Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions

Exploring Dark Matter at the CEPC via Loop Effects

Qian-Fei Xiang (向仟飞)

Institute of High Energy Physics, CAS



Beijing, Peking University, Dec 17, 2016



Outline	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Outlin	ie				



2 DM Models

3 Constrains

Orrections to Zh production cross sections

5 Conclusions

Outline	Introduction: Dark matter ●○○	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Evide	ences of Dark	Matter			



Rotation curves

	Introduction: Dark matter ●○○	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Evide	ences of Dark I	Matter			



Rotation curves



Coma Cluster

	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections	Conclusion
Evid	ances of Dark	Mattar			

Evidences of Dark Matter



Rotation curves



Coma Cluster



The Bullet Cluster







Large-Scale Structure Formation

	Introduction: Dark matter ○●○	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Dark	Matter				



The temperature power spectrum. Planck 2015, arXiv:1502.01589



13.7 BILLION YEARS AGO (Universe 380,000 years old)

12%

















Singlet-Doublet Fermionic Dark Matter (SDFDM)

Introduce left-handed Weyl fermions in the dark sector:

$$S \in (\mathbf{1}, 0), \quad D_{1} = \begin{pmatrix} D_{1}^{0} \\ D_{1}^{-} \end{pmatrix} \in (\mathbf{2}, -1/2), \quad D_{2} = \begin{pmatrix} D_{2}^{+} \\ D_{2}^{0} \end{pmatrix} \in (\mathbf{2}, +1/2)$$
$$\mathcal{L}_{S} = iS^{\dagger}\bar{\sigma}^{\mu}\partial_{\mu}S - \frac{1}{2}(m_{S}SS + h.c.)$$
$$\mathcal{L}_{D} = iD_{1}^{\dagger}\bar{\sigma}^{\mu}D_{\mu}D_{1} + iD_{2}^{\dagger}\bar{\sigma}^{\mu}D_{\mu}D_{2} - (m_{D}\epsilon_{ij}D_{1}^{i}D_{2}^{j} + h.c.)$$

Yukawa couplings: $\mathcal{L}_{HSD} = \mathbf{y}_1 H_i S D_1^i - \mathbf{y}_2 H_i^{\dagger} S D_2^i + h.c.$

Custodial symmetry limit $y = y_1 = y_2 \Rightarrow SU(2)_L \times SU(2)_R$ invariant form: $\mathcal{L}_D + \mathcal{L}_{HSD} = i\mathcal{D}_A^{\dagger}\bar{\sigma}^{\mu}D_{\mu}\mathcal{D}^A - \frac{1}{2}[m_D\epsilon_{AB}\epsilon_{ij}(\mathcal{D}^A)^i(\mathcal{D}^B)^j + h.c.] + [y\epsilon_{AB}(\mathcal{H}^A)_iS(\mathcal{D}^B)^j + h.c.]$ $SU(2)_R$ doublets: $(\mathcal{D}^A)^i = \begin{pmatrix} D_1^i \\ D_2^i \end{pmatrix}$, $(\mathcal{H}^A)_i = \begin{pmatrix} H_i^{\dagger} \\ H_i \end{pmatrix}$

Outline	Introduction: Dark matter	DM Models ○●○	Constrains	Corrections to Zh production cross sections	Conclusions
SDF	DM: State Mix	king			

The dark sector involves 3 Majorana fermions and 1 singly charged fermion

$$\mathcal{L}_{\text{mass}} = -\frac{1}{2} (S \quad D_1^0 \quad D_2^0) \mathcal{M}_N \begin{pmatrix} S \\ D_1^0 \\ D_2^0 \end{pmatrix} - m_D D_1^- D_2^+ + \text{h.c.} = -\frac{1}{2} \sum_{i=1}^3 m_{\chi_i^0} \chi_i^0 \chi_i^0 - m_{\chi^\pm} \chi^- \chi^+ + \text{h.c.}$$

$$\mathcal{M}_N = \begin{pmatrix} m_S & \frac{1}{\sqrt{2}} y_1 \nu & \frac{1}{\sqrt{2}} y_2 \nu \\ \frac{1}{\sqrt{2}} y_1 \nu & 0 & -m_D \\ \frac{1}{\sqrt{2}} y_2 \nu & -m_D & 0 \end{pmatrix}, \quad \begin{pmatrix} S \\ D_1^0 \\ D_2^0 \end{pmatrix} = \mathcal{N} \begin{pmatrix} \chi_1^0 \\ \chi_2^0 \\ \chi_3^0 \end{pmatrix}$$

$$\mathcal{N}^T \mathcal{M}_N \mathcal{N} = \text{diag}(m_{\chi_1^0}, m_{\chi_2^0}, m_{\chi_3^0}), \quad \chi^+ = D_2^+, \quad \chi^- = D_1^-$$
Couplings of the DM candidate χ_1^0 to the Higgs and Z bosons:
$$\mathcal{L}_Q = \frac{1}{2} g_1 + 0 \cdot 0 h \bar{\chi}^0 \chi^0 + \frac{1}{2} g_2 + 0 \cdot 0 Z_1 \bar{\chi}^0 \chi^0 + \chi_2^- \chi^0$$

$$g_{h\chi_{1}^{0}\chi_{1}^{0}} = -\sqrt{2}(y_{1}\mathcal{N}_{21} + y_{2}\mathcal{N}_{31})\mathcal{N}_{11}, \quad g_{Z\chi_{1}^{0}\chi_{1}^{0}} = -\frac{g}{2c_{W}}(|\mathcal{N}_{21}|^{2} - |\mathcal{N}_{31}|^{2})$$

