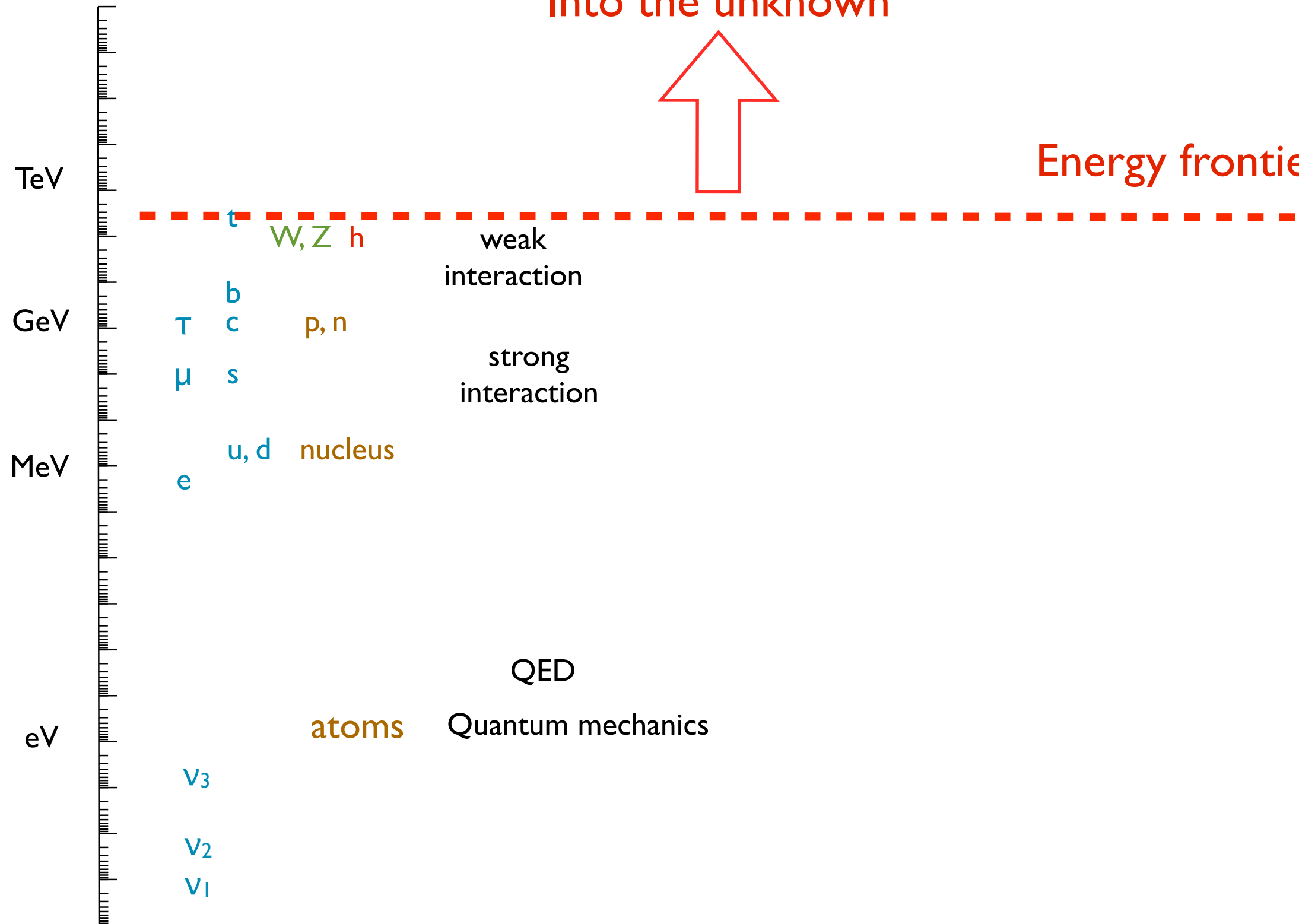
Difficult? Yes.

Prepare to stop? Absolutely not!

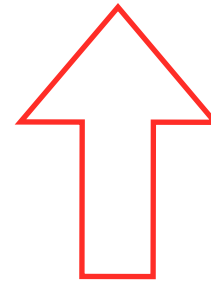
Each new step $\equiv$ new territory
Into the unknown



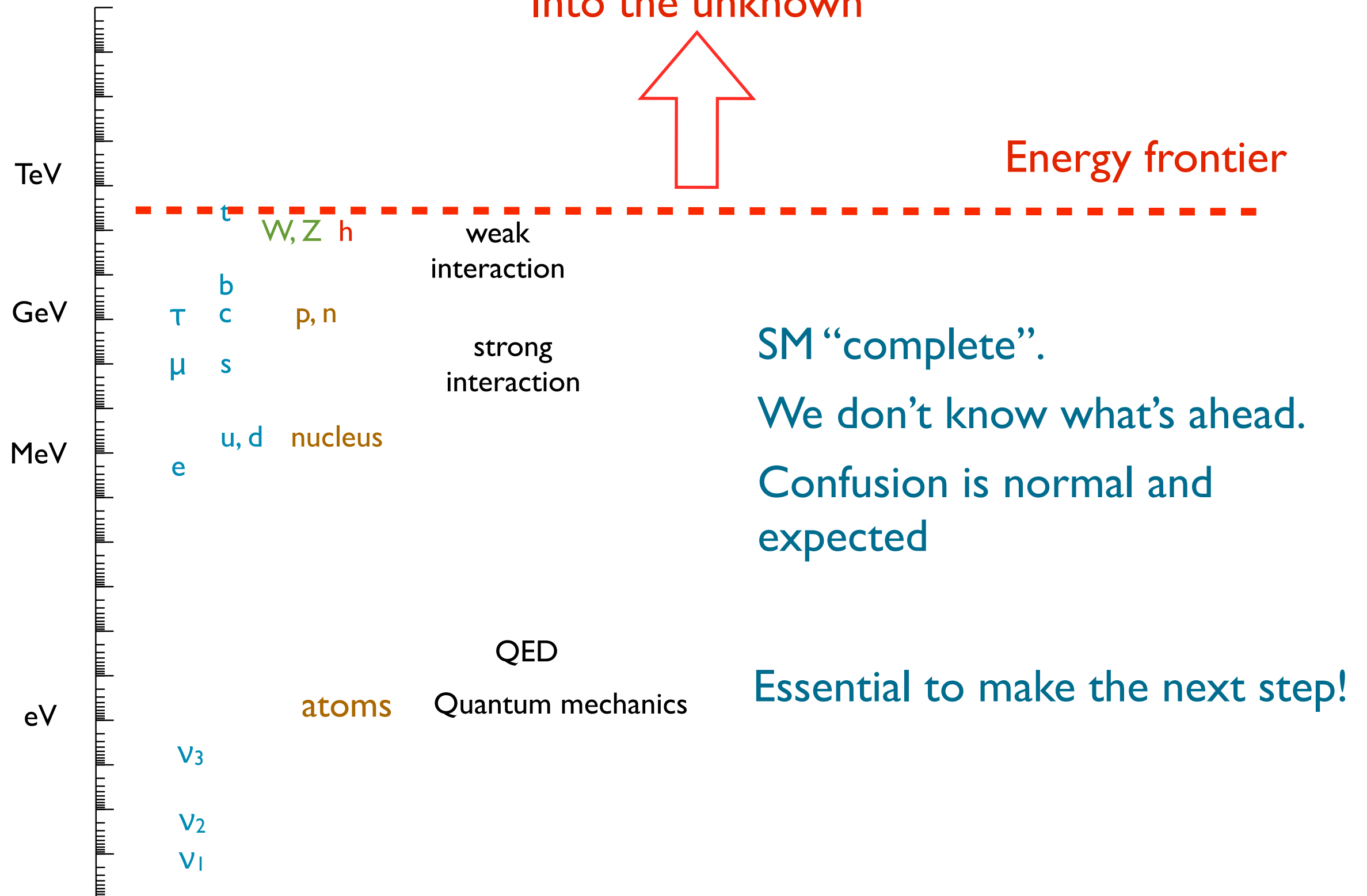
Energy frontier



Each new step $\equiv$ new territory
Into the unknown



Energy frontier



SM has big mysteries.

- Mechanism of electroweak symmetry breaking.
- Dark matter.
- Matter anti-matter asymmetry in the universe
- Flavor....
- We must keep looking!
 - ▶ At the LHC and beyond.

What are we looking for?

What are we looking for?

- We want to search for clues to solve the big mysteries of the SM. None has been answered yet.

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- Our wish list has not changed much from 10 years ago.
 - ▶ All the papers we have written before 2009 (except technicolor). Keep looking for them!

What are we looking for?

- We want to search for clues to solve the big mysteries of the SM. None has been answered yet.
- Our wish list has not changed much from 10 years ago.
 - ▶ All the papers we have written before 2009 (except technicolor). Keep looking for them!
- At the same time, new physics may show up in unexpected places.

Where is new physics

Where is new physics

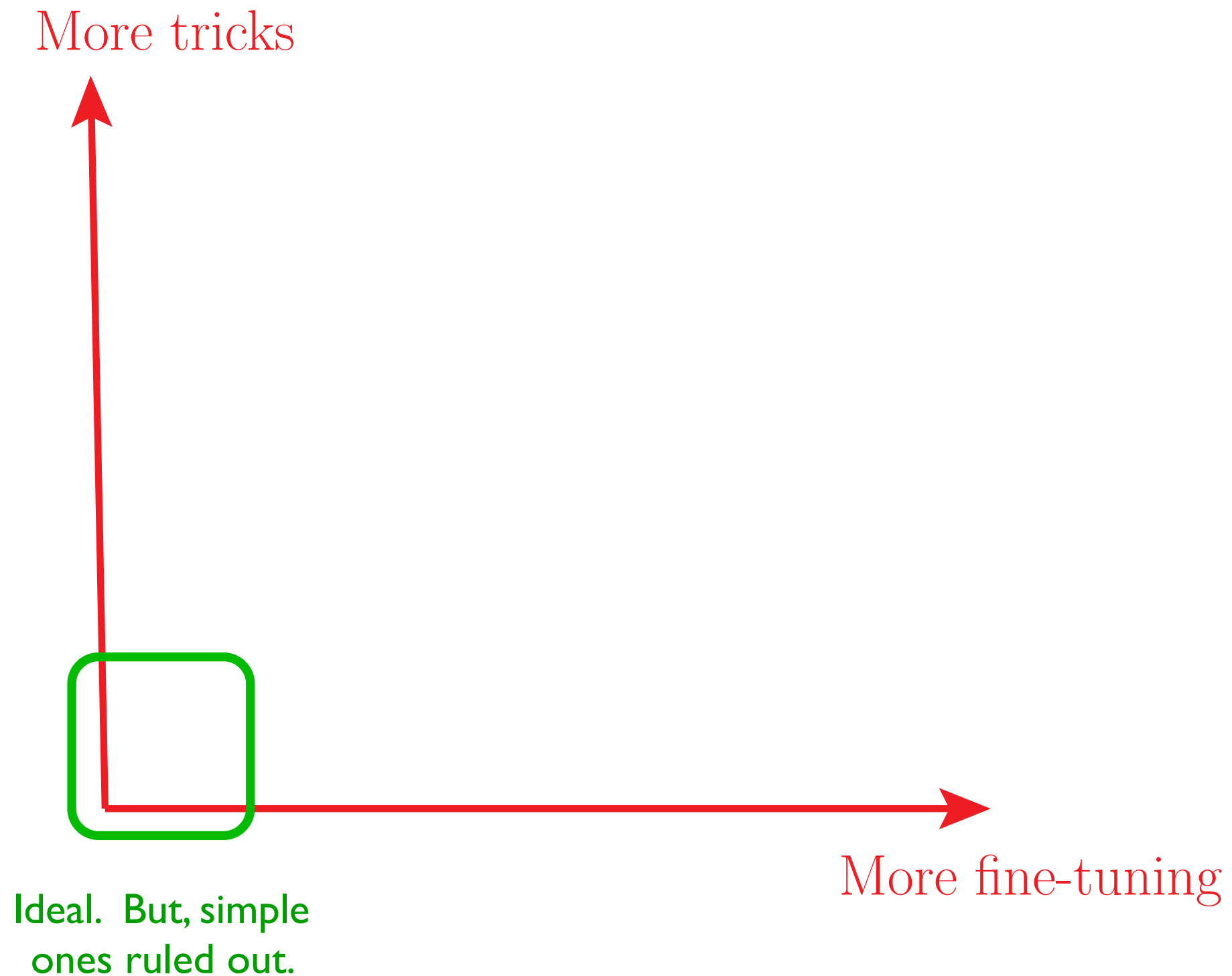
- We often ask (for the last 20 years): is X dead ($X \approx \text{SUSY}$)?

Where is new physics

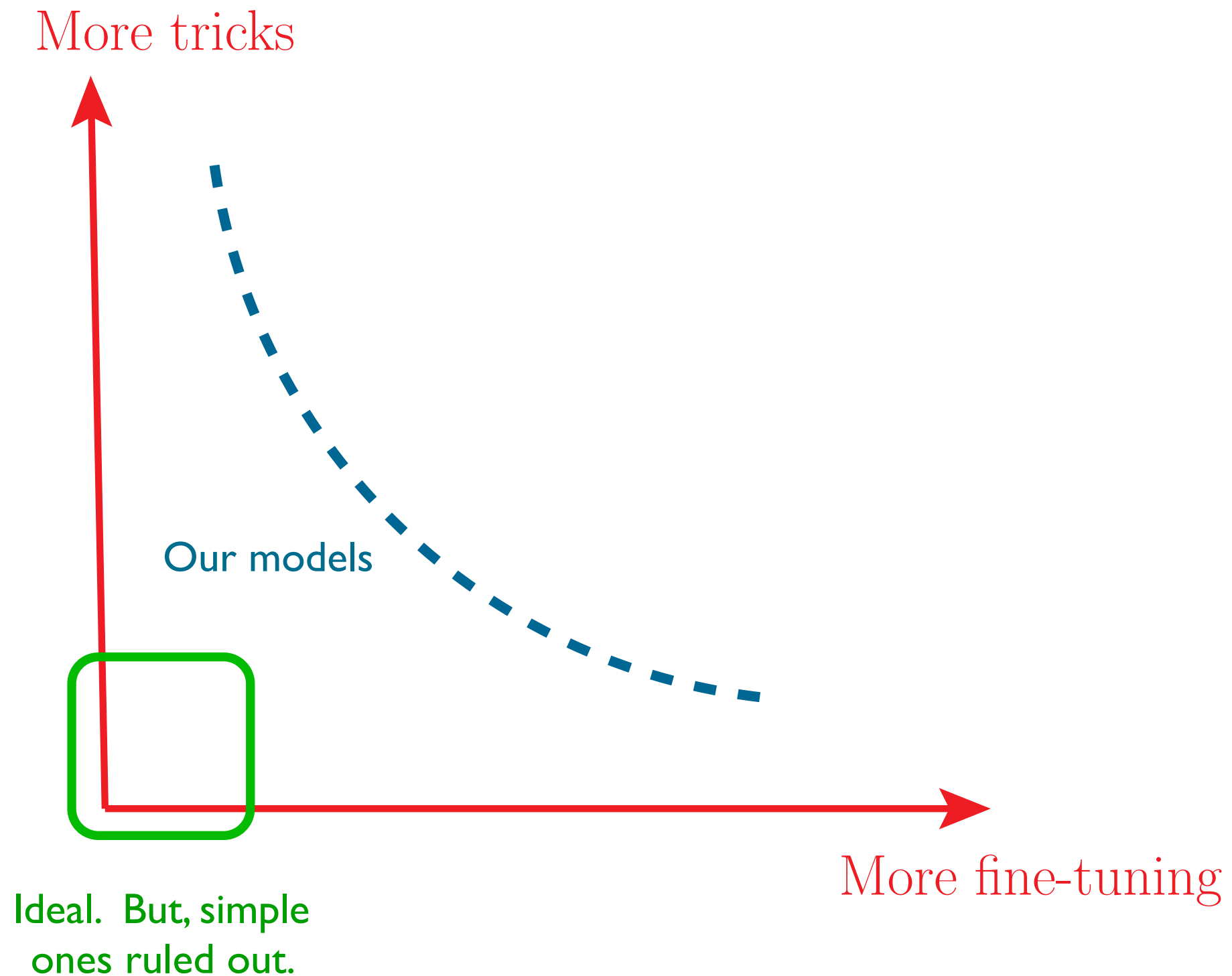
- We often ask (for the last 20 years): is X dead ($X \approx \text{SUSY}$)?
- The answer is of course: no.

