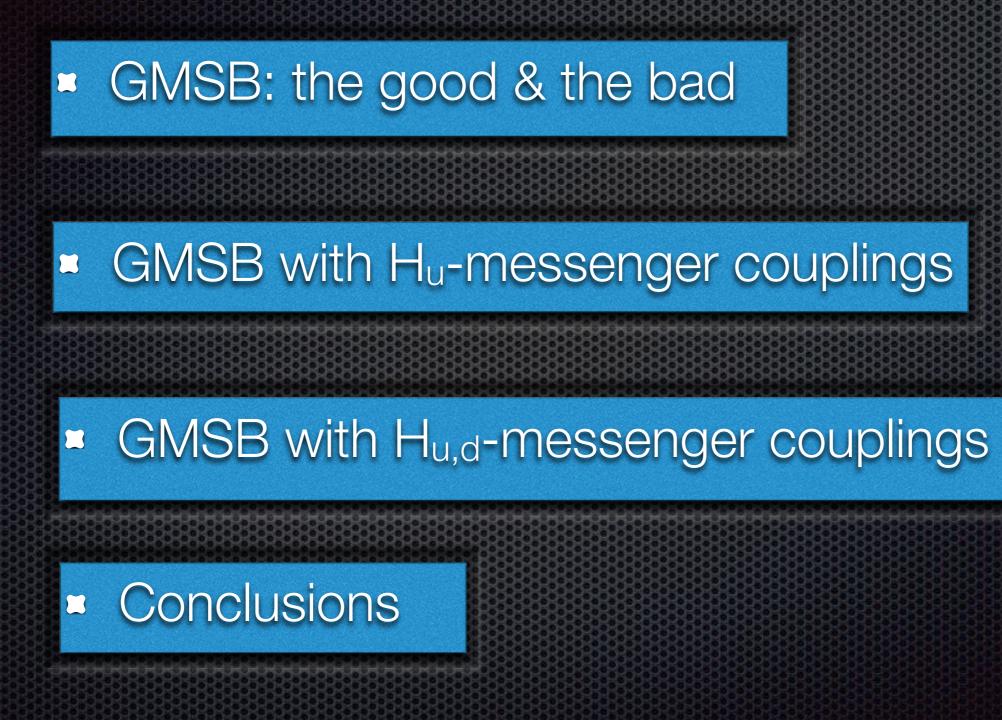
#### 第二届中国LHC物理研讨会/北京大学

# Higgs-messenger coupling extended Gauge mediated SUSY-breaking

based on arXiv:1610.06024 by ZF Kang & Phys. Rev. D 86, 095020 (2012) by ZF Kang, T. Li, T. Liu, C. Tong and J. M. Yang.

Zhaofeng Kang(康昭峰), KIAS(韩国高等研究院), 10/25/2016





# GMSB: gauge mediated SUSY-breaking

GMSB utilizes flavor universal of gauge interactions

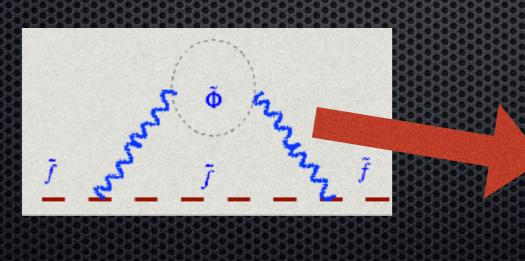
The best idea to overcome the flavor problem

