A visualization of particle tracks in a detector, showing several tracks originating from a central point and extending outwards. The tracks are colored in shades of blue and green, with some points highlighted in red. The background is a dark blue gradient.

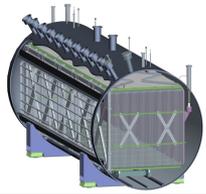
# LArTPC Calibration

**Michael Mooney**

**Brookhaven National Laboratory (BNL)**

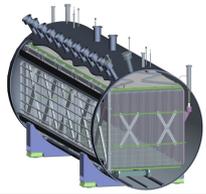
*NNN16, IHEP, Beijing, China*

*November 3<sup>rd</sup>, 2016*



# Introduction

- ◆ Goal of LArTPC calibration: measure charge and position associated with ionization signal in unbiased manner and as precisely as possible
  - Noise, detector effects lead to bias, resolution loss
  - With calibrations in place, can then look at higher-level candles to study particle reconstruction (e.g. cosmic muons, Michel electrons, photons from  $\pi^0$  decays)
- ◆ Crucial for LArTPC experiments to reach physics goals!
- ◆ In what follows, will describe relevant detector effects and calibration techniques using **MicroBooNE** as an example
  - Operational single-phase LArTPC: plenty of data to begin looking at these effects
  - Also reference **test stand measurements** where relevant



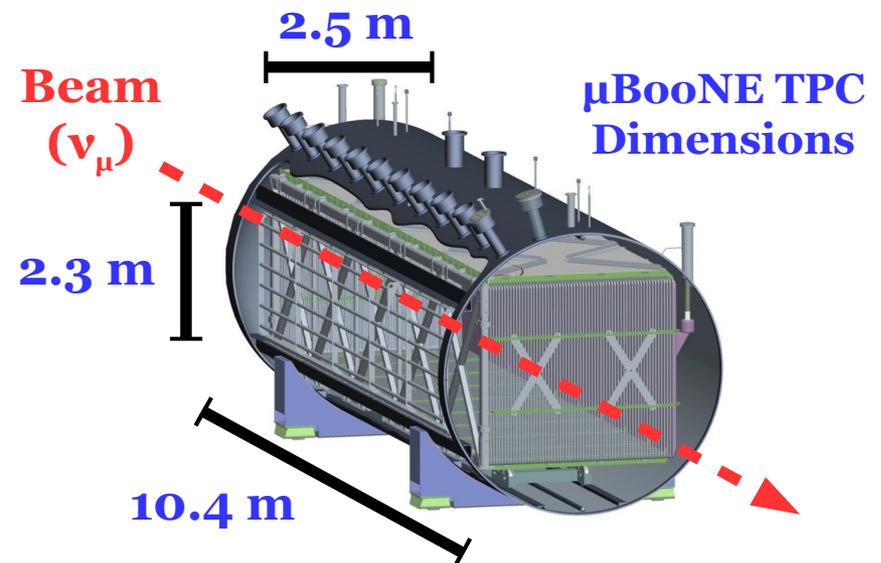
# Case Study: MicroBooNE

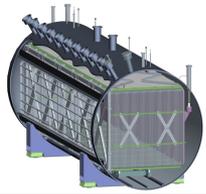
## ◆ “Micro Booster Neutrino Experiment”

- Accelerator  $\nu$  experiment @ FNAL
- LArTPC with 89 ton active mass
- Non-evacuated liquid argon fill
- Cold (in LAr) front-end electronics
- Near-surface operation
- UV laser calibration system

## ◆ Physics goals:

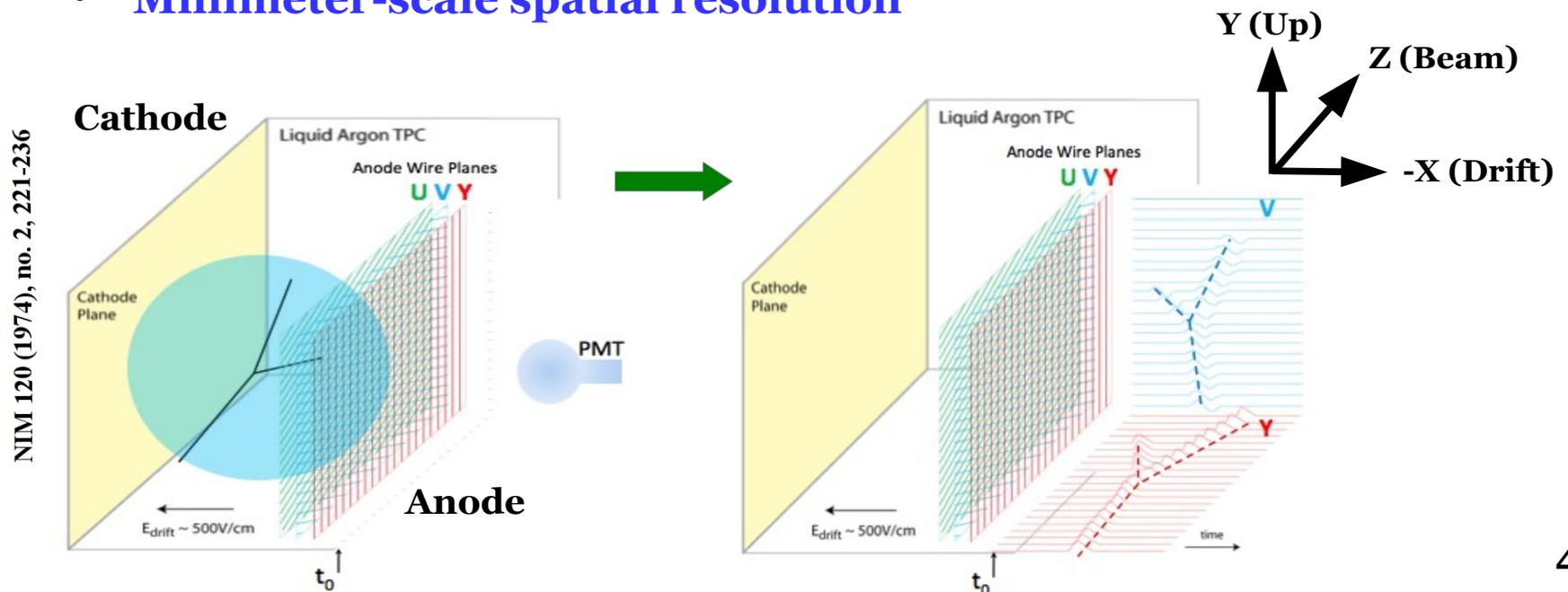
- Investigate MiniBooNE low-energy excess
- Measure first low-energy  $\nu$ -Ar cross sections
- R&D for future detectors
- Key step for Short Baseline Neutrino (SBN) program

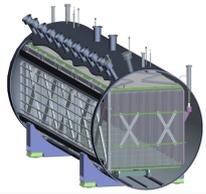




# MicroBooNE TPC

- ◆ Two **induction** planes (U, V) and one **collection** plane (Y); drifted ionization in LAr puts signal on all three
  - Drift E field at 273 V/cm, ~uniform via surrounding field cage
  - 8000+ channels in total with front-end electronics in LAr
- ◆ 3D event reconstruction by combining signals from all three planes (minimum two needed), each with 3 mm wire pitch
  - **Millimeter-scale spatial resolution**

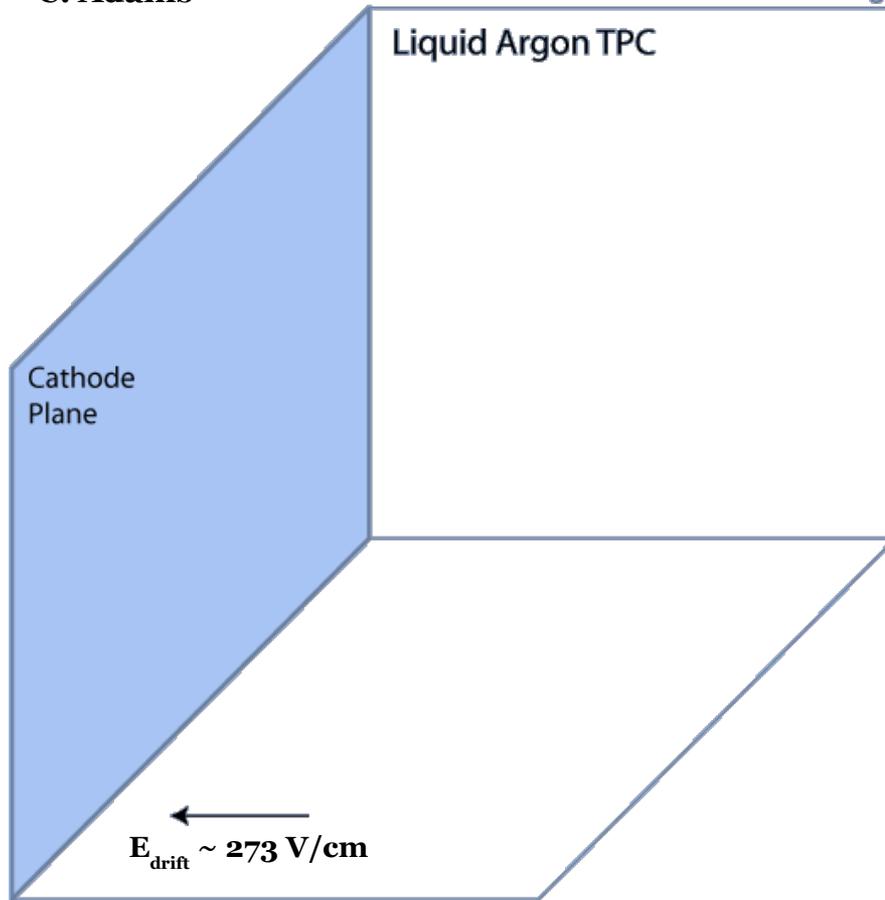


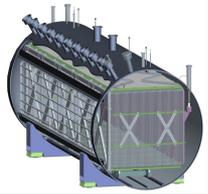


# Signal Formation

Cartoon Credit:  
C. Adams

Anode wire planes:  
U V Y

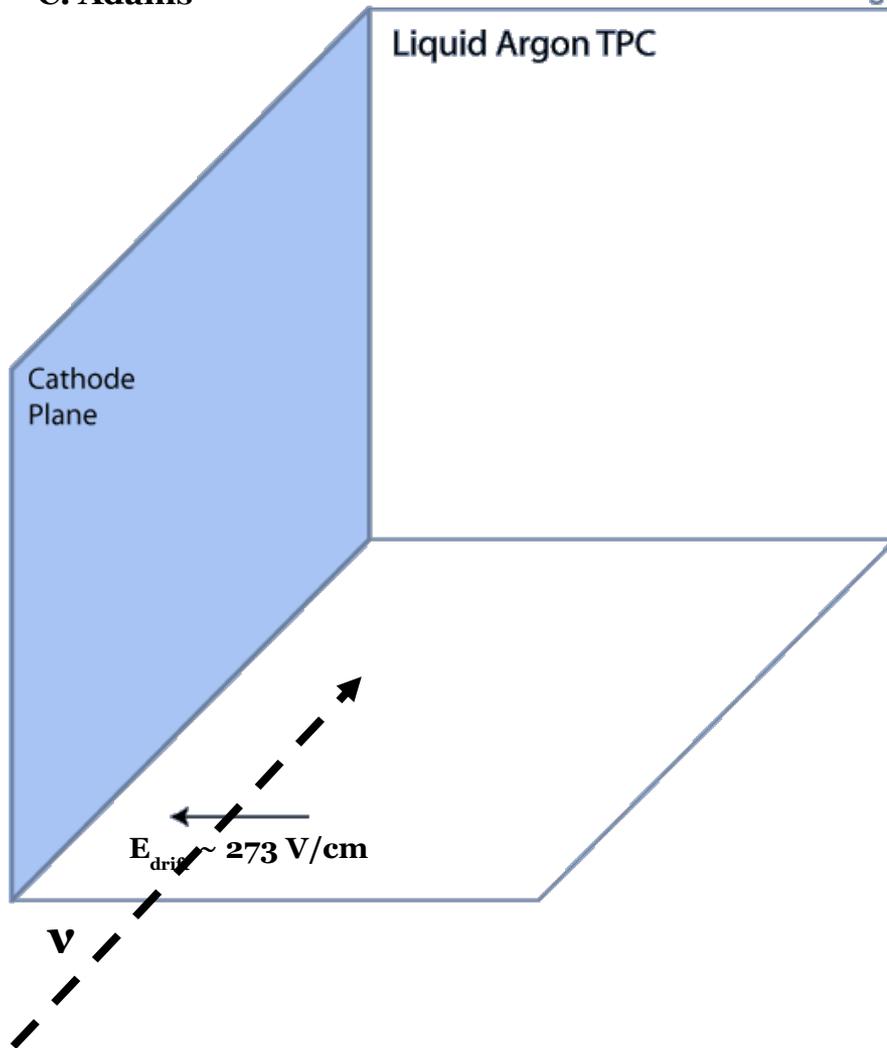


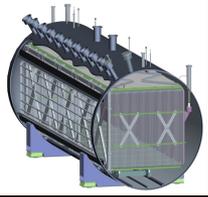


# Signal Formation

Cartoon Credit:  
C. Adams

Anode wire planes:  
U V Y

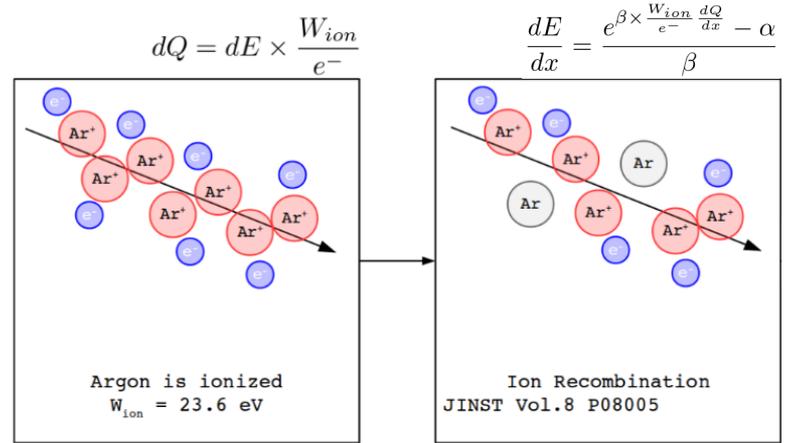
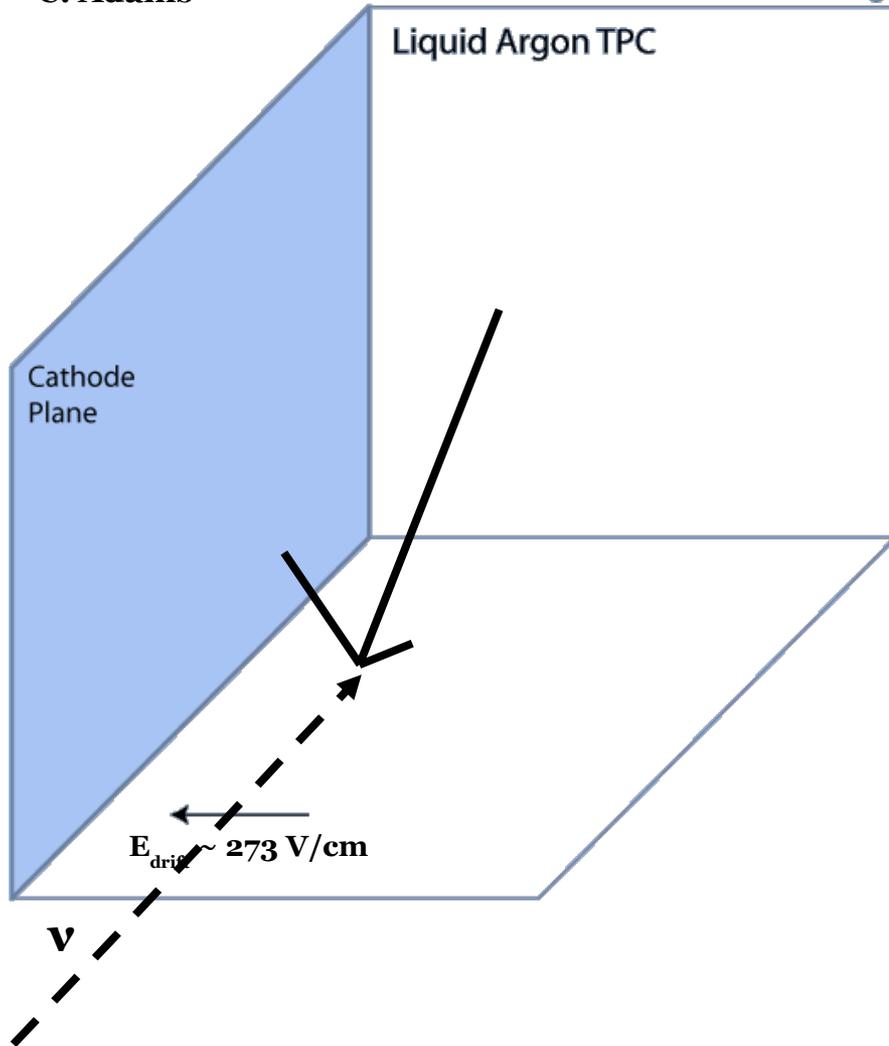


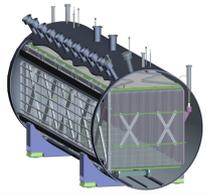


# Signal Formation

Cartoon Credit:  
C. Adams

Anode wire planes:  
U V Y

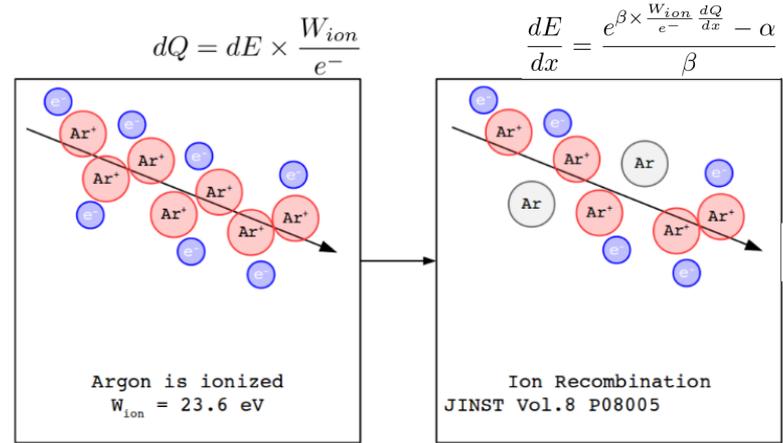
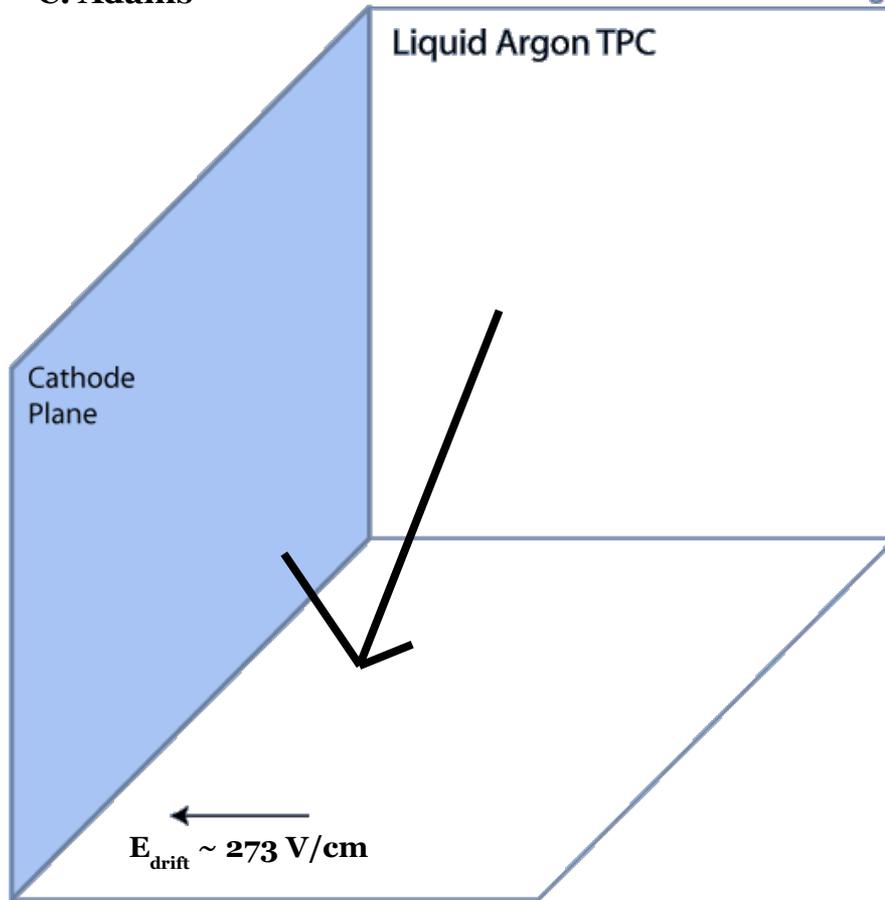


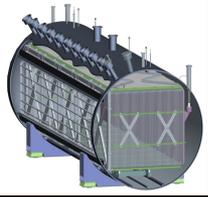


# Signal Formation

Cartoon Credit:  
C. Adams

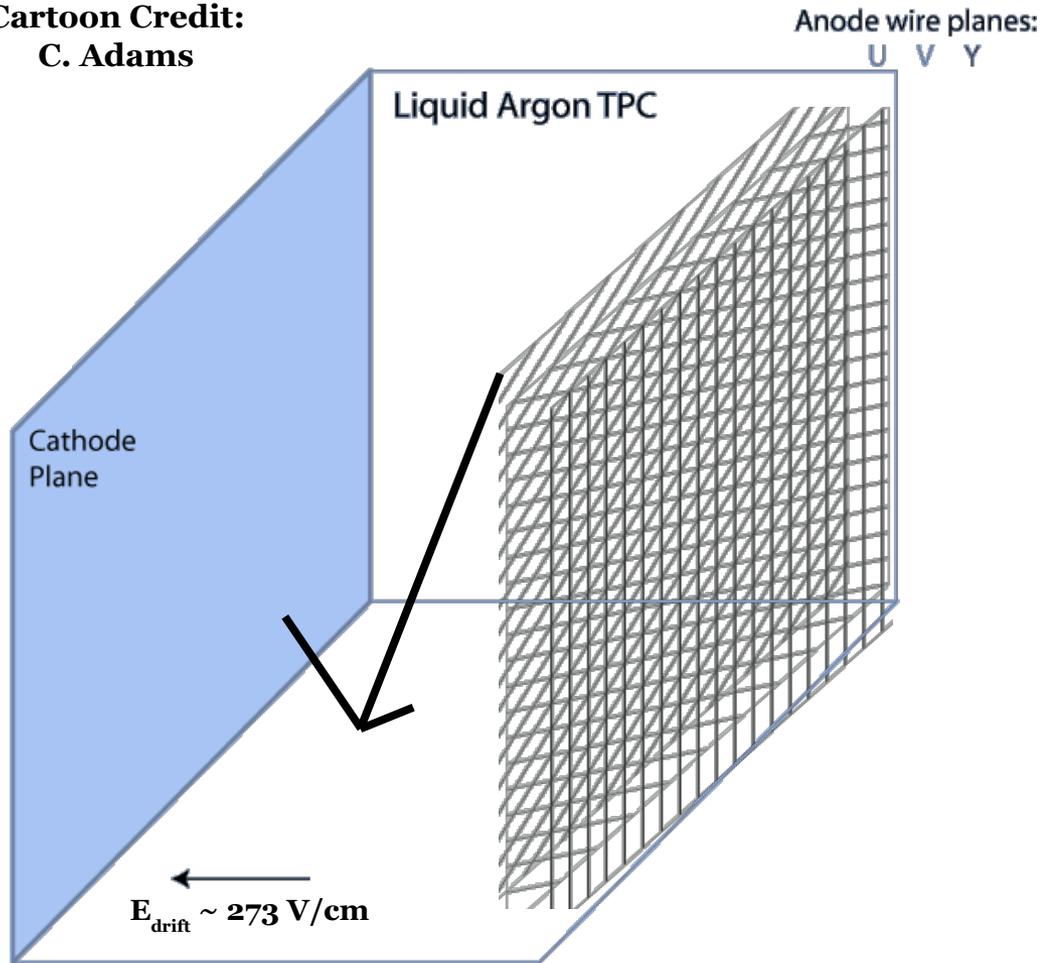
Anode wire planes:  
U V Y

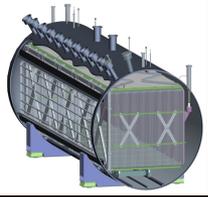




# Signal Formation

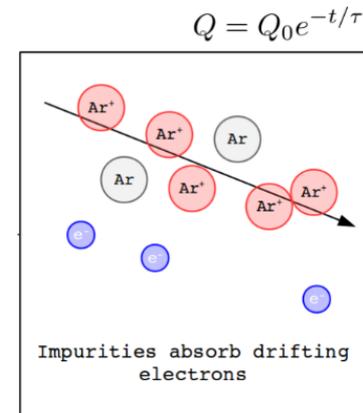
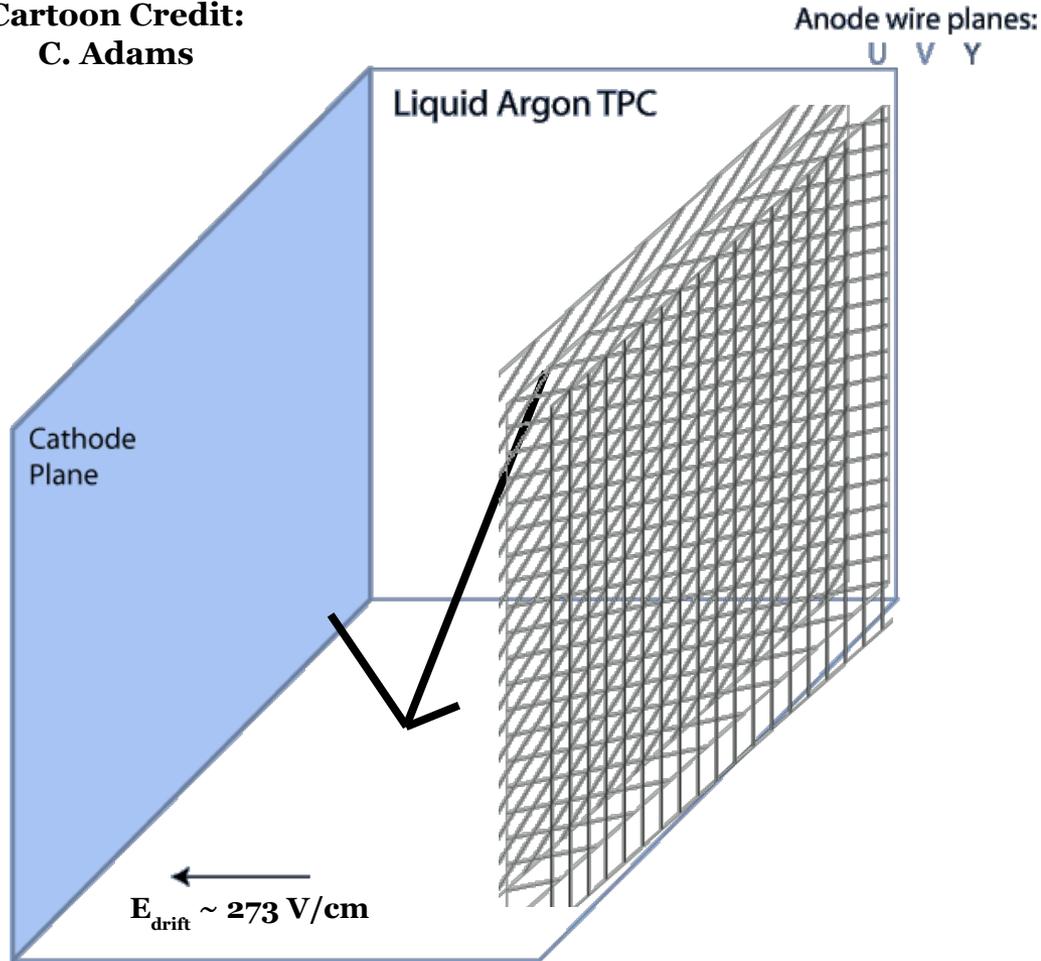
Cartoon Credit:  
C. Adams

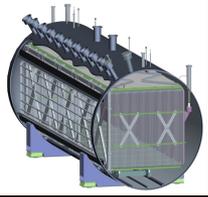




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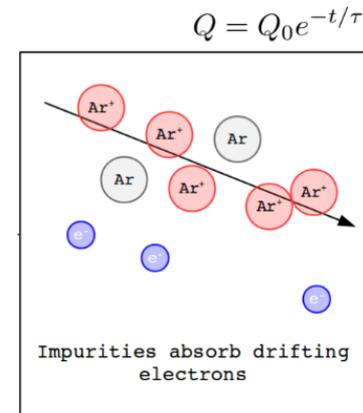
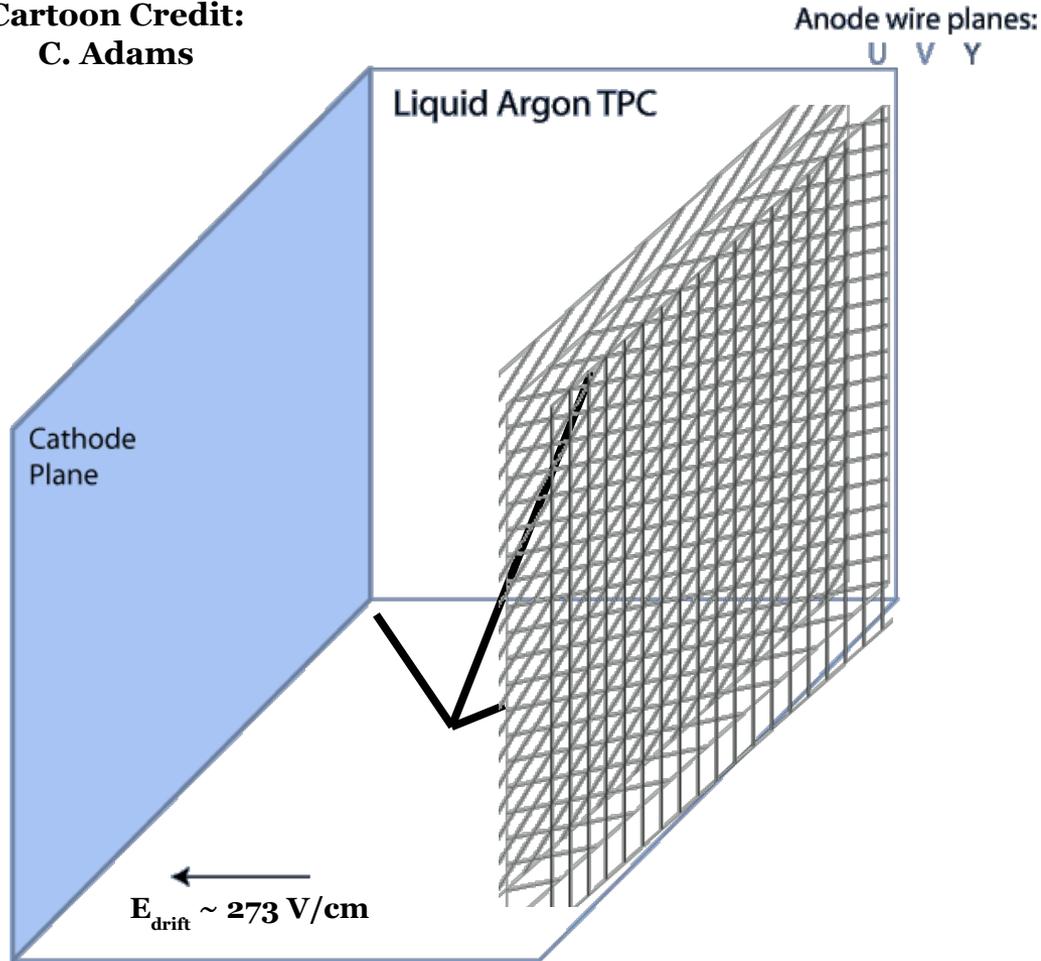
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C. Adams