New physics will never die, they just get heavier.

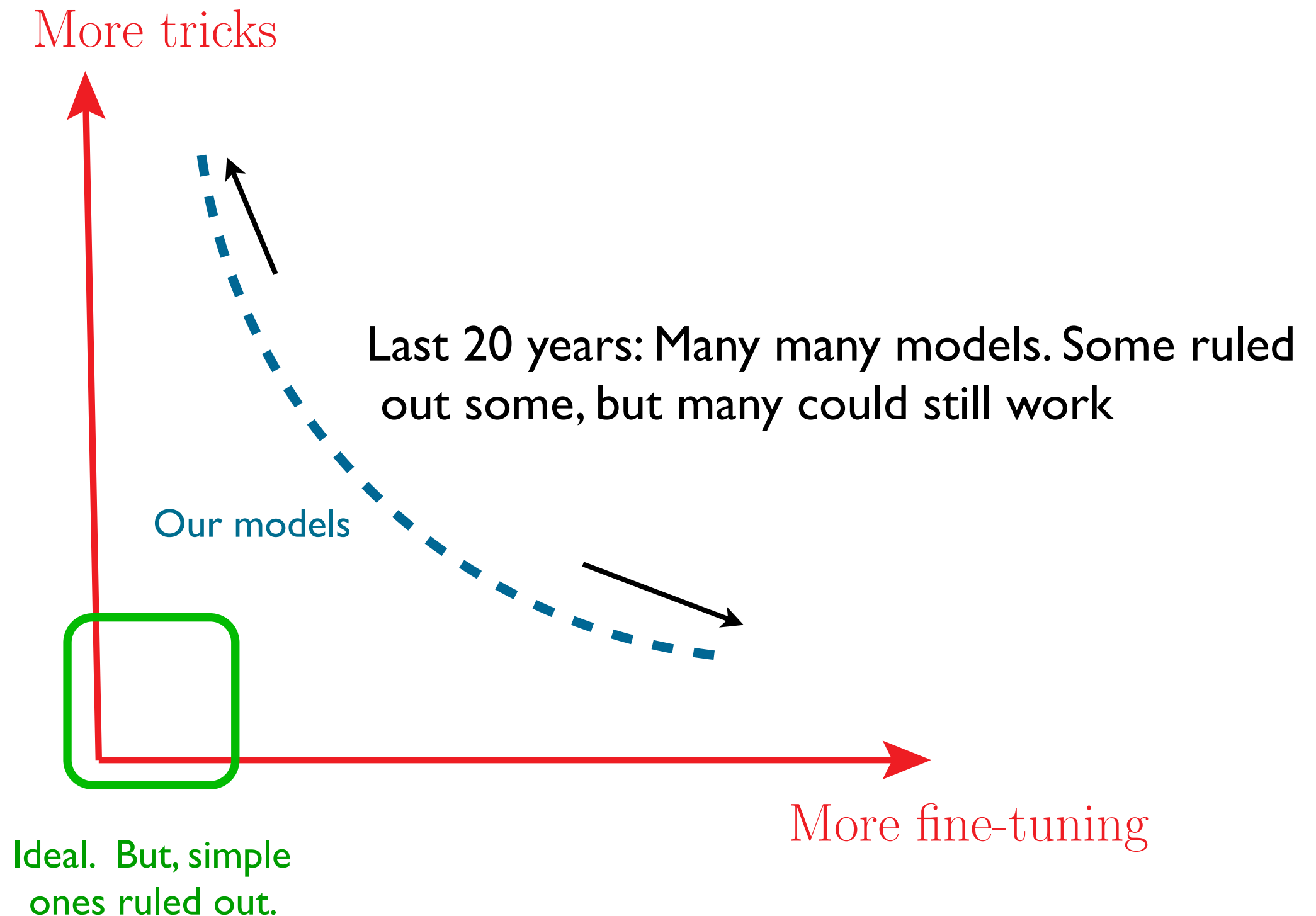
Where we are in theory space



Where we are in theory space



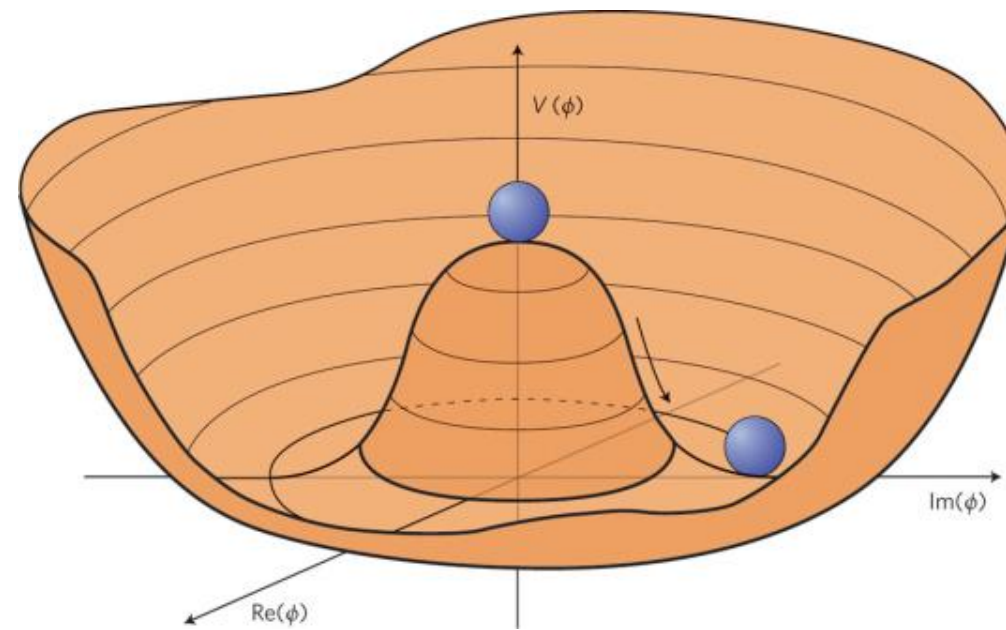
Where we are in theory space



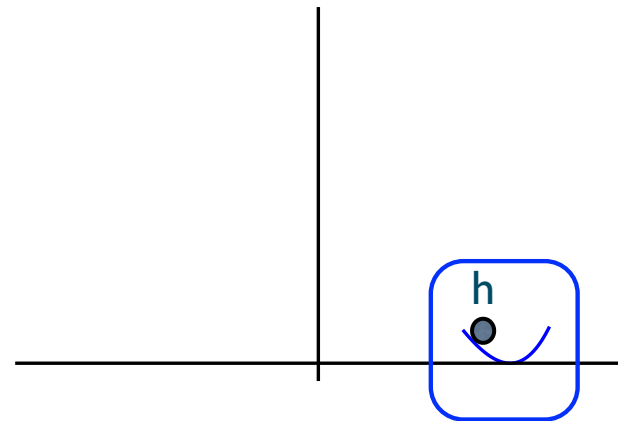
Main targets

1. Higgs and Naturalness

Mysteries of the Higgs boson



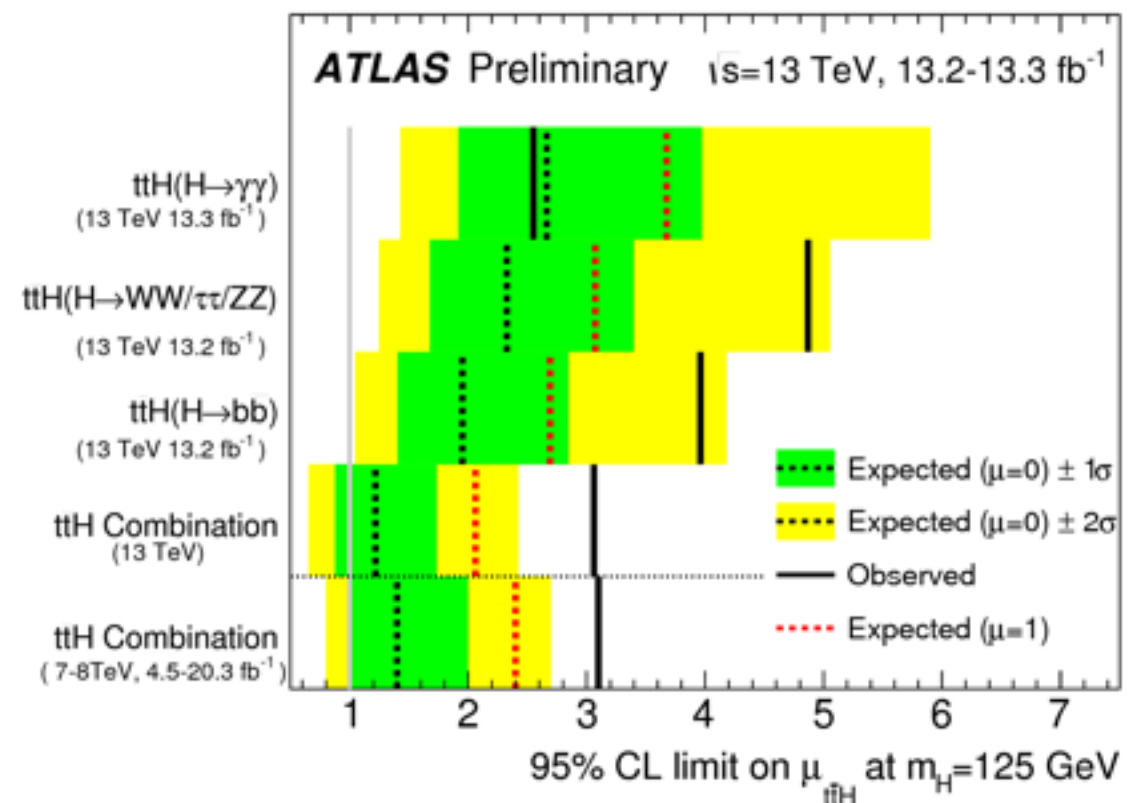
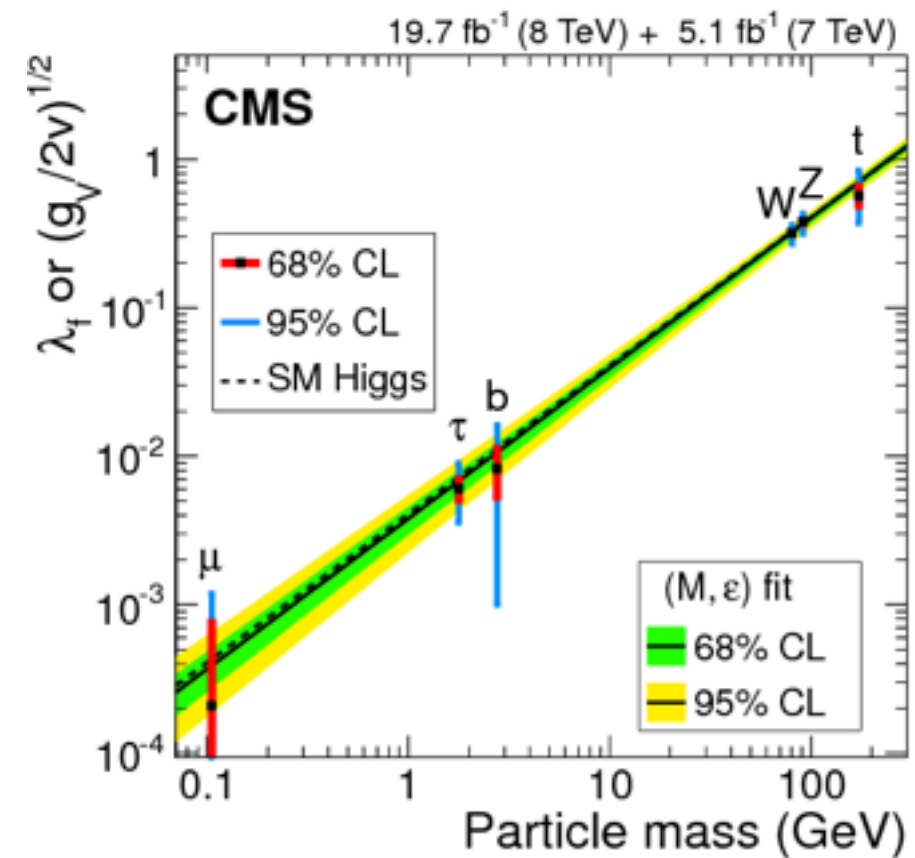
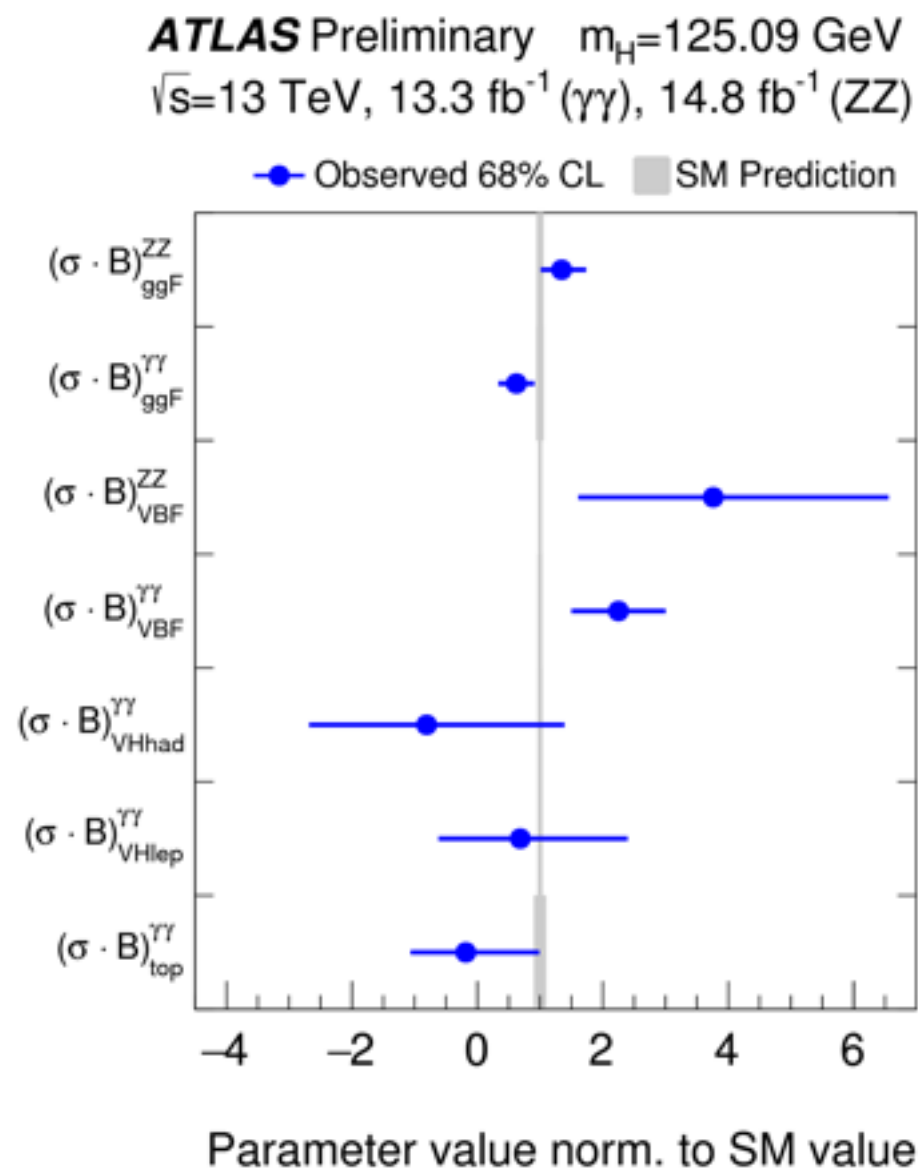
Mysteries of the Higgs boson



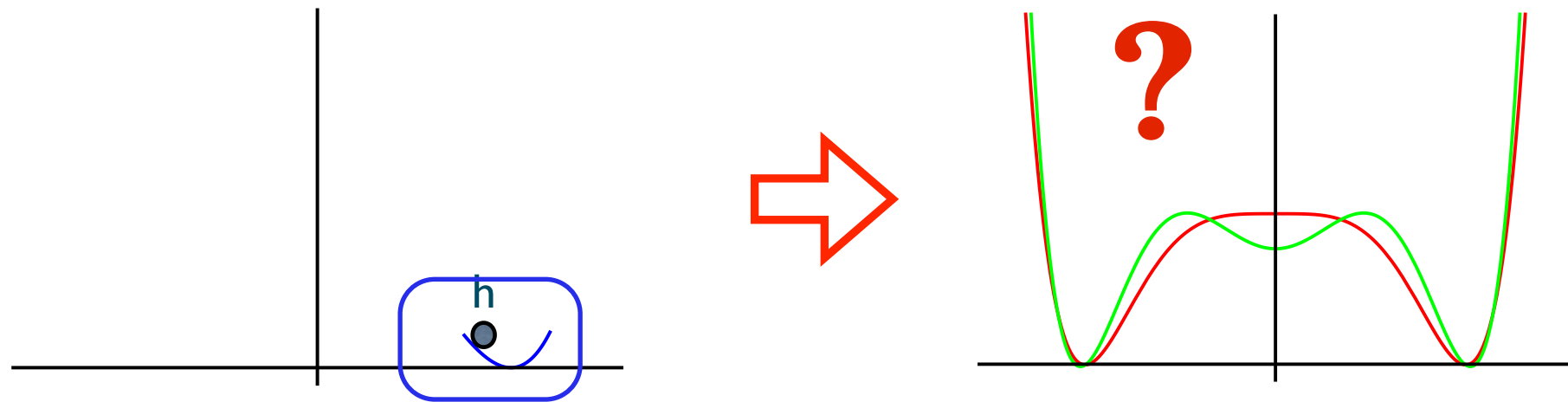
What we know now

- How to predict/calculate Higgs mass?
- What does the rest of the Higgs potential look like?
- Is it connected to the matter anti-matter asymmetry?