Custodial symmetry limit $y_1 = y_2 \implies g_{Z\chi_1^0\chi_1^0} = 0$ $y_1 = y_2$ and $m_D < m_T \implies g_{h\chi_1^0\chi_1^0} = 0$

	Introduction: Dark matter	DM Models ○○●	Constrains	Corrections to <i>Zh</i> production cross sections	Conclusions
SDF	DM spectrums				

Particle mass as functions of y_1 .



- $m_{\chi^{\pm}} = m_D$.
- $y_1 = y_2$, or $y_1 = -y_2$, the mass of one neutral particle equals to the mass of χ^{\pm} . Hard to detect at colliders directly.

Outline	Introduction: Dark matter	DM Models	Constrains ●000	Corrections to Zh production cross sections	Conclusions
	- ·				

Relic density

- $T >> m_{\chi_1^0}$, DM particles are in thermal equilibrium with SM particles.
- $T < m_{\chi_1^0}$, number density of DM becomes exponentially suppressed
- The number density of DM is governed by the Boltzmann equation:

$$\frac{dn_{\chi}}{dt} + 3Hn_{\chi} = - <\sigma_{\chi\chi}\nu > (n_{\chi}^2 - n_{\chi,eq}^2)$$

• Numerically, this yields an abundance of:

$$\Omega h^2 = 0.11 \times \left(\frac{2.2 \times 10^{-26} cm^3/s}{<\sigma_{\chi\chi} \nu >_{freeze}} \right)$$

	Introduct		DM Models	Constrains ●○○○	Corrections to Zh production cross sections	Conclusions
- ··		•.				

Relic density

- $T >> m_{\chi_1^0}$, DM particles are in thermal equilibrium with SM particles.
- $T < m_{\chi_1^0}$, number density of DM becomes exponentially suppressed
- The number density of DM is governed by the Boltzmann equation:

$$\frac{dn_{\chi}}{dt} + 3Hn_{\chi} = - <\sigma_{\chi\chi}\nu > (n_{\chi}^2 - n_{\chi,eq}^2)$$

• Numerically, this yields an abundance of:

$$\Omega h^2 = 0.11 \times \left(\frac{2.2 \times 10^{-26} \text{cm}^3/\text{s}}{<\sigma_{\chi\chi} \nu >_{freeze}} \right)$$

• require $\Omega h^2 \leq 0.1186$. Planck[1502.01589] $m_{\chi_0^1}$ can only contribute to part of observed DM abundance, or some non-thermal production mechanics is needed.

Qian-Fei Xiang (IHEP) Exploring Dark Matter at the CEPC via Loop Effects



	Introduction: Dark matter	DM Models	Constrains ○●○○	Corrections to Zh production cross sections	Conclusions
Direc	t detection				

The couplings of the DM candidate χ_1^0 to the Higgs and Z bosons

$$\mathcal{L} \supset \frac{1}{2} g_{h\chi_1^0\chi_1^0} h \bar{\chi}_1^0 \chi_1^0 + \frac{1}{2} g_{Z\chi_1^0\chi_1^0} Z_{\mu} \bar{\chi}_1^0 \gamma^{\mu} \gamma_5 \chi_1^0$$

induce spin-independent (SI) and spin-dependent (SD) scatterings between DM and nuclei, respectively

Qian-Fei Xiang

(IHEP)



D:					
Outline	Introduction: Dark matter	DM Models	Constrains ○●○○	Corrections to Zh production cross sections	Conclusions

Direct detection

The couplings of the DM candidate χ^0_1 to the Higgs and Z bosons

$$\mathcal{L} \supset \frac{1}{2} g_{h\chi_1^0\chi_1^0} h \bar{\chi}_1^0 \chi_1^0 + \frac{1}{2} g_{Z\chi_1^0\chi_1^0} Z_\mu \bar{\chi}_1^0 \gamma^\mu \gamma_5 \chi_1^0$$

induce **spin-independent (SI)** and **spin-dependent (SD)** scatterings between DM and nuclei, respectively

Most stringent constraints from current direct detection experiments:

- **SI:** PandaX-II [1607.07400], LUX [1608.07648]
- SD: PandaX-II (proton)[1611.06553], LUX (neutron) [1602.03489], PICO (proton) [1503.00008, 1510.07754]