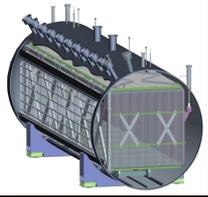




# Signal Formation

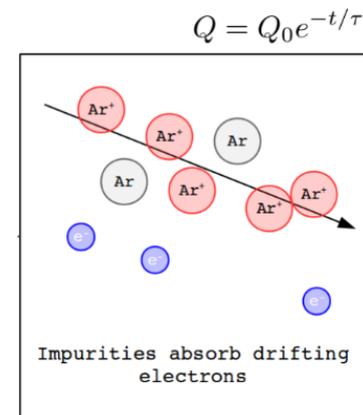
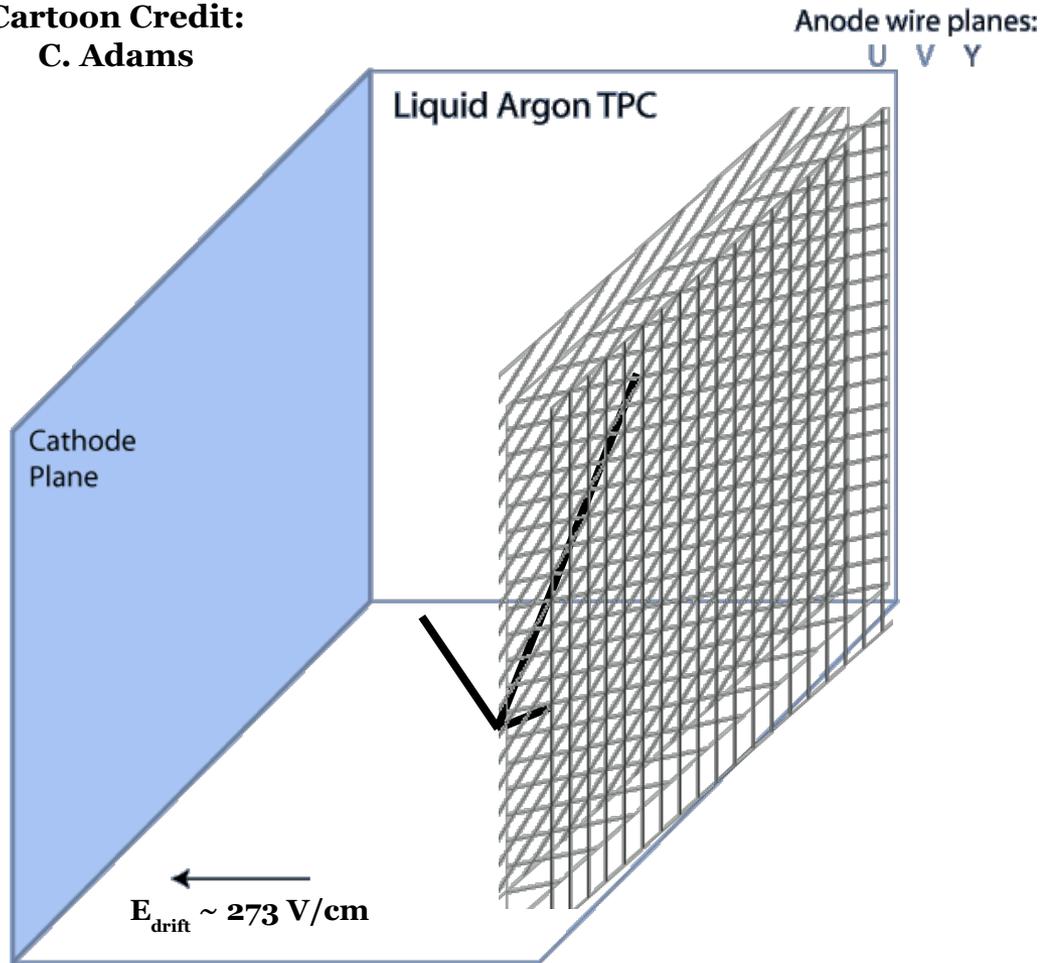
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C. Adams

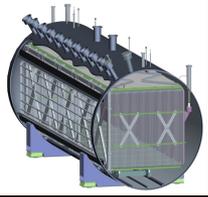




# Signal Formation

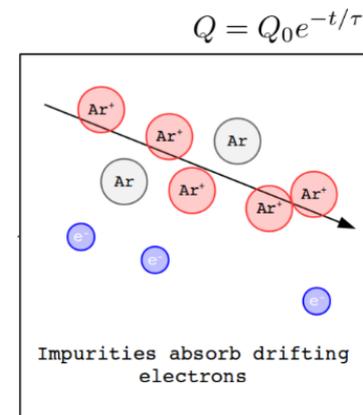
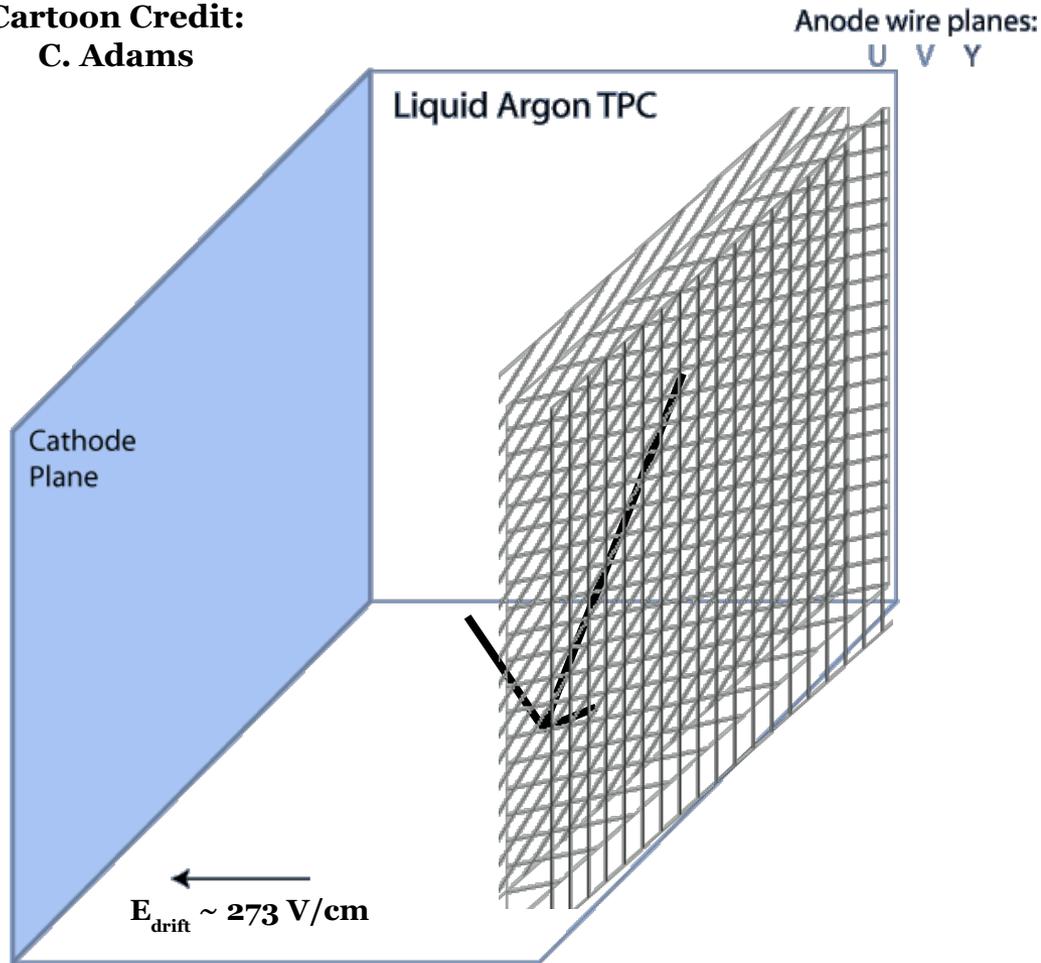
Cartoon Credit:  
C. Adams

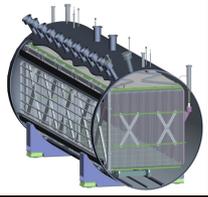




# Signal Formation

Cartoon Credit:  
C. Adams

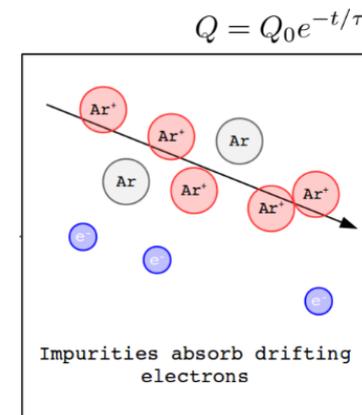
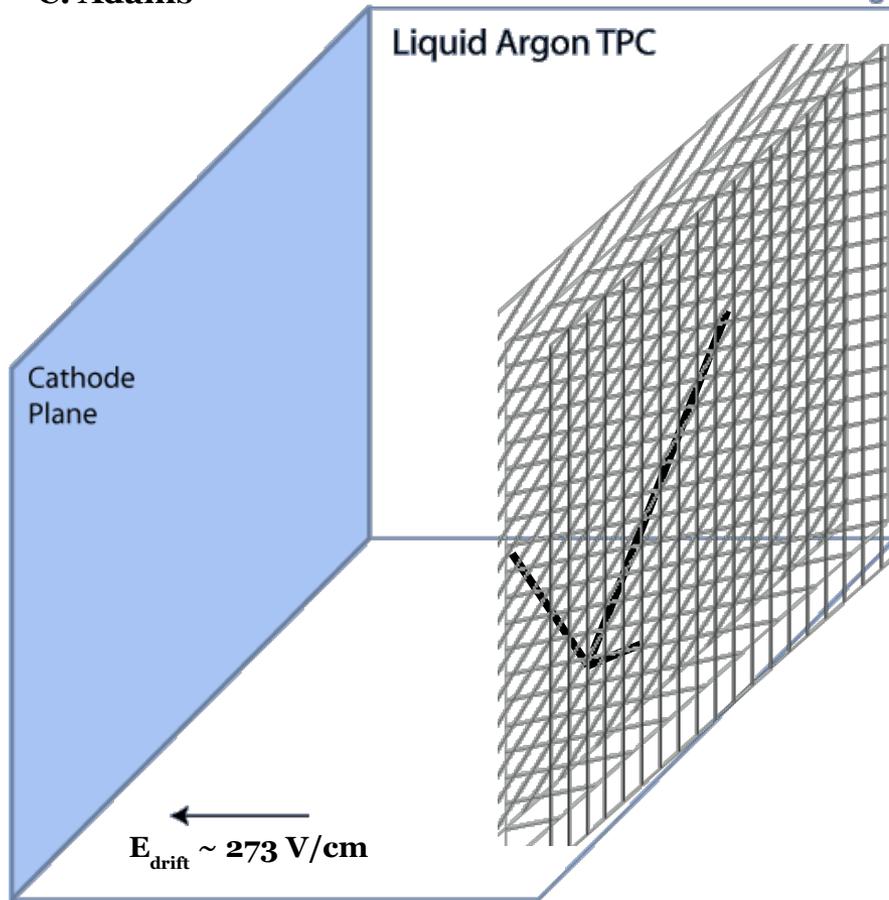


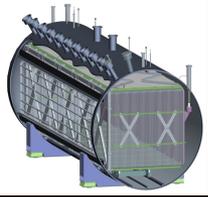


# Signal Formation

Cartoon Credit:  
C. Adams

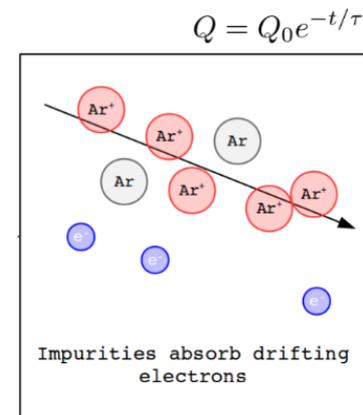
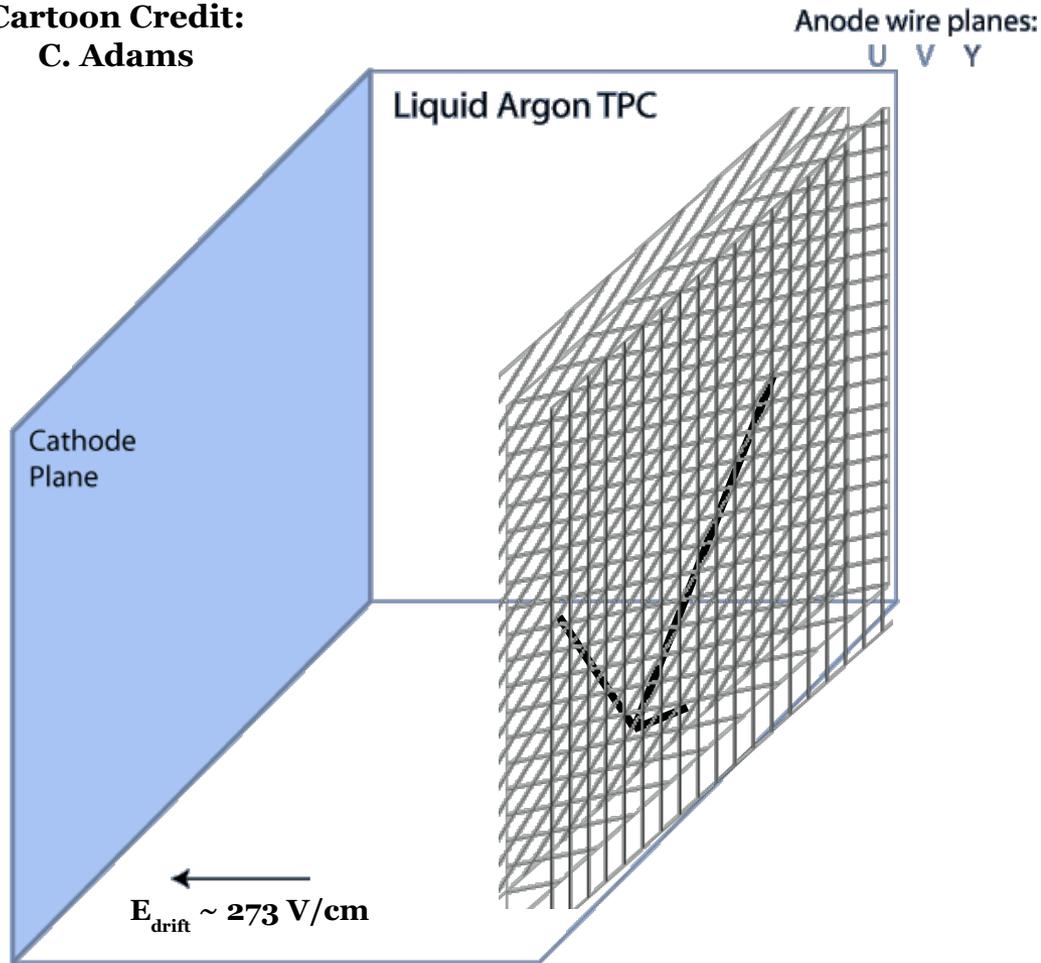
Anode wire planes:  
U V Y

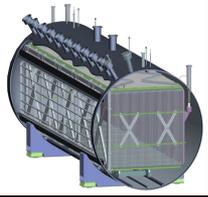




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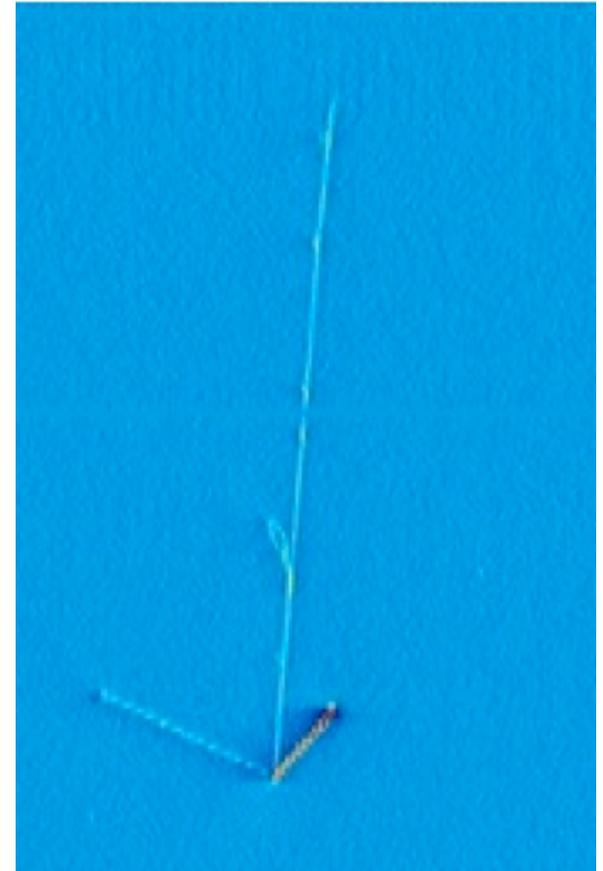
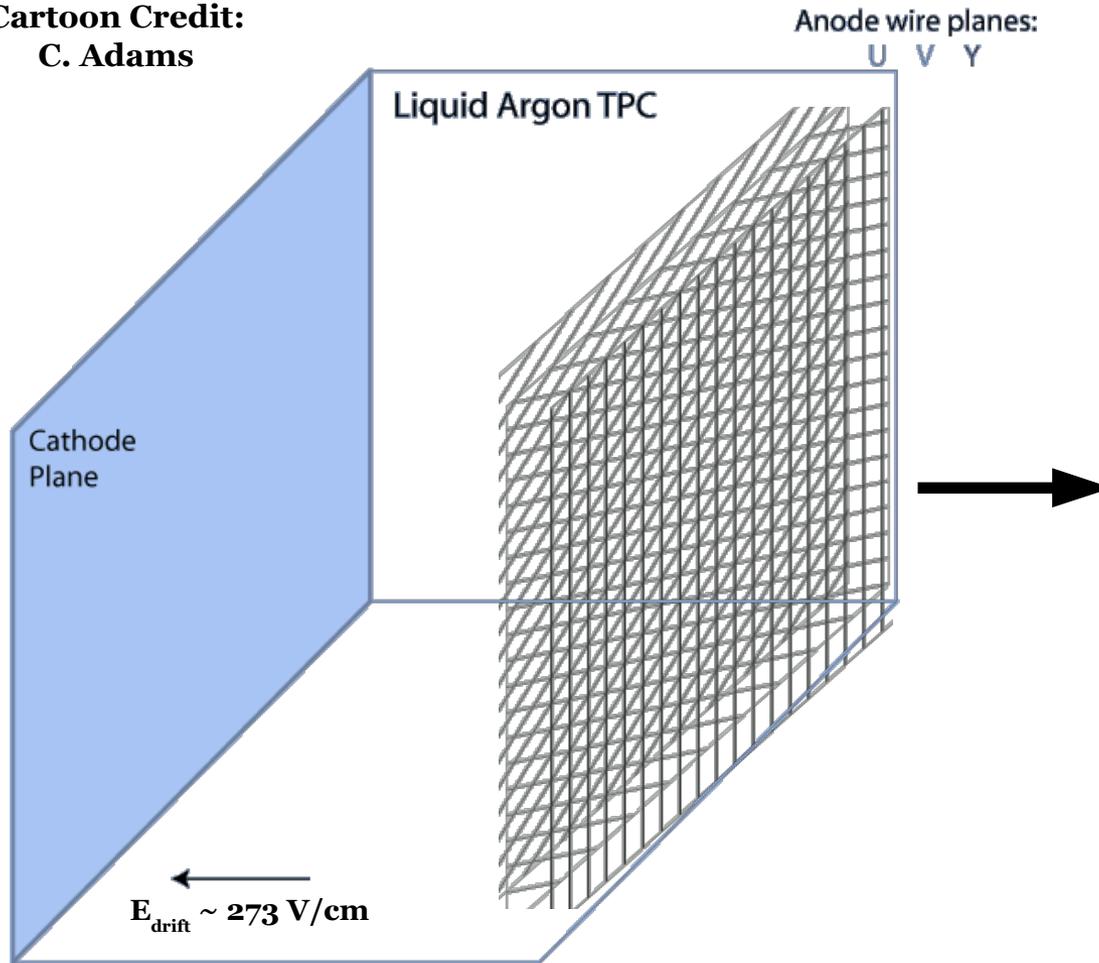
Cartoon Credit:  
C. Adams

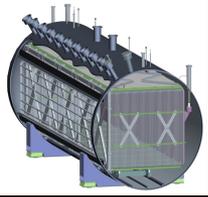




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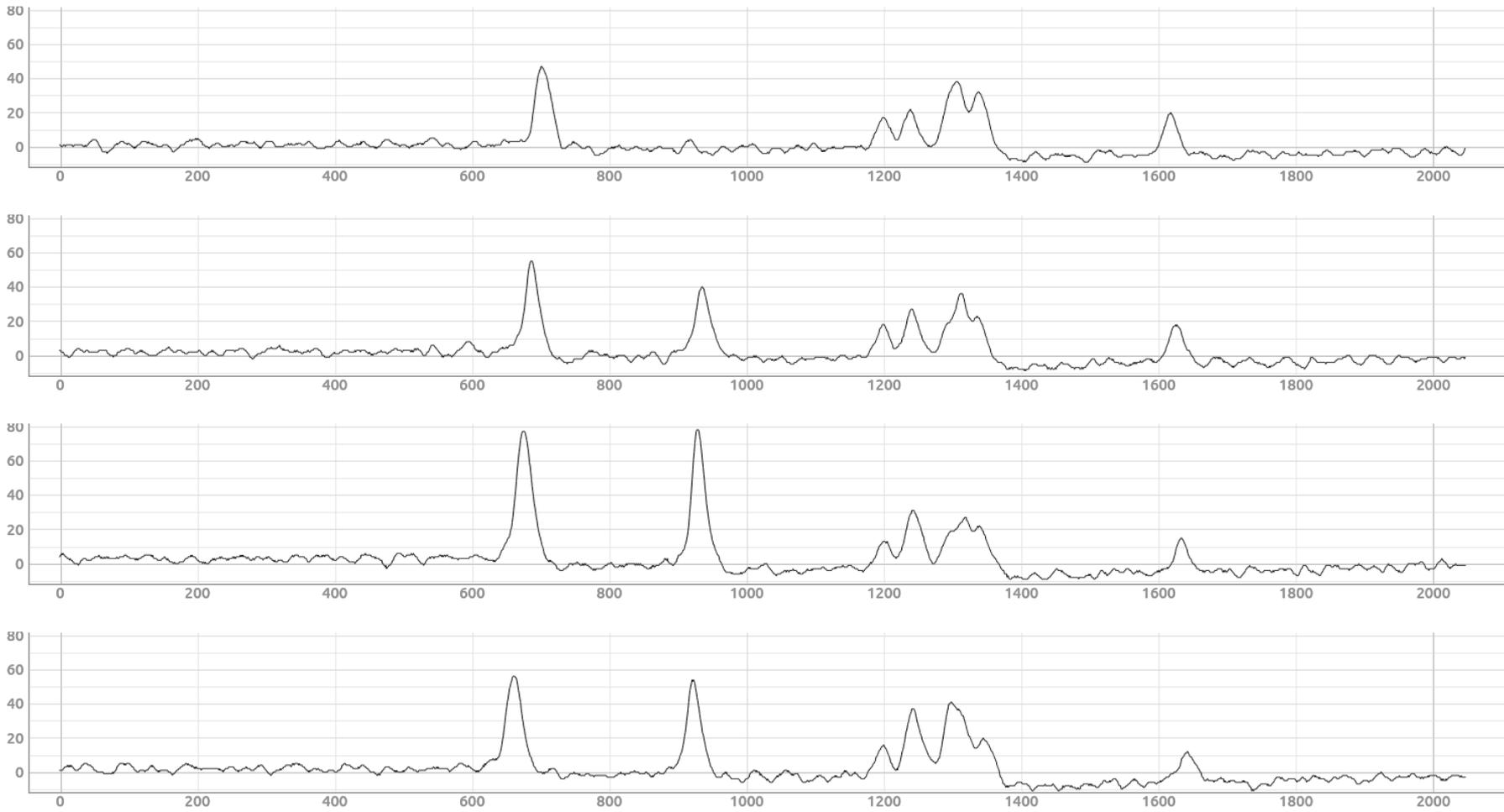
Cartoon Credit:  
C. Adams





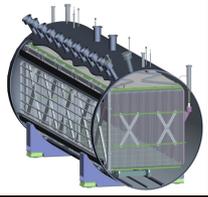
# Raw Waveform Output

Image Credit: C. Adams



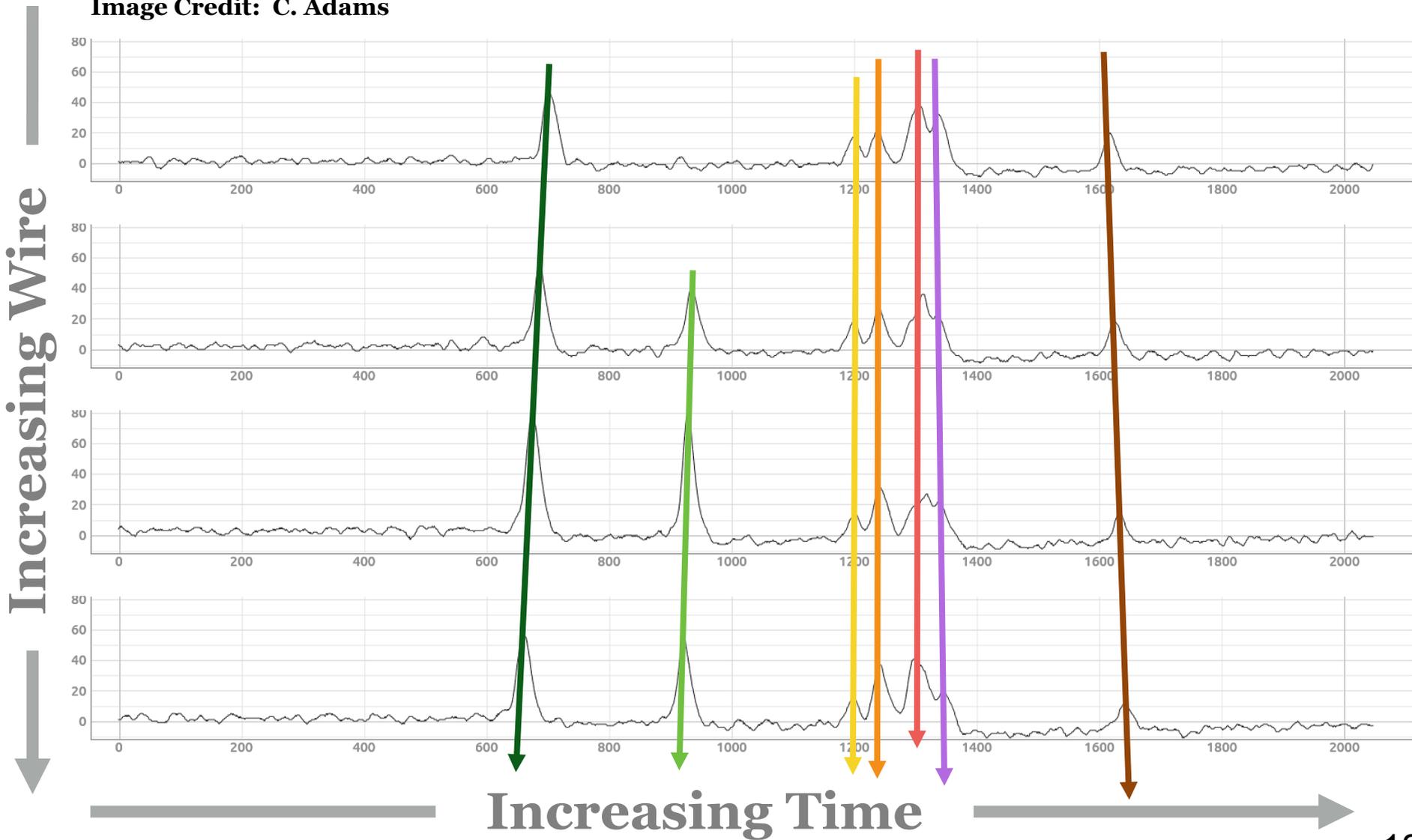
Increasing Wire

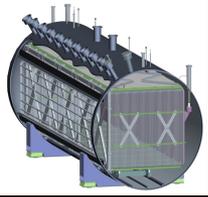
Increasing Time



# Raw Waveform Output

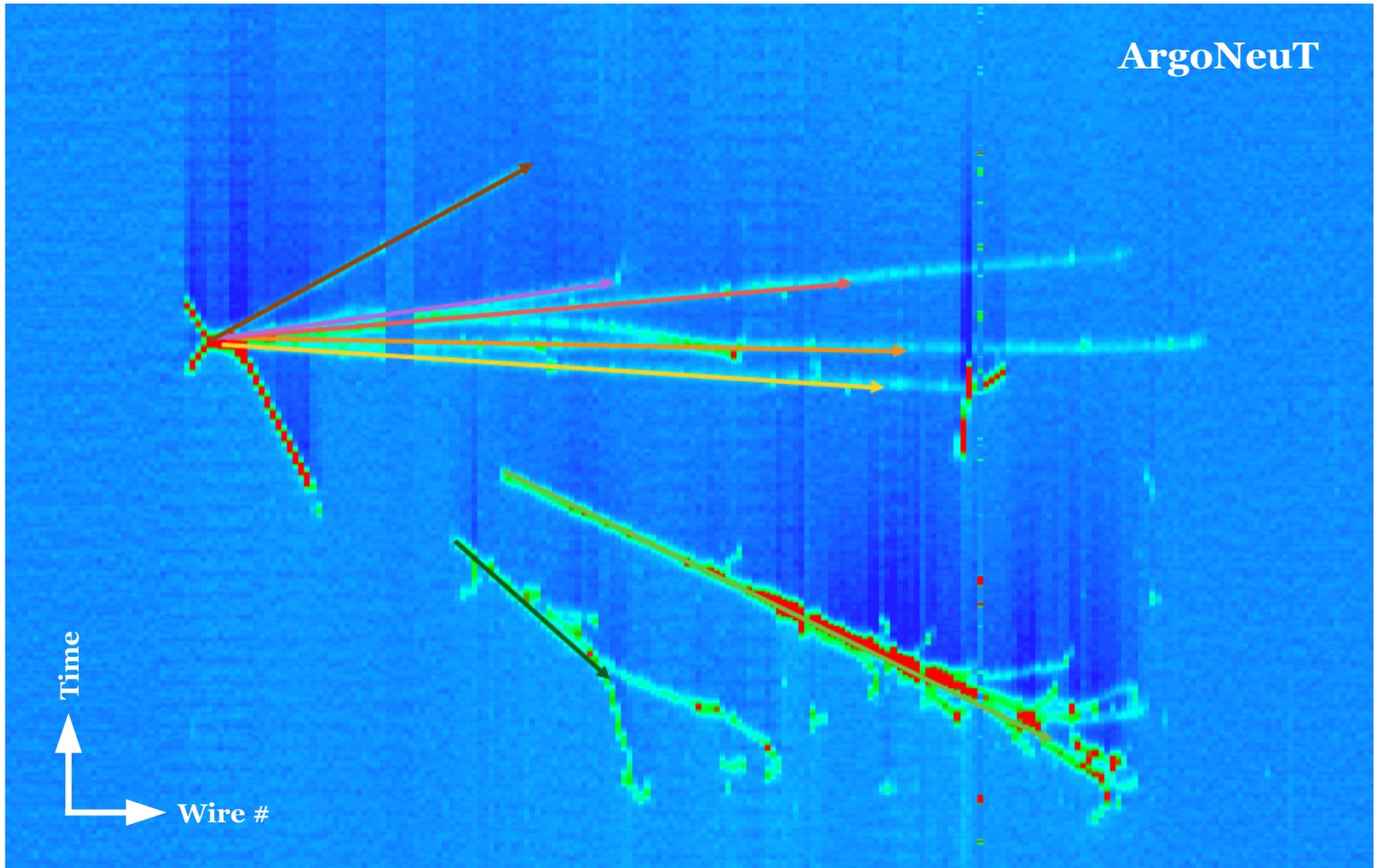
Image Credit: C. Adams

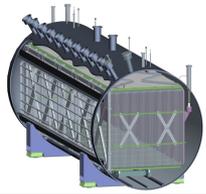




# LArTPC Imaging

Image Credit: C. Adams

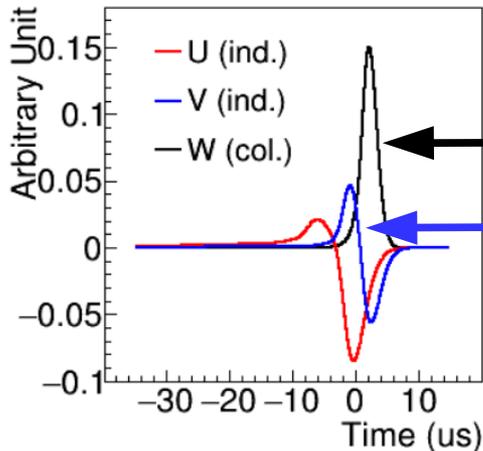
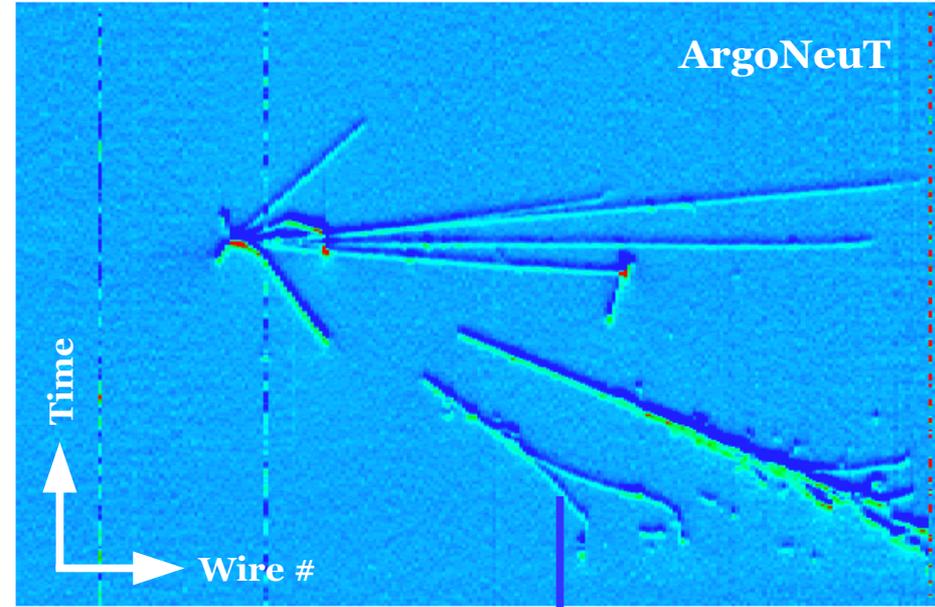
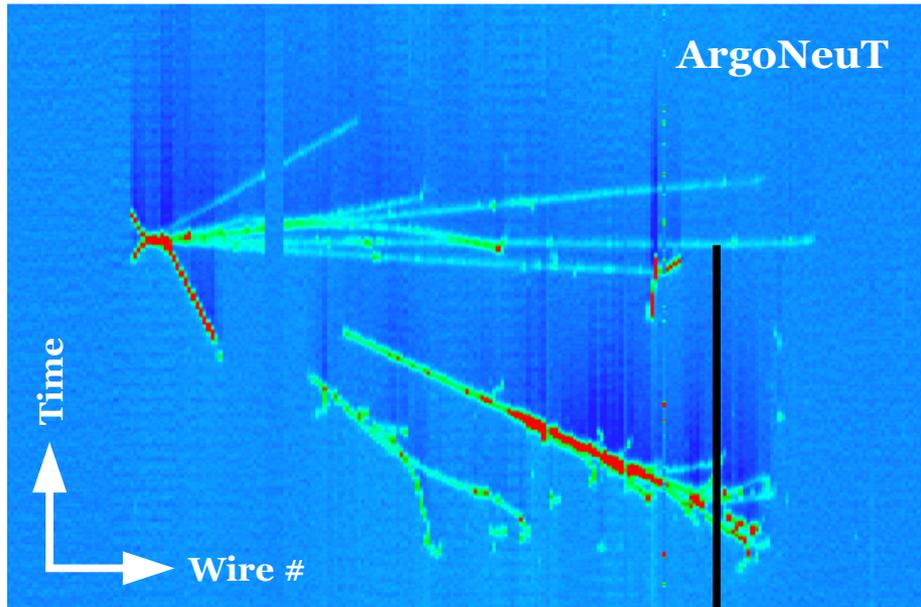