Understanding the Higgs better



Nature of EW phase transition



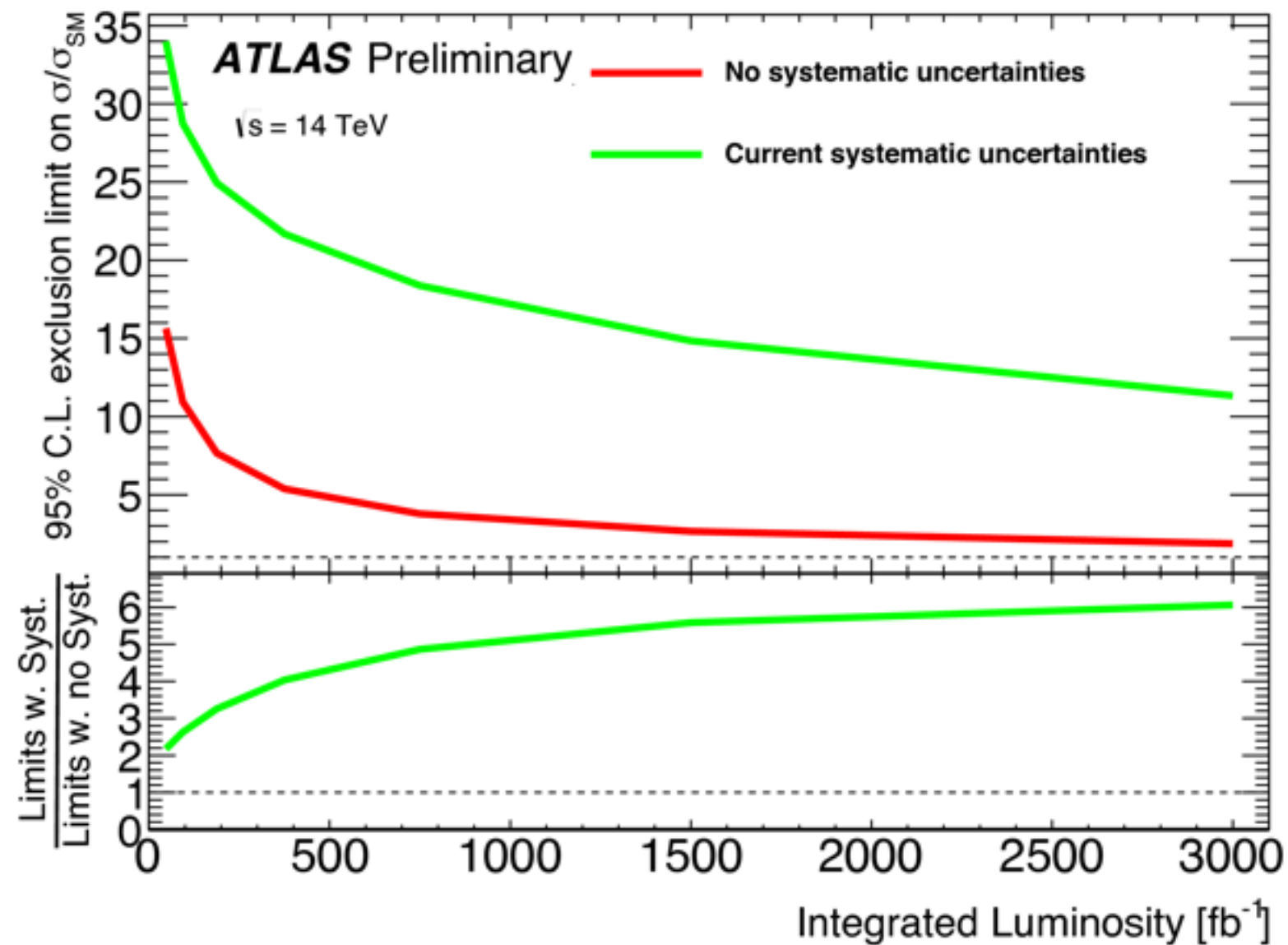
What we know from

Related to the question of matter anti-matter
asymmetry in the universe

- Generic predictions of non-minimal electroweak phase transition.

Order 1 deviation in triple Higgs

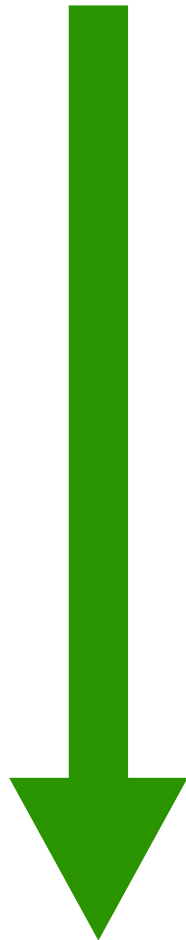
Probing self coupling with di-Higgs



Difficult. But we need to try harder.
Could give us surprises

Explaining EWSB: naturalness

..... The energy scale of new physics
responsible for EWSB



Electroweak scale, 100 GeV.
 m_h , m_W ...

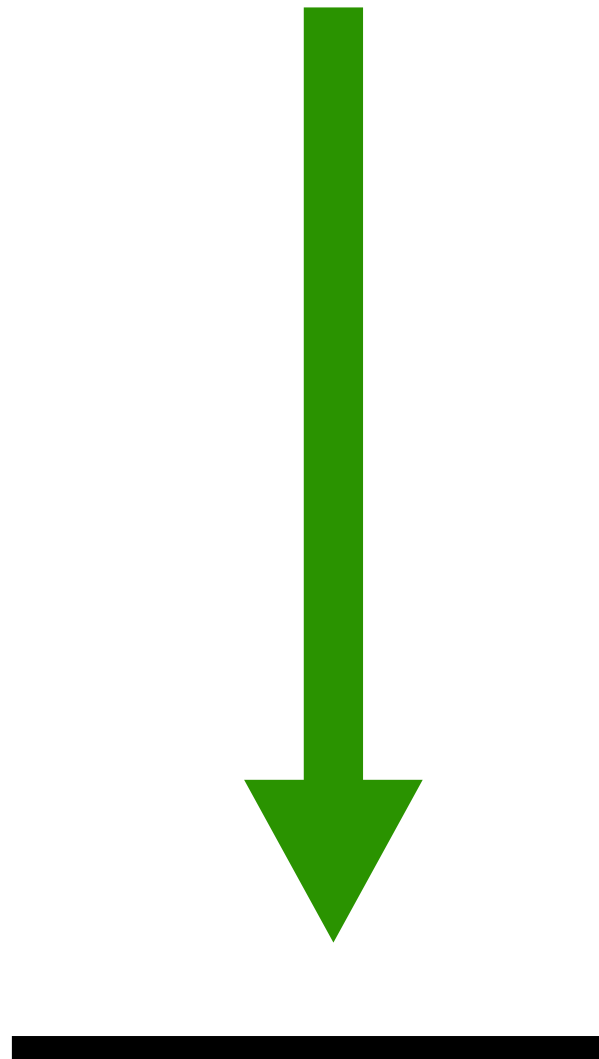
Explaining EWSB: naturalness

..... The energy scale of new physics
responsible for EWSB

What is this energy scale?

$M_{\text{Planck}} = 10^{19} \text{ GeV}, \dots?$

If so, why is so different from 100 GeV?



Electroweak scale, 100 GeV.

$m_h, m_W \dots$

Naturalness of electroweak symmetry breaking

• • • • •



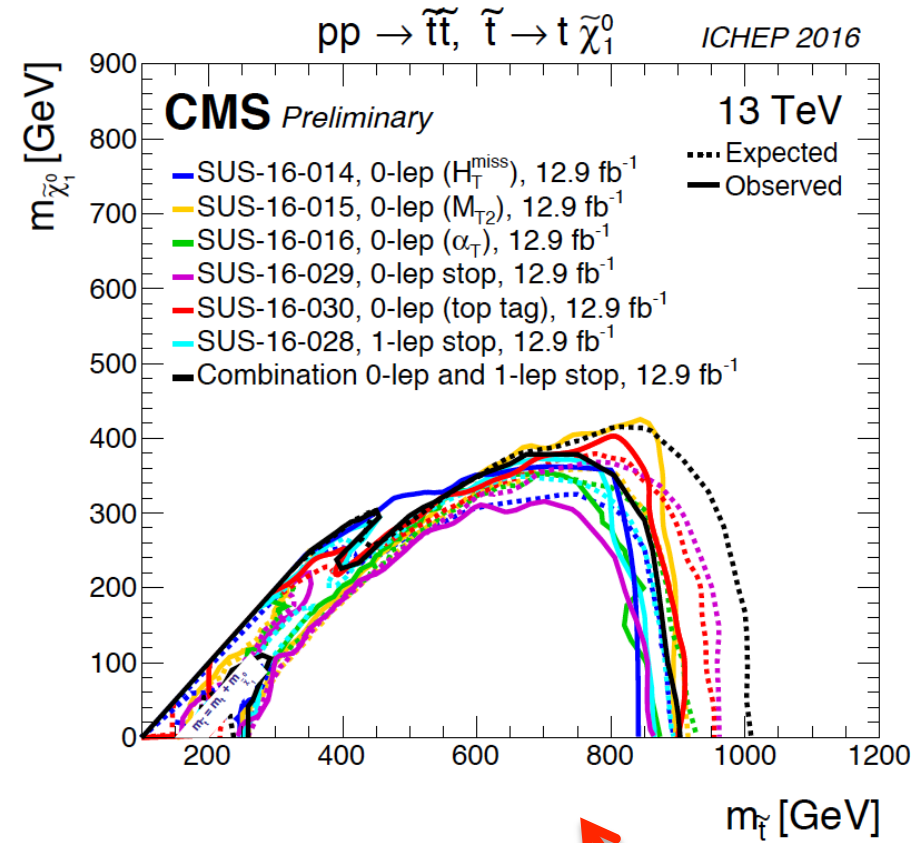
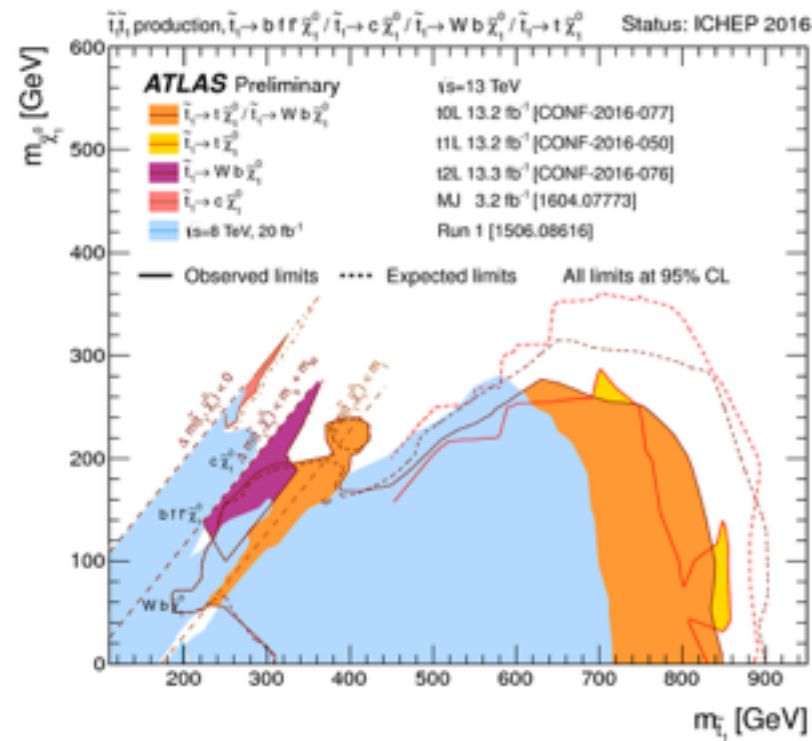
The energy scale of new physics
responsible for EWSB

TeV new physics.
Naturalness motivated



Electroweak scale, 100 GeV.
 m_h , m_W ...