hidden sector breaking SUSY, described by the spurion X=M+F $\theta^2$ 

messenger sector  $X\Phi\overline{\Phi}$ 

A soft SUSY breaking sector in the MSSM automatically respects MFV

 $C_1(i)$ 



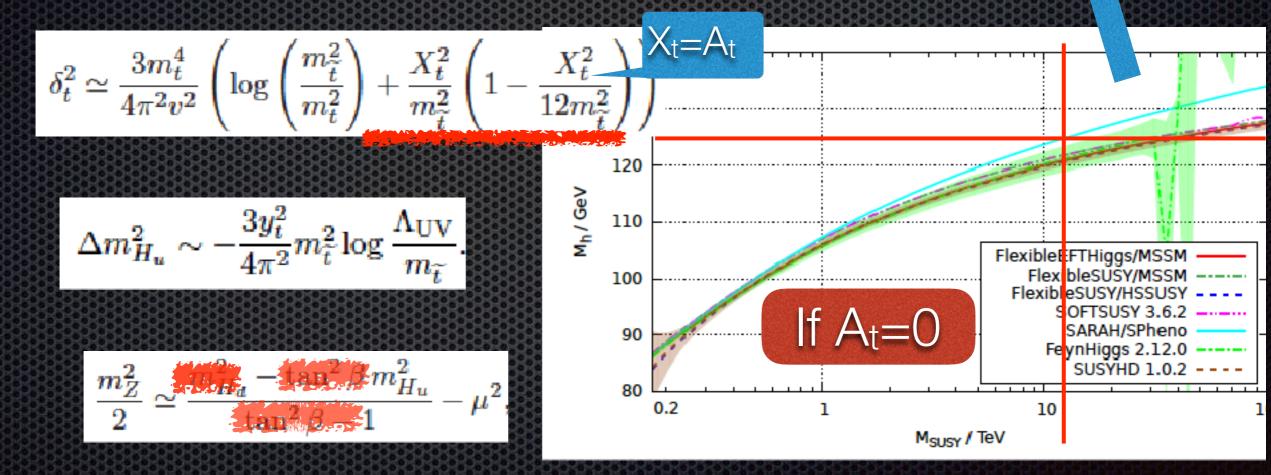
$$\begin{split} M_a &= \frac{\alpha_a}{4\pi} \Lambda, \qquad (a = 1, 2, 3), \Lambda = \frac{F}{M} \\ m_{\phi_i}^2 &= \frac{2\Lambda^2}{2} \left[ \left( \frac{\alpha_3}{4\pi} \right)^2 C_3(i) + \left( \frac{\alpha_2}{4\pi} \right)^2 C_2(i) + \left( \frac{\alpha_1}{4\pi} \right) \right] \\ 0 \widetilde{Q} H_u \widetilde{U}^c + 0 \widetilde{Q} H_d \widetilde{D}^c + 0 \widetilde{L} H_d \widetilde{E}^c \end{split}$$

# GMSB: gauge mediated SUSY-breaking

#### Post h(125), awkward....

From P. Athron, J. h. Park, T. Steudtner, D. Stckinger and A. Voigt, 1609.00371

VERY serious fine-tuning (at least in the MSSM)!



Fail in explaining  $(g-2)_{\mu}$ , because sleptons are very heavy!

# GMSB: gauge mediated SUSY-breaking

#### Post h(125), awkward....

NOT mentioning to the long-standing µ/Bµ problem, that will be presented latter

Want a large  $A_t$ ? Coupling  $H_u$  to messengers:  $\lambda_u H_u \mathcal{O}_u$ 

• choosing messenger  $10 = (Q_{\Phi}, U_{\Phi}, E_{\Phi})$  of SU(5) with couplings

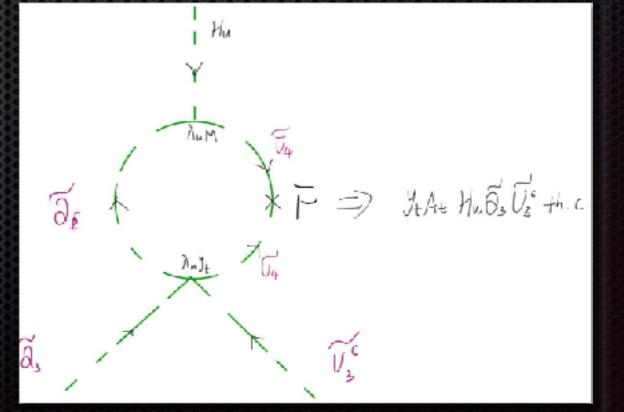
$$\begin{split} W_{10} = &\lambda_u Q_{\Phi} H_u U_{\Phi} + \lambda_{\Phi} Q_{\Phi} U_{\Phi} + \frac{V_{\Phi}}{2} Q_{\Phi} U_{\Phi} + \frac{V_{\Phi}}{2} Q_$$

N<sub>f</sub> sets of messengers

Z. Kang, T. Li, T. Liu, C. Tong and J. M. Yang, Phys.Rev. D 86, 095020 (2012).

 Feymann digram demonstration of one-loop At

$$A_t = -N_m \alpha_u \Lambda \qquad N_m \equiv 3N_f$$
$$\Lambda \equiv F/(4\pi M) \sim \mathcal{O}(10 \text{TeV})$$



Higgs-messenger VS matter-messenger coupling

J. L. Evans, M. Ibe, S. Shirai and T. T. Yanagida, Phys. Rev. D 85, 095004 (2012); A. Basirnia, D. Egana-Ugrinovic, S. Knapen and D. Shih, JHEP 1506, 144 (2015)...

Although At also arises if the matters Q<sub>3</sub> /U<sub>3</sub> couple to messengers, HGMSB **naturally maintains MFV**, the most important feature of GMSB.