# Different Plane Views

“Collection” Plane (Y)

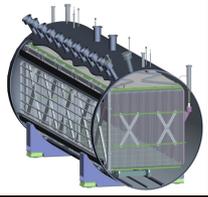
“Induction” Plane (U, V)



**Collection Wire Response:  
Unipolar Signal**

**Induction Wire Response:  
Bipolar Signal**

In order efficiently find ionization signal and correctly determine charge, must first *remove detector response* → **deconvolution**

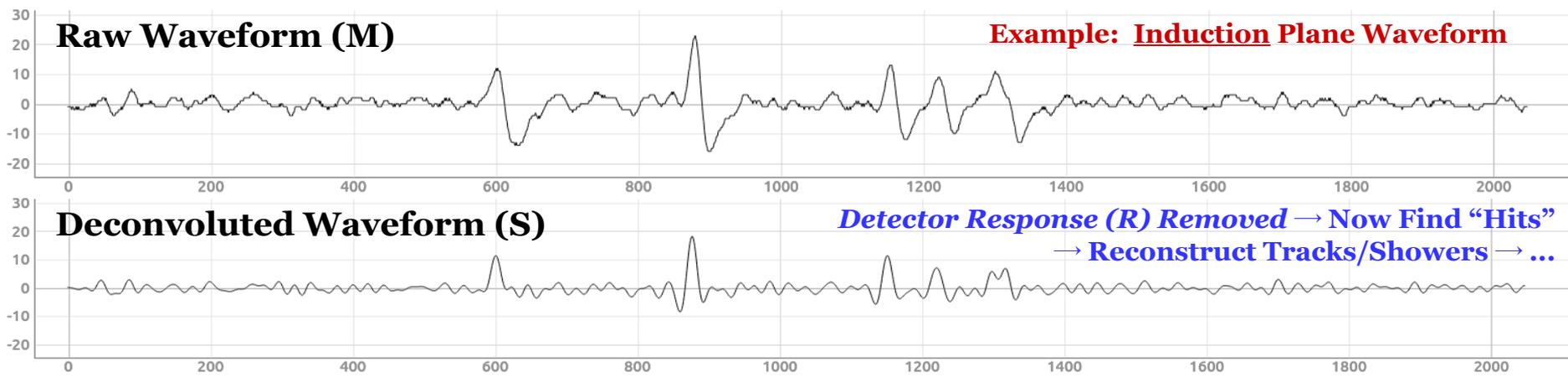
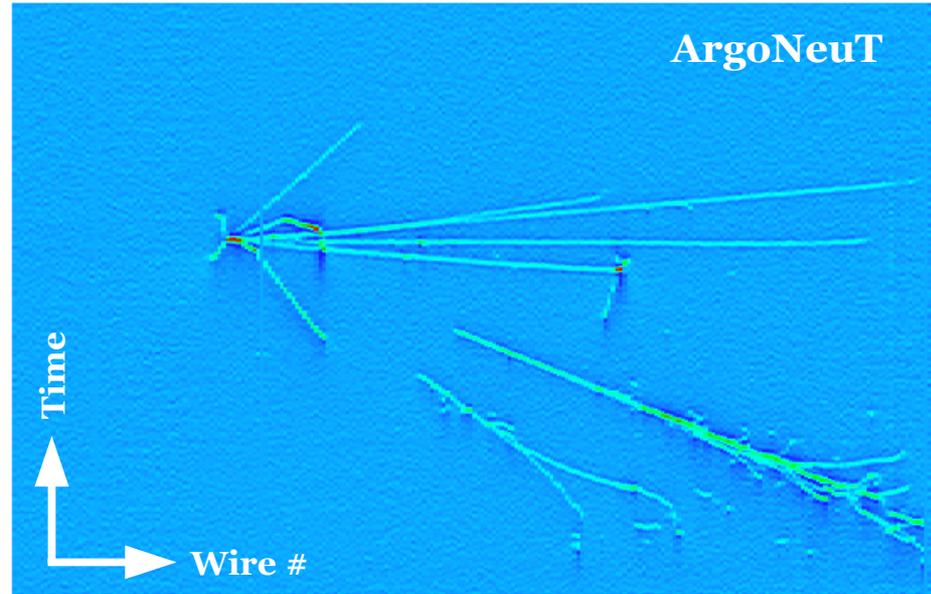
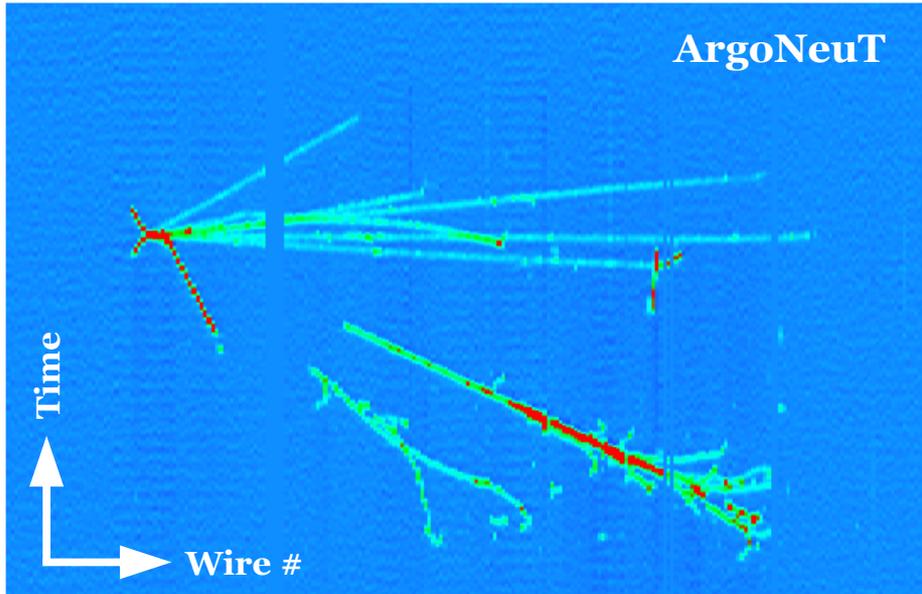


# Deconvolution:

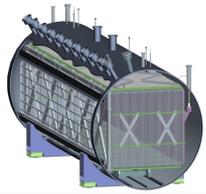
$$S(\omega) = \frac{M(\omega)}{R(\omega)} \cdot F(\omega)$$

## “Collection” Plane (Y)

## “Induction” Plane (U, V)

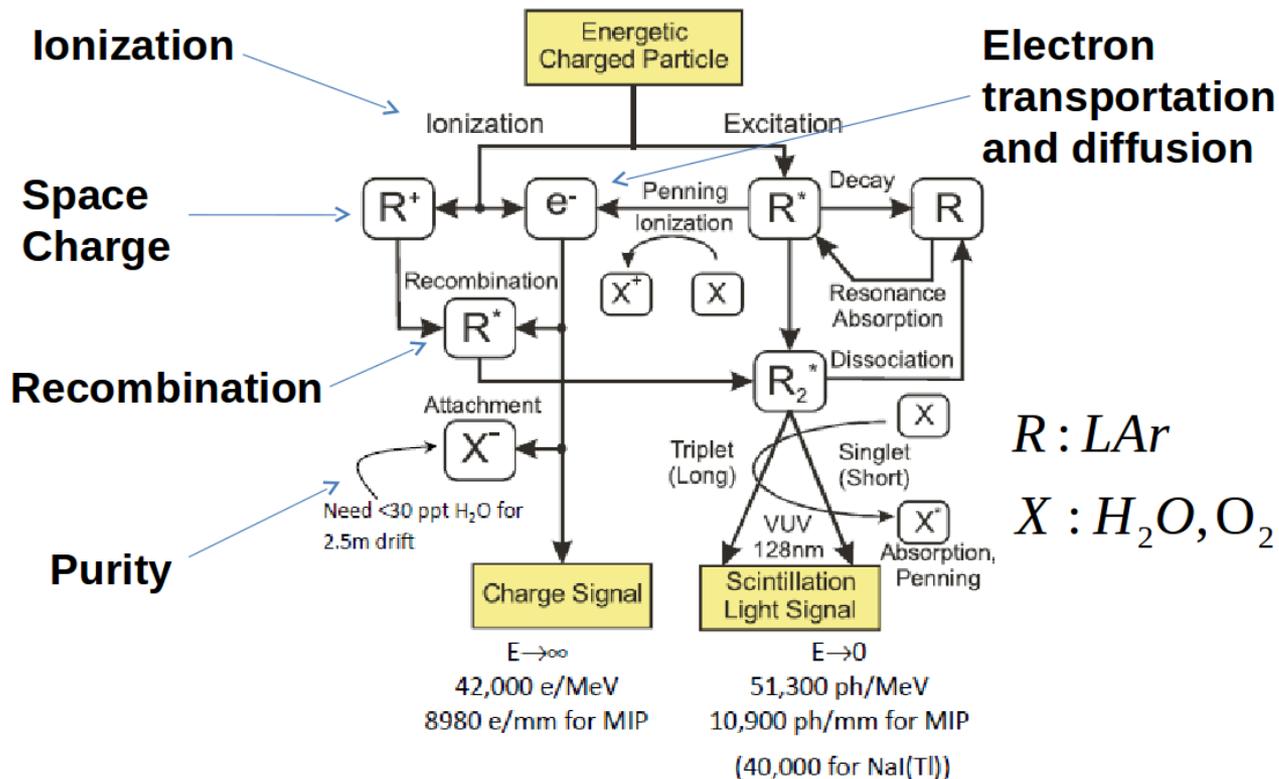


**Filter (F): Prevents Blow-up of Noise During Deconvolution**

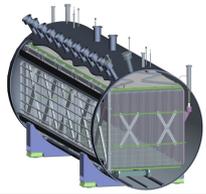


# Calibration Scheme

- ◆ Must understand detector effects to develop LArTPC technology
  - Essential for SBN and DUNE
  - Noise removal, space charge effects (SCE), wire response, energy scale, diffusion,  $e^-$  lifetime, etc.

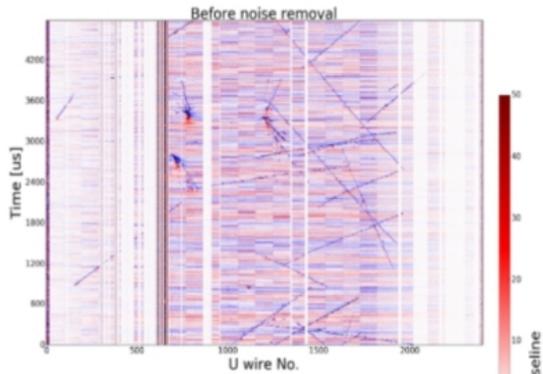


**Important to understand detector effects and develop calibration scheme for unbiased, precise determination of ionization charge.**

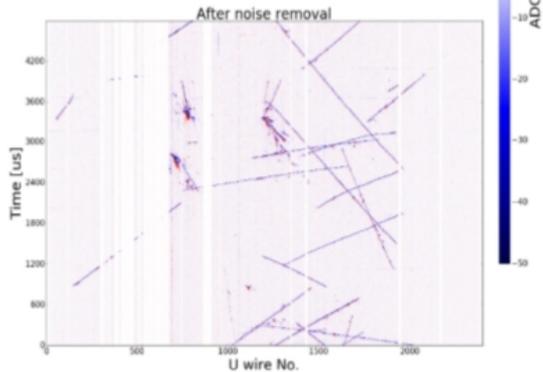


# Reducing Noise Levels

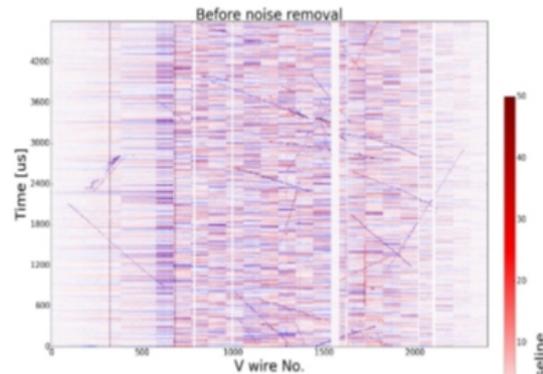
## First Induction (U)



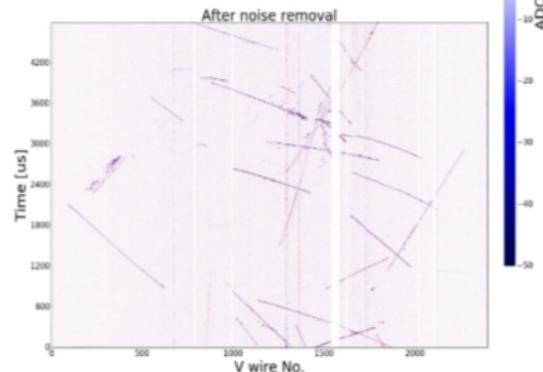
MicroBooNE Preliminary



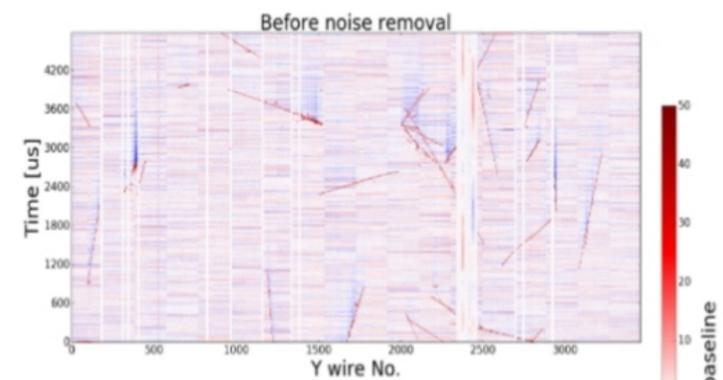
## Second Induction (V)



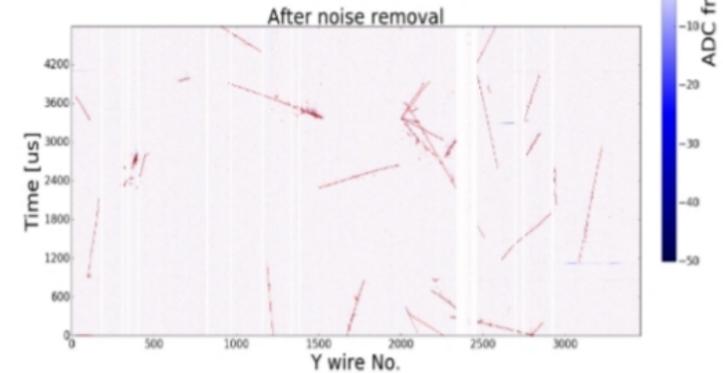
MicroBooNE Preliminary



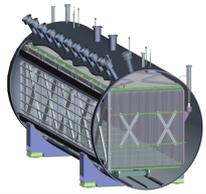
## Collection (Y)



MicroBooNE Preliminary

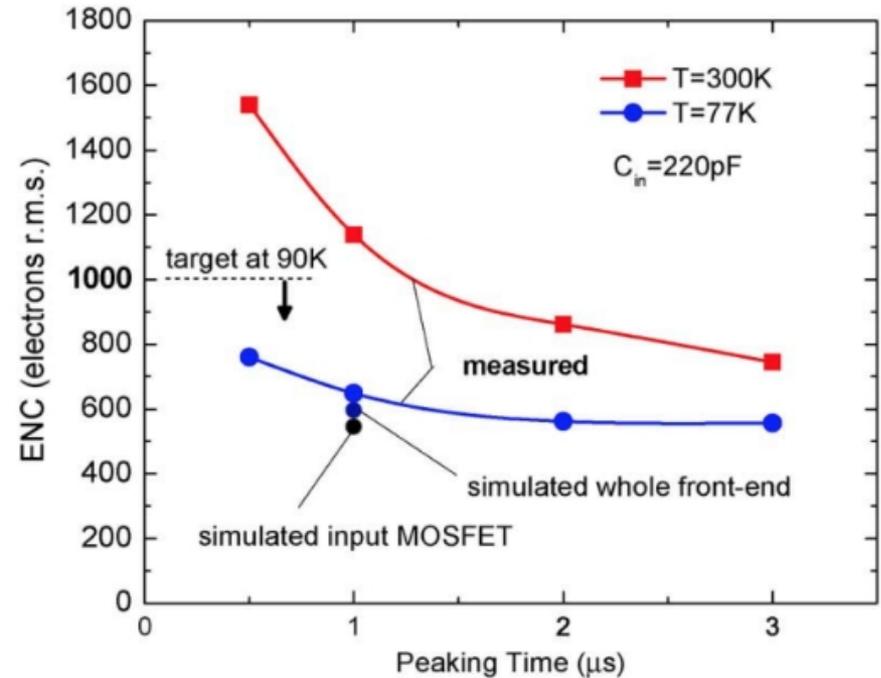
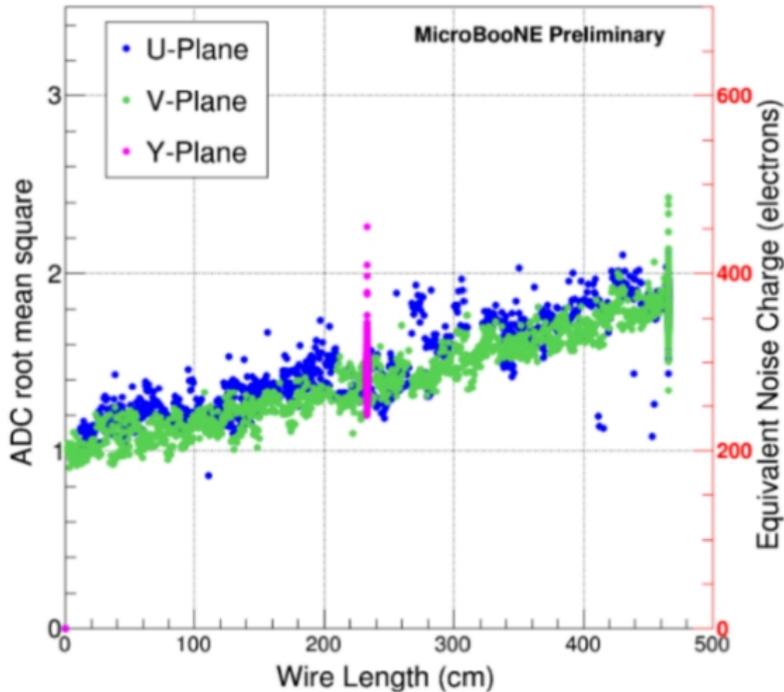


- ◆ Significantly more noise “out of the box” on induction planes (top row)
- ◆ All planes look very clean after software noise filtering (bottom row)

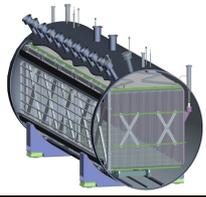


# Final Noise Levels

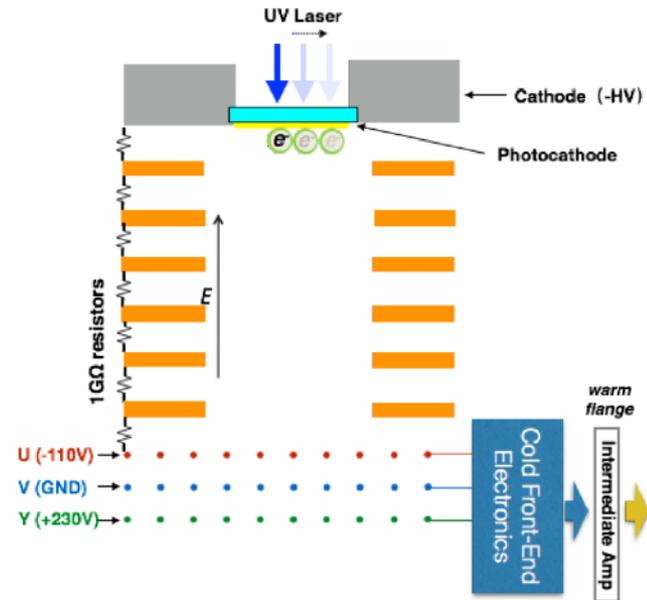
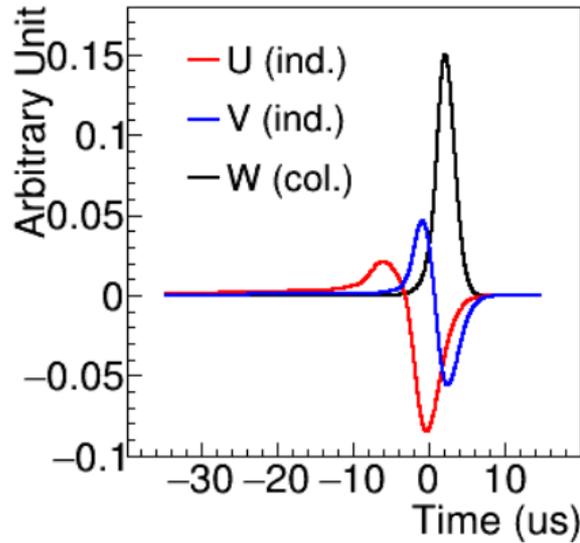
Wire Noise Level in MicroBooNE



- ◆ After software noise filtering on MicroBooNE data, see noise levels expected from bench measurements of cold front-end electronics
  - Scales linearly with wire length (capacitance)
  - Thanks to cold front-end electronics and noise-filtering techniques, low ENC achievable in 100-ton-scale LArTPC: **ENC < 400 e<sup>-</sup>**

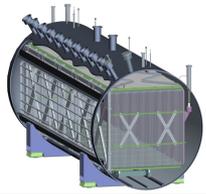


# Field Response Measurement



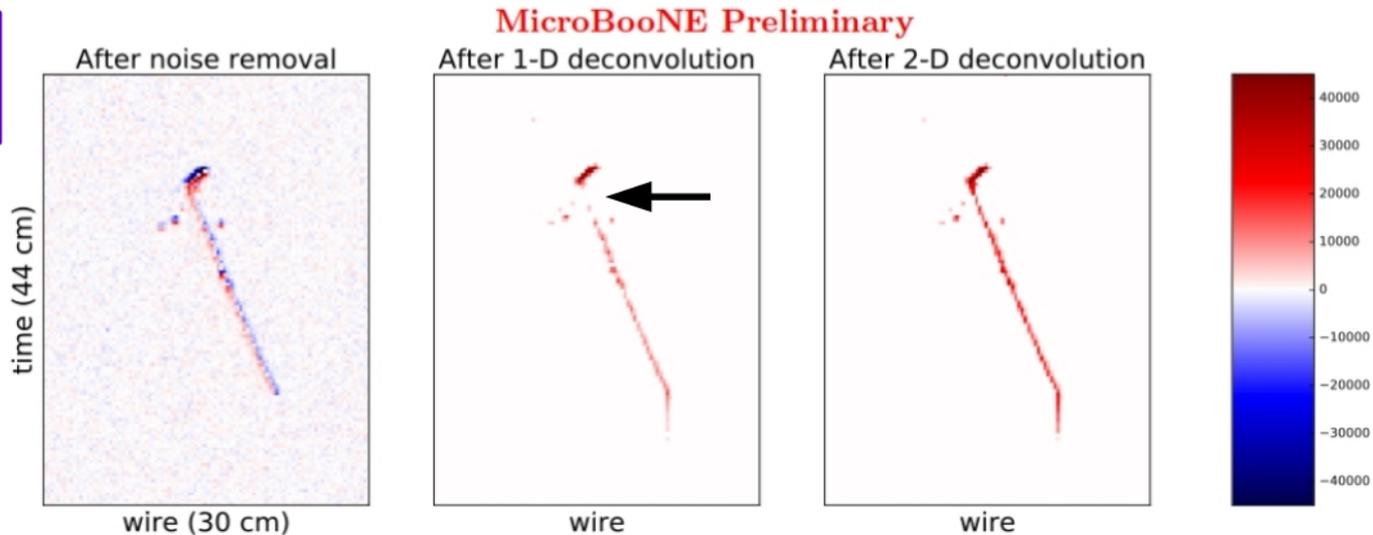
LAr Field Calibration System (“LARFCS”)

- ◆ Must remove correct field response of wires in deconvolution to enable unbiased charge estimation
- ◆ Simulated field response (Garfield) needs verification with **data**
- ◆ Measurement with MIPs in situ folds in track extent across wire pitch
- ◆ **BNL test stand** aiming to make measurement with point-like source from laser pulsed on photocathode – “LARFCS”

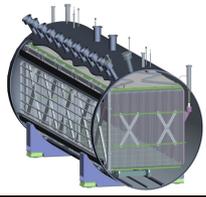


# Dynamic Induced Charge

## Second Induction Plane

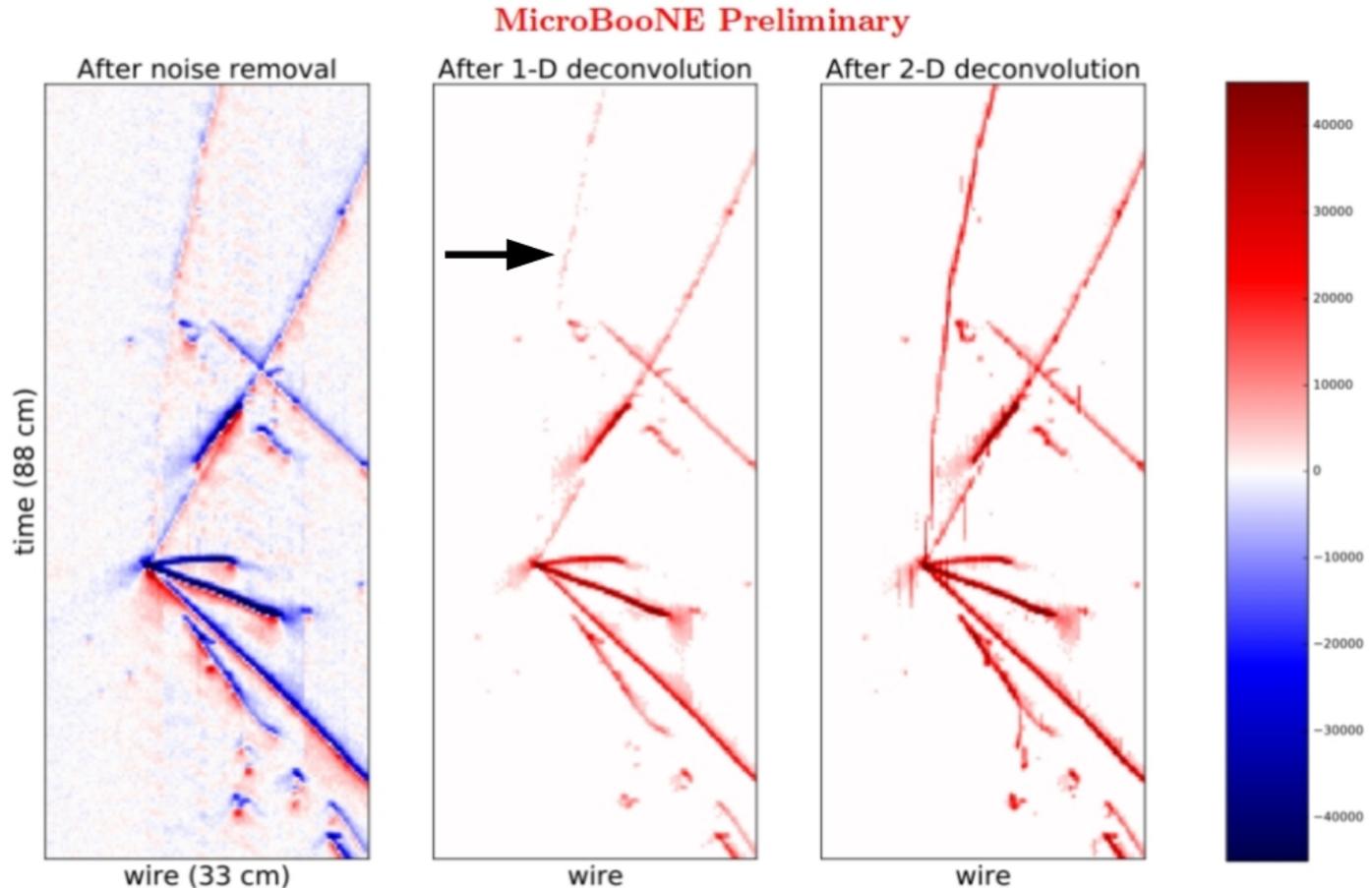


- ◆ Nominally assume ionization leads to signal on only one wire
- ◆ In reality, nearby wires also see some signal
  - Characteristics of this induced signal dynamically dependent on track angle – “Dynamic Induced Charge” or DIC
  - Effect leads to cancellation of signals on waveform for tracks at high angles → **hits lost** → **problems in track/shower reconstruction**
- ◆ Solution is to account for DIC when removing detector response in deconvolution to extract charge – improves imaging