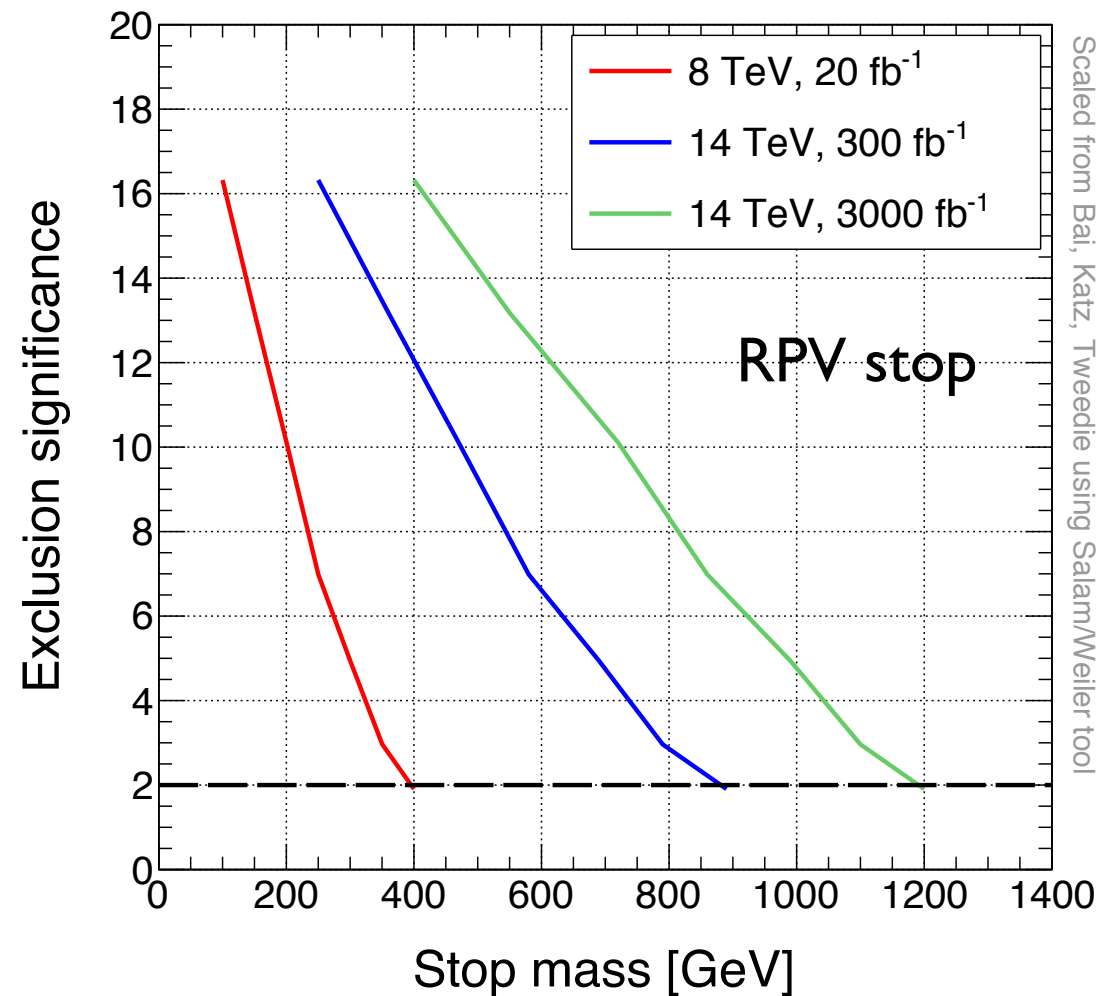
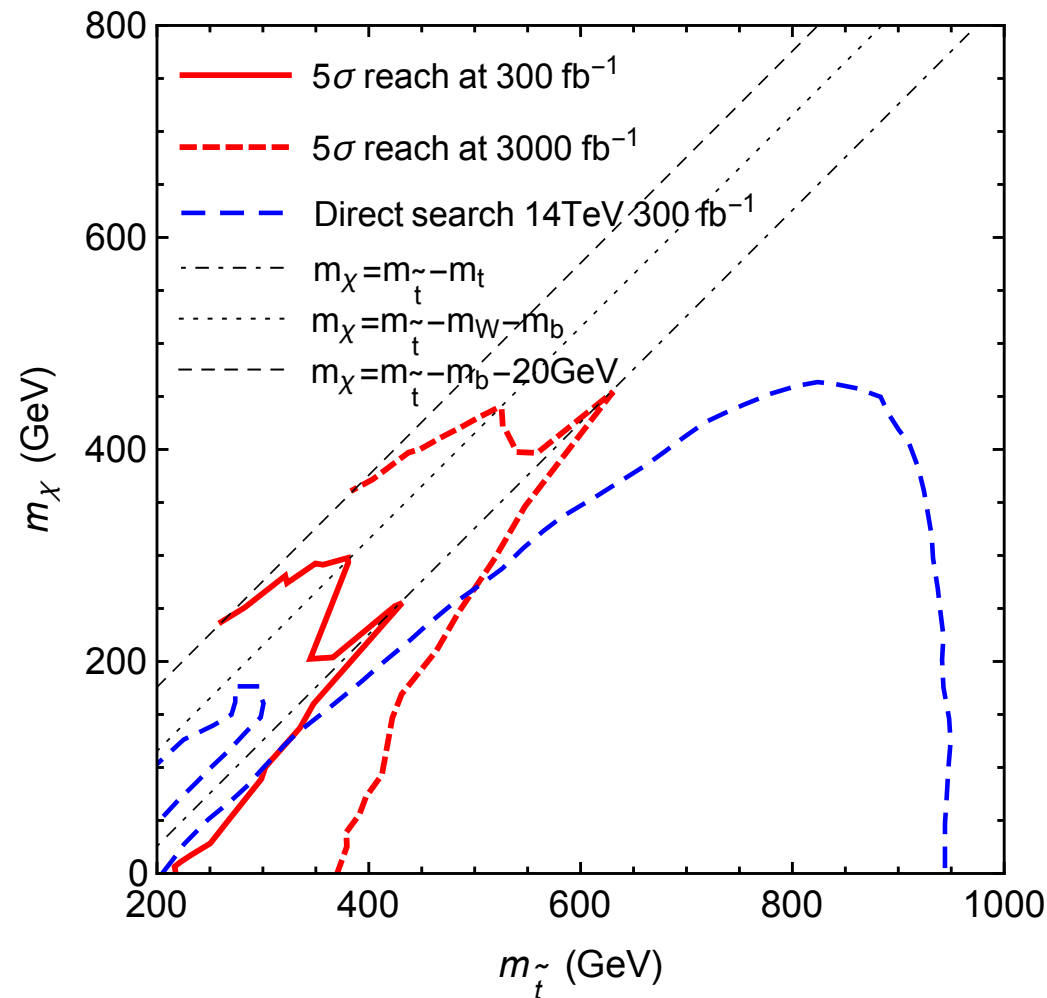
For example: top partner (SUSY)



- Naturalness $\propto m_{\text{stop}}^2$
- LHC will teach us a lot about naturalness!

Cover gaps, look for alternatives

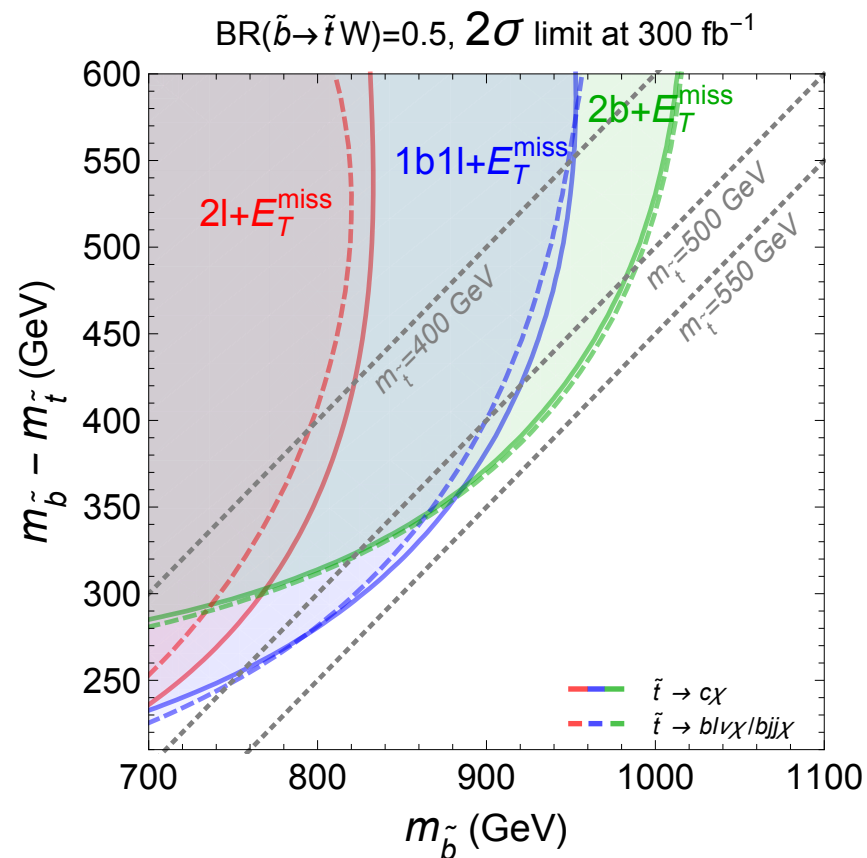
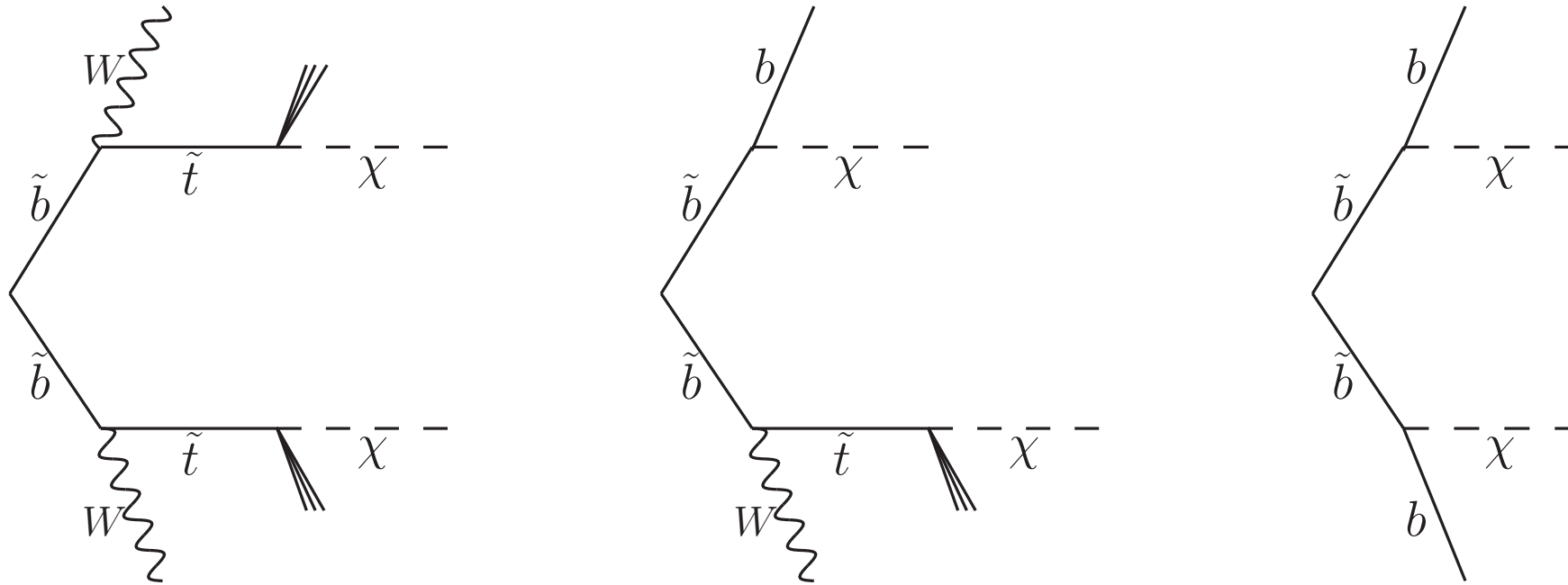
An and LTW, I506.00653



- Compressed, stealth, RPV...
- Run 2 will leave no stone unturned.

Sbottom helps stop search

Haipeng An, Jayin Gu and LTW



Promising reach
for stop nearly
degenerate with LSP

My favorite channel.

Effective, Mini-split, spread, zprime-mediation, ...

Heavy squarks $m_{\tilde{t},\tilde{b}} < m_{\tilde{u},\tilde{d}}$

10s - 100s TeV

$\tilde{u}, \tilde{d}, \dots$
 $\tilde{t}, \tilde{b}$

Light gaugino

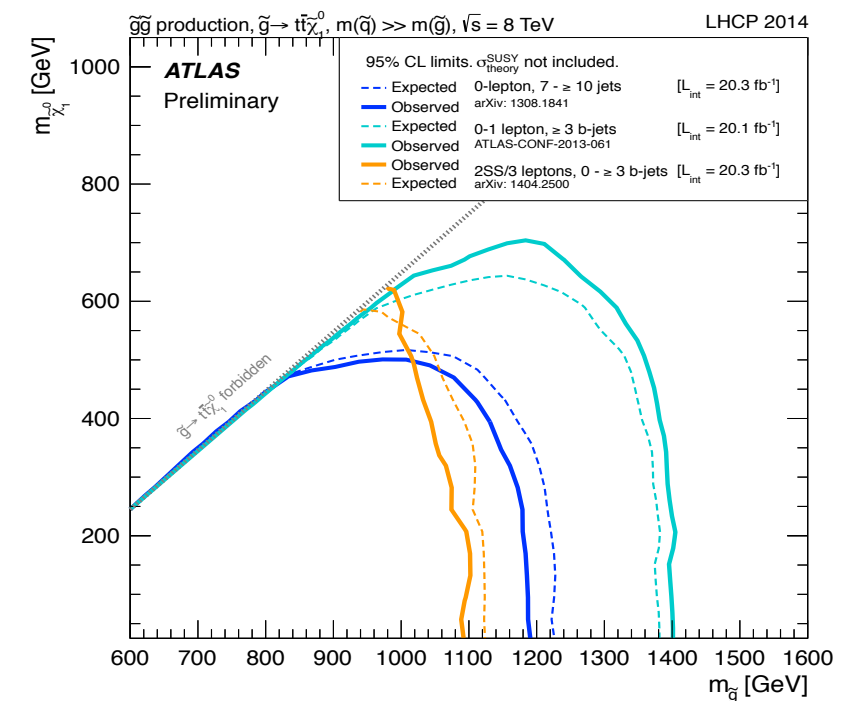
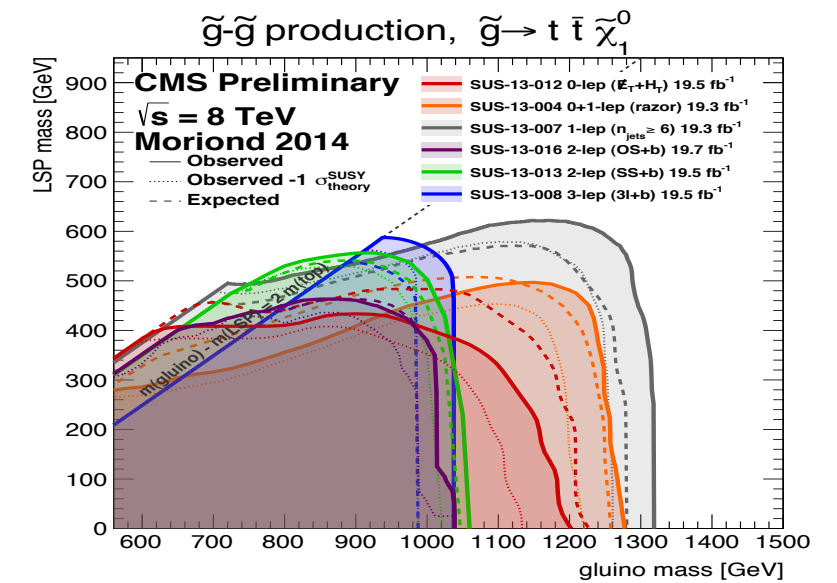
TeV-ish

$\tilde{g}$
 $\tilde{N} \tilde{C}^\pm$

Fermionic partners still tend to be light.