Moreover, HGMSB has the potential to be embedded in the framework which dynamically generates the  $\mu$  term by coupling Higgs to the messengers, though the  $\mu$ /B $\mu$  problem

At/mHu problem: Not only a large Au but also large mHu

- Soft parameters of Higgs and any MSSM fields coupling to Higgs acquire Yukawa-mediated SUSY-breaking effects
- Systematic calculation: wave function renormalization method
  G. F. Giudice and R. Rattazzi, Nucl. Phys. B 511 (1998)

$$\begin{split} \Delta m_{\widetilde{Q}_3}^2 &= -\frac{N_m}{3} \left( 3\alpha_t \alpha_u + 3\alpha_b \alpha_d + \alpha_b \alpha'_d \right) \Lambda^2, \\ \Delta m_{\widetilde{U}_3^c}^2 &= -2N_m \alpha_t \alpha_u \Lambda^2, \\ \Delta m_{\widetilde{D}_3^c}^2 &= -2\frac{N_m}{3} \alpha_b (3\alpha_d + \alpha'_d) \Lambda^2. \end{split}$$

In particular, H<sub>u</sub> receives a large & positive contribution

 $\Delta m_{H_u}^2 = N_m \alpha_u \left[ (N_m + 3) \alpha_u - \frac{16}{3} \alpha_3 - 3\alpha_2 - \frac{13}{15} \alpha_1 \right] \Lambda^2$ 

 radiative EWSB probably fails
originally, cancelation between Yukawa and QCD gauge parts is the reason for introducing messenger 10

# What if includes $\lambda_d H_d \mathcal{O}_d$ ?

The model with H<sub>u,d</sub>-messenger couplings: 

$$\begin{split} W_{10} = &\lambda_u Q_{\Phi} H_u U_{\Phi} + \lambda_d \overline{Q}_{\Phi} H_d \overline{U}_{\Phi} + V_{\Phi} \\ &+ X \left( \lambda_Q Q_{\Phi} \overline{Q}_{\Phi} + \lambda_U \overline{U}_{\Phi} U_{\Phi} \right) + W_{\text{MSSM}}, \end{split}$$

turned off before to avoid µ/Bµ problem

Dynamically generate µ Hu Hd & Bµ Hu Hd at one loop with 

> $\mu = f(\lambda_Q / \lambda_U) N_m \sqrt{\alpha_u \alpha_d} \Lambda,$  $B_{\mu} = f(\lambda_Q / \lambda_U) N_m 4\pi \times \sqrt{\alpha_u \alpha_d} \Lambda^2,$

Great! It explains the the origin of  $\mu$ , the unique massive coupling in MSSM superpotential

 $\sin 2\beta = \frac{2B_{\mu}}{m_{H_{\pi}}^2 + m_{H_{d}}^2 + 2\mu^2}$ 

But Bµ has dimension 2! TOO BIG TO FILL the EWSB equation — — the  $\mu/B\mu$  problem

Nonradiative EWSB as a common solution

 In the At/mHu problem we encounter a too large mHu>0, which hampers the conventional radiative EWSB

 Similarly, in the µ/Bµ problem above we encounter a too large Bµ and thus EWSB fails:

$$\sin 2\beta = \frac{2B_{\mu}}{m_{H_u}^2 + m_{H_d}^2 + 2\mu^2}$$

However, what if  $m_{H_d}^2$  is NOT negligible ? ! !

C. Csaki, A. Falkowski, Y. Nomura and T. Volansky, Phys. Rev. Lett. 102 (2009) 111801.

Both equations can be satisfied by this large & positive  $m_{H_d}^2$ 

Generate a huge 
$$m_{H_d}^2$$
 !!

Include the "full" H<sub>u,d</sub>-messenger couplings:

$$\begin{split} W_{10} = &\lambda_u Q_{\Phi} H_u U_{\Phi} + \lambda_d \overline{Q}_{\Phi} H_d \overline{U}_{\Phi} + \frac{\lambda'_d}{2} E_{\Phi} H_d H_d + \\ &+ X \left( \lambda_Q Q_{\Phi} \overline{Q}_{\Phi} + \lambda_U \overline{U}_{\Phi} U_{\Phi} \right) + W_{\text{MSSM}}, \end{split} \qquad \text{messenger 10} \\ \text{admits such a term} \\ \Delta m_{H_d}^2 \approx \frac{N_m}{3} \alpha'_d \left( N_m \alpha'_d + 2 N_m \alpha_d + 3 \alpha_b - 3 \alpha_2 \right) \Lambda^2 \end{split}$$

Features of nonradiative EWSB

■ A large m<sub>Hd</sub>~10 TeV thus heavy H/A/H±

 $m_{H_d}^2 \simeq 2B_\mu / \sin 2\beta \approx \tan \beta B_\mu$ 



- Tanβ is favored to be not very large, thus alleviating the need for a superhuge  $m_{H_d}^2$   $m_{H_d}^2/\tan^2\beta m_{H_u}^2 \sim \mu^2$
- A small μ is strongly favored, but:

1. a small  $\mu$  term does not guarantee naturalness 2. cancelation is hidden between  $m_{H_u}^2 > 0$  and  $m_{H_u}^2$ 