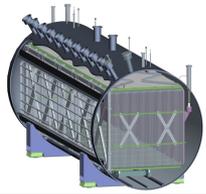


# Another DIC Example

First  
Induction  
Plane

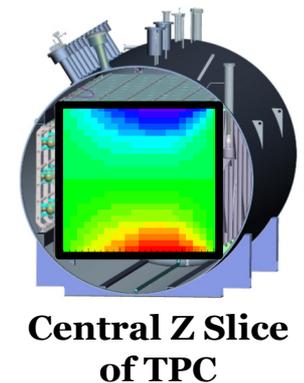
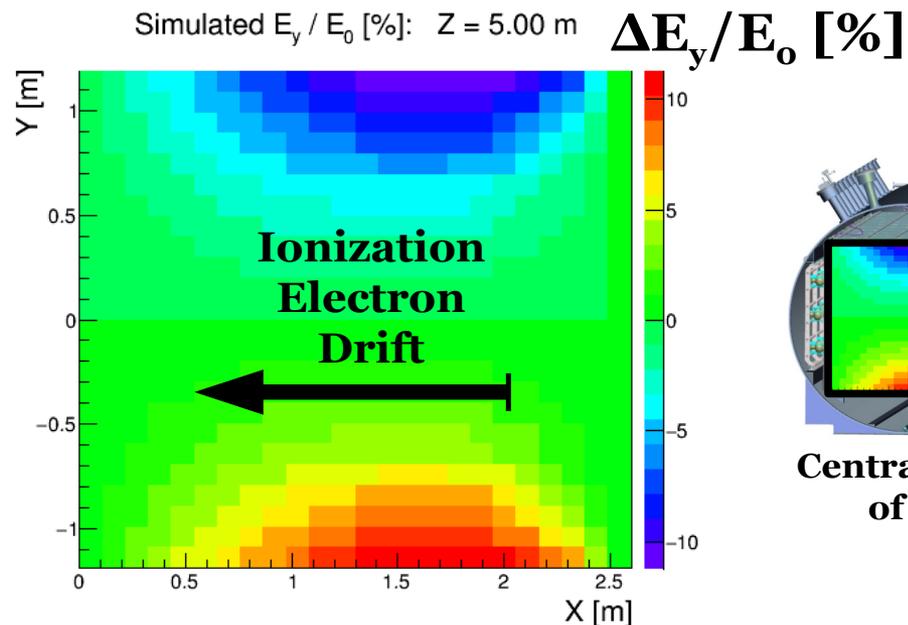
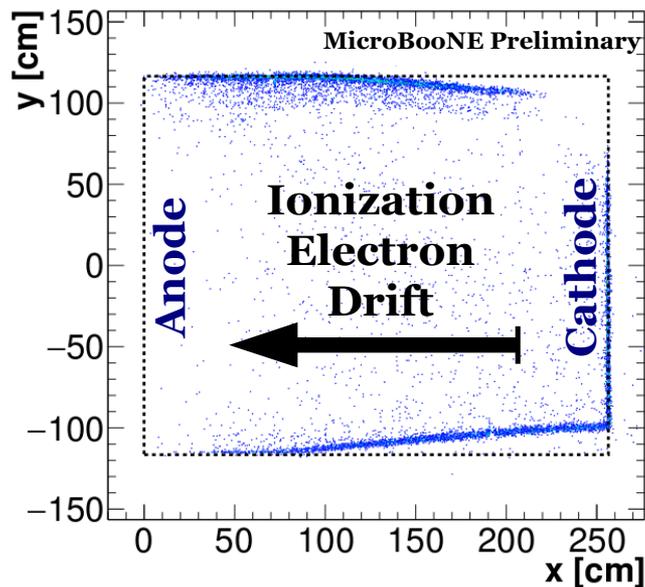


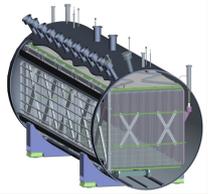
- ◆ Effect can be studied in depth at **LArFCS** with point-like source
- ◆ Smaller effect for larger wire spacing (e.g. 5 mm for DUNE)



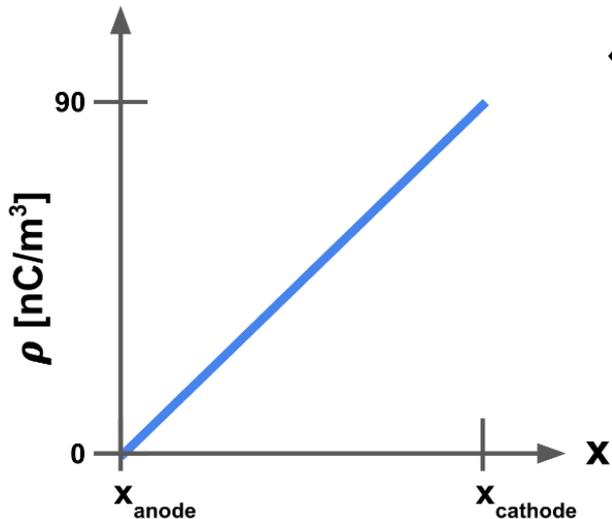
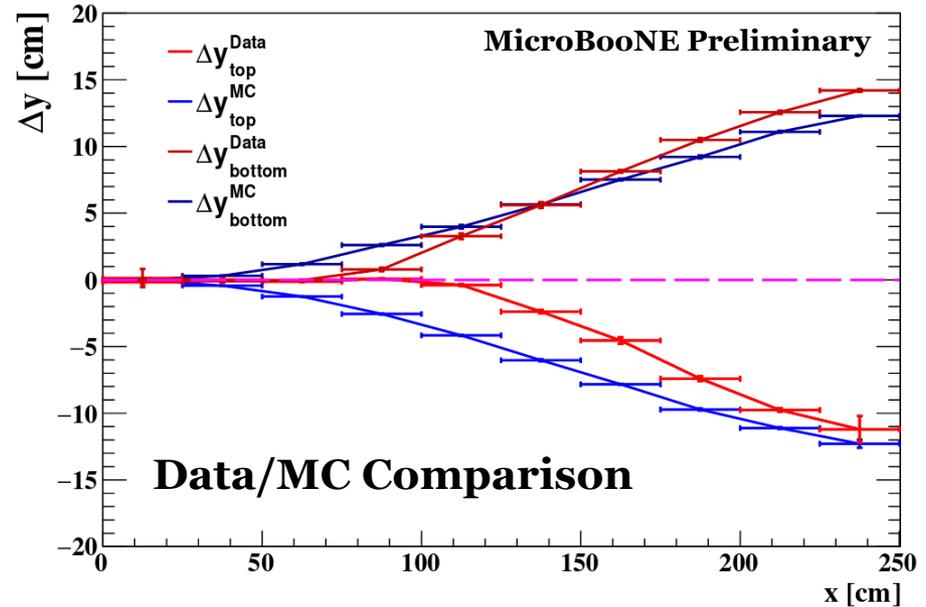
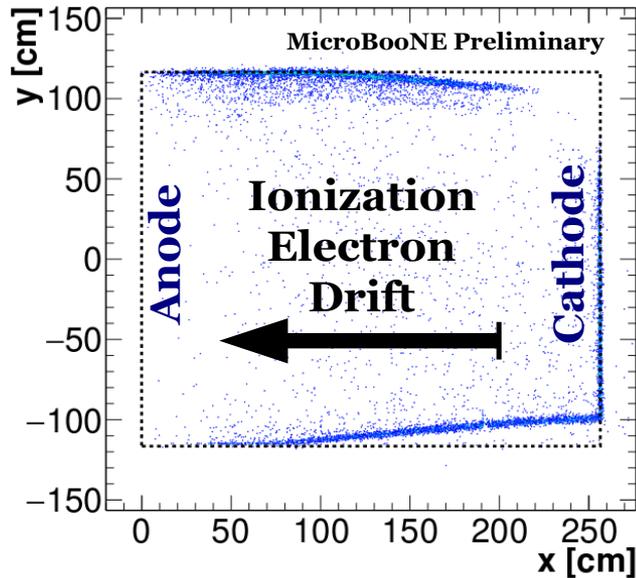
# Space Charge Effects

- ◆ Looking at cosmic data, noticed offsets in track start/end points from top/bottom of TPC
  - Very suggestive of space charge effects (SCE) at MicroBooNE, a near-surface experiment (20-30 cosmits per 4.8 ms readout window)
  - **Space charge:** build-up of slow-moving  $\text{Ar}^+$  ions due to e.g. cosmic muons impinging active volume of TPC (via ionization)
  - Leads to E field distortions, spatial distortions in ionization position

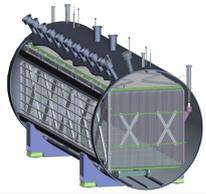




# SCE Data/MC Comparison

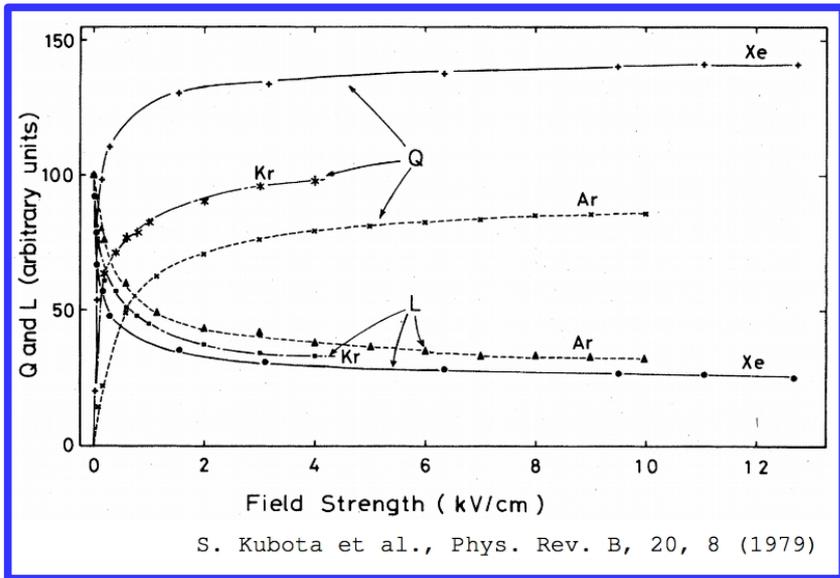
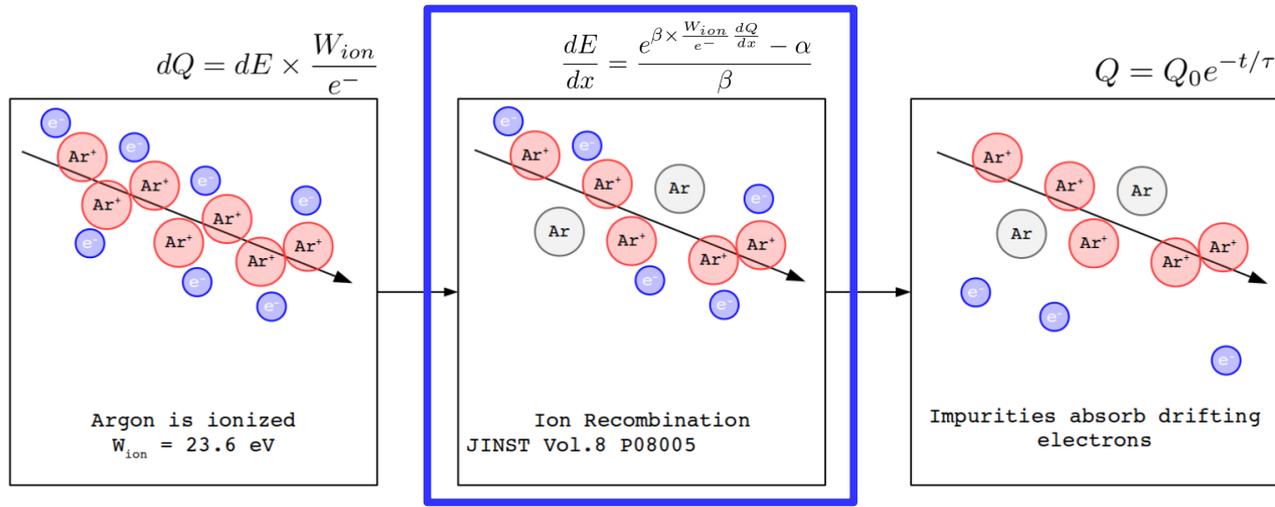


- ◆ SCE simulation qualitatively reproduces effect
  - Assumes linear space charge profile
  - Agreement in normalization, basic shape features, but offset near anode in data... consistent with impact from **liquid argon flow**
  - Can impact track/shower reconstruction and calorimetry – calibrate out **in 3D** using **UV laser system**, cosmic muon tracks

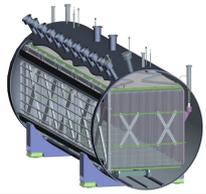


# Ion Recombination

Image Credit:  
D. Caratelli

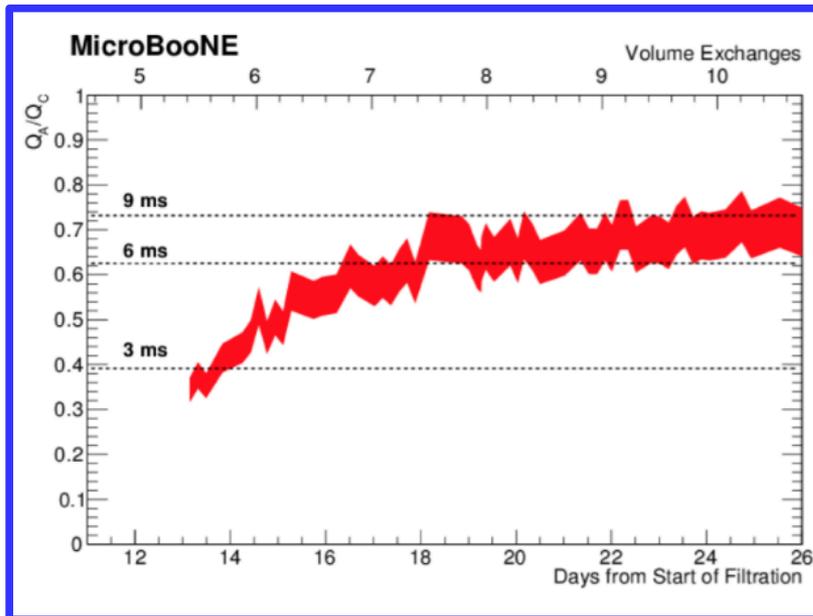
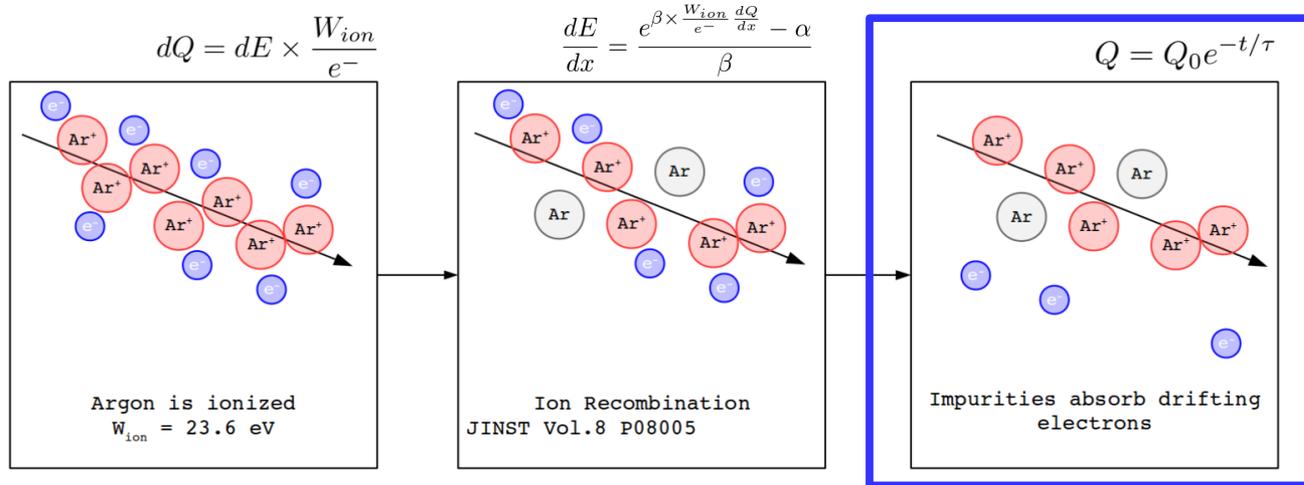


- ◆ Charge quenching from prompt recombination of ionization electrons with argon ions leads to charge loss
  - Sizable: ~**50%** Q loss (@ 273 V/cm)
- ◆ Correction depends on E field, dE/dx
  - Tracks (muon, proton, etc.): **simple**
  - Electromagnetic showers: apply to individual charge depositions based on dE/dx – **more complicated**

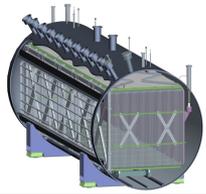


# Electron Lifetime

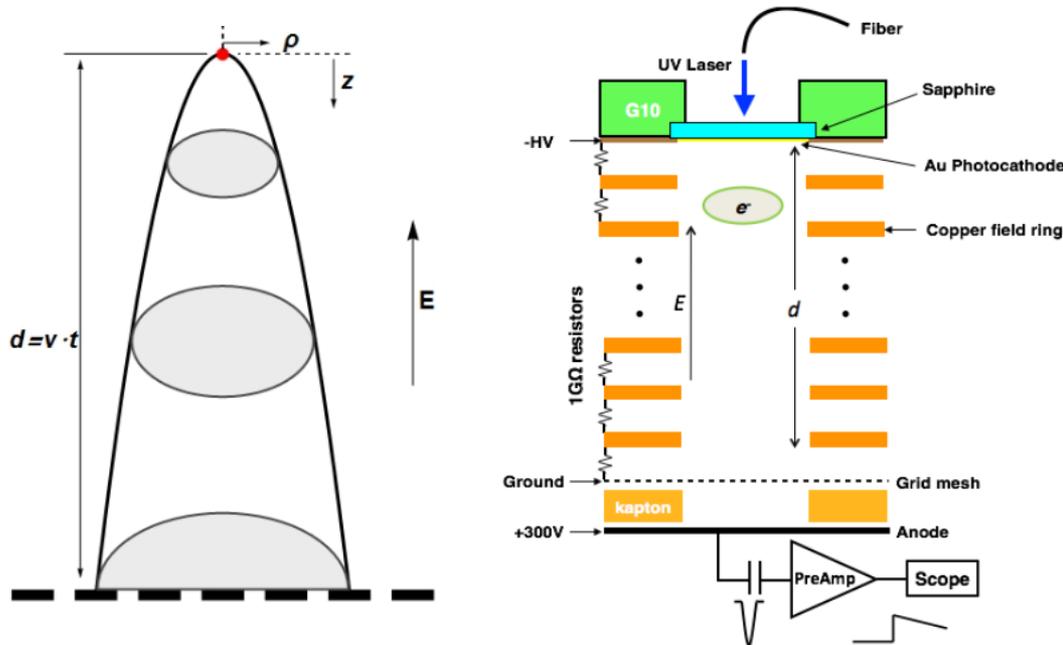
Image Credit:  
D. Caratelli



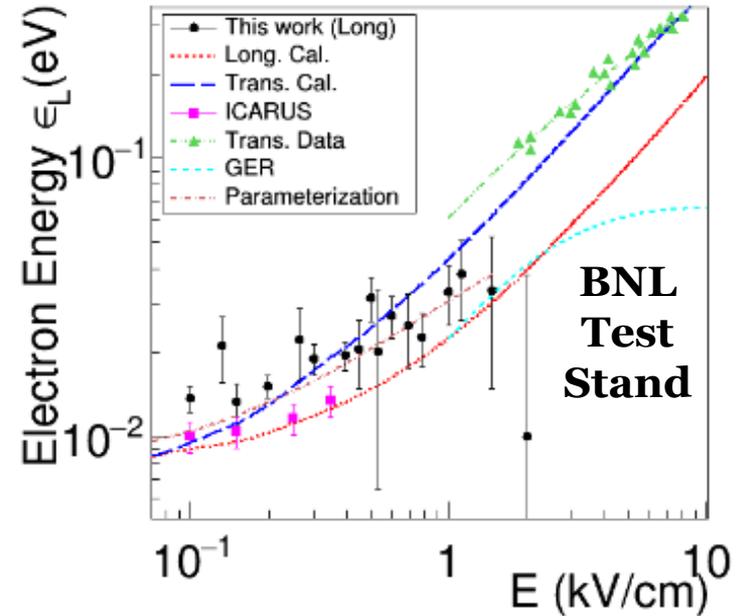
- ◆ Electron lifetime, as measured by purity monitors, consistently **above 6 ms** for majority of run thus far – design: **3 ms**
  - **6 ms**: conservative lower bound
- ◆ Important conclusion: can operate LArTPCs in non-evacuated cryostats with high electron lifetime
- ◆ Calibrate out via measurement of charge from TPC tracks vs. drift distance



# Diffusion



See Y. Li et al. NIMA 816 160 (2016)



- ◆ Diffusion can reduce the spatial resolution of reconstructed particle trajectories, especially for longer drift times → **must measure, simulate**
- ◆ Important for e.g. supernova neutrinos vs. alpha/beta decays (track-like vs. point-like)
- ◆ Measure longitudinal diffusion at **BNL test stand**

$$\sigma_{LIT} = \sqrt{\frac{2 \cdot \epsilon_{LIT} \cdot d}{E}} \cdot \frac{t}{d}$$

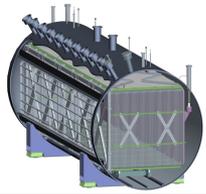
$E$ : electric field

$d$ : drift distance

$t$ : drift time

$\sigma$ : width of electron cloud

$\epsilon$ : electron energy



# Diffusion

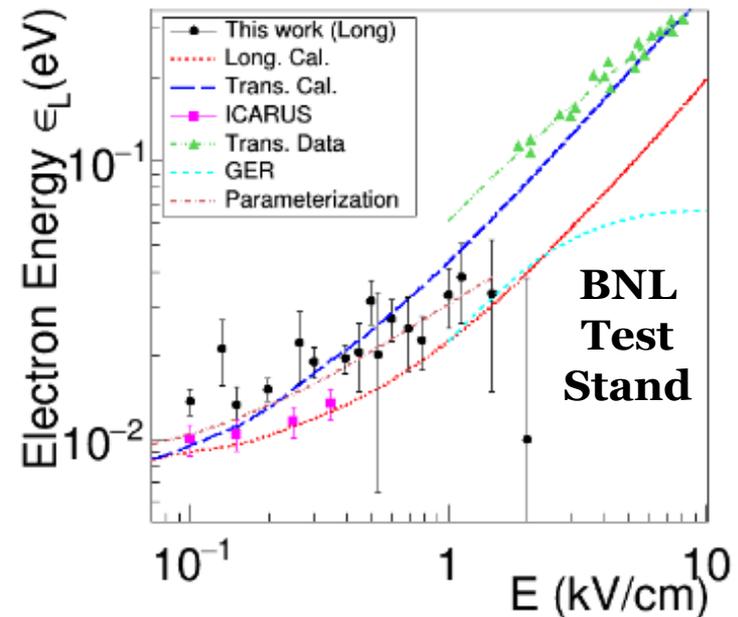
$$\epsilon_L = \frac{b_0 + b_1 E + b_2 E^2}{1 + (b_1/b_0)E + b_3 E^2} \left( \frac{T}{T_1} \right)$$

$$\begin{aligned} b_0 &= 0.0075 \\ b_1 &= 742.9 \\ b_2 &= 3269.6 \\ b_3 &= 31678.2 \end{aligned}$$

**500  
V/cm  
Drift  
Field**

<b>Drift Length</b>	$\sigma_L$	$\sigma_T$
<b>1.8 m</b>	<b>1.2 mm</b>	<b>1.7 mm</b>
<b>3.6 m</b>	<b>1.8 mm</b>	<b>2.5 mm</b>
<b>6.0 m</b>	<b>2.3 mm</b>	<b>3.1 mm</b>
<b>20.0 m</b>	<b>4.2 mm</b>	<b>5.7 mm</b>

See Y. Li et al. NIMA 816 160 (2016)



- ◆ Diffusion can reduce the spatial resolution of reconstructed particle trajectories, especially for longer drift times → **must measure, simulate**
- ◆ Important for e.g. supernova neutrinos vs. alpha/beta decays (track-like vs. point-like)
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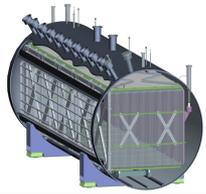
$E$ : electric field

$d$ : drift distance

$t$ : drift time

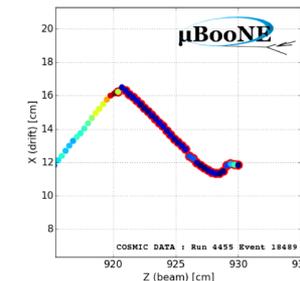
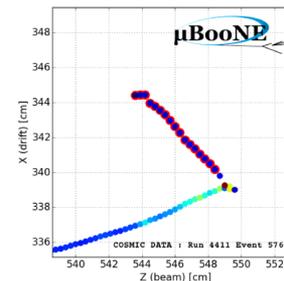
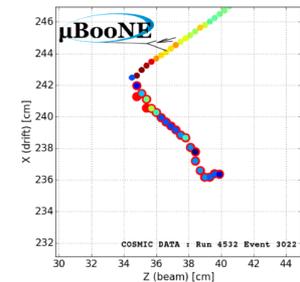
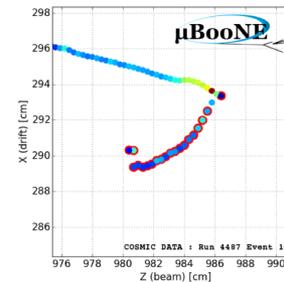
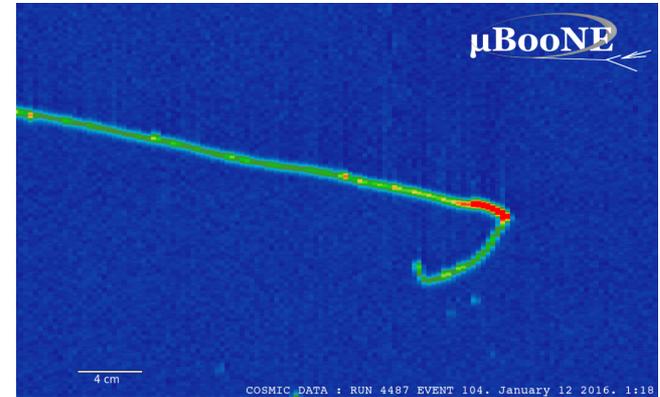
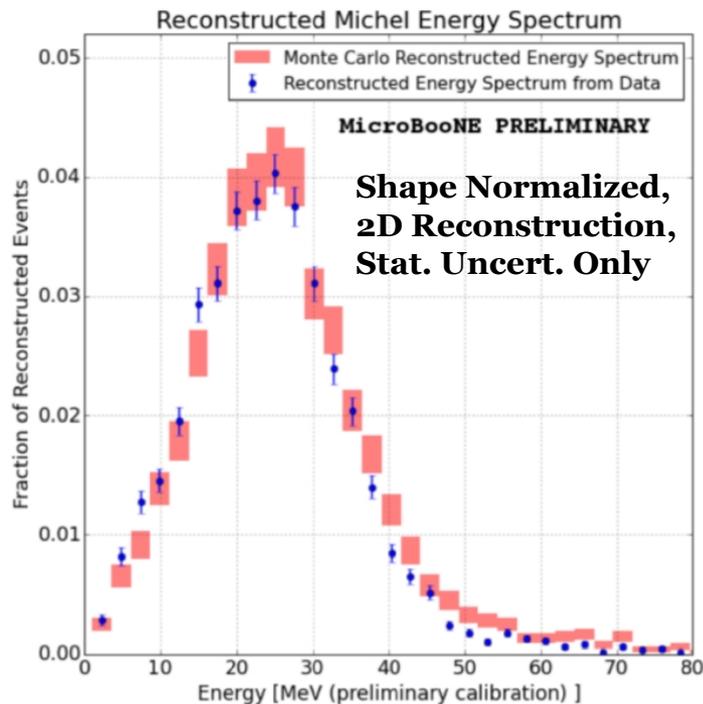
$\sigma$ : width of electron cloud

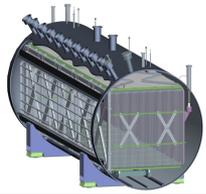
$\epsilon$ : electron energy



# Michel $e^-$ Spectrum

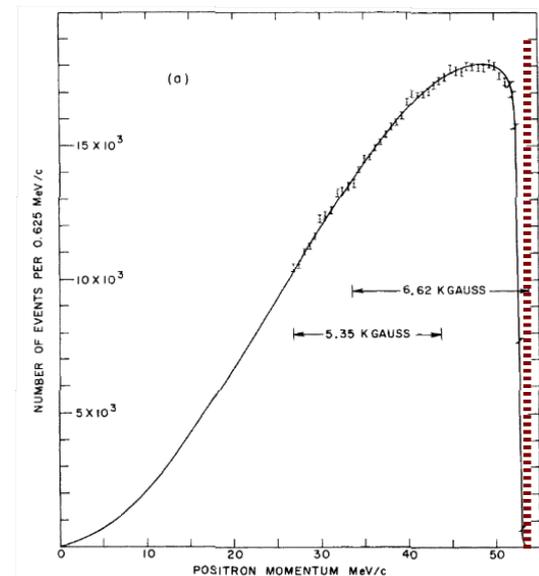
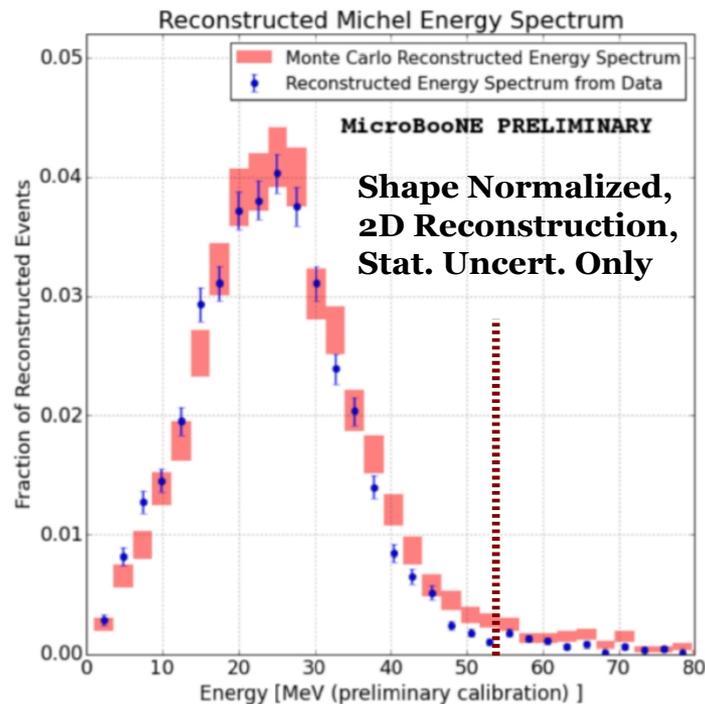
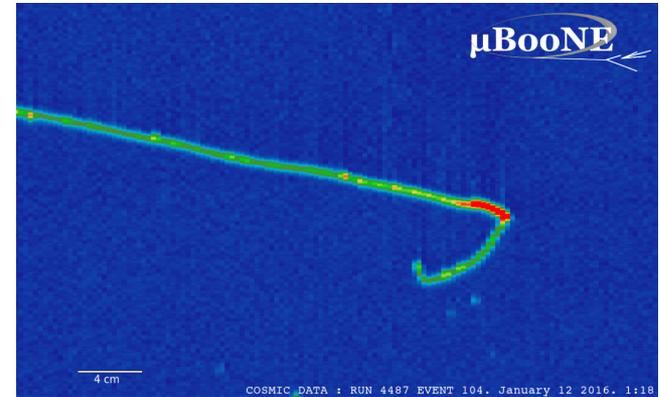
- ◆ Tag Michel electrons from cosmic muon decay using “kink” topology and muon Bragg peak
  - Uses automated reconstruction
  - Important calibration sample for **energy scale**, tuning  $e^-$ ,  $\gamma$  **reconstruction** (charge clustering)

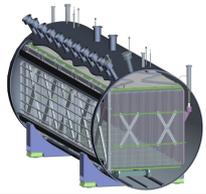




# Michel $e^-$ Spectrum

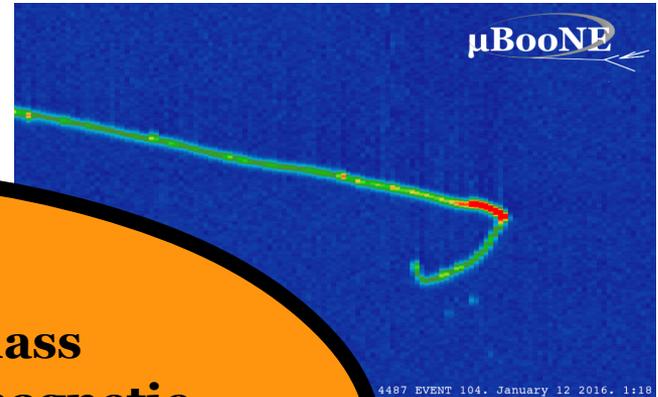
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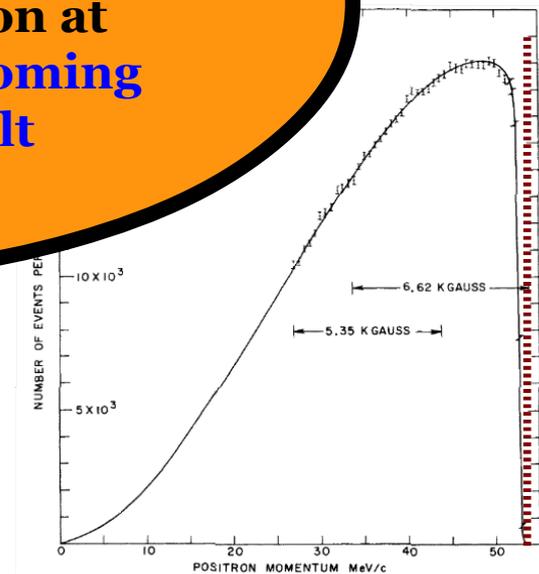
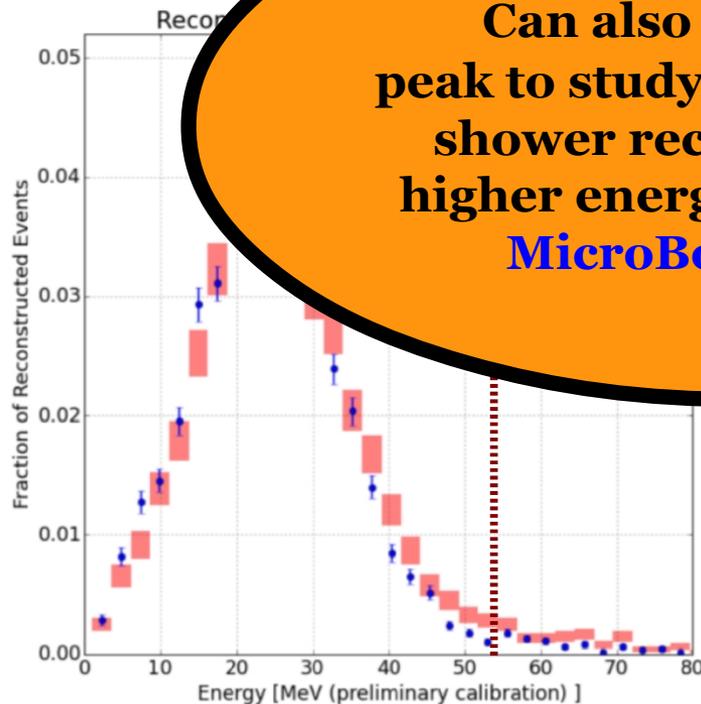


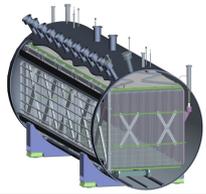
# Michel $e^-$ Spectrum

- ◆ Tag Michel electrons from cosmic muon decay using “kink” topology and muon Bragg peak
  - Uses automated reconstruction
  - Important calibration sample for **energy scale**, **timing**, and **reconstruction**



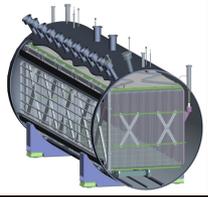
Can also use  $\pi^0$  mass peak to study electromagnetic shower reconstruction at higher energies – **upcoming MicroBooNE result**



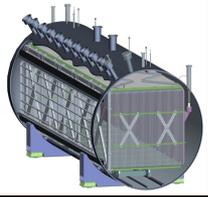


# Summary

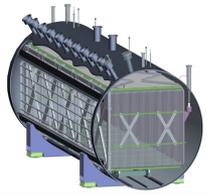
- ◆ LArTPC calibration essential for unbiased, precise determination of ionization charge
- ◆ This requires first **removing noise** to **find signals**
- ◆ Then **account for detector effects**, including wire response, to **obtain charge information** correctly
- ◆ Finally use **high-level candles** (e.g. Michel electrons) to **tune particle trajectory/energy reconstruction**
- ◆ Extensive process, but necessary before producing robust physics measurements



Thanks!