LHC signal

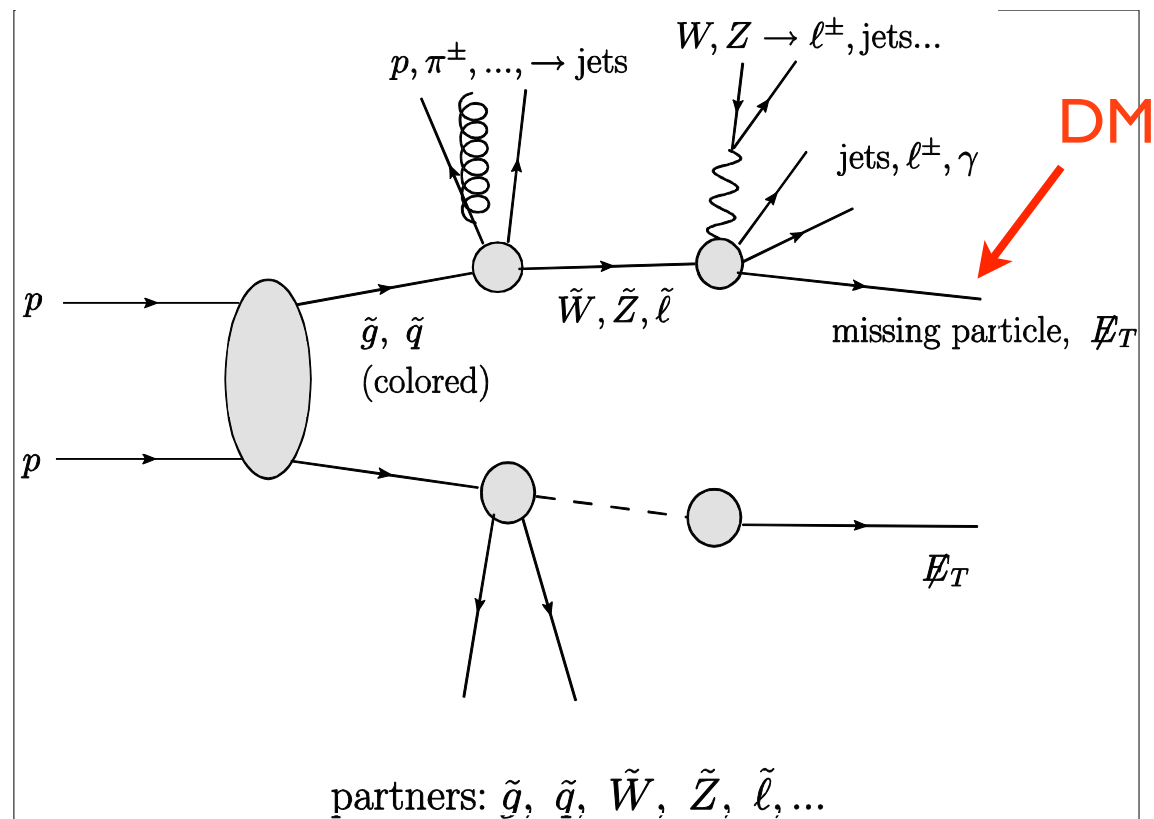
$pp \rightarrow \tilde{g}\tilde{g} \rightarrow t\bar{t}t\bar{t}, t\bar{t}b\bar{b}, t\bar{t}t\bar{b} \dots$



Main targets

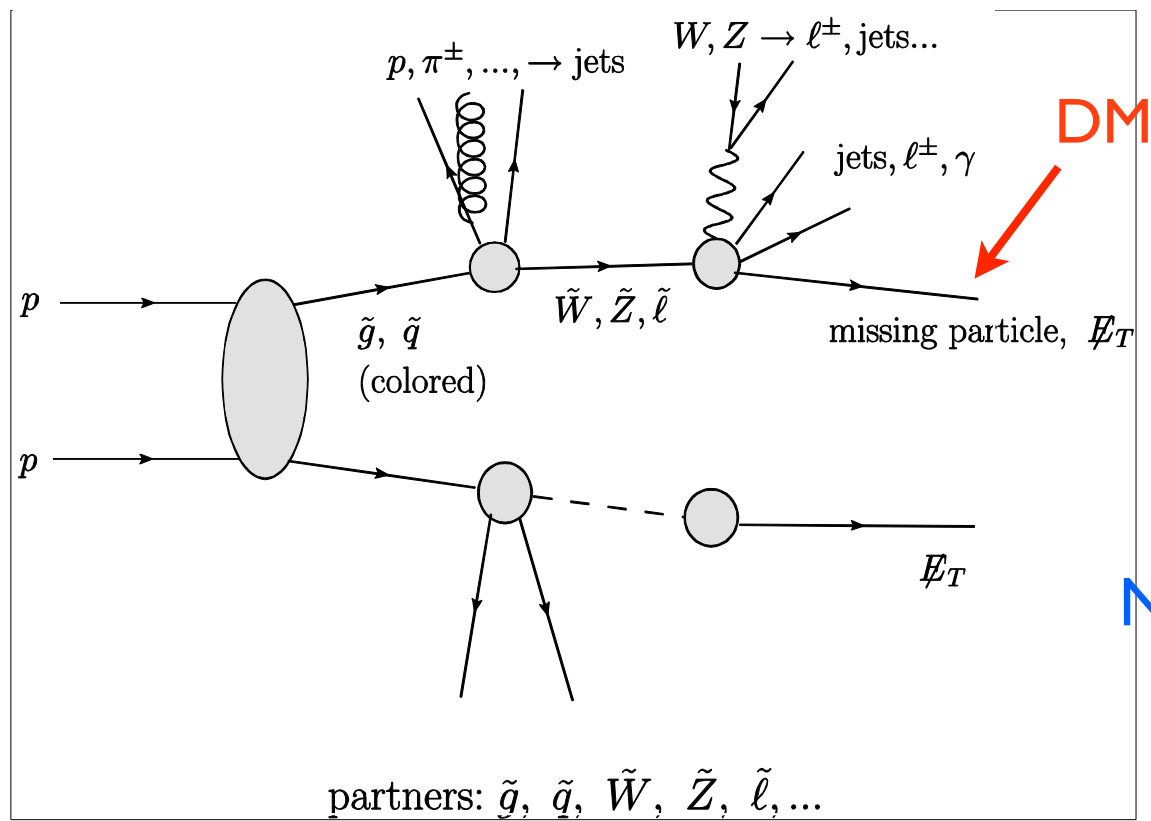
2. dark matter

“standard” story.

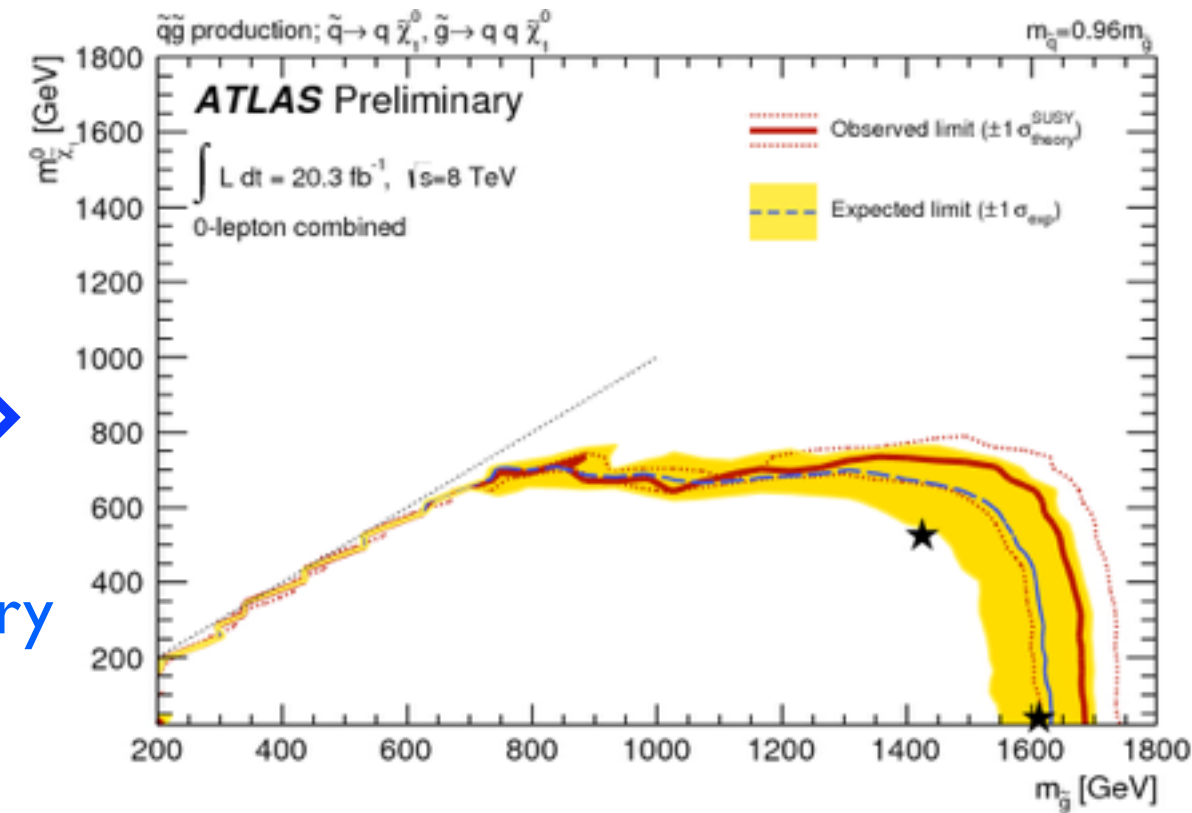


- WIMP is part of a complete model at weak scale.
- It's produced as part of the NP signal, shows up as missing energy.

“standard” story.

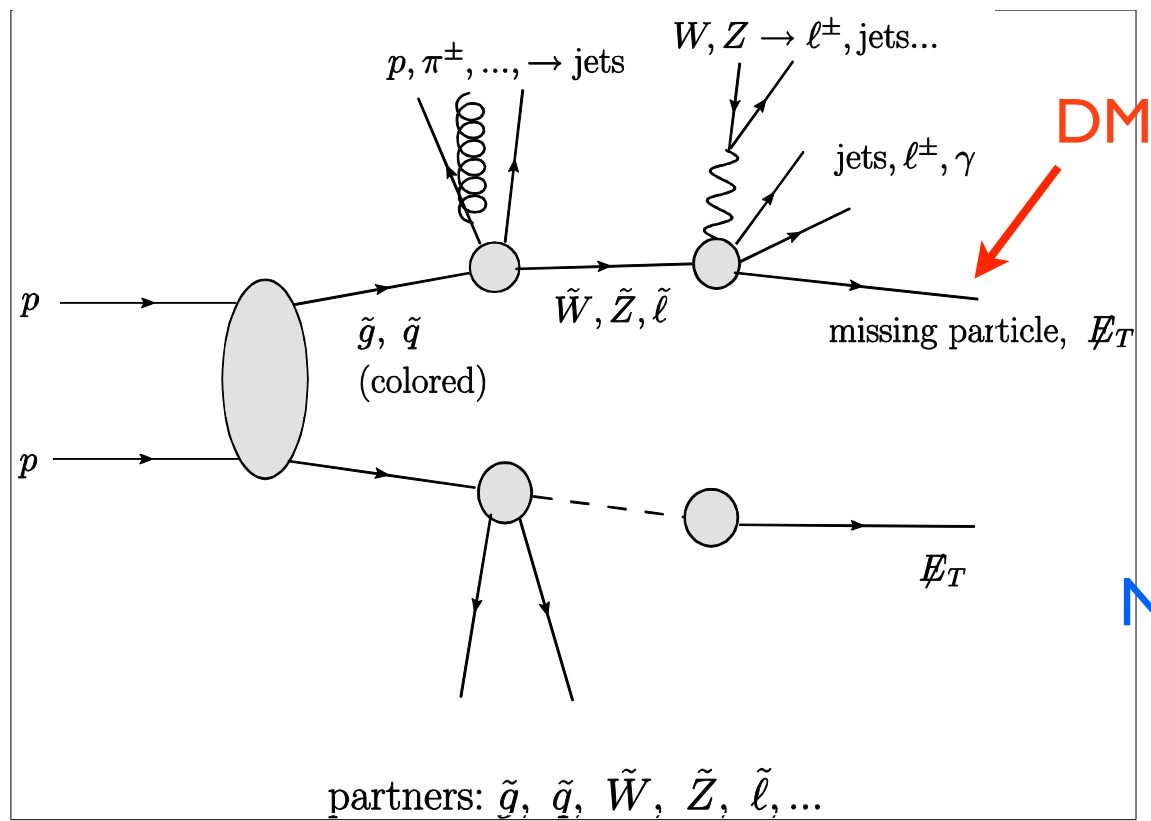


No discovery yet

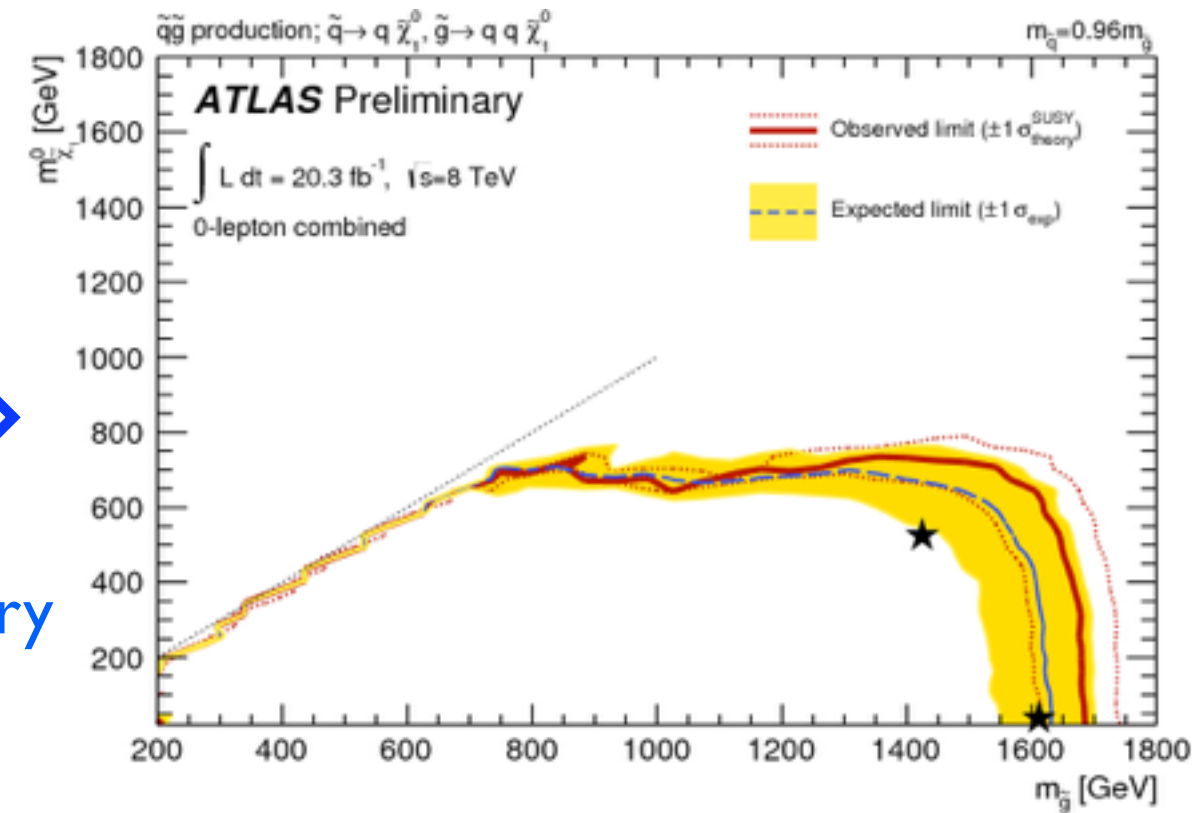


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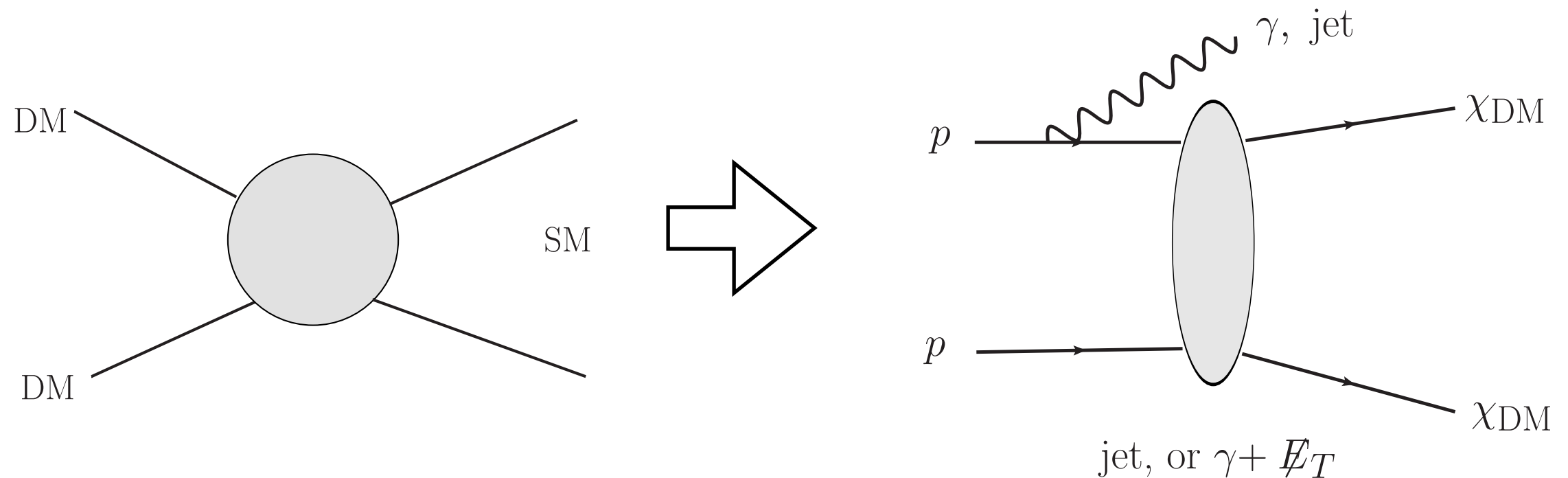
No discovery
yet



Of course, most WIMP parameter space not covered yet.
Still plausible at the LHC, will keep looking.

