







# **Gd** -doping and the impact on



Pablo Fernández, Univ. Autónoma Madrid for the Super Kamiokande Collaboration

2016/11/05 — NNN16, Beijing

## Index

#### **Gadolinium in Water-Cerenkov Detectors**

- Gadolinium Neutron Tagging
- The SuperK-Gd Project
- Gadolinium Addition to SuperK
- Gadolinium Neutron Tagging Reconstruction in SuperK

## Index

#### **Gadolinium in Water-Cerenkov Detectors**

- Gadolinium Neutron Tagging
- The SuperK-Gd Project
- Gadolinium Addition to SuperK
- Gadolinium Neutron Tagging Reconstruction in SuperK

#### **Gadolinium Impact on Physics**

- Low Energy Neutrinos in SuperK-Gd
  - ★ Supernova Diffuse Neutrino Background, Early Warning, Burst
  - **★** Reactor Neutrinos
  - **★** Solar neutrinos
- High Energy Physics in SuperK-Gd
  - ★ Proton decay
  - ★ Long Baseline Neutrinos
  - **★** Atmospheric Neutrinos

## Index

#### **Gadolinium in Water-Cerenkov Detectors**

- Gadolinium Neutron Tagging
- The SuperK-Gd Project
- Gadolinium Addition to SuperK
- Gadolinium Neutron Tagging Reconstruction in SuperK

#### **Gadolinium Impact on Physics**

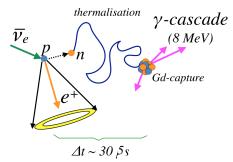
- Low Energy Neutrinos in SuperK-Gd
  - ★ Supernova Diffuse Neutrino Background, Early Warning, Burst
  - **★** Reactor Neutrinos
  - ★ Solar neutrinos
- High Energy Physics in SuperK-Gd
  - ★ Proton decay
- ★ Long Baseline Neutrinos
- **★** Atmospheric Neutrinos

#### **Conclusions**

## **Gadolinium** in Water-Cerenkov Detectors

#### **Gadolinium Neutron Tagging**

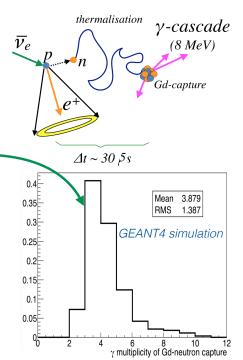
- ★ Gadolinium has the largest neutron capture cross-section of all stable nuclei in nature, around 48800 barn on average given the natural abundances of its isotopes
- ★ The Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub> salt was finally decided to be the best option in terms of transparency and corrosion



#### **Gadolinium Neutron Tagging**

- ★ Gadolinium has the largest neutron capture cross-section of all stable nuclei in nature, around 48800 barn on average given the natural abundances of its isotopes
- ★ The Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub> salt was finally decided to be the best option in terms of transparency and corrosion
- \* After the neutrino interaction neutrons may be emitted and thermalised (~15μs mean time). Once the neutron is thermal, it is captured by a Gd nucleus (~20μs mean lifetime), resulting in an excited Gd nucleus, then this nucleus deexcites emitting a γ-ray cascade of ~8 MeV of 3 to 5 photons

$$n + {}^{X}Gd \rightarrow {}^{X+1}Gd^* \rightarrow {}^{X+1}Gd + \gamma s (8 MeV)$$

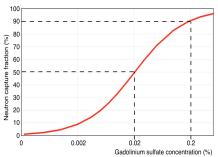


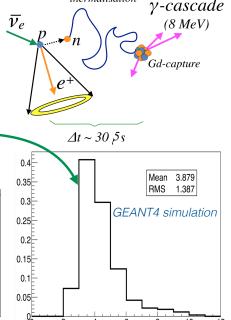
#### **Gadolinium Neutron Tagging**

- ★ Gadolinium has the largest neutron capture cross-section of all stable nuclei in nature, around 48800 barn on average given the natural abundances of its isotopes
- ★ The Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub> salt was finally decided to be the best option in terms of transparency and corrosion
- \* After the neutrino interaction neutrons may be emitted and thermalised (~15μs mean time). Once the neutron is thermal, it is captured by a Gd nucleus (~20μs mean lifetime), resulting in an excited Gd nucleus, then this nucleus deexcites emitting a γ-ray cascade of ~8 MeV of 3 to 5 photons

$$n + {}^{X}Gd \rightarrow {}^{X+1}Gd^* \rightarrow {}^{X+1}Gd + \gamma s (8 MeV)$$

★ For 0.2% Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub> concentration by mass dissolved in water, 90% of neutrons produced will be captured by Gd and 50% with 0.02% concentration

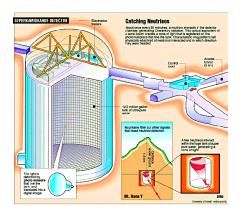




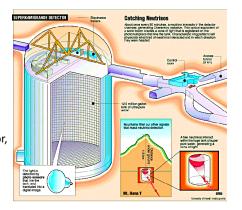
γ multiplicity of Gd-neutron capture

thermalisation

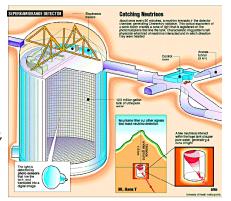
- ★ Super-Kamiokande (SK) is a 50,000 ton water Cherenkov detector located in the Kamioka mine under 1000 m of rock
- ★ SK celebrates this year its <u>20<sup>th</sup> anniversary</u> since the start of its measurements