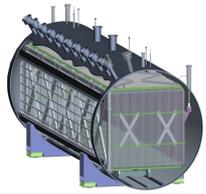
# Backup



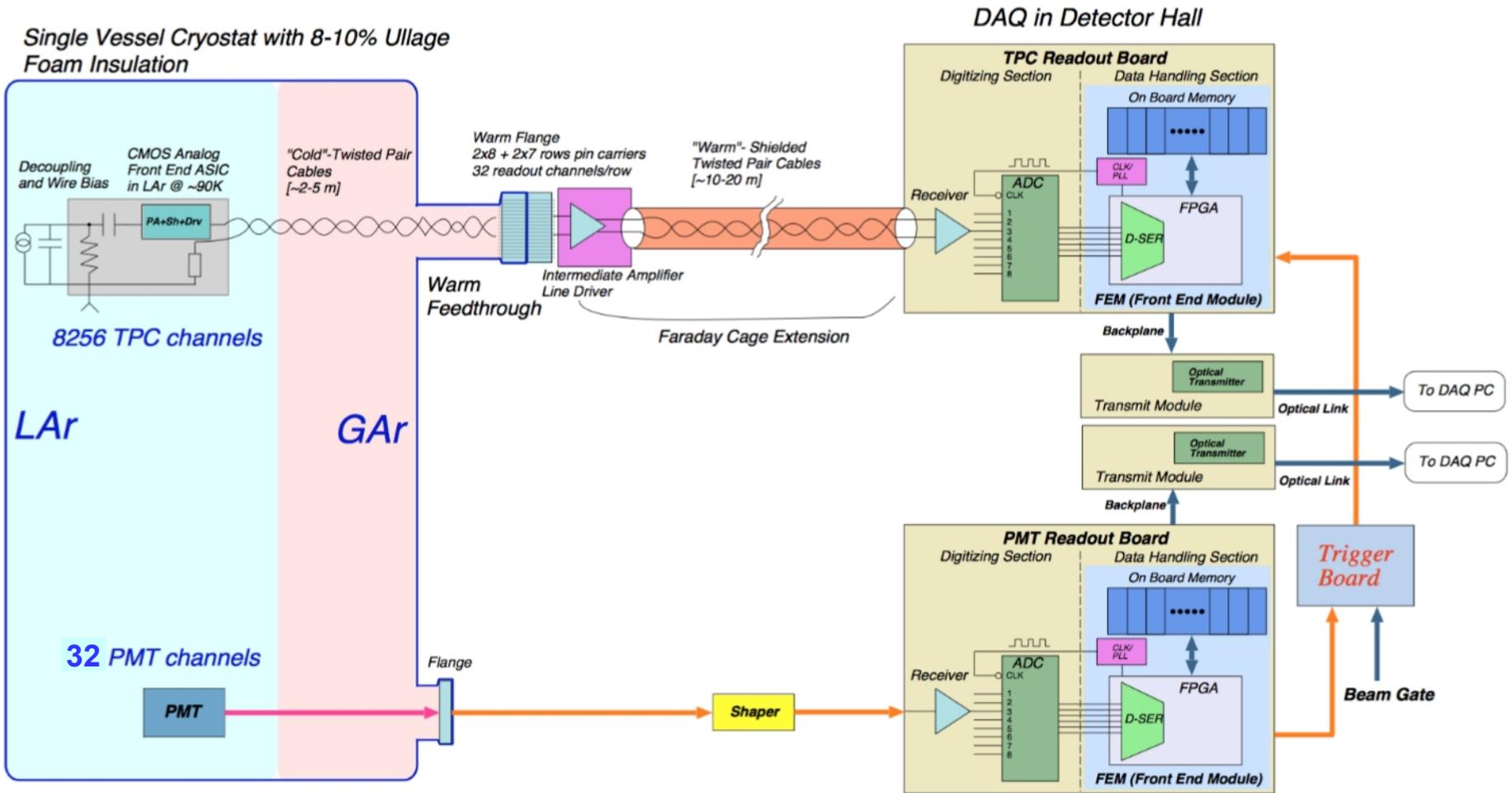
# Why Liquid Argon?

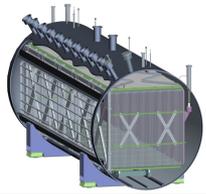
	He	Ne	Ar	Kr	Xe	Water
Boiling Point [K] @ 1atm	4.2	27.1	87.3	120	165	373
Density [g/cm <sup>3</sup> ]	0.125	1.2	1.4	2.4	3	1
Radiation Length [cm]	755.2	24	14	4.9	2.8	36.1
dE/dx [MeV/cm]	0.24	1.4	2.1	3	3.8	1.9
Scintillation [ $\gamma$ /MeV]	19,000	30,000	40,000	25,000	42,000	
Scintillation $\lambda$ [nm]	80	78	128	150	175	
Approx. Cost [\$/kg]	52	330	5	330	1200	

- ◆ Argon is cheap: ~1% of atmosphere
- ◆ Dense target (more  $\nu$ -N interactions per unit time)
- ◆ High scintillation light yield, argon transparent to own light
- ◆ Relatively small radiation length for EM shower containment



# Electronics Chain

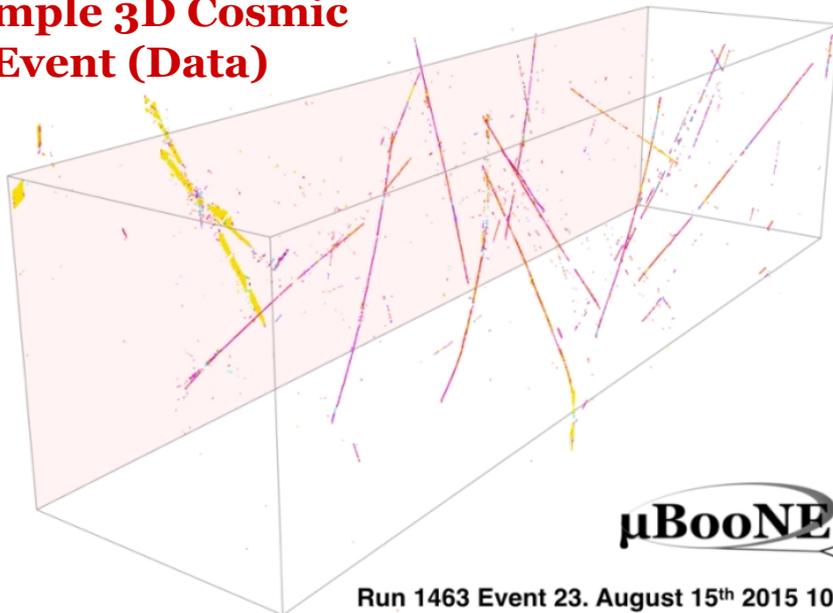




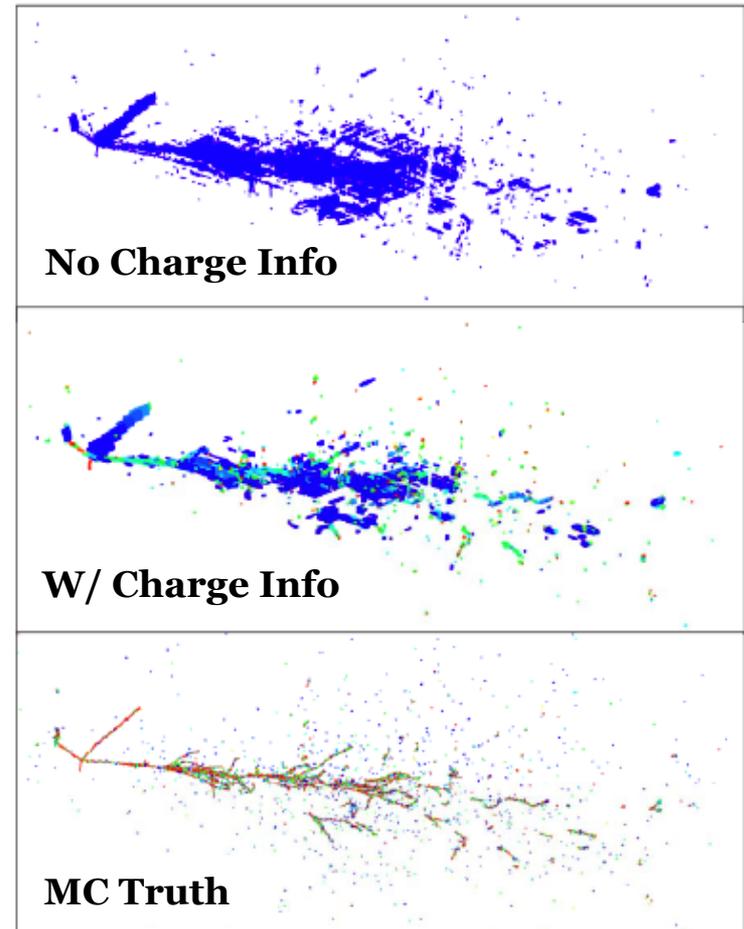
# Event Reconstruction

- ◆ Multiple ways to get to 3D:
  - Identify clustered tracks/showers in 2D, match across planes
  - Create 3D hits from wire triplets (matching charge) and directly cluster tracks/showers
    - “Wire-Cell” method (see images)

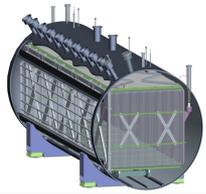
## Example 3D Cosmic Event (Data)



## Example MC Interaction Event (2D Projection of 3D)

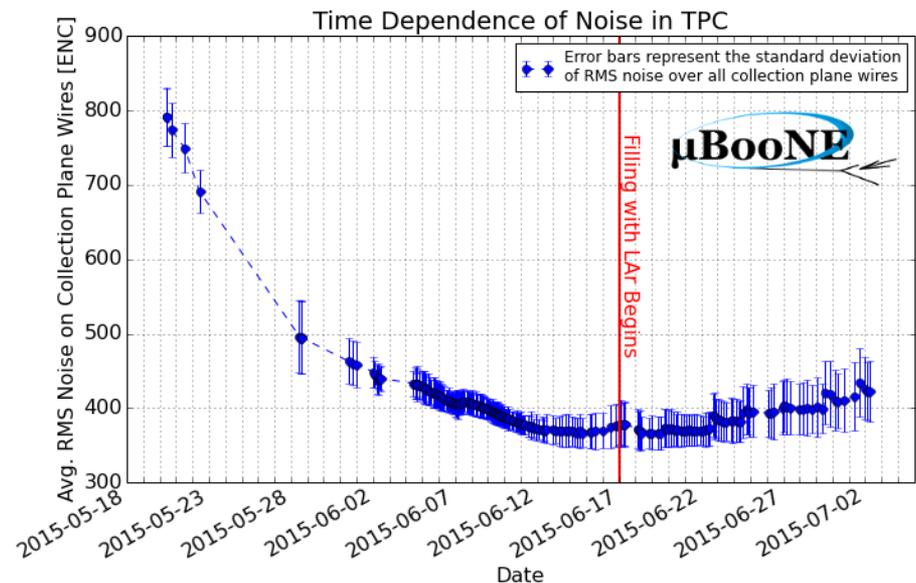
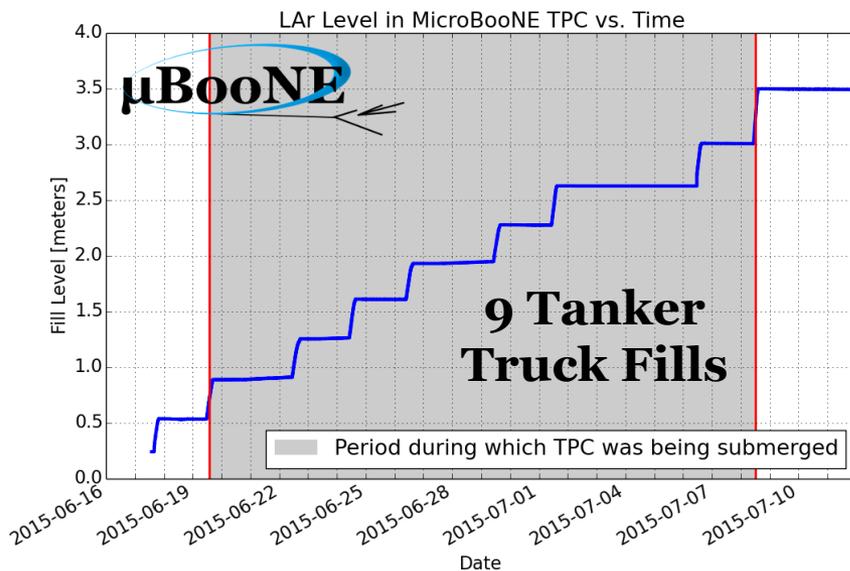


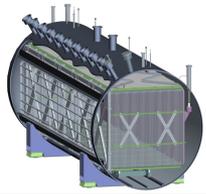
<http://www.phy.bnl.gov/wire-cell/bee/>



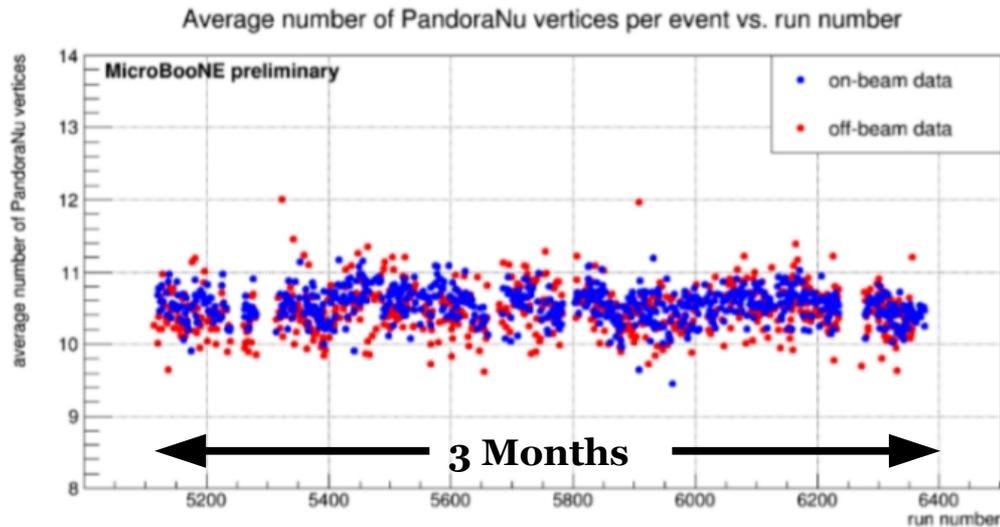
# Looking at Noise Data

- ◆ First look at **noise data**: during TPC and cryogenics commissioning (April-July 2015)
- ◆ TPC noise level dropped during purge/cool-down
  - Expected (desired) feature of cold electronics; noise level as expected from design
- ◆ TPC noise level slowly rose with LAr fill
  - Increased capacitance w/ LAr



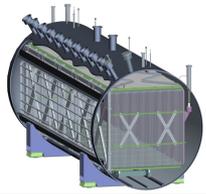


# Detector Stability

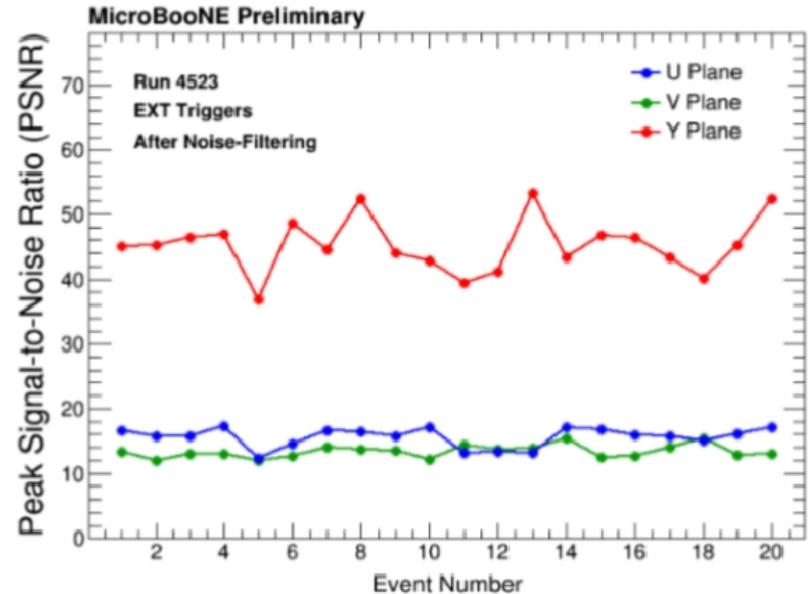
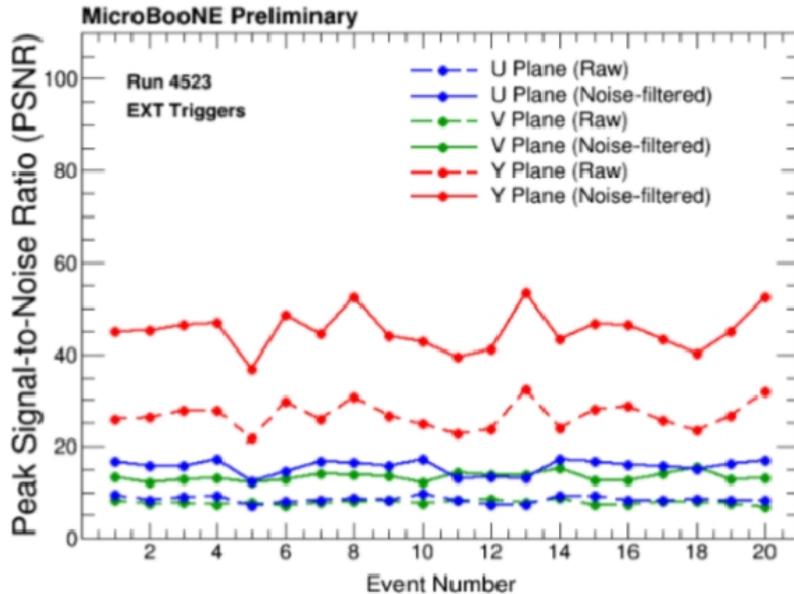


Issue	# Channels Affected
Unresponsive ASICs	~300
Shorted Wires	~400
Noisy Channels	~50
Uninstrumented	~100

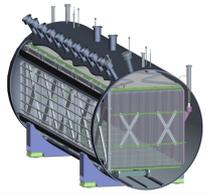
- ◆ High detector uptime – only handful of cathode, pump trips in first year of operations
  - Gaining operational experience with large LArTPCs – essential for running **future LArTPC experiments**
- ◆ Both high-level and low-level features in data stable over time
- ◆ Number of unresponsive/noisy channels very stable w.r.t. time
  - 10% unresponsive/noisy, but **97%** of detector volume has **at least two planes operational** (minimum needed for 3D reco.)



# Peak Signal-to-Noise Ratio

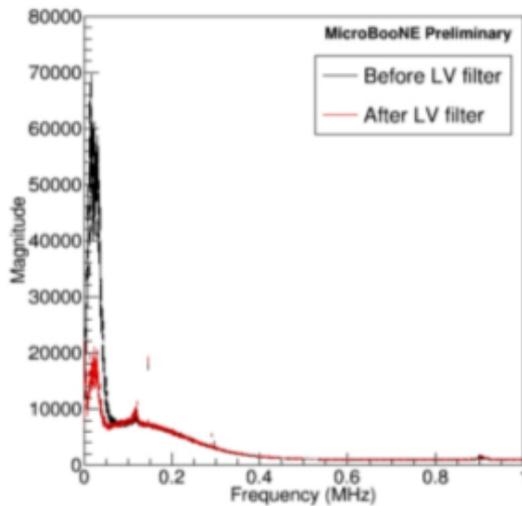


- ◆ Peak Signal-to-Noise Ratio (PSNR – signal height divided by noise RMS) very high after software noise filtering
  - Note: here calculated for all signals in event (not just MIPs)
  - Collection plane: **PSNR > 40**
  - Induction planes: **PSNR > 12** (note bipolar nature of signal)
- ◆ Higher PSNR post-filtering → charge resolution improves

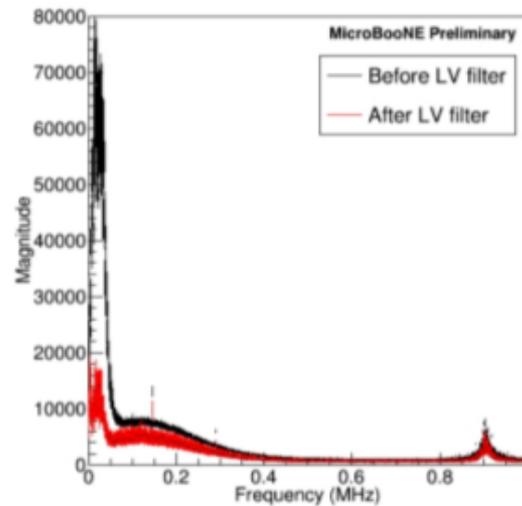


# Low-Frequency Noise

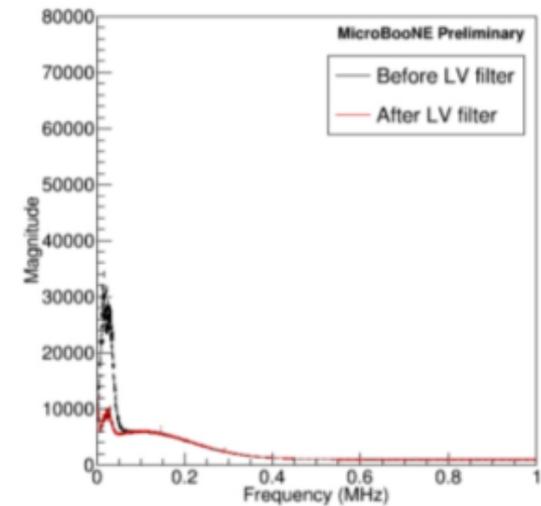
## First Induction (U)



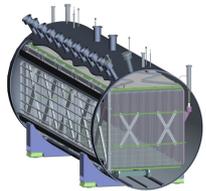
## Second Induction (V)



## Collection (Y)

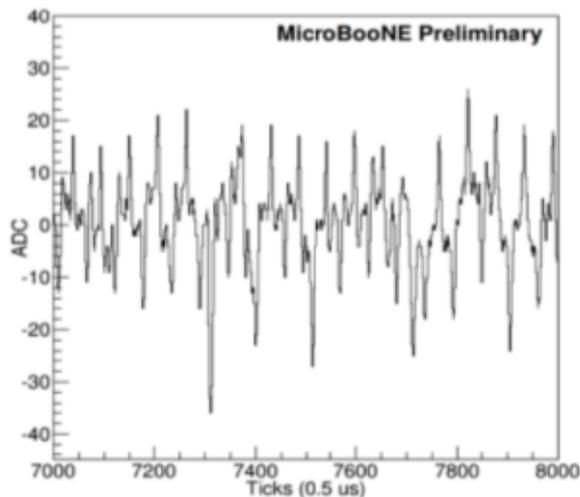


- ◆ Majority of noise present before filtering is due to low-frequency (10-30 kHz) noise coherent across all channels on a cold motherboard pair – thought to be associated with voltage regulators
- ◆ Almost completely filtered out with software algorithm that takes advantage of coherent nature of noise
- ◆ Hardware fix: upgrade service boards (this summer)

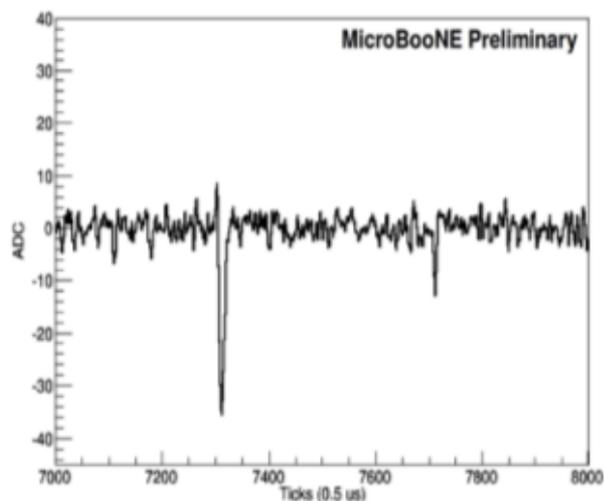


# Other TPC Noise Features

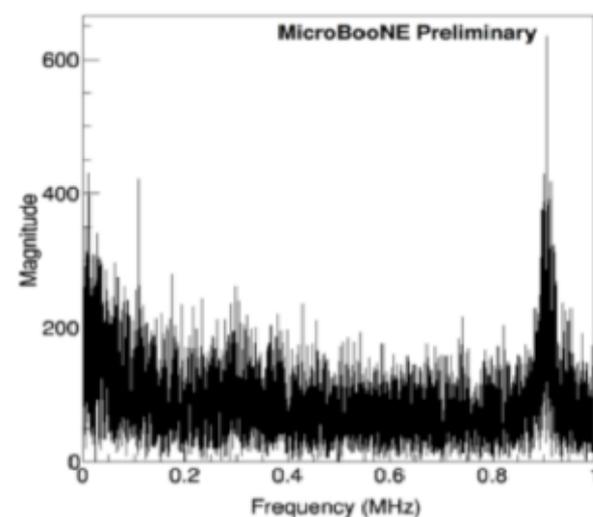
Zoomed Raw Signal



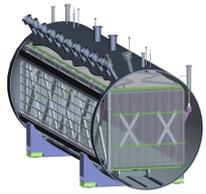
Signal After Full Noise Filtering



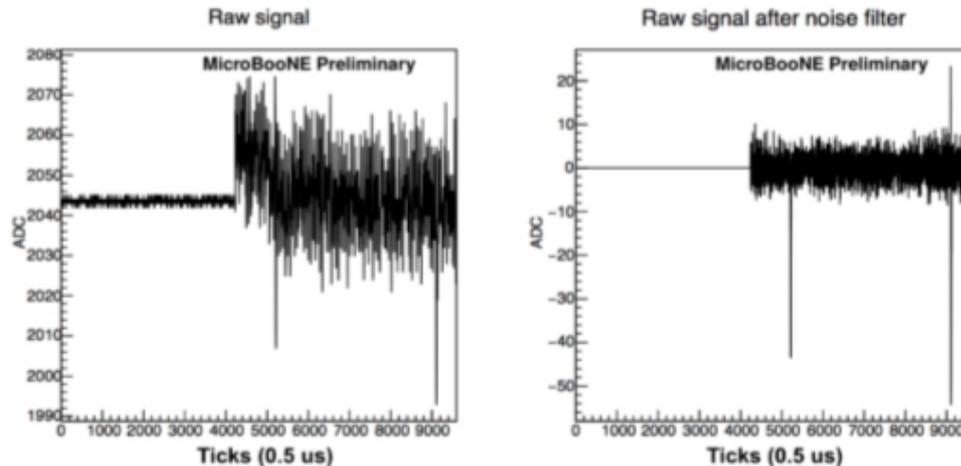
FFT of Raw waveform



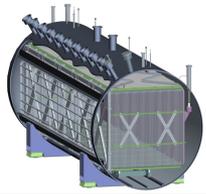
- ◆ Another major noise source primarily on first induction plane: narrow-band noise associated with cathode HV power supply
  - Can filter out easily in software (compare left to middle)
  - Hardware fix: install second filter pot for cathode HV (this summer)
- ◆ High-frequency pick-up noise on downstream side of TPC (right)
  - Suppressed by higher shaping time, easily filtered out with low-pass filter