Bonus: filling the  $(g-2)_{\mu}$  discrepancy by light stau

• Nonradiative EWSB with a huge  $m_{H_d}^2$  leads to

$$S = \text{Tr}(Y_f \hat{m}_{\tilde{f}}^2) \sim m_{H_d}^2 / 2 > 0 \Longrightarrow \frac{dm_{\tilde{L}}^2}{dt} = -\frac{1}{16\pi^2} \frac{3}{5} g_1^2 S + \dots$$

Sparticles with negative hyper charge such as left-handed sleptons will be decreased significantly

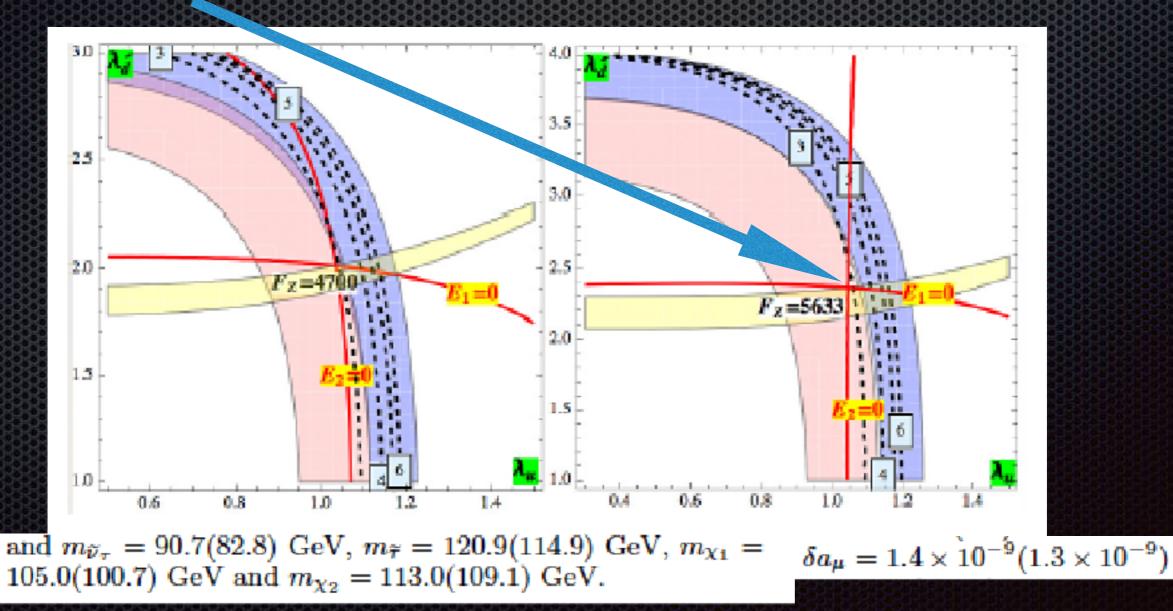
 $\tilde{v}_{\tau}$  tends to the NLSP

$$\delta a_{\mu} \approx \frac{\alpha_2}{2\pi} \frac{m_{\mu}^2 M_{\lambda_2} \mu \tan \beta}{m_{\widetilde{\nu}_{\mu}}^4} F_a \left( \frac{M_{\lambda_2}^2}{m_{\widetilde{\nu}_{\mu}}^2}, \frac{\mu}{m_{\widetilde{\nu}_{\mu}}^2} \right)$$

SC

Two benchmarks with M=10<sup>12</sup> GeV and 10<sup>7</sup> GeV

The cross points of two red lines are the solutions to EWSB



# A light Higgsino & slepton world for $(g\text{-}2)_{\mu}$

Light but quite dengnerate Higgsinos currently are hidden at LHC (missing energy+soft leptons/jets), even forever?

C. Han, D. Kim, S. Munir and M. Park, JHEP 1504,132 (2015).

Light sleptons are even less hopeful. But in one scenario where the sneutrino is not far above  $m_Z/2$ , the relatively hard leptons says from charginos still can be hunted, e.g.,

$$pp \to \chi^{\mp} \chi^{\pm} \to \tau^{+} \tau^{-} + \text{MET}(= \widetilde{\nu}_{\tau} \widetilde{\nu}_{\tau}^{*}).$$

#### Conclusions

GMSB offer the best idea to organize soft SUSY, but it suffers the  $\mu/B\mu$  problem, and serious fine-tuning problem post h(125)

HGMSB including both  $H_u$ &H<sub>d</sub> coupling to messengers can address these problems, via nonradiative EWSB; as a bonus, the (g-2)<sub>µ</sub> puzzle can be addressed, too.

The light Higgsino & slepton world satisfying  $(g-2)_{\mu}$  can be tested in certain scenarios

Thank you for your attention!