Qian-Fei Xiang (IHEP) Exploring Dark Matter at the CEPC via Loop Effects



Dec 17. 2016



New light particles could be produced directly at high energy colliders:



We consider all possible processes:

Qian-Fei Xiang (IHEP)

$$pp \rightarrow \chi^0_i \chi^0_j, \chi^0_i \chi^\pm, \chi^\pm \chi^\pm$$





Colliders constrains

New light particles could be produced directly at high energy colliders:



We consider all possible processes:

Qian-Fei Xiang (IHEP)

$$pp \rightarrow \chi^0_i \chi^0_j, \chi^0_i \chi^\pm, \chi^\pm \chi^\pm$$

LHC monojet search at 8 TeV with 20.3 data. [1502.01518]





Colliders constrains

New light particles could be produced directly at high energy colliders:



We consider all possible processes:

$$pp \rightarrow \chi^0_i \chi^0_j, \chi^0_i \chi^\pm, \chi^\pm \chi^\pm$$

LHC monojet search at 8 TeV with 20.3 data. [1502.01518]

In addition, LEP charged particles searches require $m_{\gamma^{\pm}} \leq 103.5$ [0509008]. In our model SFFDM, $m_{\gamma^{\pm}} = m_D$.



Qian-Fei Xiang (IHEP) Exploring Dark Matter at the CEPC via Loop Effects



Dec 17. 2016

Outline	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Higgs	and Z decays	6			

- If the kinematics is allowed, the Higgs and Z could decay into DM particles. Such **invisible decays** are constrained by the relevant searches at colliders.
- If new charged particles couple to the Higgs, they could modify the Higgs decay width to di-photon via triangle loops. $\Delta\Gamma/\Gamma$ could be tested at a level of 9%.

	Introduction: Dark matter	DM Models	Constrains ○○○●	Corrections to Zh production cross sections	Conclusions
Higgs	and Z decays				

- If the kinematics is allowed, the Higgs and Z could decay into DM particles. Such **invisible decays** are constrained by the relevant searches at colliders.
- If new charged particles couple to the Higgs, they could modify the Higgs decay width to di-photon via triangle loops. $\Delta\Gamma/\Gamma$ could be tested at a level of 9%.



We require $\Gamma(Z \rightarrow inv) < 2$ MeV, LEP [0509008]. and $\Gamma(h \rightarrow inv)/\Gamma_h < 0.28\%$, [CEPC-pre-CDR].

	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections ●○○	Conclusions
C					

Corrections

We consider the **loop corrections** of new particles to the associated Zh production.

This process is affected by both the gauge and yukawa interactions.



Outline	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections ●○○	Conclusions	
Common at in the second s						

Corrections

We consider the **loop corrections** of new particles to the associated Zh production.

This process is affected by both the gauge and yukawa interactions.





Formally, the coupling of h to ZZ is :



Outline	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections ●○○	Conclusions	
Common at in the second s						

Corrections

We consider the **loop corrections** of new particles to the associated Zh production.

This process is affected by both the gauge and yukawa interactions.



Formally, the coupling of h to ZZ is :



Two cases in which $\Delta \sigma_{ZH} / \sigma_{ZH} = |\sigma_{ZH} - \sigma_{ZH,SM}| / \sigma_{ZH,SM}$ could be enhanced:

• $m_{\chi_i^0} + m_{\chi_j^0} = m_Z(m_H), \ m_{\chi_i^\pm} + m_{\chi_j^\pm} = m_Z(m_H), \ m_{\chi_i^0} + m_{\chi_j^\pm} = m_W$ • $m_{\chi_i^0} + m_{\chi_j^0} = \sqrt{s}, \ m_{\chi_i^\pm} + m_{\chi_j^\pm} = \sqrt{s}$

	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections $\bigcirc \bullet \bigcirc$	Conclusions
Result	ts				

We use **FeynArts**, **FormCalc** to calculate corrections to the Zh production cross section.

At **CEPC** with 5 ifb data, it can measure $\Delta \sigma_{ZH} / \sigma_{ZH}$ to a level of 0.51%.



Color region is can be explored by CEPC with 5 ifb data. gray region is beyond the ability of CEPC.

Introduction: Dark matter	DM Models	Constrains	Corrections to <i>Zh</i> production cross sections	Conclusions

Comparison



	Introduction: Dark matter	DM Models	Constrains	Corrections to Zh production cross sections	Conclusions
Concl	usions				

- We proposed a DM model with one Weyl $SU(2)_L$ singlet and two Weyl $SU(2)_L$ doublets. After EW vacuum breaking, three neutral particles are obtained with the lightest one could service as the DM candidate.
- We considerd the constrains from relic density, direct detections, and colliders.
- We calculated the corrections to the cross section of Zh associated production. The ability of CEPC with 5 ifb data is comparable to or great than other detections, especially when $y_1 = \pm y_2$, where other detections may lose significance.