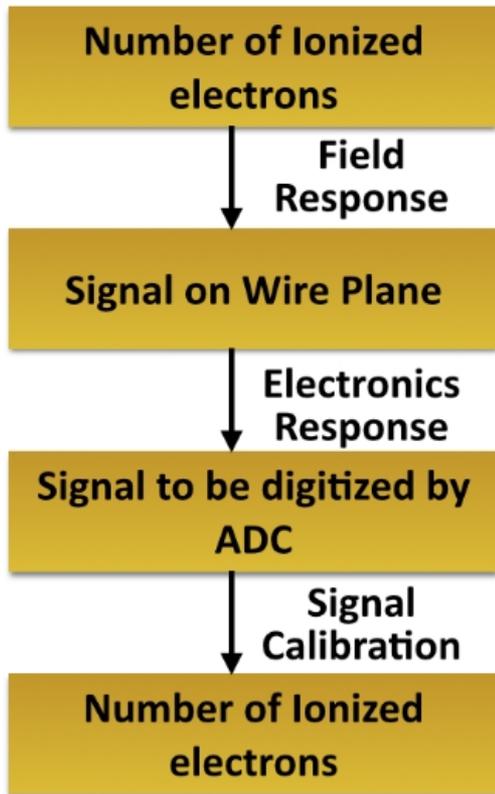
# Effects of ASIC Saturation



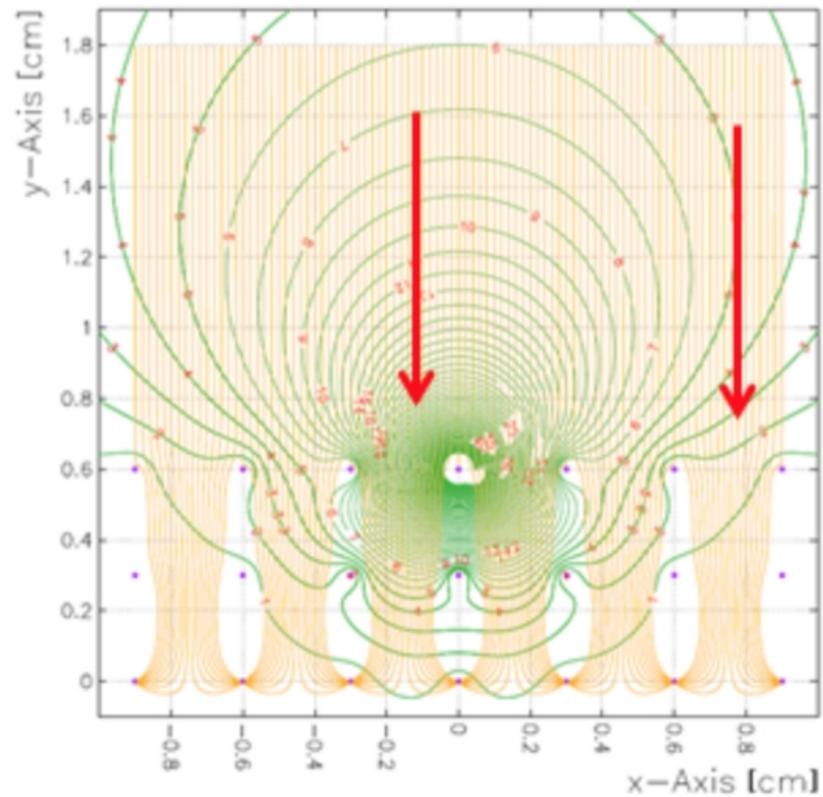
- ◆ Occasionally ASICs found to “saturate” leading to dead regions of TPC waveforms
  - Charge builds up too fast on capacitor in ASIC circuit
  - Current source believed to be from vibrating wires – worse for longer wires
- ◆ Solution is to use higher bias current (“leakage current”) setting in ASIC – occurrence small in MicroBooNE, but accounted for in software noise filtering step
  - New ASIC design includes higher leakage current settings

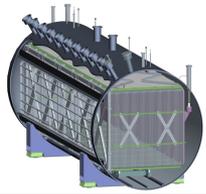


# Signal Processing Chain



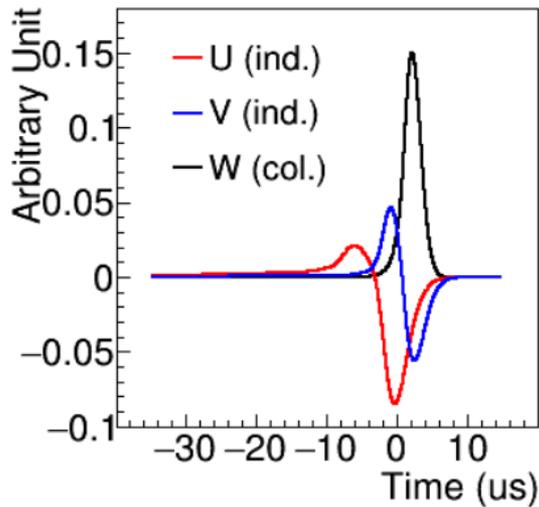
Weighting potential of a wire



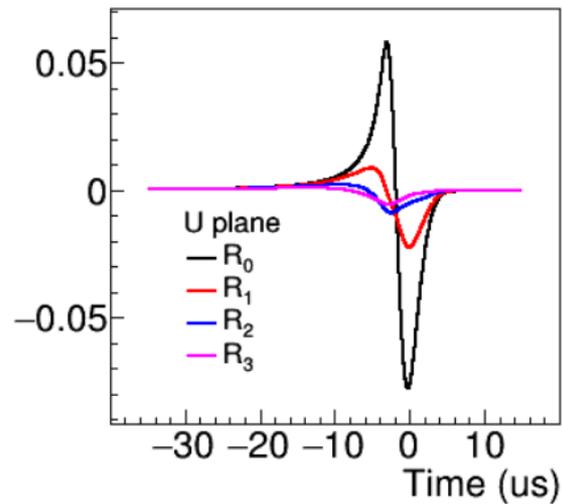


# Response Functions

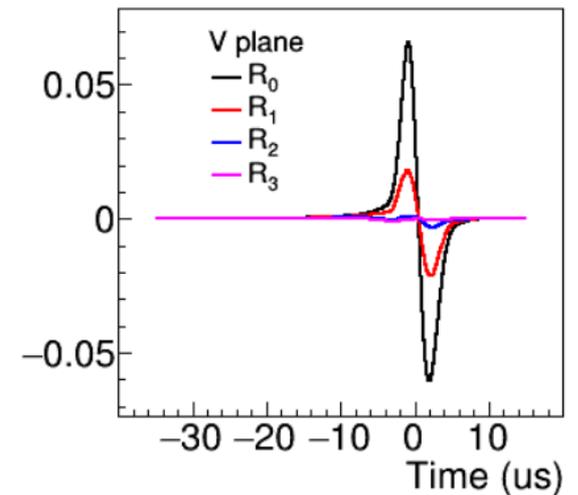
1D Data-inspired Responses

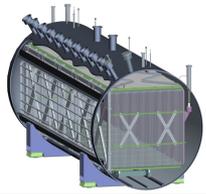


2D Garfield Responses

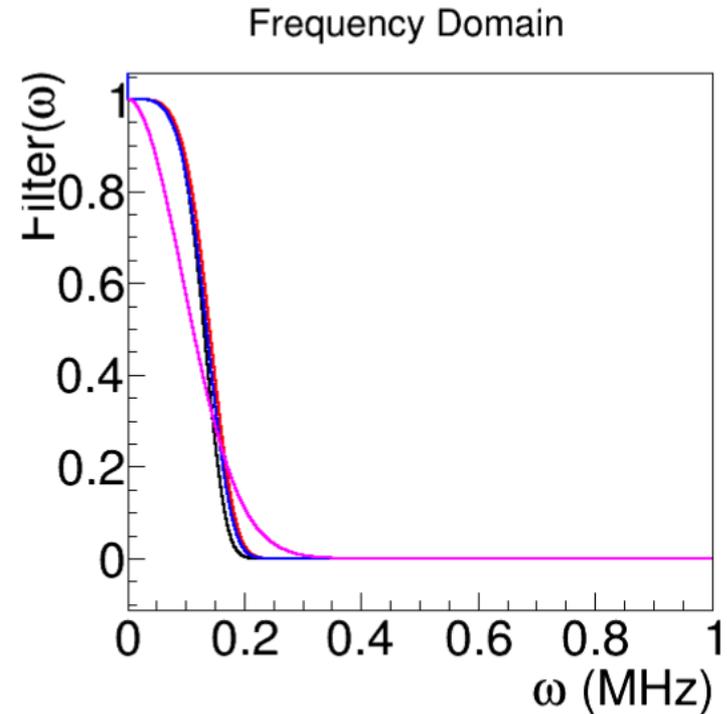
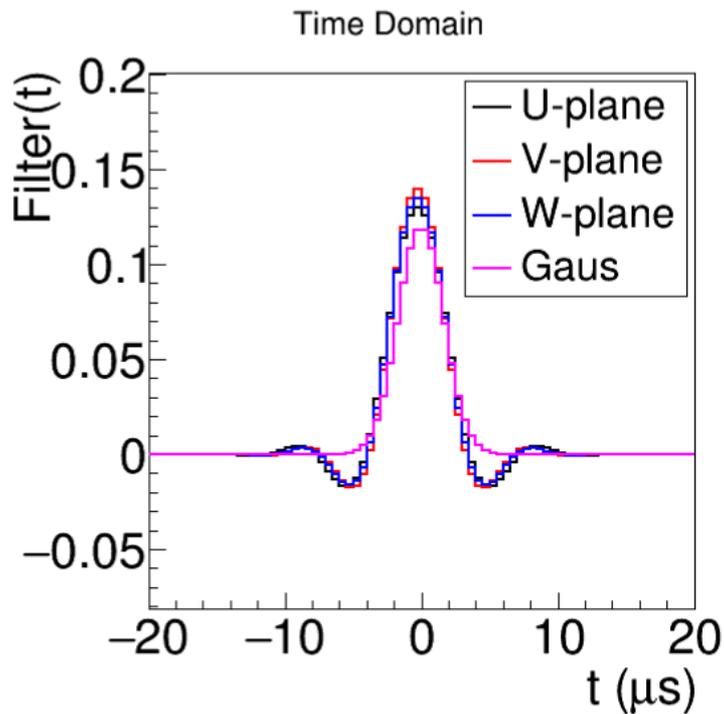


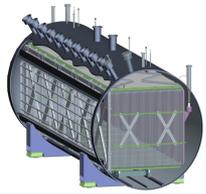
2D Garfield Responses



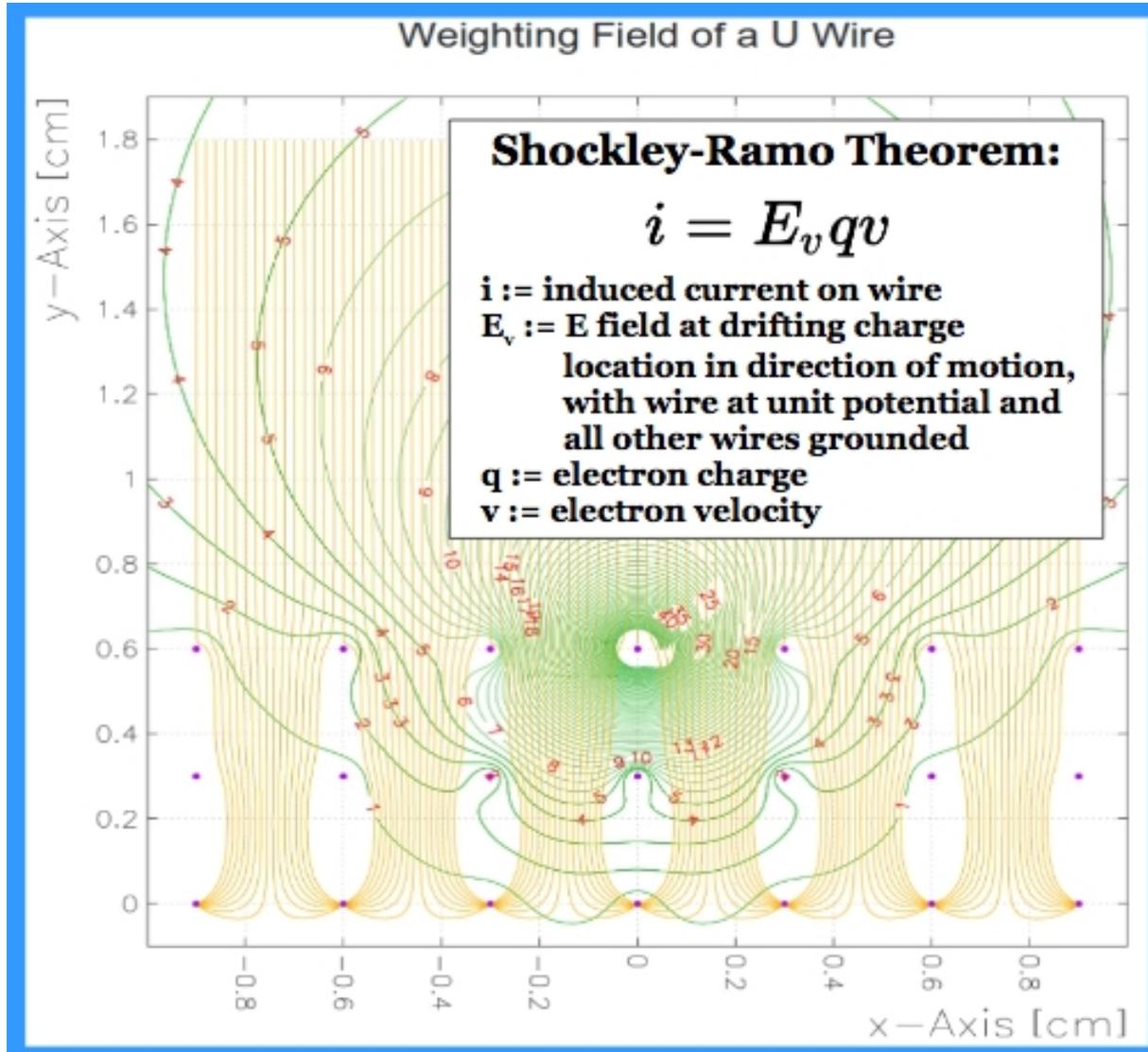


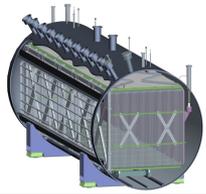
# Deconvolution Filter





# Shockley-Ramo Theorem





# Deconvolution Procedure

The deconvolution is employed to estimate the true ionization signal  $S$  from the measured signal  $M$  on the raw waveform, with

$$R(t, t_0) \equiv R(t - t_0),$$

$$M(t_0) = \int_{-\infty}^{\infty} R(t, t_0) \cdot S(t) \cdot dt.$$

This process is done in the frequency domain and utilizes a known full response  $R$  (field + electronics) and a filter  $F$ :

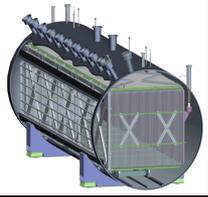
$$F(\omega) = \begin{cases} e^{-\frac{1}{2} \cdot (\frac{\omega}{\omega_0})^2} & \omega > 0 \\ 0 & \omega = 0, \end{cases}$$

$$S(\omega) = \frac{M(\omega)}{R(\omega)} \cdot F(\omega).$$

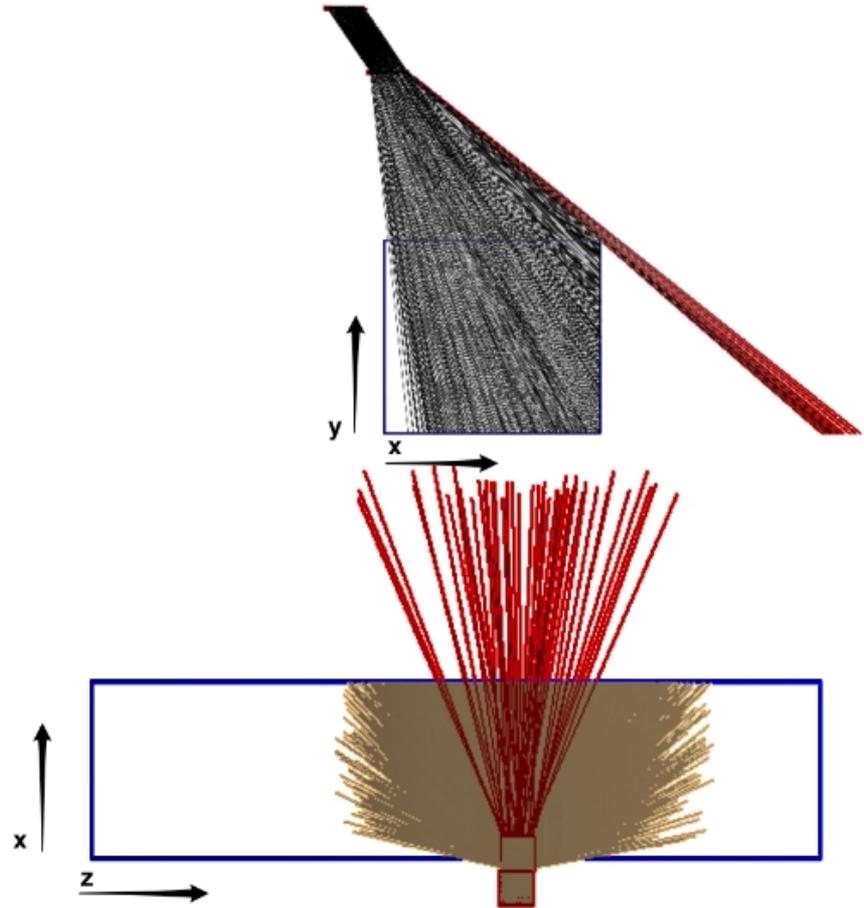
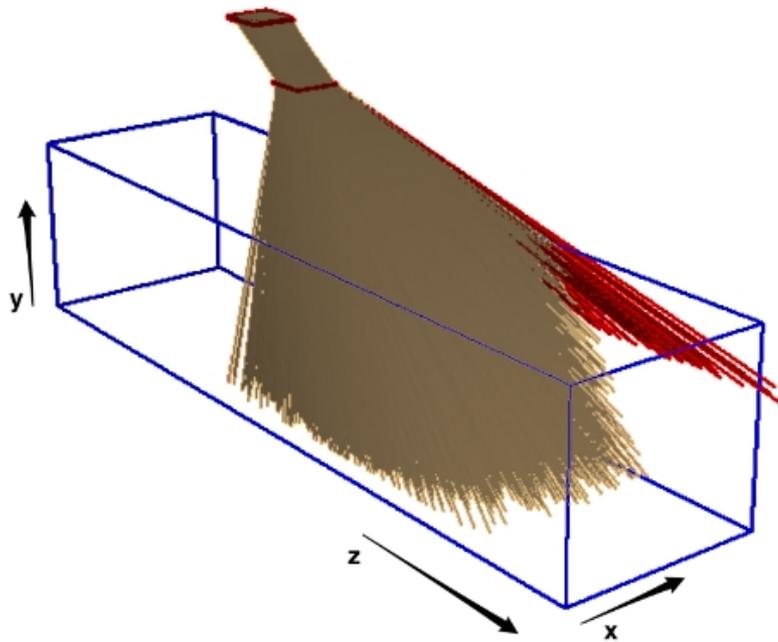
A 2D (time vs. wire) deconvolution is done for the U/V planes in order to account for different responses from nearby wires:

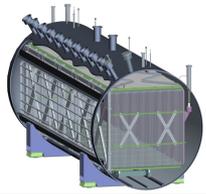
$$\begin{pmatrix} M_1(\omega) \\ M_2(\omega) \\ \vdots \\ M_{n-1}(\omega) \\ M_n(\omega) \end{pmatrix} = \begin{pmatrix} R_0(\omega) & R_1(\omega) & \dots & R_{n-2}(\omega) & R_{n-1}(\omega) \\ R_1(\omega) & R_0(\omega) & \dots & R_{n-3}(\omega) & R_{n-2}(\omega) \\ \vdots & \vdots & \ddots & \vdots & \vdots \\ R_{n-2}(\omega) & R_{n-3}(\omega) & \dots & R_0(\omega) & R_1(\omega) \\ R_{n-1}(\omega) & R_{n-2}(\omega) & \dots & R_1(\omega) & R_0(\omega) \end{pmatrix} \cdot \begin{pmatrix} S_1(\omega) \\ S_2(\omega) \\ \vdots \\ S_{n-1}(\omega) \\ S_n(\omega) \end{pmatrix}.$$

The 2D version ( $R$  matrix inversion) recovers reconstructed tracks at high angles with respect to the anode plane.

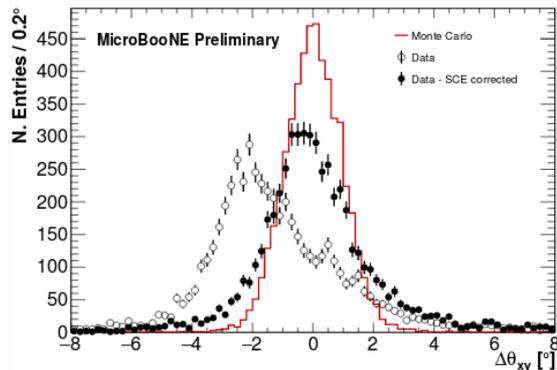
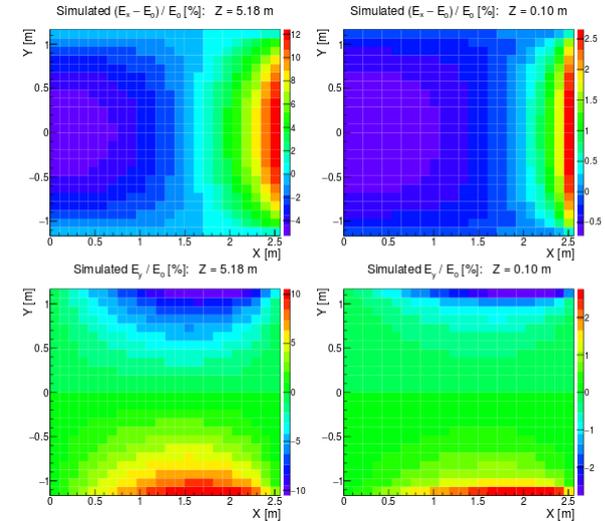
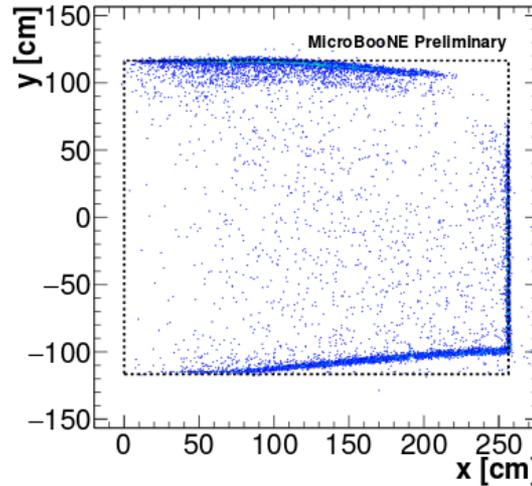
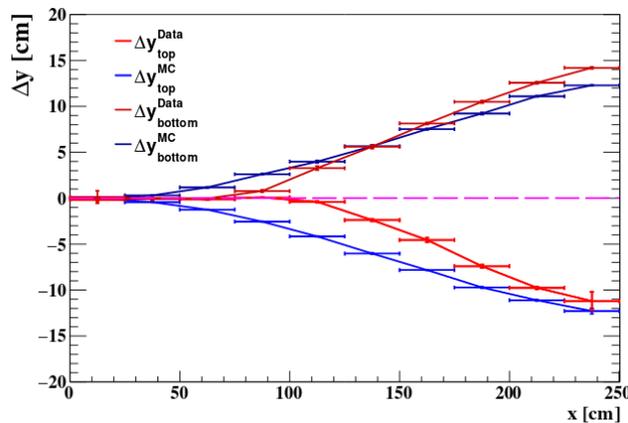
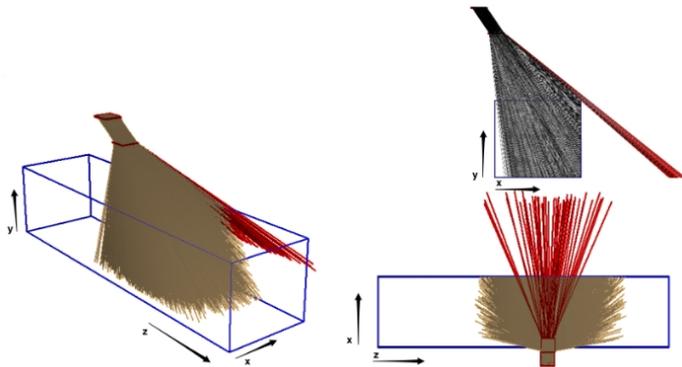


# Muon Counter System

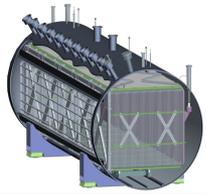




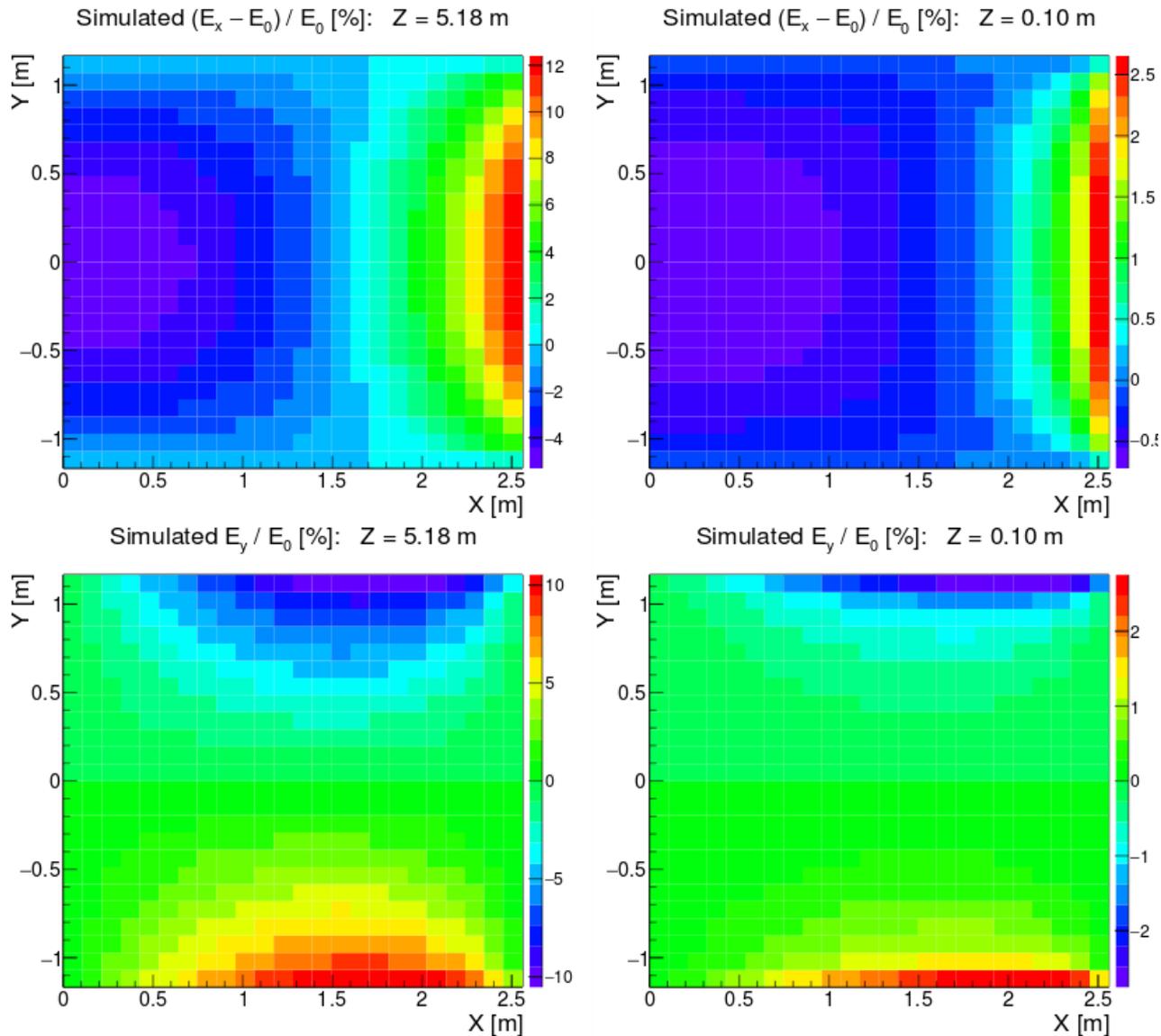
# SCE Analysis Overview

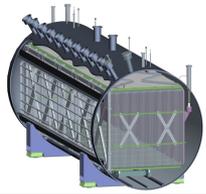


- ◆ Utilize MuCS for track  $t_0$  tags
  - Probe features in drift direction
- ◆ See strange feature in data
- ◆ Seems like SCE feature, so take a look at SCE simulation
- ◆ Data vs. MC: similar, but differences
- ◆ Attempt partial correction – can reduce impact of effect in data



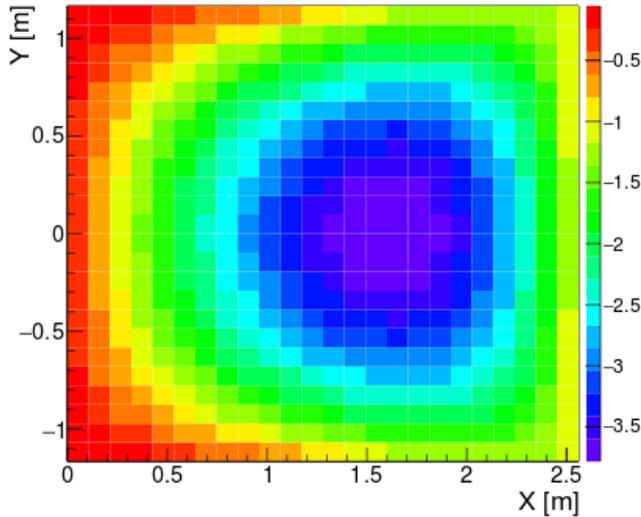
# SCE E Field Distortions



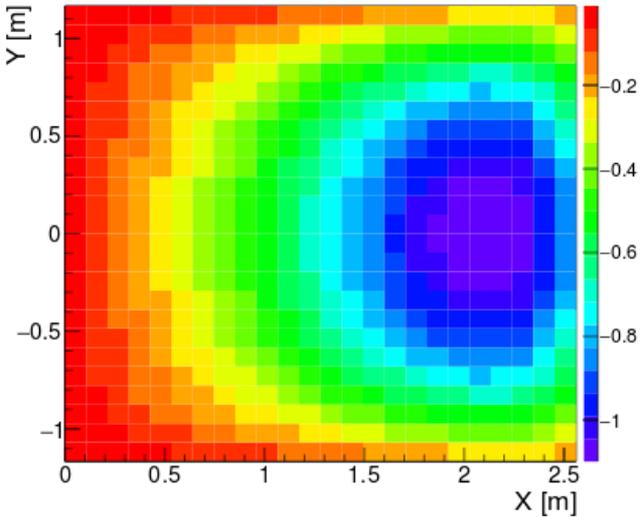


# SCE Spatial Distortions

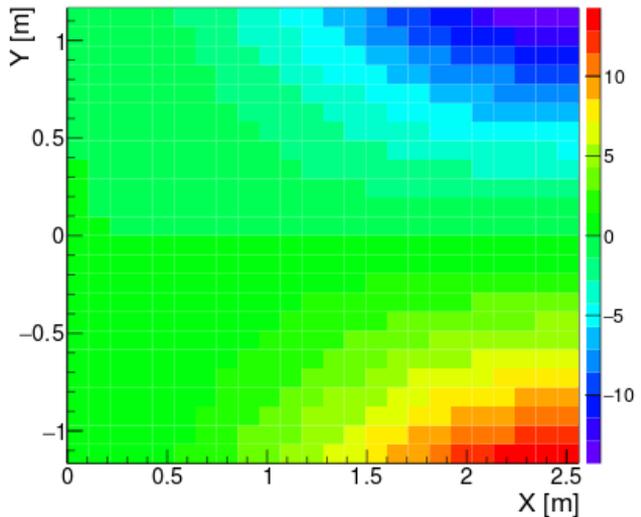
$X_{\text{reco}} - X_{\text{true}} [\text{cm}]$ :  $Z = 5.18 \text{ m}$



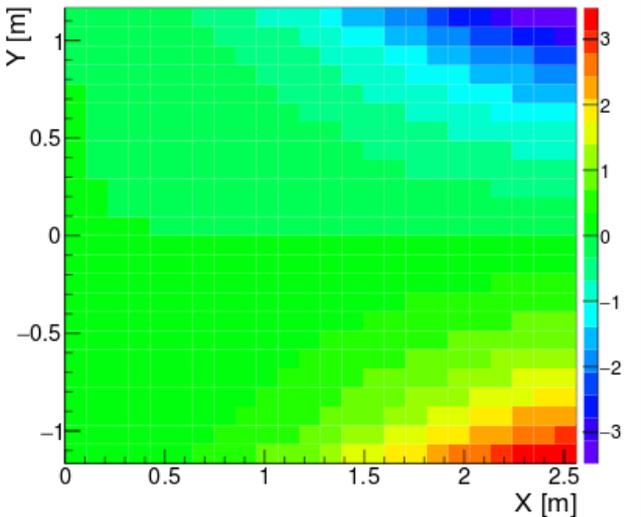
$X_{\text{reco}} - X_{\text{true}} [\text{cm}]$ :  $Z = 0.10 \text{ m}$

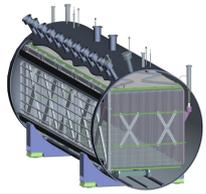


$Y_{\text{reco}} - Y_{\text{true}} [\text{cm}]$ :  $Z = 5.18 \text{ m}$



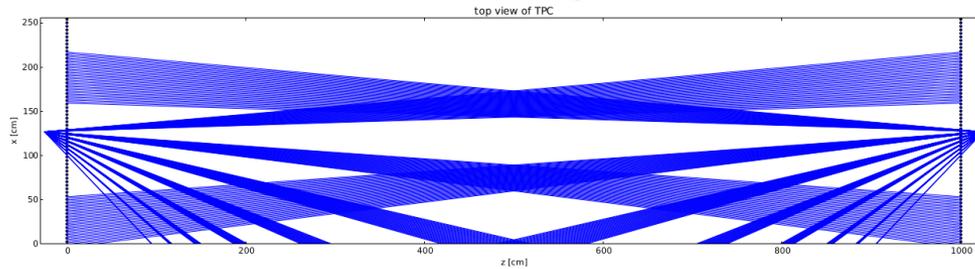
$Y_{\text{reco}} - Y_{\text{true}} [\text{cm}]$ :  $Z = 0.10 \text{ m}$