Focus more on basic channel

- pair production + additional radiation.



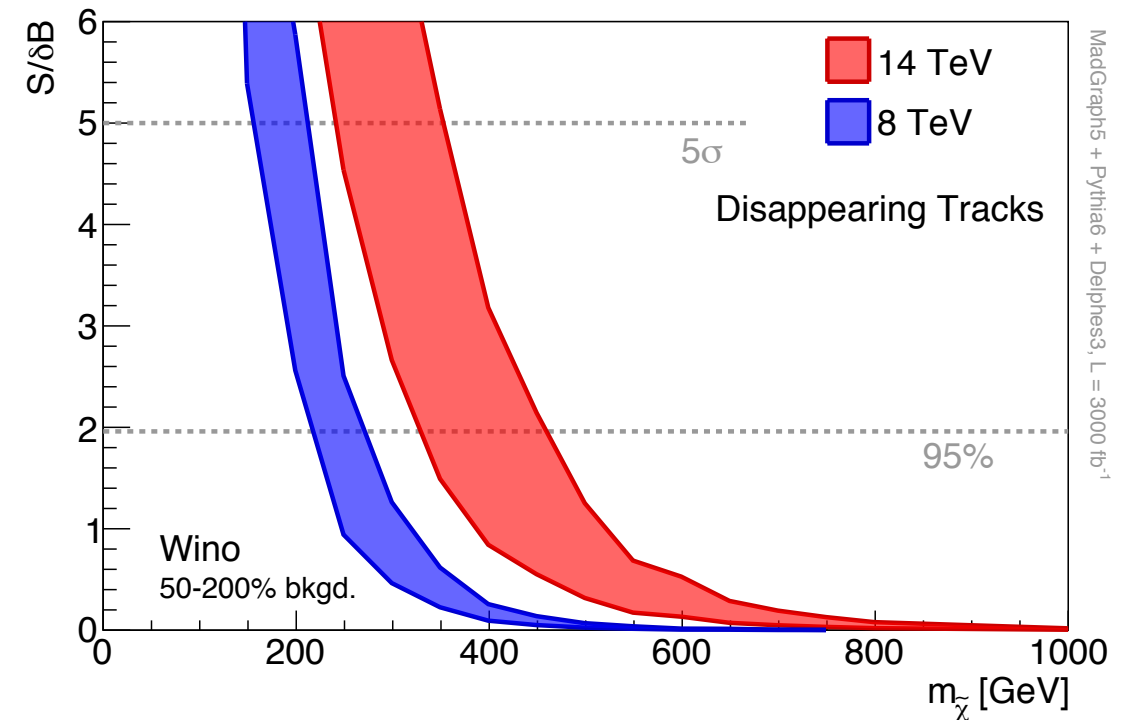
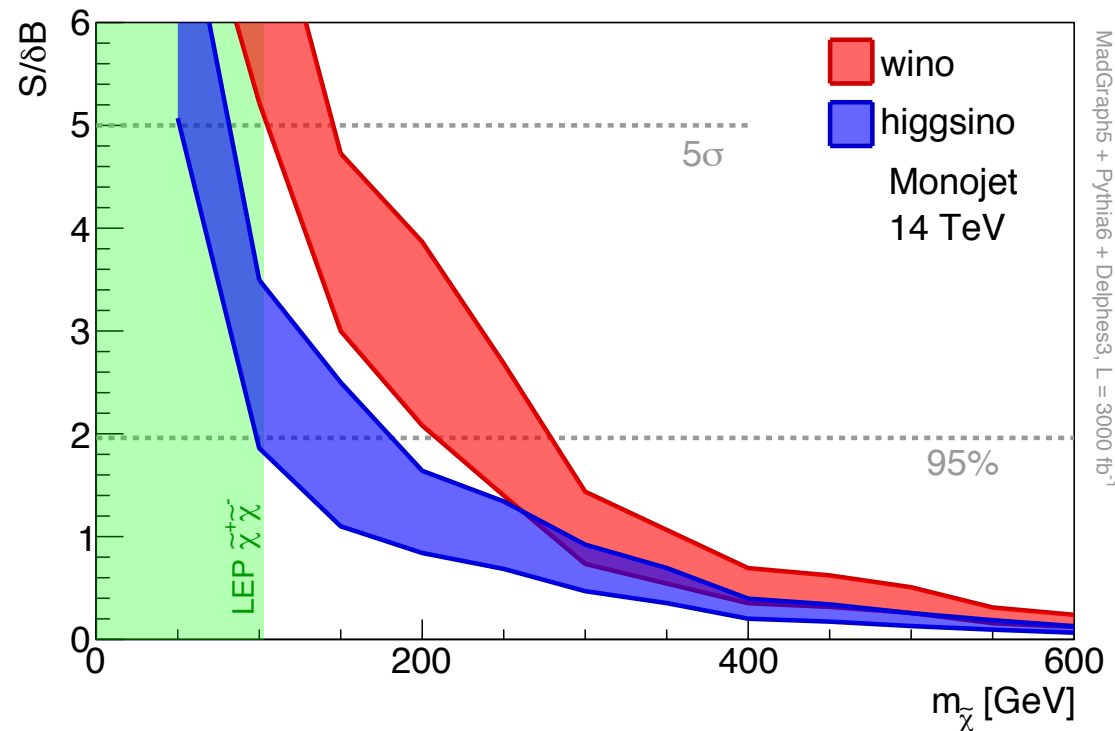
- Mono-jet, mono-photon, mono-Higgs...
- Have become “Standard” LHC searches.

Beltran, Hooper, Kolb, Krusberg, Tait, 1002.4137

Goodman, Ibe, Rajaraman, Shepherd, Tait, Yu, 1005.1286

Bai, Fox, Harnik, 1005.3797

Focusing on more basic channel

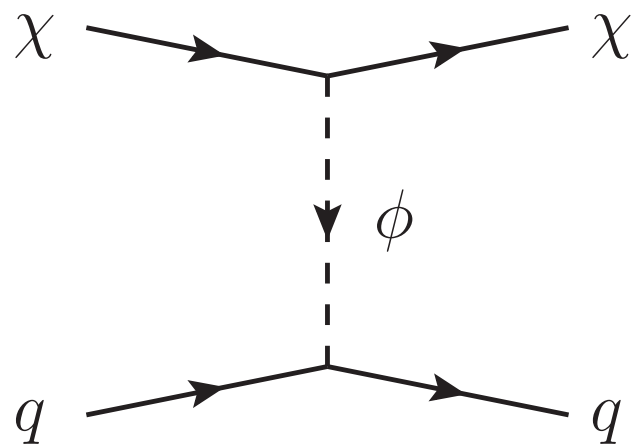


- Very challenging. Systematics dominated
 - Weak discovery reach at 14 TeV, 3 ab^{-1} .
- Disappearing track promising!

Consider new mediators

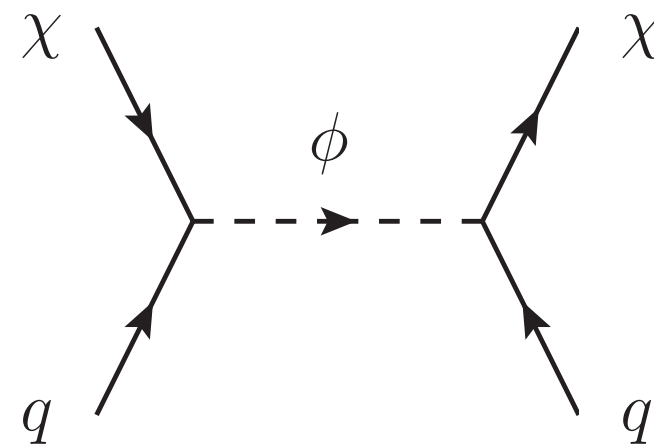
direct detection $\rightarrow$

collider detection $\uparrow$



s - channel

ϕ can be scalar or Z'



t - channel

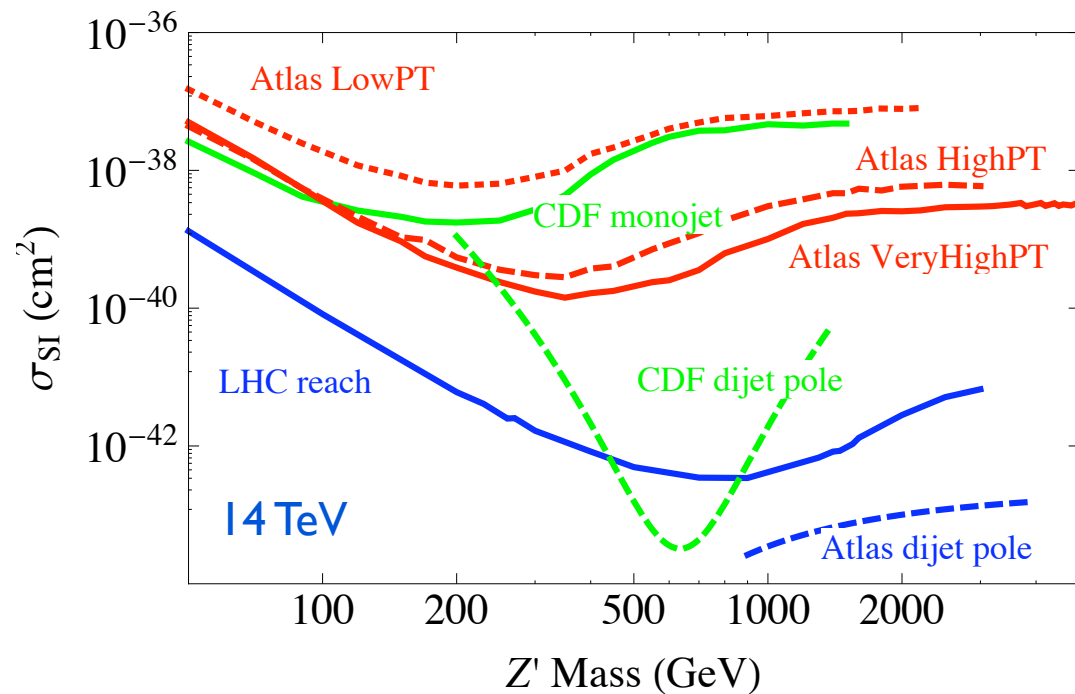
ϕ squark like

indirect detection $\rightarrow$

MonoX vs direct search for mediator

An, Ji, LTW, I202.2894

s-channel

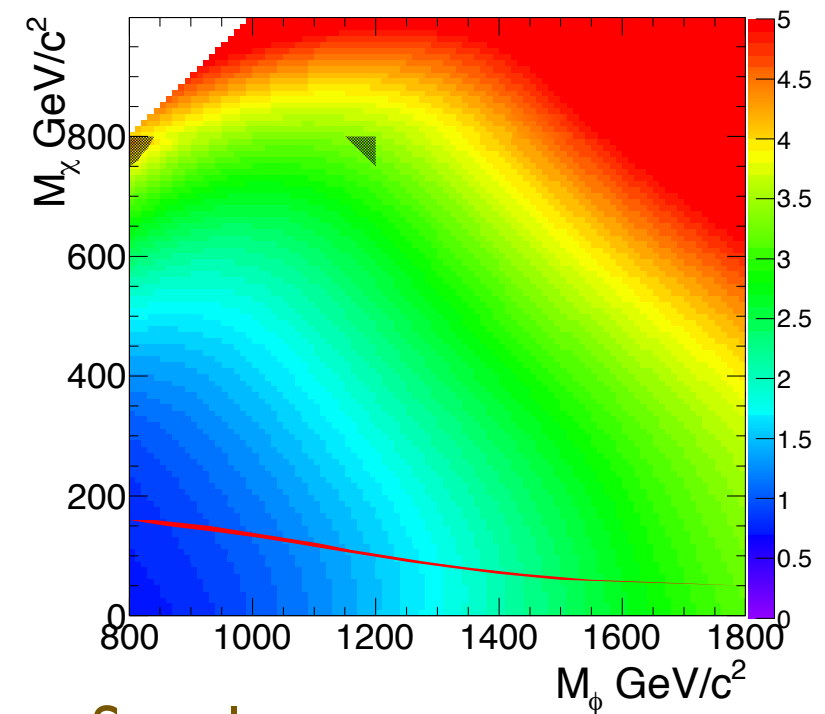


See also

Busoni, De Simone, Jacques, Morgante, Riotto, I405.3101

Haipeng An, Hao Zhang, LTW, I308.0592

t-channel



See also:

Chang, Edezhath, Hutchinson, Luty, I307.8120

Bai, Berger, I308.0612

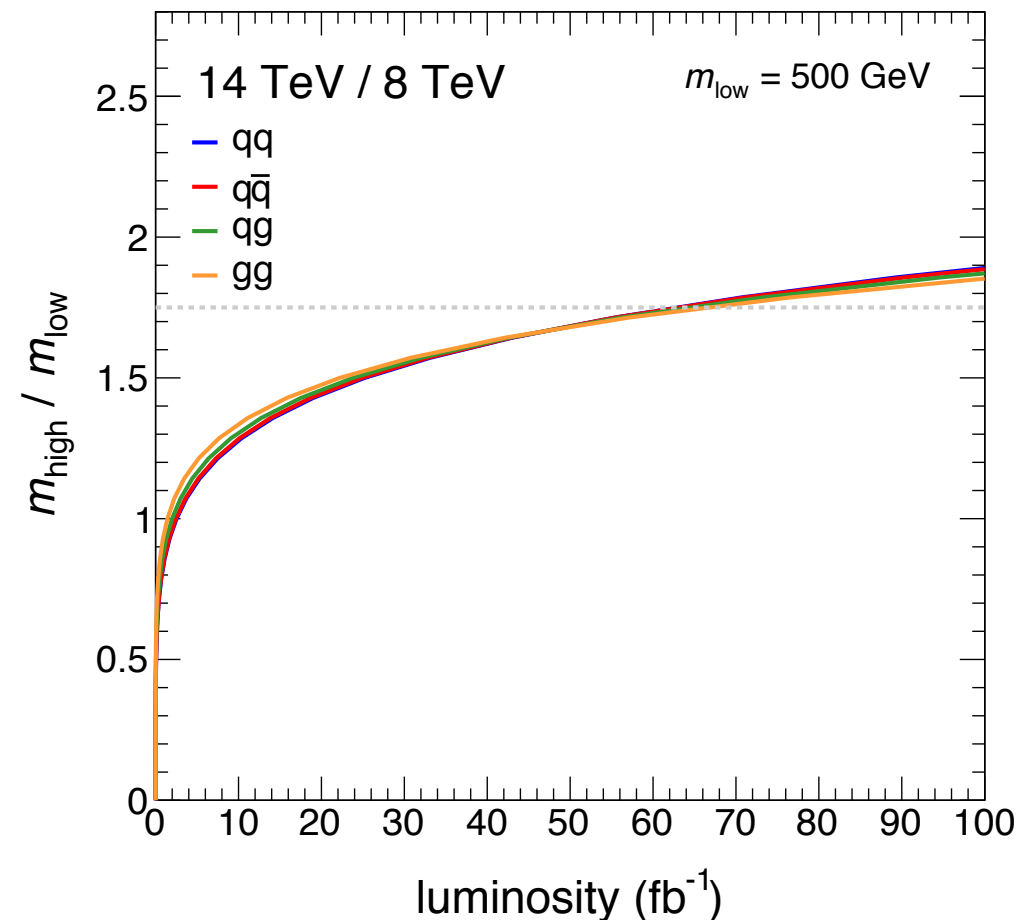
DiFranzo, Nagao, Rajaraman, Tait, I308.2679