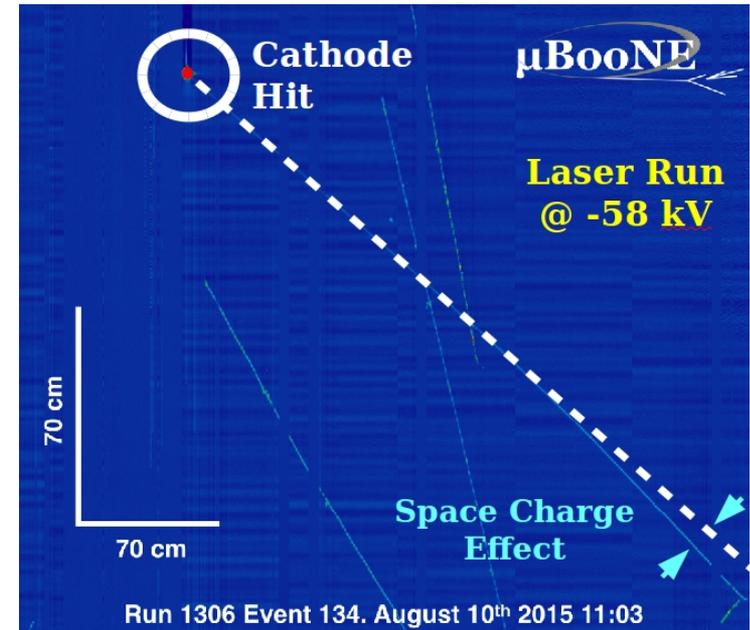
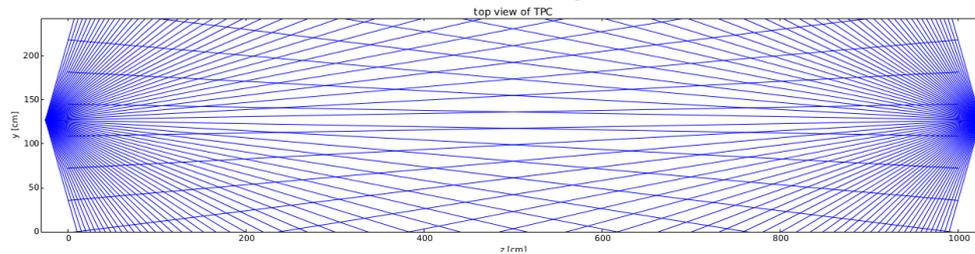


# SCE Laser Calibration

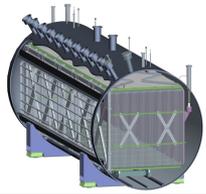
Simulated Laser Coverage: **X-Z Plane**



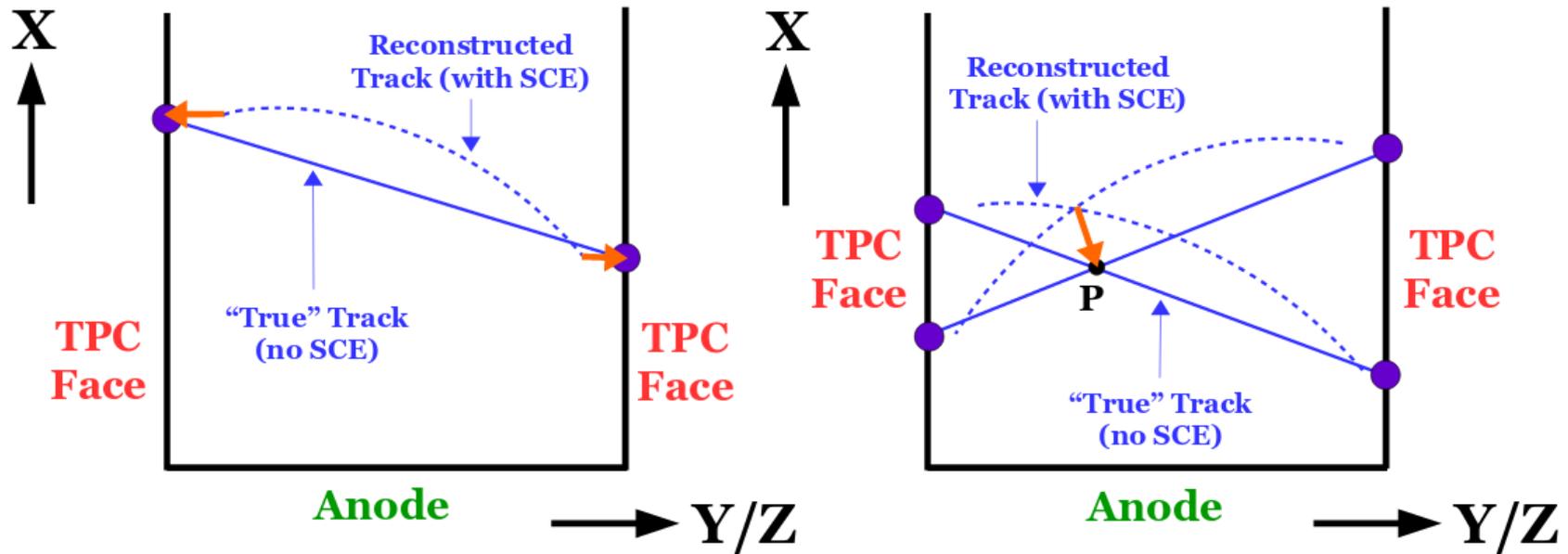
Simulated Laser Coverage: **Y-Z Plane**



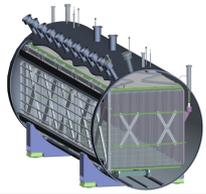
- ◆ Qualitatively, SCE very clear in laser event displays
- ◆ Can make point-to-point SCE correction throughout TPC using crossing point of two laser tracks



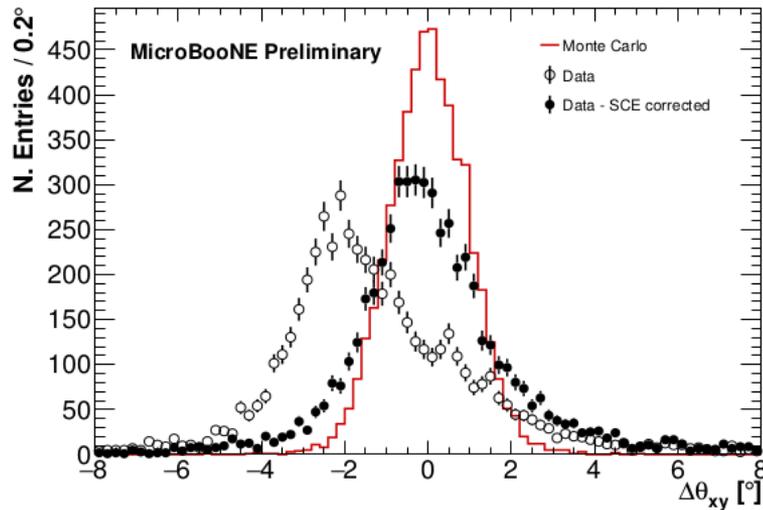
# SCE Cosmic Muon Calibration



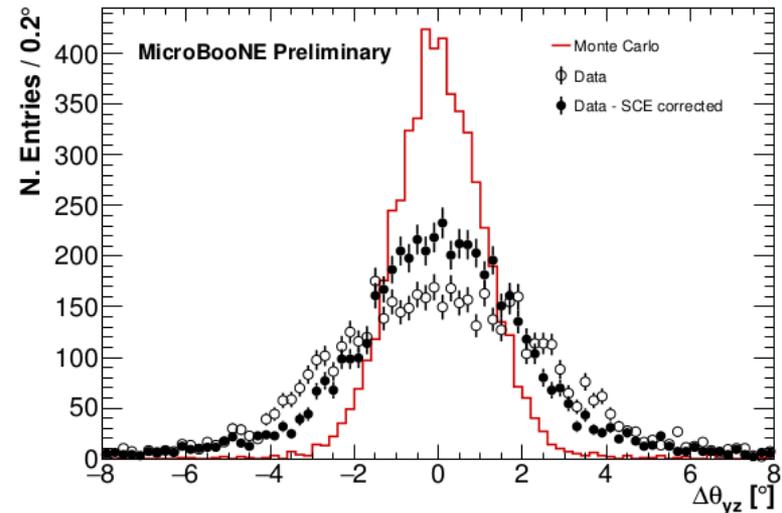
- ◆ Uses  $t_0$ -tagged tracks: anode-cathode crossing tracks, anode/cathode-piercing tracks and MuCS-tagged tracks
- ◆ Calibrate points in TPC using single tracks (TPC faces) and pairs of tracks (TPC bulk) – utilize  $\sim$ straight tracks using MCS measurement (high momentum  $\rightarrow$   $\sim$ straight)



# SCE Corr. Validation

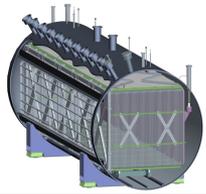


(a)  $\Delta\theta_{xy}$  residual distribution

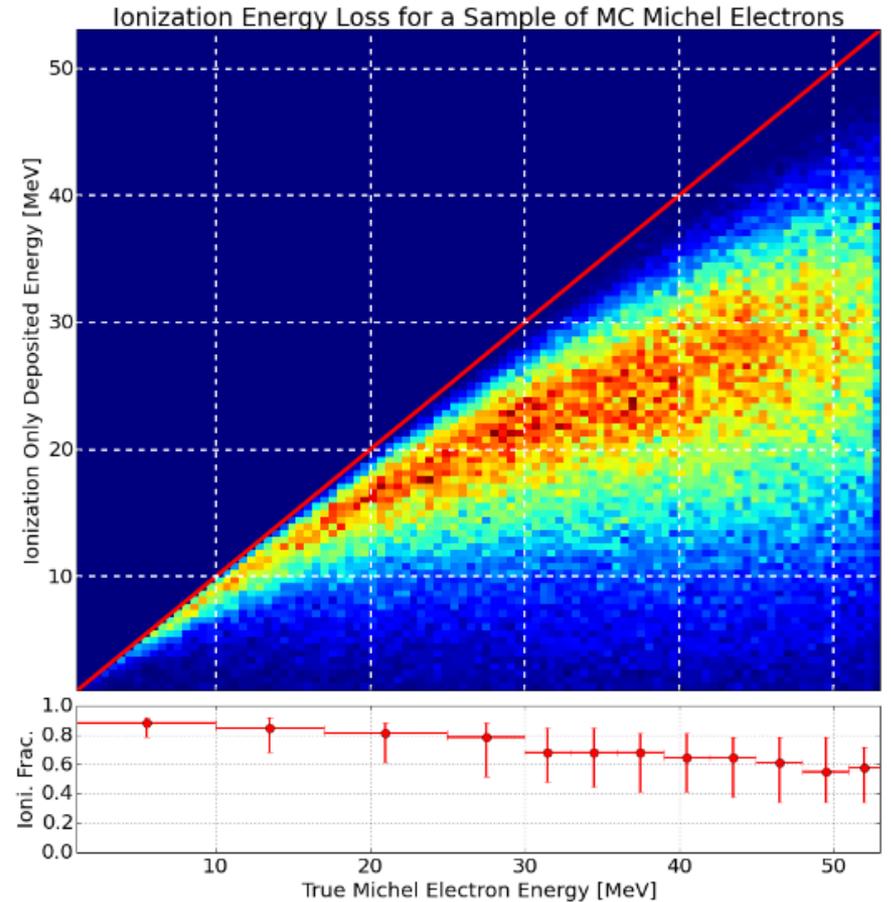
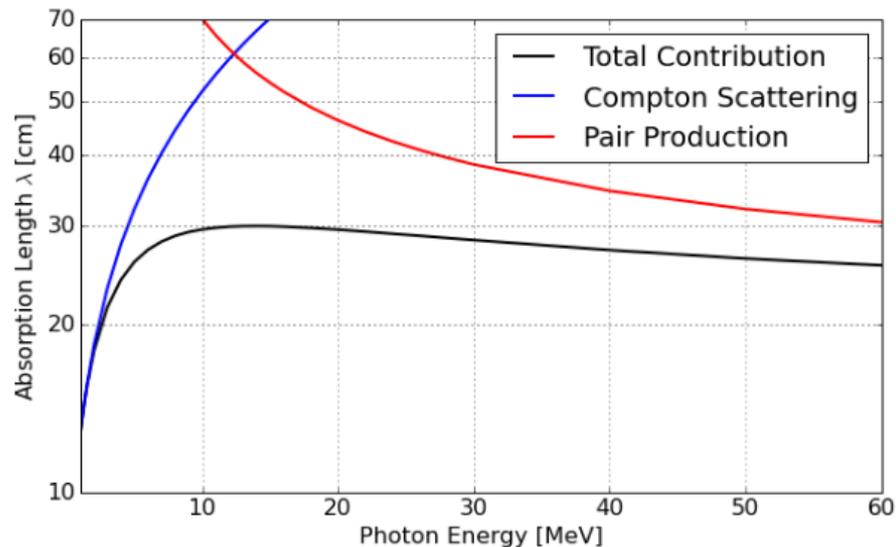
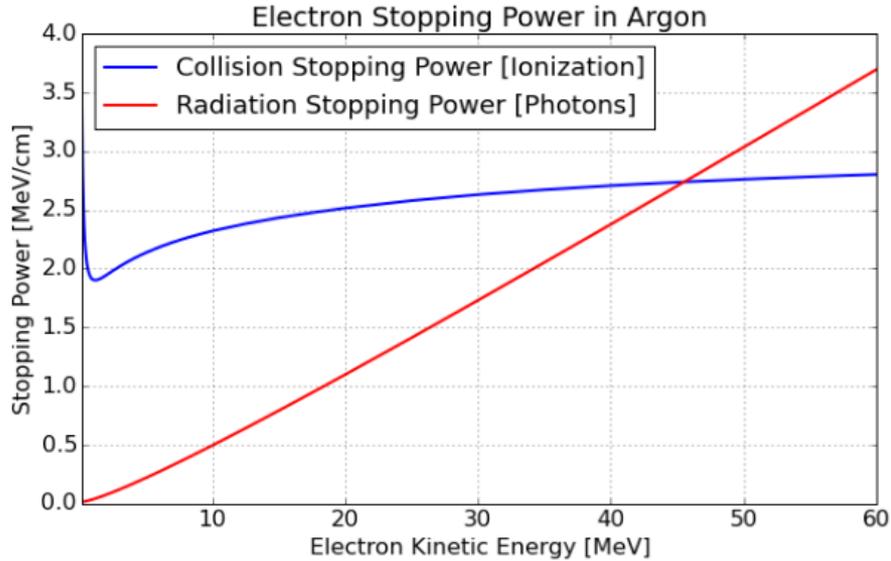


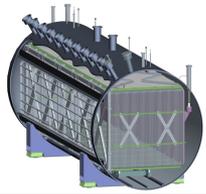
(b)  $\Delta\theta_{yz}$  residual distribution

- ◆ Validate SCE calibration using separate sample of  $t_0$ -tagged tracks
  - Look at track angles, track hit density, etc.
  - Also characterize time-dependence of effect – **important!**
- ◆ MicroBooNE SCE public note gives example of this type of validation using MuCS-tagged tracks (angular residuals)



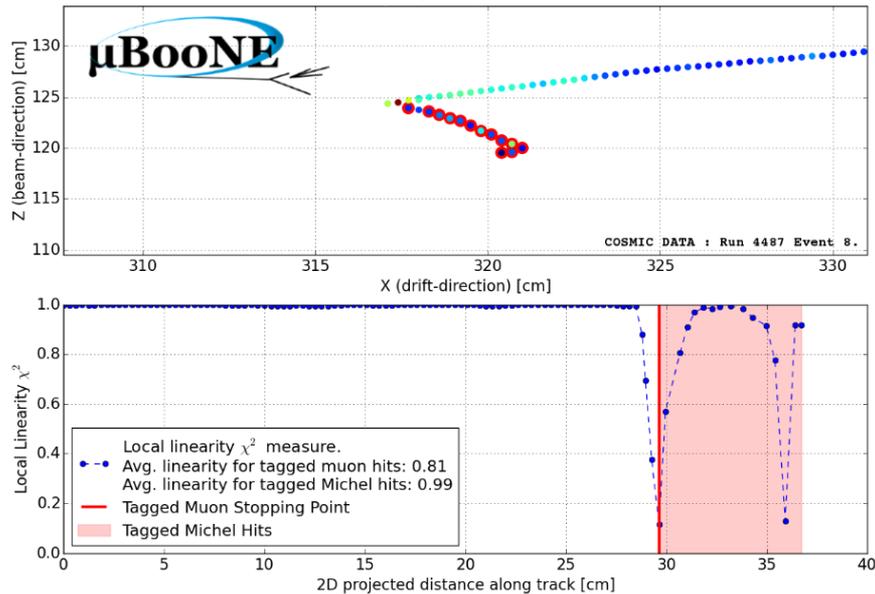
# Michel $e^-$ Energy Loss



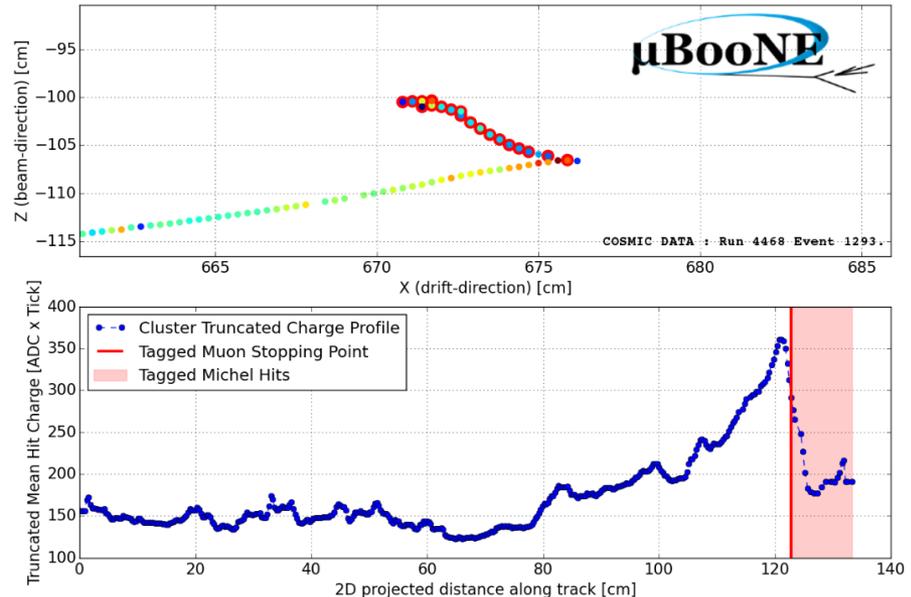


# Michel $e^-$ Reconstruction

## Topology (“Kink”)

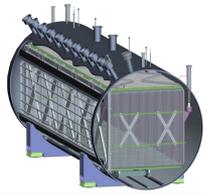


## Calorimetry (Bragg Peak)

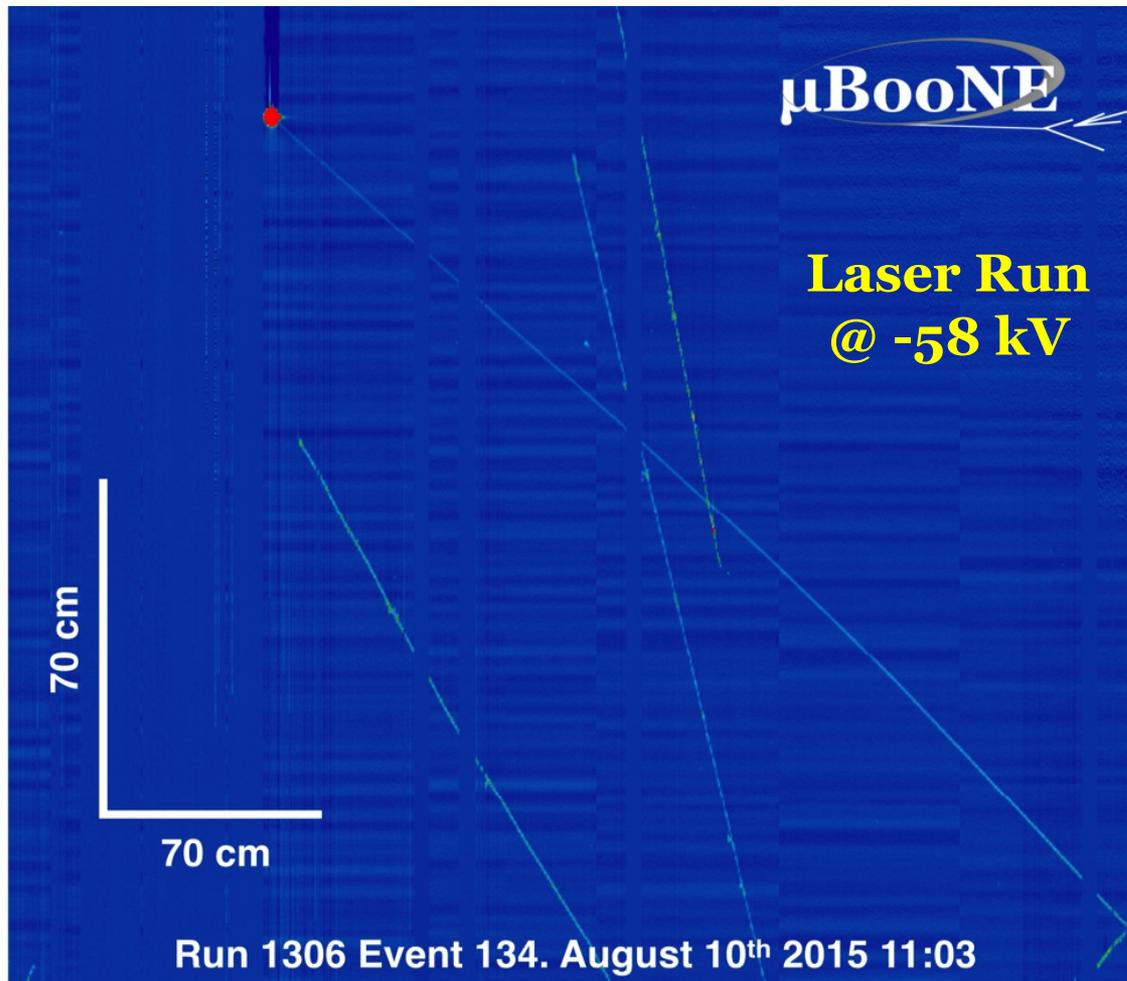


### ◆ Tag Michel electrons from cosmic muon decay using characteristic topology and calorimetric information

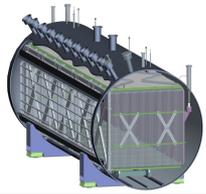
- 2D reconstruction for now (collection plane only)
- Yields a high purity (80-90%) and low efficiency (2-3%) sample of Michel electrons



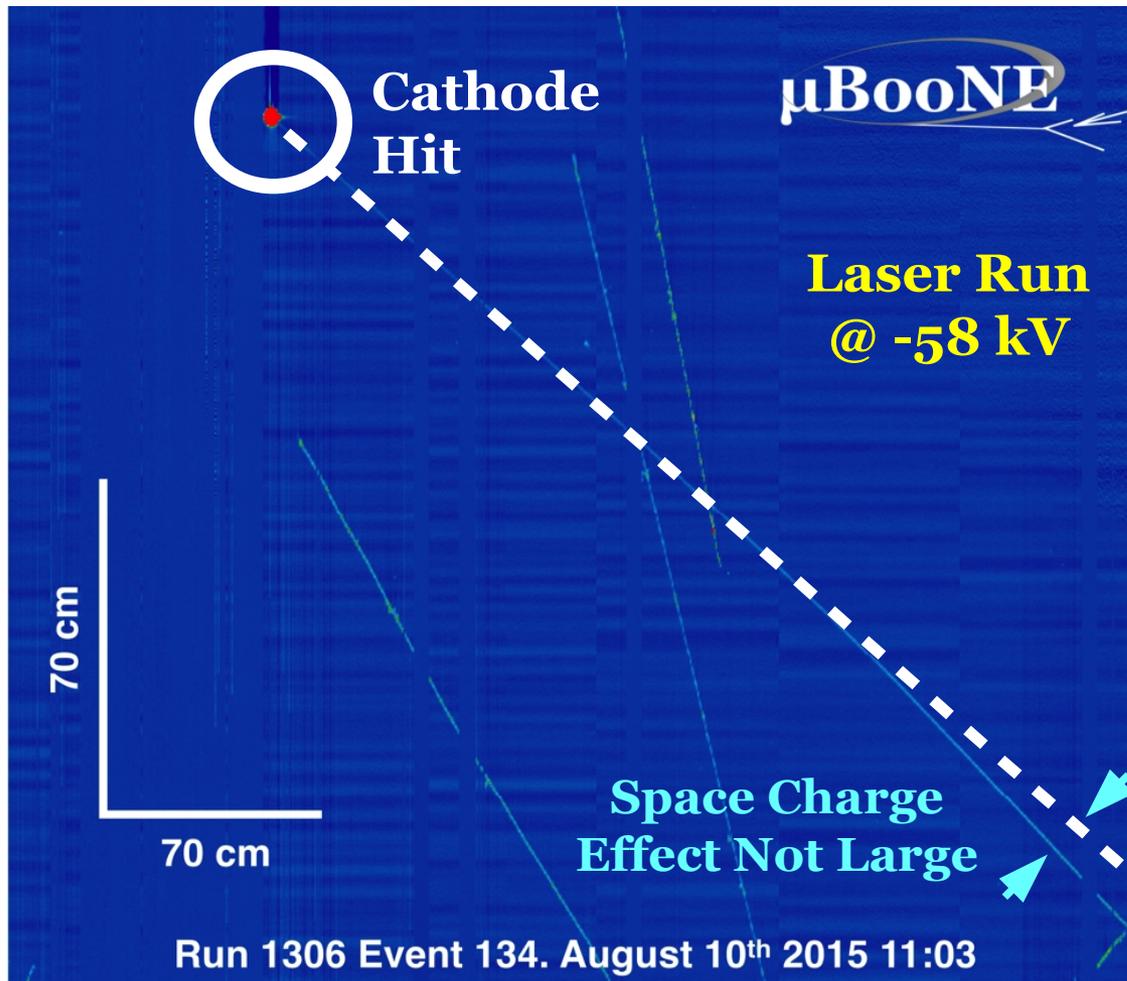
# Special Runs



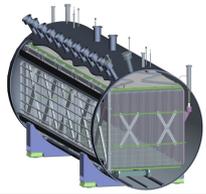
- ◆ Also have taken special runs for calibrations and detector physics – laser, cosmic, special ASIC settings, etc.



# Special Runs

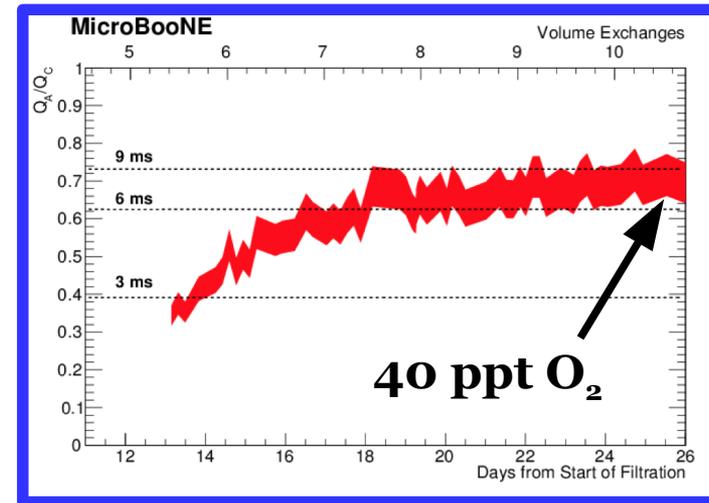
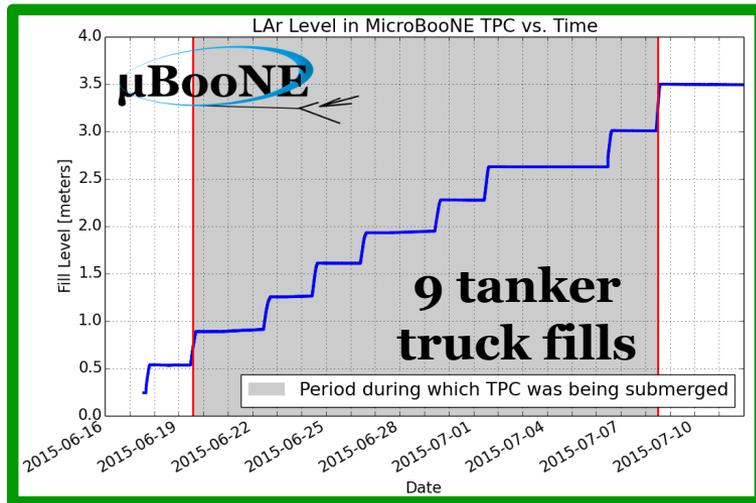
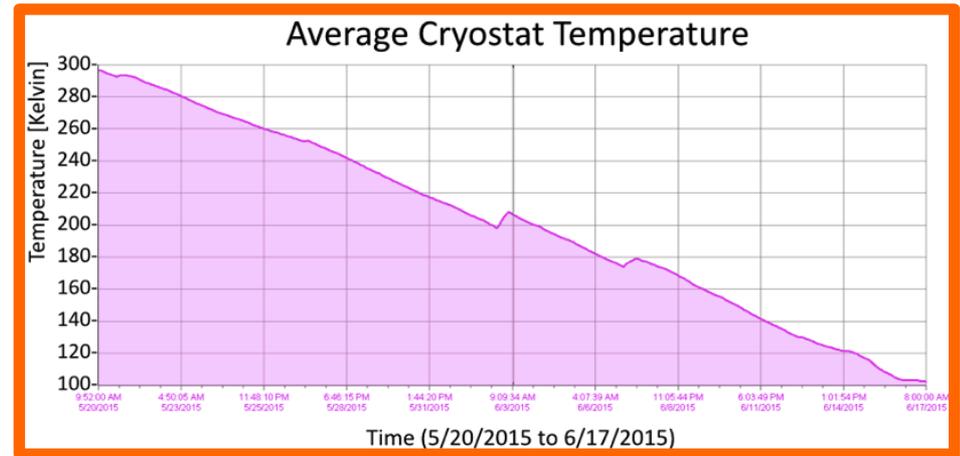
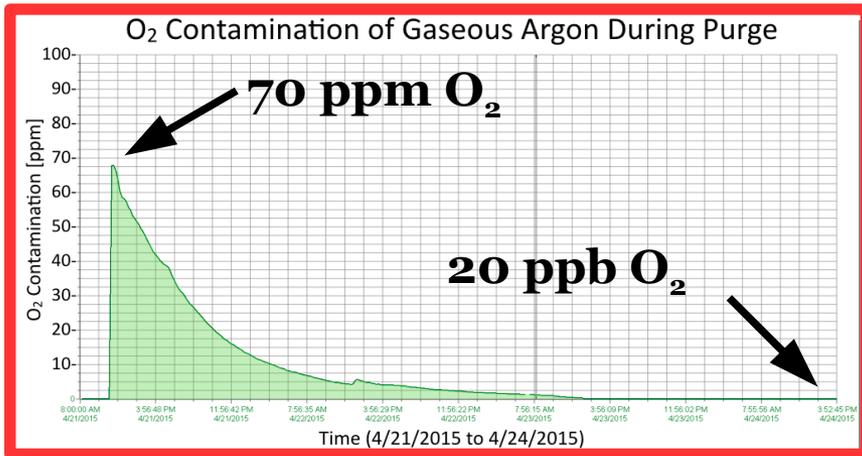


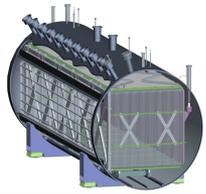
- ◆ Also have taken special runs for calibrations and detector physics – laser, cosmic, special ASIC settings, etc.



# Cryogenics Commissioning

- ◆ Four stages: (1) **purge** (2) **cool-down** (3) **LAr fill** (4) **recirculation and purification** → **operating at least 2-3 times design purity!**





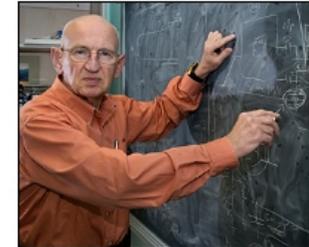
# LArTPC: Early History

## Early History of the Development of LArTPC

- W. Willis and V. Radeka, Liquid argon ionization chambers as total absorption detector, NIMA 120:221 (1974)
- D. R. Nygren, The Time Projection Chamber: A New  $4\pi$  Detector for Charged Particles. eConf. C740805:58 (1974)
- H. H. Chen et al. A Neutrino detector sensitive to rare process. I. A study of neutrino electron reactions. FNAL-Proposal-0496 (1976)
- C. Rubbia, The liquid argon time projection chamber: a new concept for neutrino detector, CERN-EP/77-08 (1977)



William Willis



V. Radeka



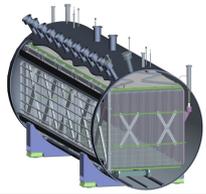
D. R. Nygren



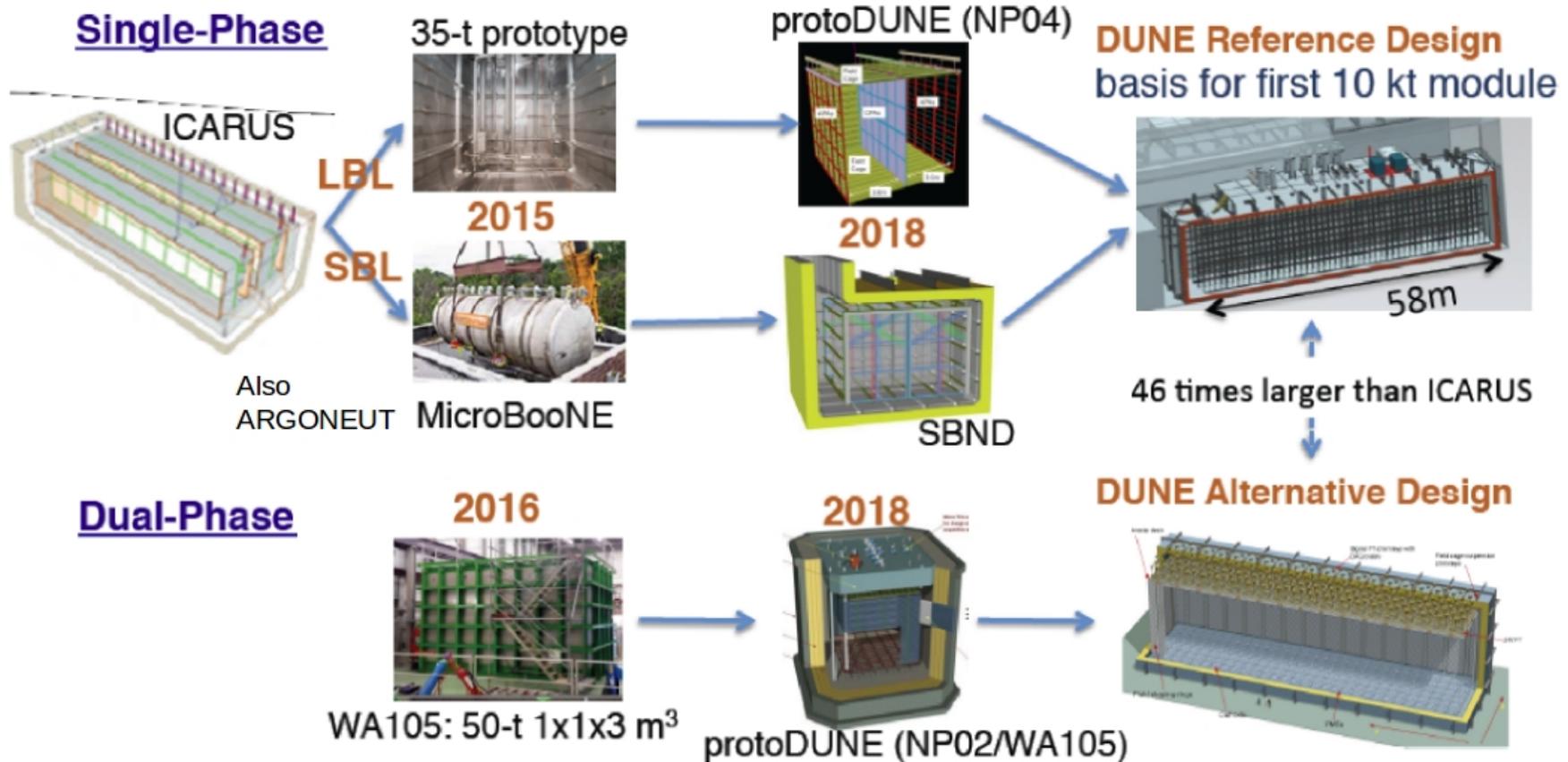
H. H. Chen

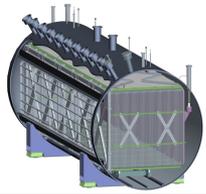


C. Rubbia



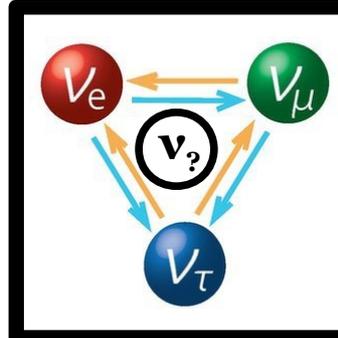
# LArTPC Experiments



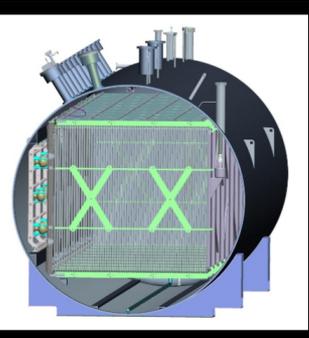
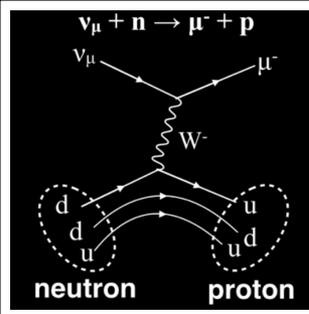


# MicroBooNE Physics Goals

**Oscillation Physics  
(Investigate MiniBooNE Low-Energy Excess)**



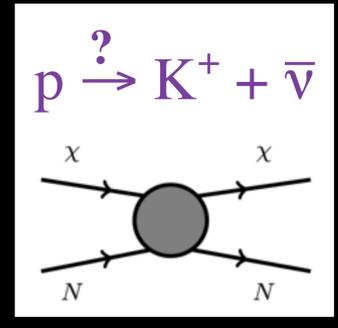
**Cross-Section Measurements**



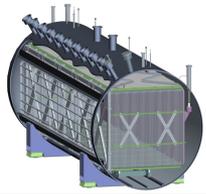
**Detector Physics for LArTPC R&D**



**Supernova Neutrino Detection**



**Exotic Physics**

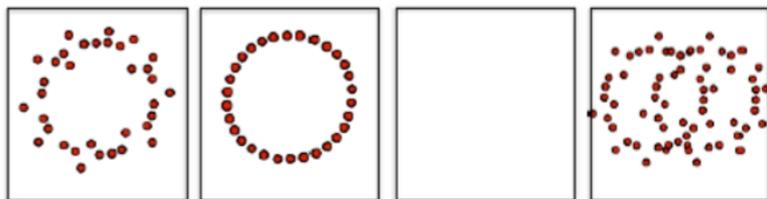
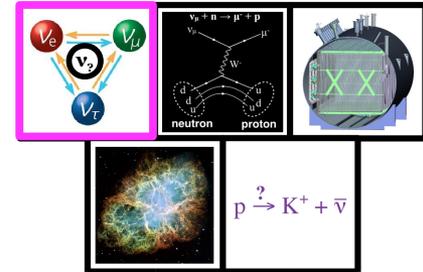


# Low-Energy Excess

◆ Low-energy  $\nu_e/\bar{\nu}_e$  candidate excess seen at MiniBooNE

- MiniBooNE: Cherenkov detector (also on BNB)
- Baseline too short (541 m) for 3-flavor  $\nu_\mu \rightarrow \nu_e$  oscillation

◆ No  $e^\pm/\gamma$  separation... is excess misunderstood background, sterile neutrino, or... ?



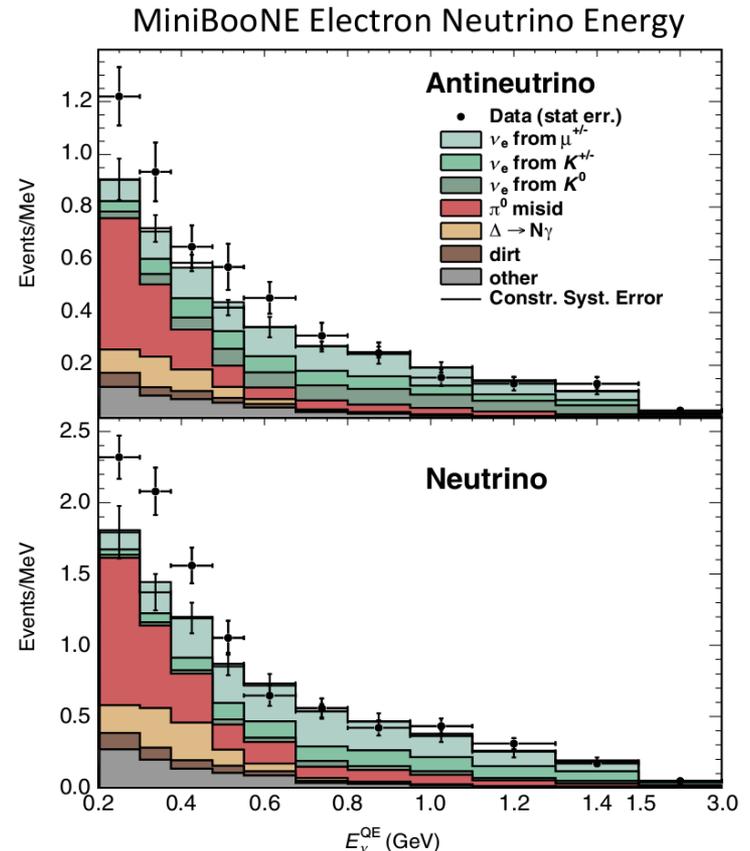
Electron,  
Photon

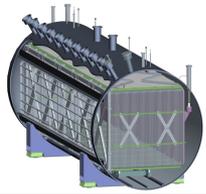
Muon

Proton

$\pi^0 \rightarrow \gamma + \gamma$

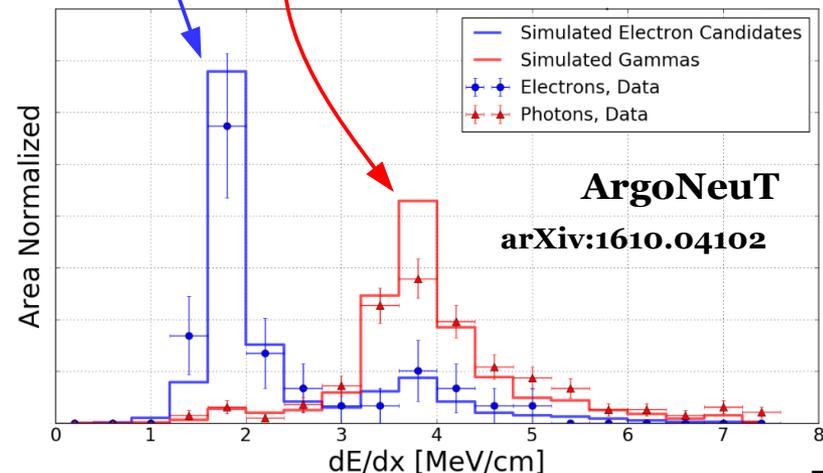
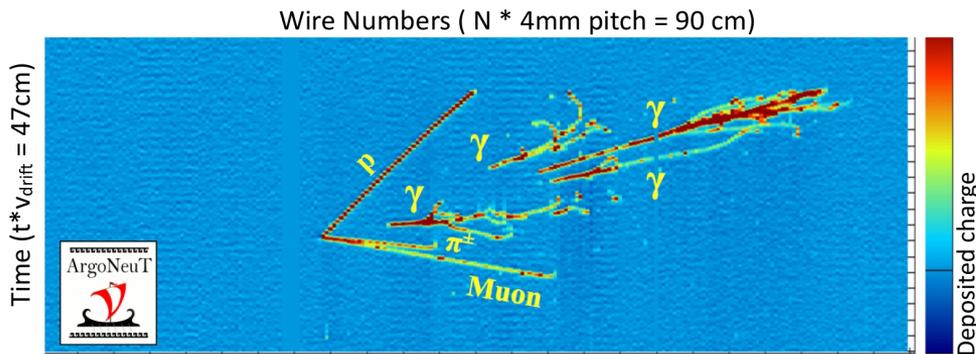
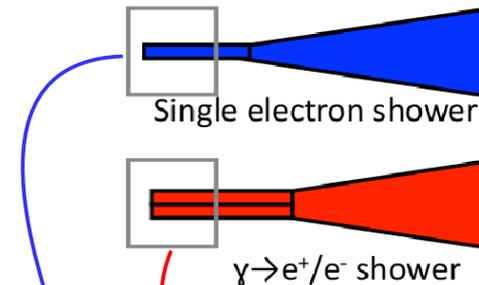
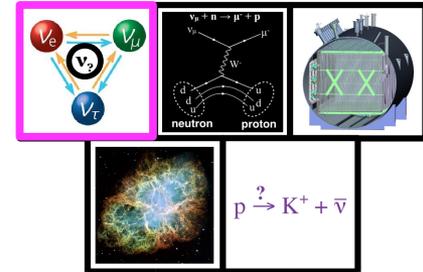
(Cherenkov Detector)

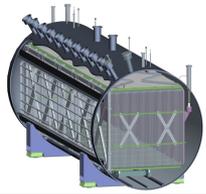




# Low-Energy Excess (cont.)

- ◆ Can discriminate  $e^\pm/\gamma$  with MicroBooNE's LArTPC
  - Shower displacement from vertex (“gap”) for  $\gamma$  also provides separation
  - Separation with  $dE/dx$
- ◆ End result: either discover new particle or improve MC for future experiments

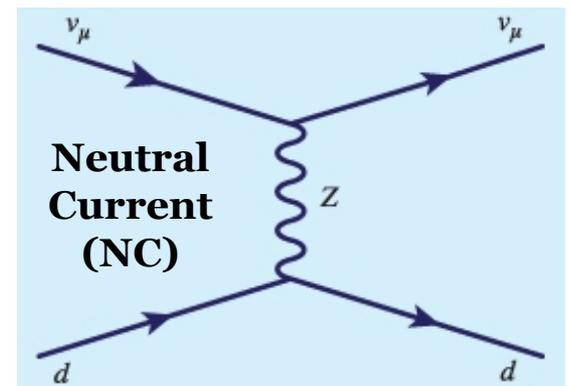
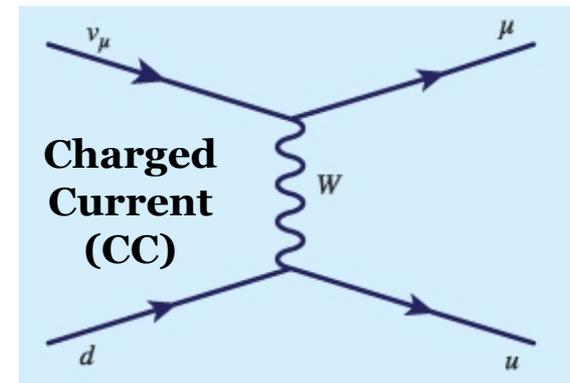
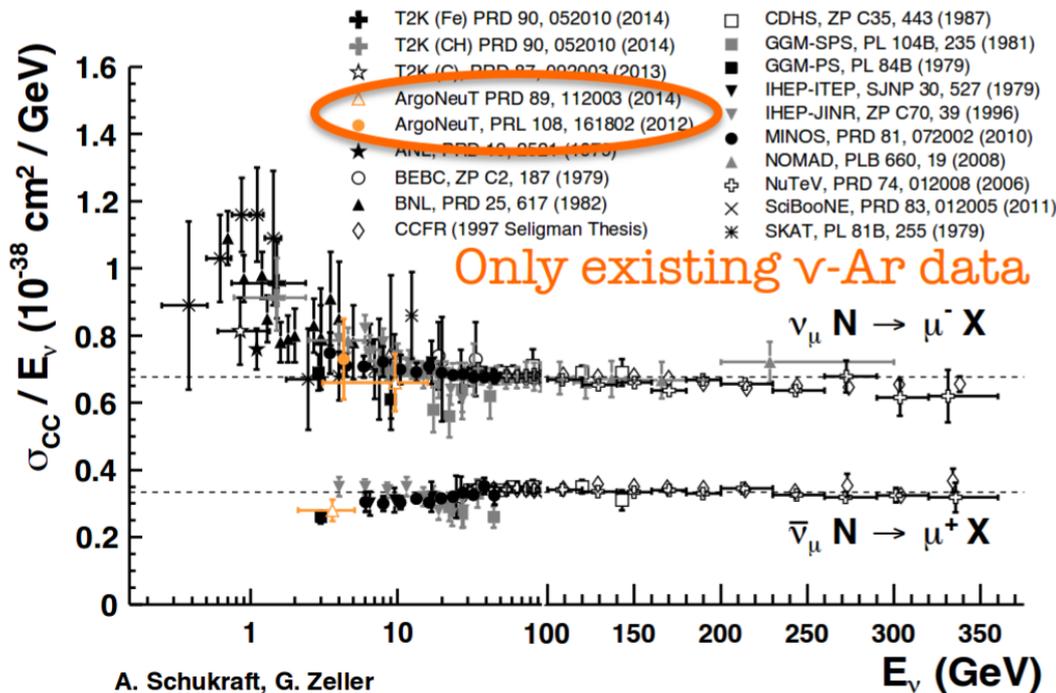
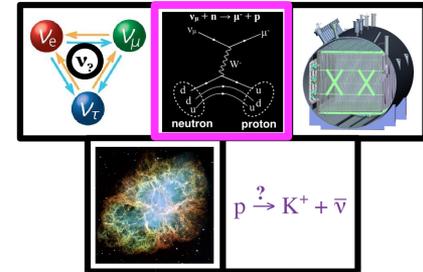


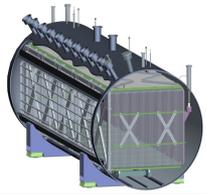


# Cross Section Measurements

◆ Cross-section measurements at MicroBooNE will teach us more about nuclear effects, neutrino energy reconstruction, etc.

- e.g. nucleon-nucleon correlations

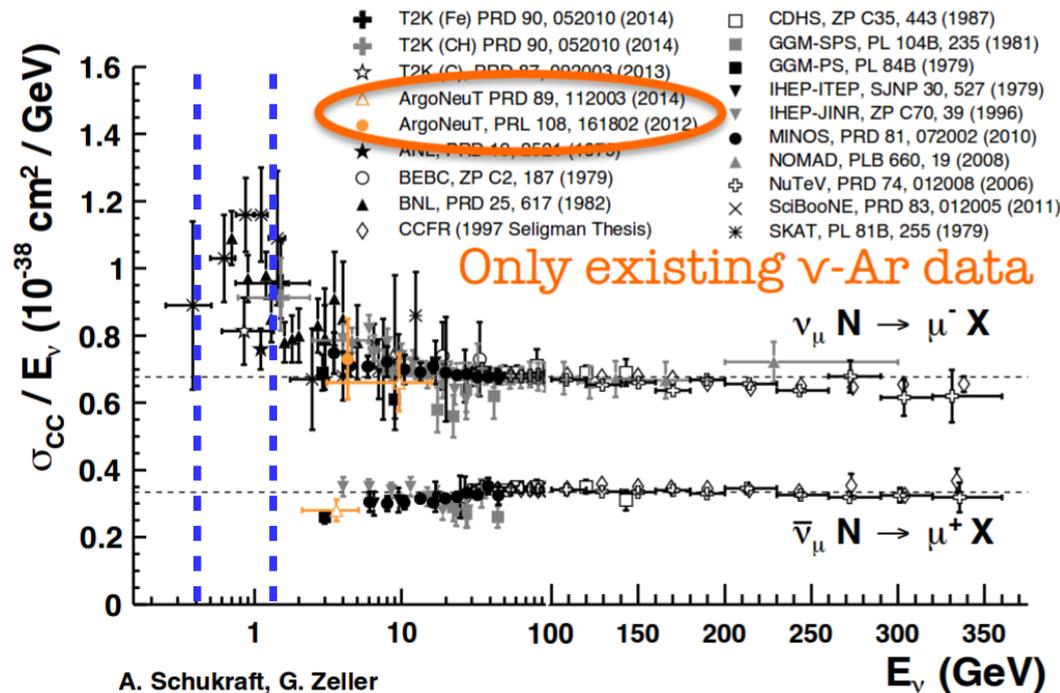
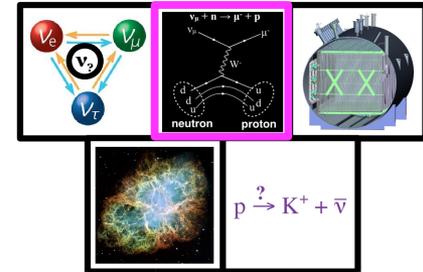




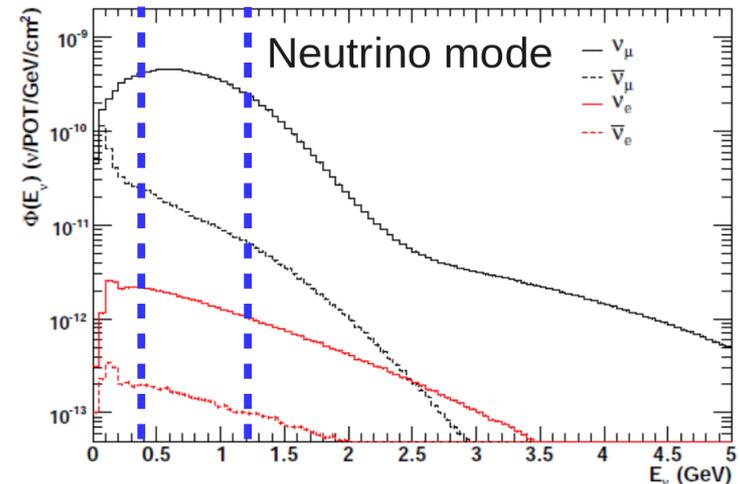
# Cross Section Measurements

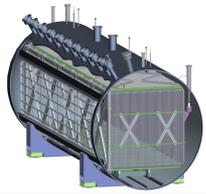
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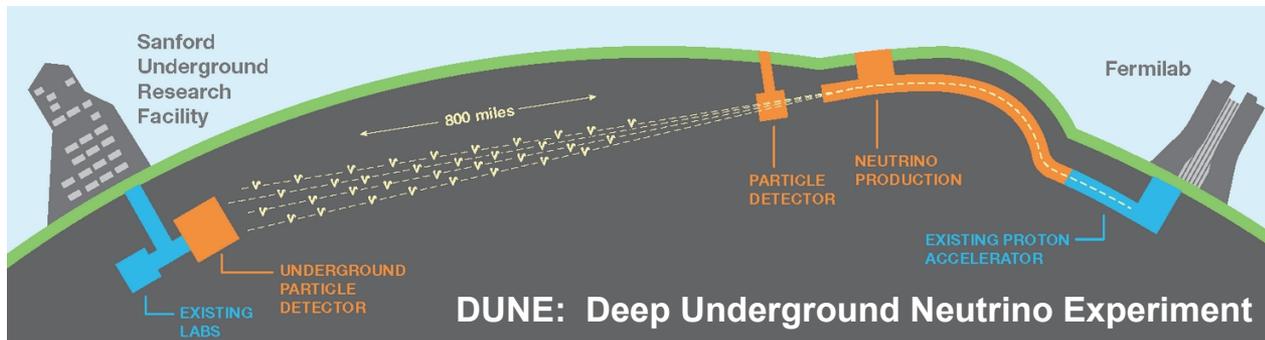
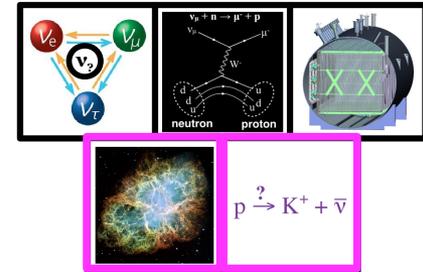
Energy Range:  
DUNE's 2<sup>nd</sup> Osc.  
Maximum

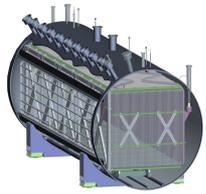




# Supernova $\nu$ , Exotic Physics

- ◆ Additional topics include studies related to supernova neutrinos and exotic physics
  - If we're lucky, supernova neutrinos ( $\sim 10$  MeV) captured using continuous readout stream and SNEWS alert system
    - Also study zero suppression, triggering schemes
  - Can study proton decay backgrounds in MicroBooNE's LArTPC
    - Signal:  $\mathbf{p} \rightarrow \mathbf{K}^+ + \mathbf{\nu}$
    - Background (cosmogenic):  $\mathbf{K}_L^0 + \mathbf{p} \rightarrow \mathbf{K}^+ + \mathbf{n}$
- ◆ Both helpful to DUNE physics program

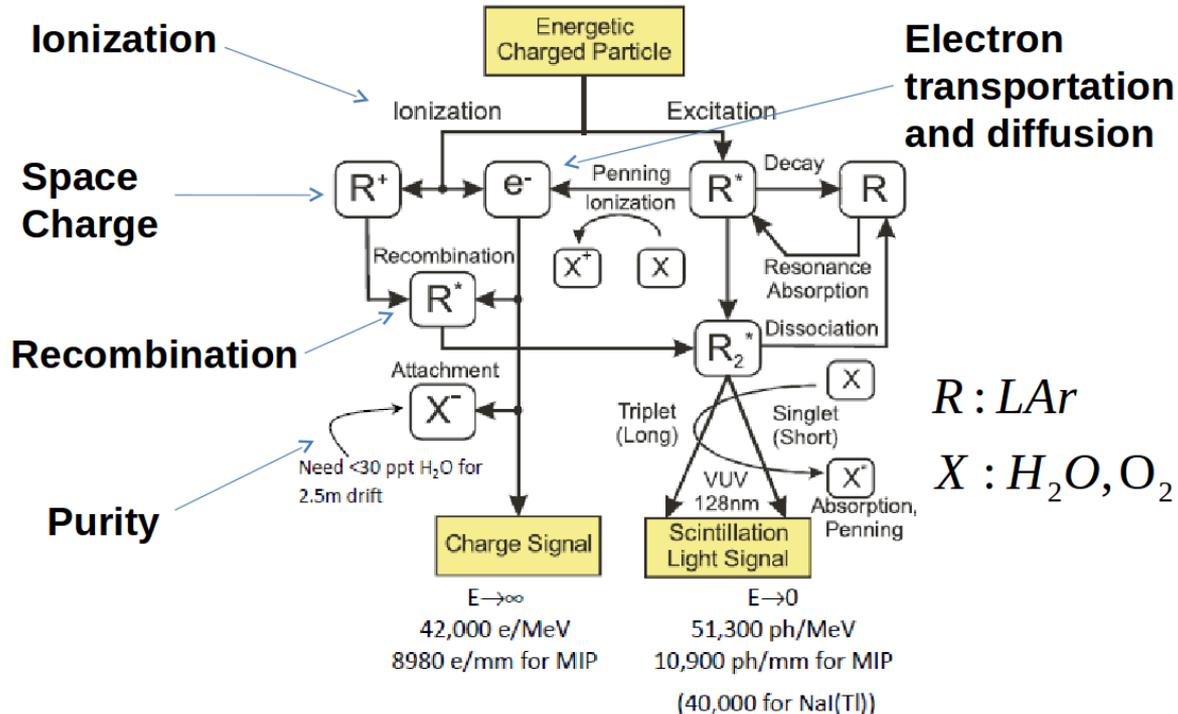
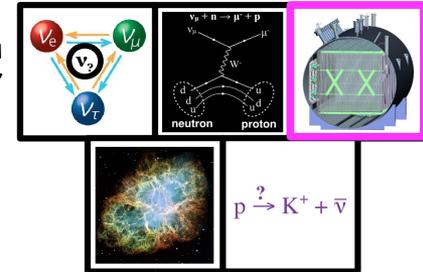




# Detector Physics

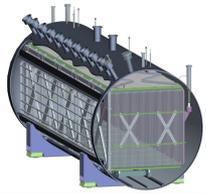
◆ Must understand detector effects to develop LArTPC technology

- Essential for SBN and DUNE
- Space charge effects (SCE), wire response, energy scale, noise studies, diffusion, e- lifetime, etc.



**Important to understand detector effects and develop calibration scheme for unbiased, precise determination of ionization charge.**

$R : \text{LAr}$   
 $X : \text{H}_2\text{O}, \text{O}_2$



# Booster Neutrino Beam

## Fermilab Neutrino Experiments

**Booster  $\nu$  beam**

*MicroBooNE, SBN program*

**Booster Neutrino Beam: “BNB”**

- Receives 8 GeV Protons from Booster
- $\nu_\mu$  ( $\bar{\nu}_\mu$ ) beam

**MicroBooNE @ BNB:**

- On-axis at 470 m baseline
- First three years in  $\nu_\mu$  mode (pre-SBN)

**NuMI  $\nu$  beam**

*NOVA, MINERvA, MINOS+*

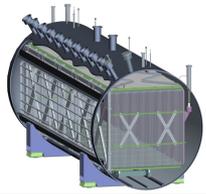
**Booster**

proton energy: 8 GeV

**Main Injector**

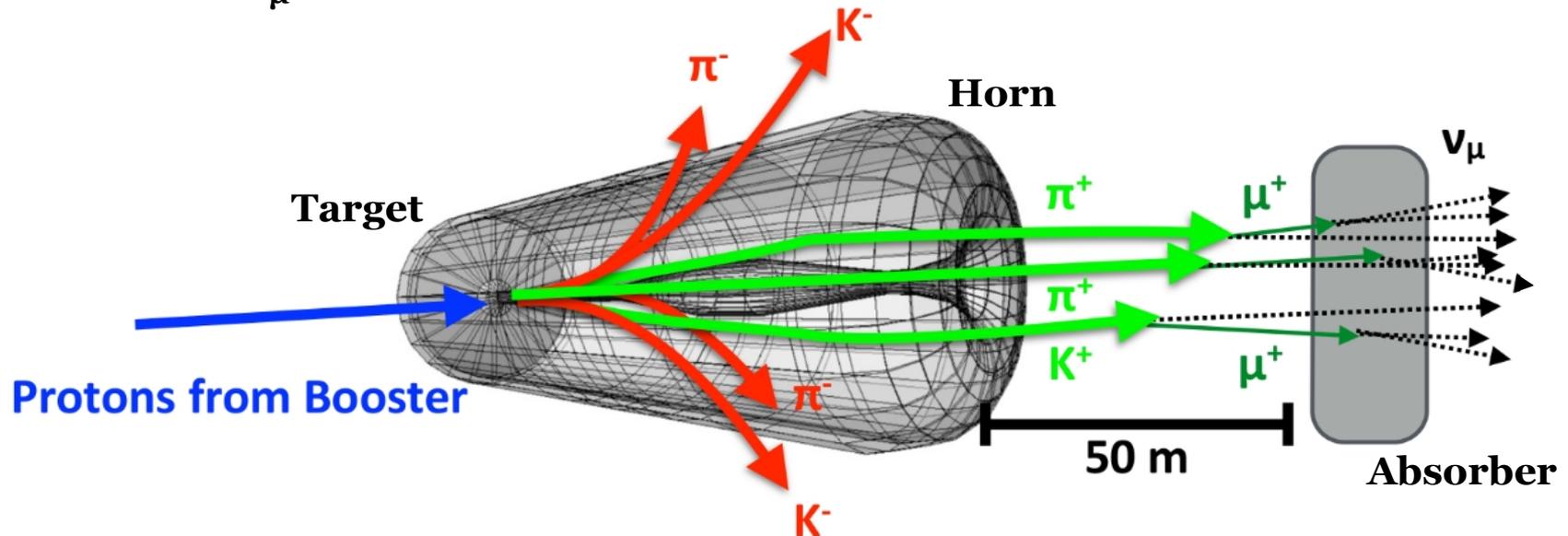
proton energy: 120 GeV

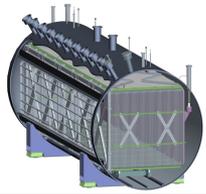
**DUNE  $\nu$  beam**



# BNB Overview

- ◆ Protons hit beryllium target producing mesons
- ◆ Magnetic field of horn focuses positive mesons, defocuses negative mesons
- ◆ 50 m decay pipe for  $\pi^+$  and  $K^+$  decay to primarily  $\mu^+$  and  $\nu_\mu$
- ◆ Layers of steel and concrete absorb charged particles
- ◆ Result:  $\nu_\mu$  beam





# BNB Overview

- ◆ Protons hit beryllium target producing mesons
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