Papucci, Vichi, Zurek, I402.2285

For a “to-do-list”, see dark matter forum publication: I507.00966

Channel with most “growth” potential

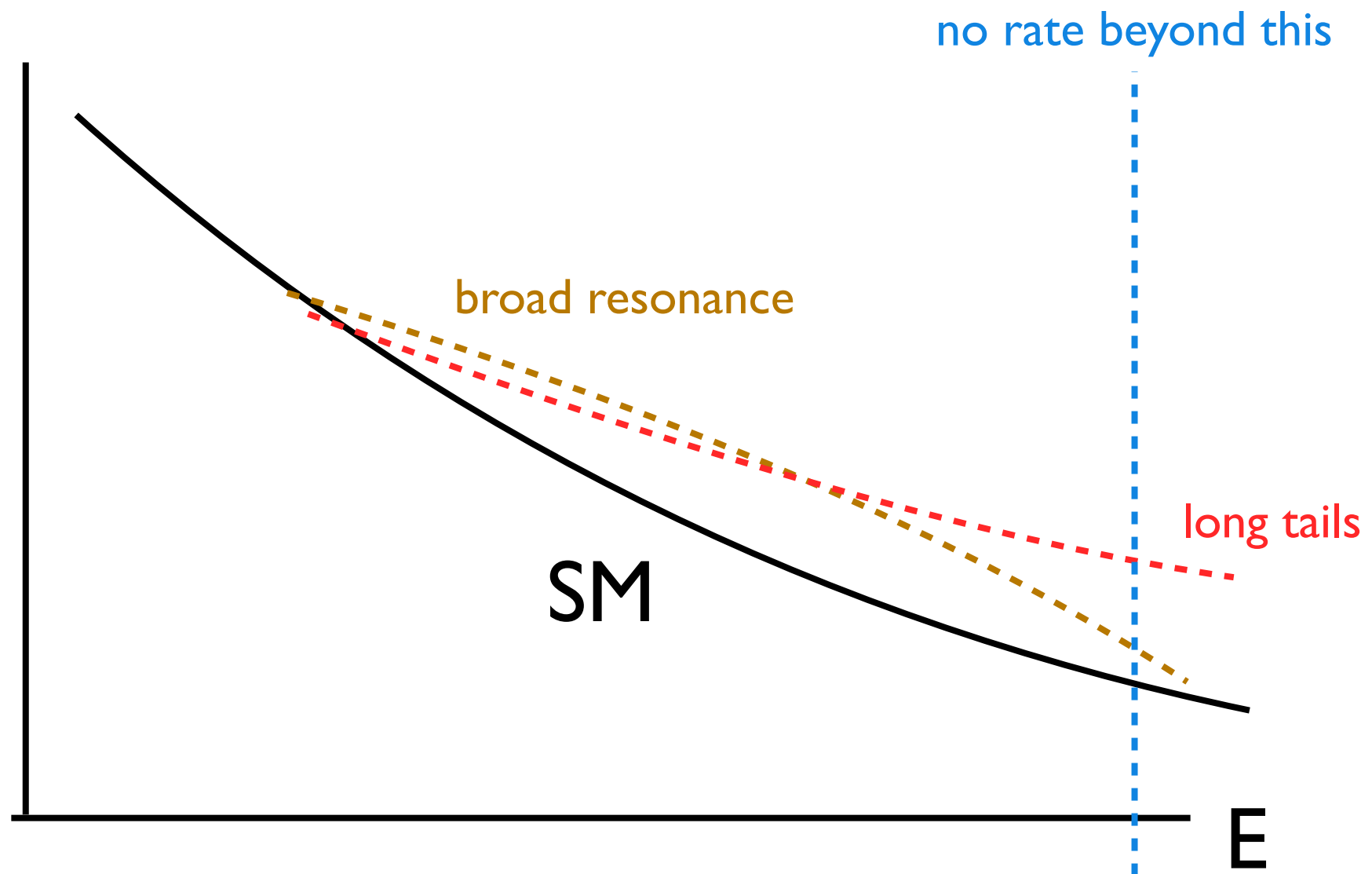
500 GeV, e.g. pair of 250 GeV electroweak-ino



Weakly coupled, back ground limited (small S/B)

Examples: electroweak-ino pair production
dark matter. Weak boson fusion, etc.

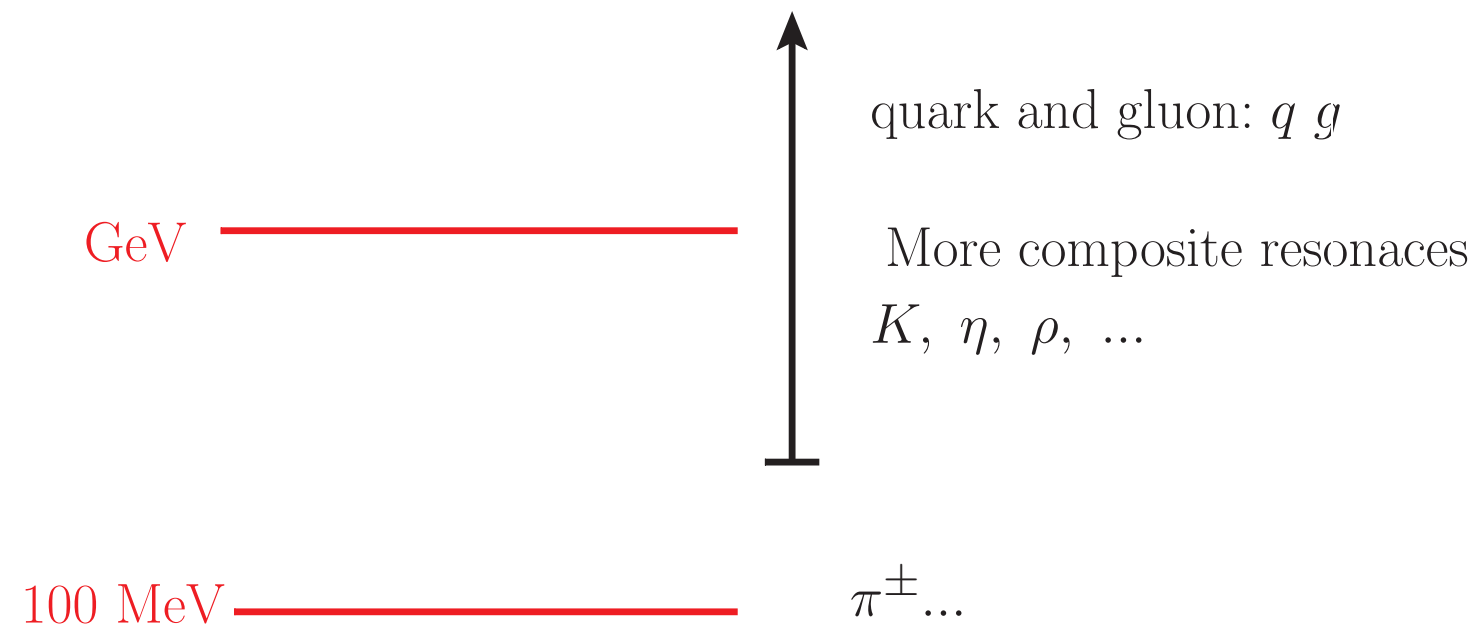
Example



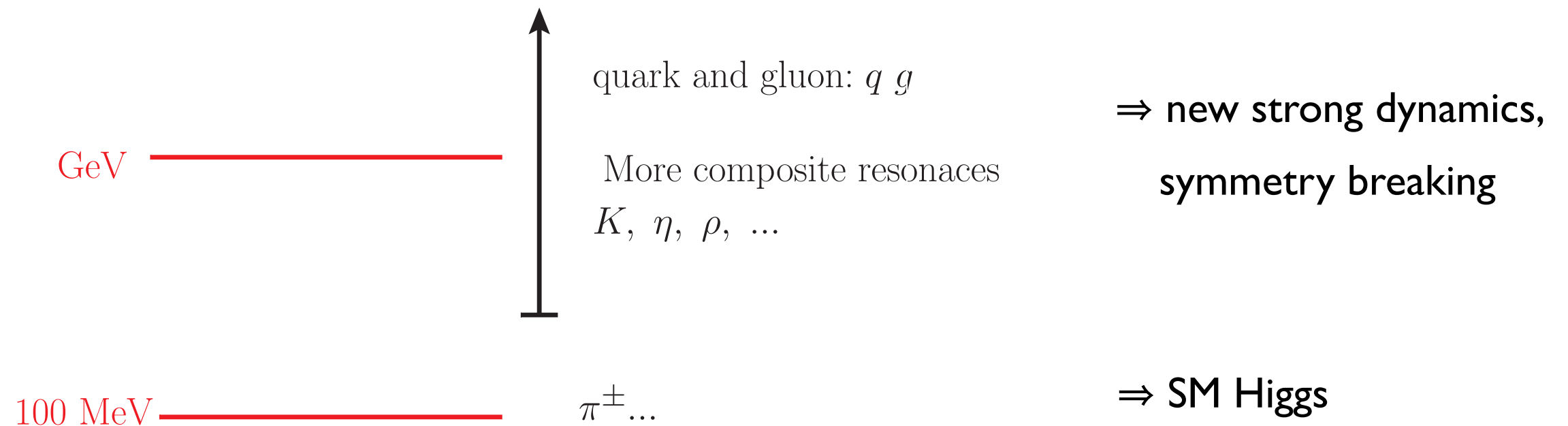
— Motivation? Strongly coupled heavy new physics

e.g. Greco, Liu, Pomarol, Rattazzi

“Learning” from QCD

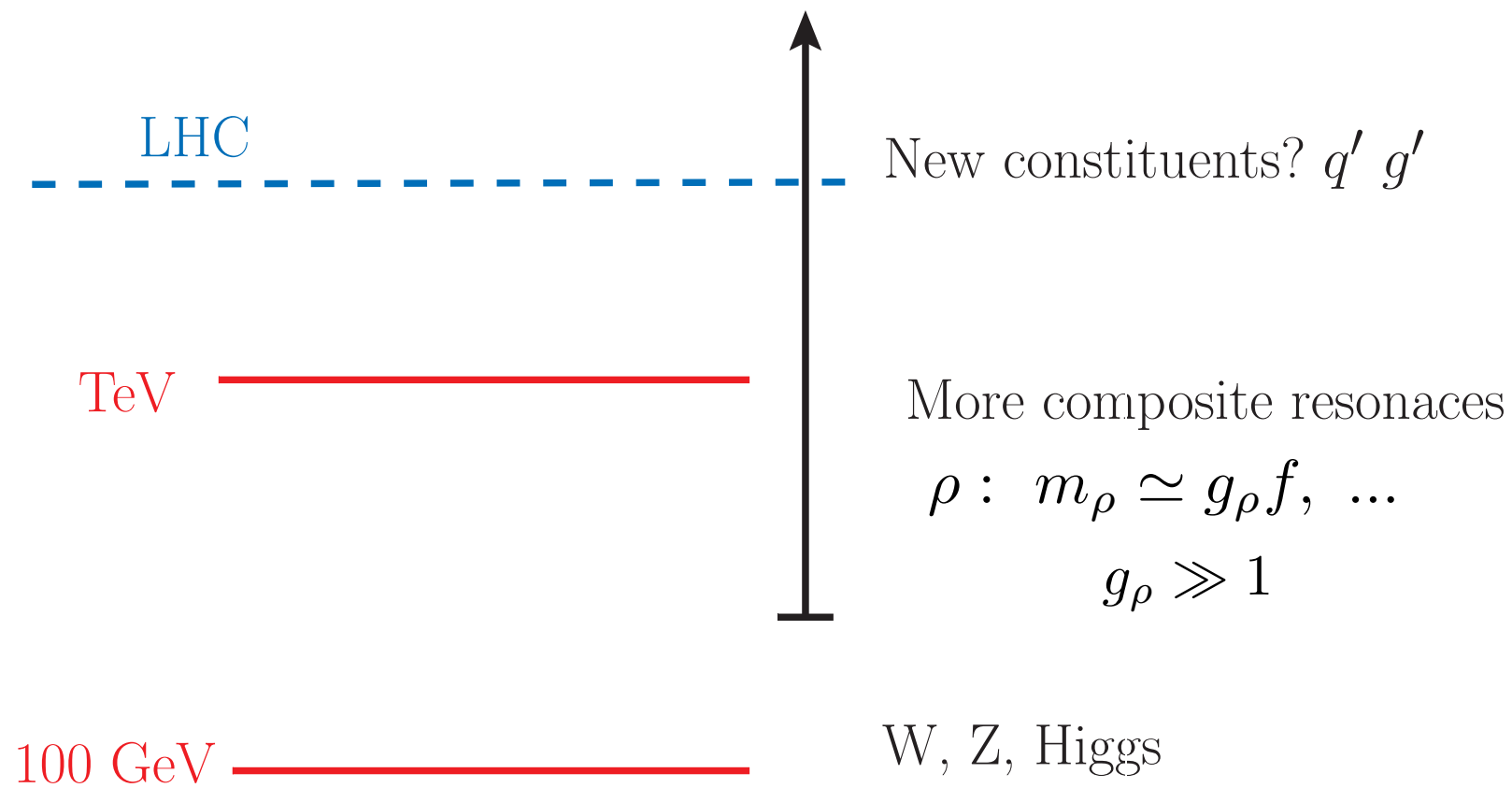


“Learning” from QCD



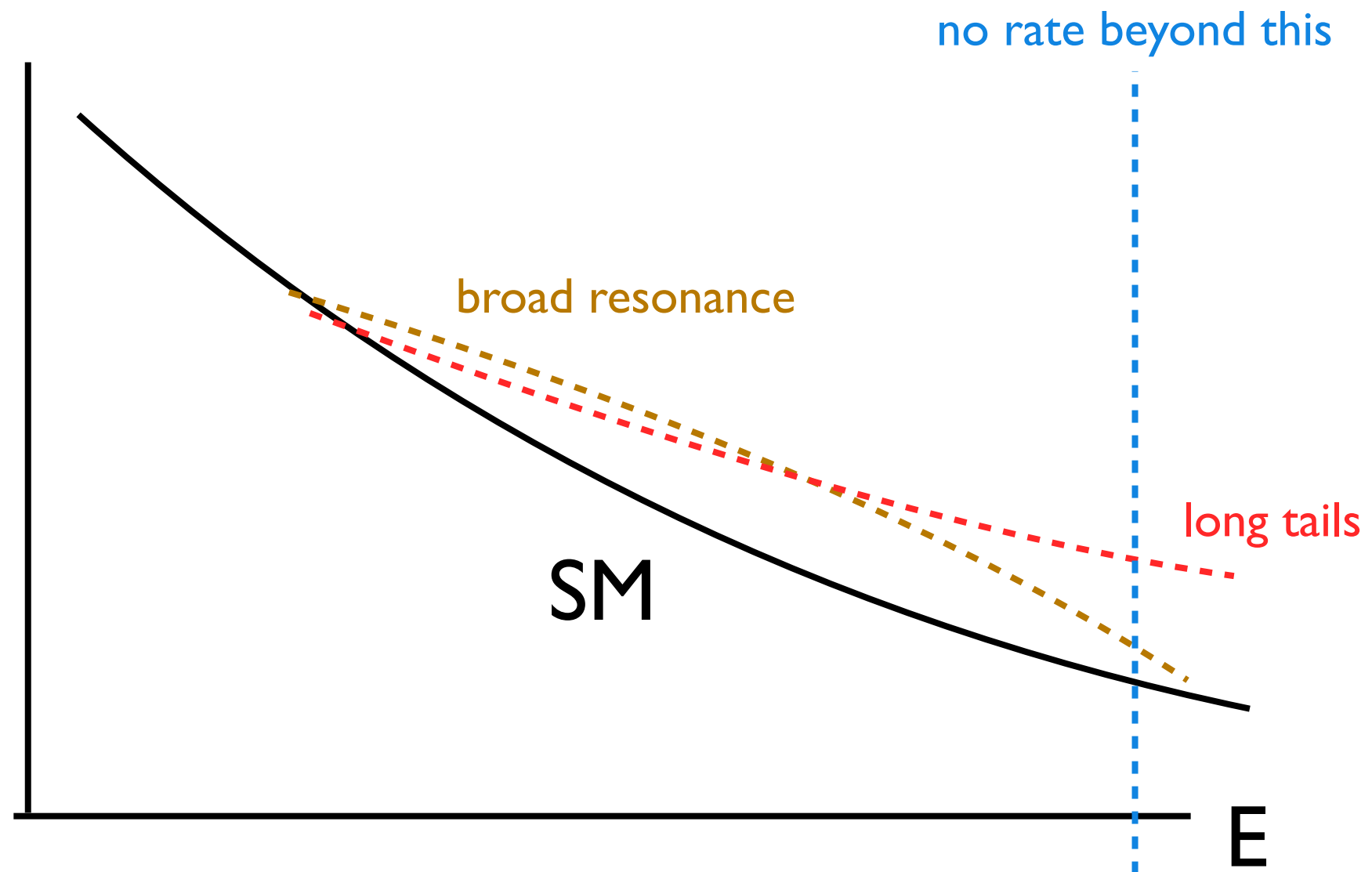
- Construct a new strong dynamics in which the low lying states will be the SM Higgs.
- Composite Higgs models. Still a natural theory.
- Nature may be more interesting, but it could also just repeat itself.

Composite Higgs



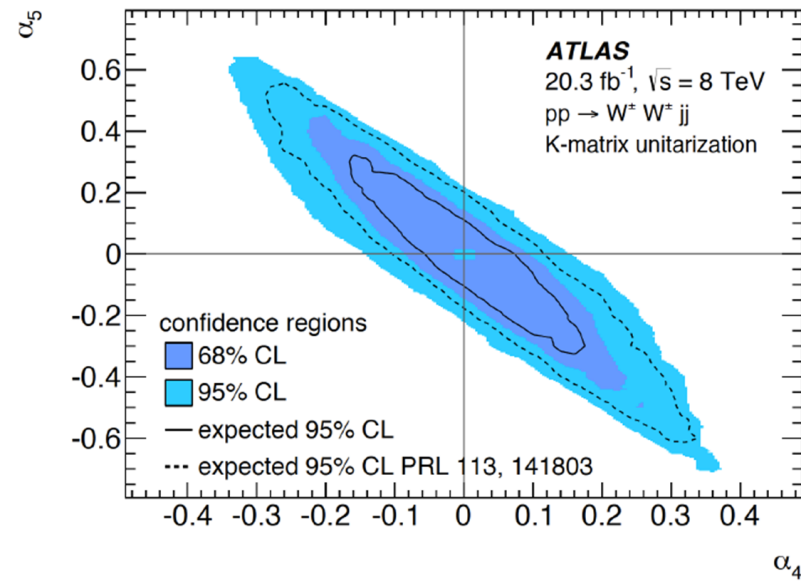
- Many models in this class.
- Similar scenarios: Randall–Sundrum...

Example



- Final states can be di-boson, $t\bar{t}$, etc.
- Closely related to electroweak symmetry

On-going EFT style searches



$$\begin{aligned}\mathcal{O}_{WWW} &= \text{Tr}[W_{\mu\nu}W^{\nu\rho}W_{\rho}^{\mu}] \\ \mathcal{O}_W &= (D_{\mu}\Phi)^{\dagger}W^{\mu\nu}(D_{\nu}\Phi) \\ \mathcal{O}_B &= (D_{\mu}\Phi)^{\dagger}B^{\mu\nu}(D_{\nu}\Phi)\end{aligned}$$

$$L_{S,0} = \left[(D_{\mu}\Phi)^{\dagger} D_{\nu}\Phi \right] \times \left[(D_{\mu}\Phi)^{\dagger} D_{\nu}\Phi \right]$$

$$L_{S,1} = \left[(D_{\mu}\Phi)^{\dagger} D^{\mu}\Phi \right] \times \left[(D_{\nu}\Phi)^{\dagger} D_{\nu}\Phi \right]$$

$$L_{M,0} = \text{Tr}[\hat{W}_{\mu\nu}\hat{W}^{\mu\nu}] \times \left[(D_{\beta}\Phi)^{\dagger} D^{\beta}\Phi \right]$$

$$L_{M,1} = \text{Tr}[\hat{W}_{\mu\nu}\hat{W}^{\nu\beta}] \times \left[(D_{\beta}\Phi)^{\dagger} D^{\mu}\Phi \right]$$

$$L_{M,6} = \left[(D_{\mu}\Phi)^{\dagger} \hat{W}_{\beta\nu}\hat{W}^{\beta\nu} D^{\mu}\Phi \right]$$

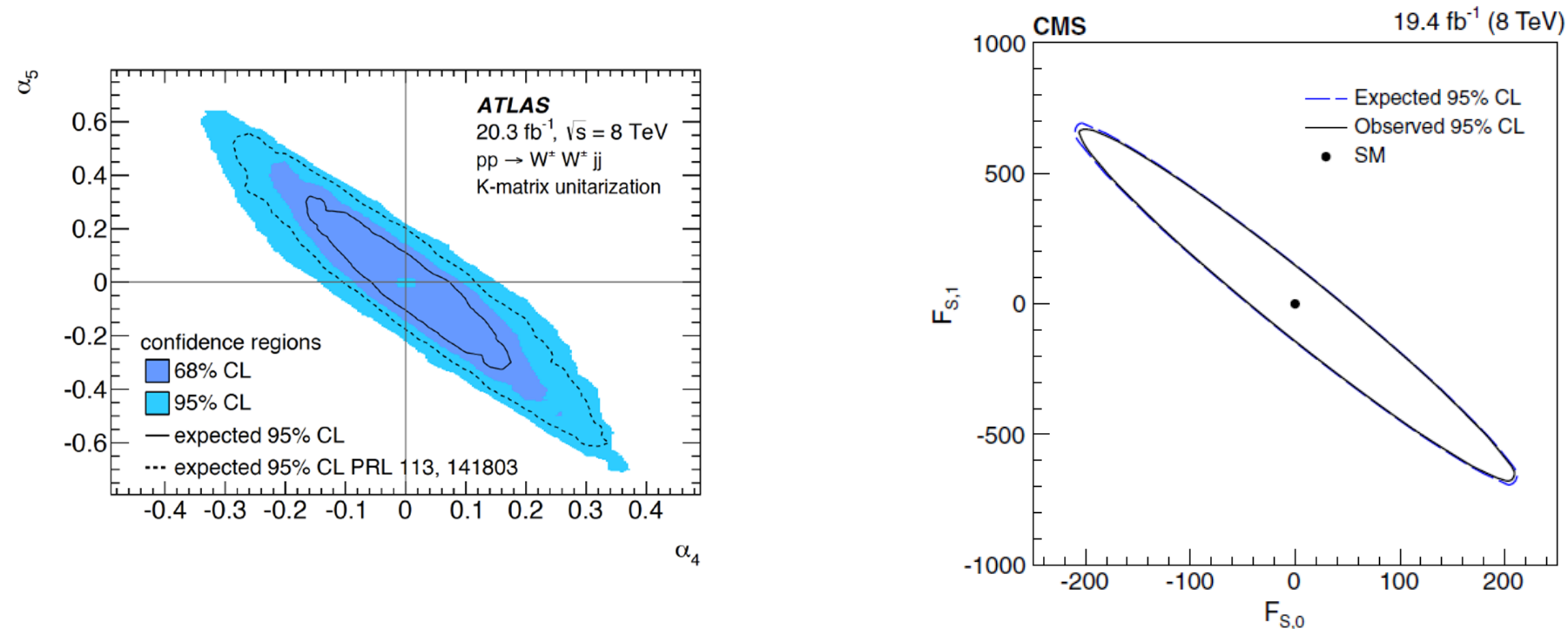
$$L_{M,7} = \left[(D_{\mu}\Phi)^{\dagger} \hat{W}_{\beta\nu}\hat{W}^{\beta\mu} D^{\nu}\Phi \right]$$

$$L_{T,0} = \text{Tr} [W_{\mu\nu}W^{\mu\nu}] \times \text{Tr} [W_{\alpha\beta}W^{\alpha\beta}]$$

$$L_{T,1} = \text{Tr} [W_{\alpha\nu}W^{\mu\beta}] \times \text{Tr} [W_{\mu\beta}W^{\alpha\nu}]$$

$$L_{T,2} = \text{Tr} [W_{\alpha\mu}W^{\mu\beta}] \times \text{Tr} [W_{\beta\nu}W^{\nu\alpha}]$$

Comment



- EFT may not model this kind of physics well. Similar to the case of dark matter.
- May need UV complete simplified models, mediators...
- Tagging the polarization of W/Z can be useful.

A word on the future

A word on the future

- If LHC discover some new physics

The next step will be clear.

A word on the future

- If LHC discover some new physics

The next step will be clear.

- Without LHC discovery.
 - ▶ Physics case for a 100 TeV pp collider stronger than HE-LHC at 28 TeV.
 - ▶ Cost+technological challenge. Perhaps easier to “sell” only as a second step of a circular Higgs factory in longer term.
 - ▶ **CEPC is the obvious next step.**

Conclusion

- LHC run 2 is pursuing a comprehensive program which covers the ground pretty well.
- Moriond 2017 can be an interesting juncture.

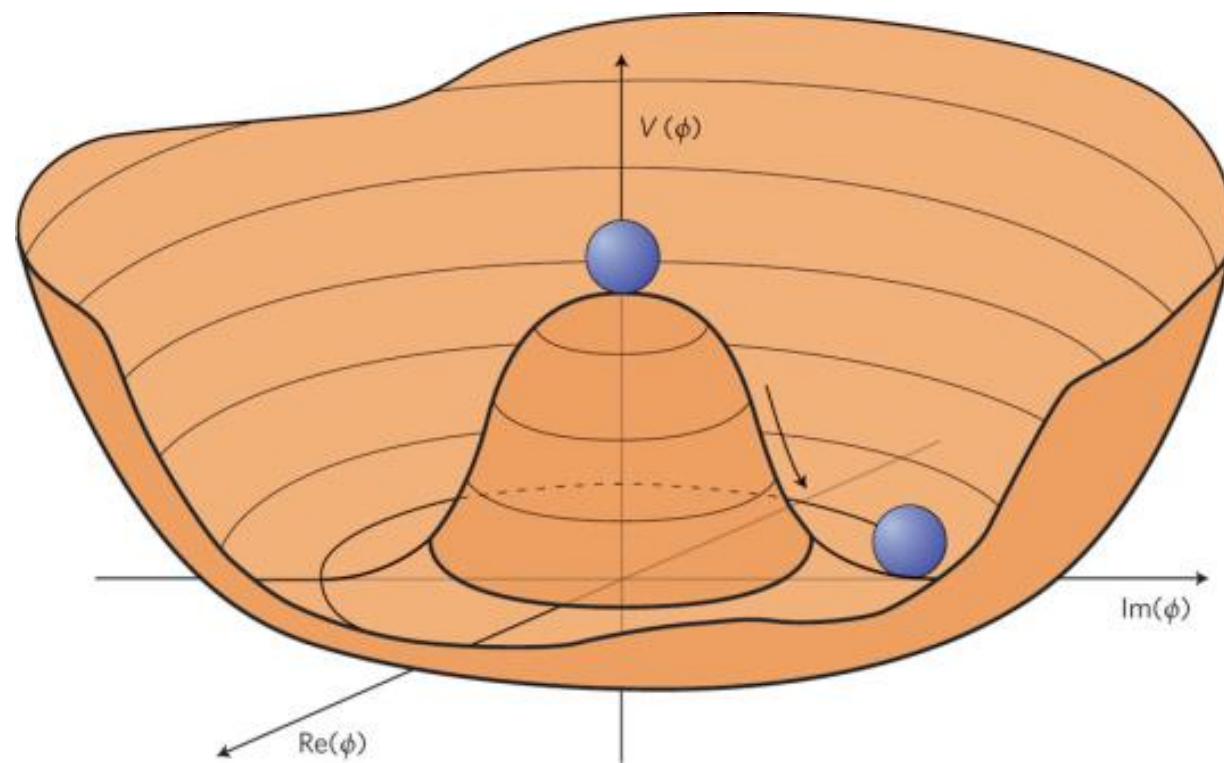
Between early discovery vs slow gain with lumi

- In the longer run, focusing on weakly coupled or broad features. Di-boson, $t\bar{t}$, etc.



A lot to look forward to!

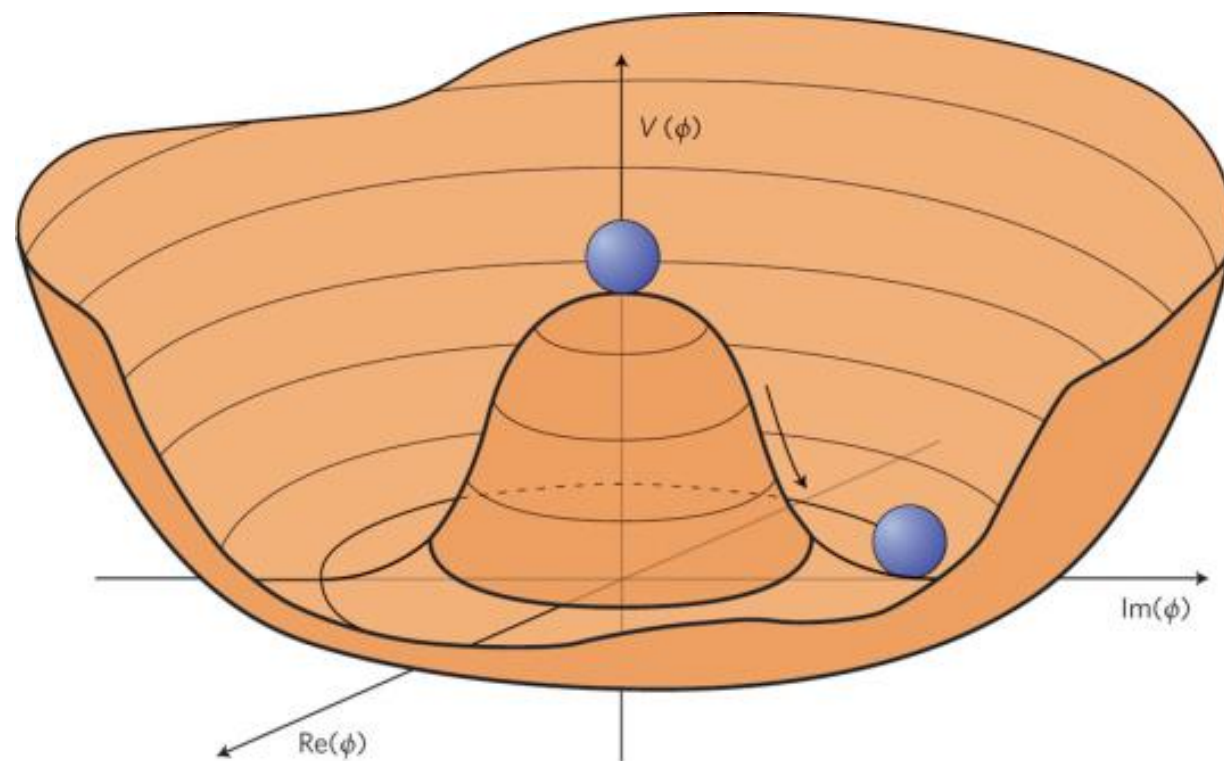
“Simple” picture: Mexican hat



$$V(h) = \frac{1}{2}\mu^2 h^2 + \frac{\lambda}{4}h^4$$
$$\langle h \rangle \equiv v \neq 0 \rightarrow m_W = g_W \frac{v}{2}$$

Similar to, and motivated by
Landau-Ginzburg theory
of superconductivity.

“Simple” picture: Mexican hat



$$V(h) = \frac{1}{2}\mu^2 h^2 + \frac{\lambda}{4}h^4$$
$$\langle h \rangle \equiv v \neq 0 \rightarrow m_W = g_W \frac{v}{2}$$

Similar to, and motivated by
Landau-Ginzburg theory
of superconductivity.

However, this simplicity is deceiving.
Parameters not predicted by theory. Can not be the complete picture.

Higgs is special

particle	spin
quark: u, d,...	1/2
lepton: e...	1/2
photon	1
W,Z	1
gluon	1
Higgs	0

h: a new kind of
elementary particle