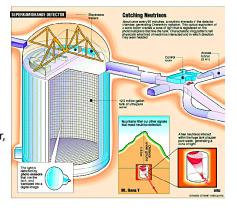
- ★ Super-Kamiokande (SK) is a 50,000 ton water Cherenkov detector located in the Kamioka mine under 1000 m of rock
- ★ SK celebrates this year its <u>20<sup>th</sup> anniversary</u> since the start of its measurements
- ★ SuperK-Gd will consist in the addition of 100 tons of Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub>, dissolved into the SK water
- ★ By adding this 0.2% by mass of the Gd salt, 90% of the produced neutrons in the detector will be captured by Gd after they thermalise
- ★ SuperK-Gd will be able to detect ~80% of all the neutrons produced in the detector, providing a much more complete information of the final state of the interaction



- ★ Super-Kamiokande (SK) is a 50,000 ton water Cherenkov detector located in the Kamioka mine under 1000 m of rock
- ★ SK celebrates this year its <u>20<sup>th</sup> anniversary</u> since the start of its measurements
- ★ SuperK-Gd will consist in the addition of 100 tons of Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub>, dissolved into the SK water
- ★ By adding this 0.2% by mass of the Gd salt, 90% of the produced neutrons in the detector will be captured by Gd after they thermalise
- ★ SuperK-Gd will be able to detect ~80% of all the neutrons produced in the detector, providing a much more complete information of the final state of the interaction
- ★ The EGADS R&D program has already show the <u>feasibility and performance</u> of adding Gd to water Cherenkov detectors



- ★ Super-Kamiokande (SK) is a 50,000 ton water Cherenkov detector located in the Kamioka mine under 1000 m of rock
- ★ SK celebrates this year its <u>20<sup>th</sup> anniversary</u> since the start of its measurements
- ★ SuperK-Gd will consist in the addition of 100 tons of Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub>, dissolved into the SK water
- ★ By adding this 0.2% by mass of the Gd salt, 90% of the produced neutrons in the detector will be captured by Gd after they thermalise
- ★ SuperK-Gd will be able to detect ~80% of all the neutrons produced in the detector, providing a much more complete information of the final state of the interaction
- ★ The EGADS R&D program has already show the <u>feasibility and performance</u> of adding Gd to water Cherenkov detectors



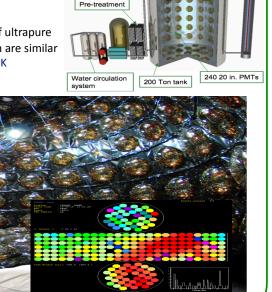
On June 27, 2015, the Super-Kamiokande collaboration approved the SK-Gd project which will enhance neutrino detectability by dissolving gadolinium in the Super-K water.

T2K and SK will jointly develop a protocol to make the decision about when to trigger the SK-Gd project, taking into account the needs of both experiments, including preparation for the refurbishment of the SK tank and readiness of the SK-Gd project, and the T2K schedule including the J-PARC MR power upgrade. Given the currently anticipated schedules, the expected time of the refurbishment is 2018.

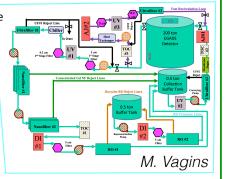
- ★ EGADS (Evaluating Gadolinium's Action on Detector Systems) was a R&D project for testing the feasibility of adding Gd in water Cherenkov detectors, specifically focused in the realisation of SuperK-Gd
- ★ EGADS detector is a scaled down Super Kamiokande, containing 200 ton of ultrapure water and instrumented with 240 PMTs (40% photocoverage), 227 of them are similar to those in the SK and the other 13 are developing photosensors for HyperK



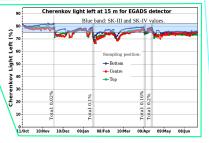
- ★ EGADS (Evaluating Gadolinium's Action on Detector Systems) was a R&D project for testing the feasibility of adding Gd in water Cherenkov detectors, specifically focused in the realisation of SuperK-Gd
- ★ EGADS detector is a scaled down Super Kamiokande, containing 200 ton of ultrapure water and instrumented with 240 PMTs (40% photocoverage), 224 of them are similar to those in the SK and the other 16 are developing photosensors for HyperK

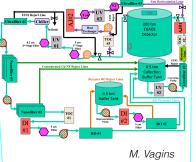


- $\star$  In this context, **EGADS** has proven that 4 key characteristics can be achieved for the realisation of SuperK-Gd
  - New water systems for dissolving Gd into water (pre-treatment) and purifying the Gd-doped water (selective filtration system). These are specifically design to match the ultrapurity requirements like the SK water and to keep all the Gd and sulphate ions dissolved



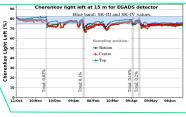
- $\star$  In this context, **EGADS** has proven that 4 key characteristics can be achieved for the realisation of SuperK-Gd
  - New water systems for dissolving Gd into water (pre-treatment) and purifying the Gd-doped water (selective filtration system). These are specifically design to match the ultrapurity requirements like the SK water and to keep all the Gd and sulphate ions dissolved
  - The sophisticated Gd water system makes the water transparency of Gd-doped water very similar to that of SK

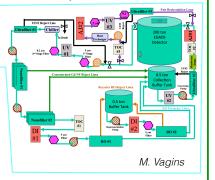


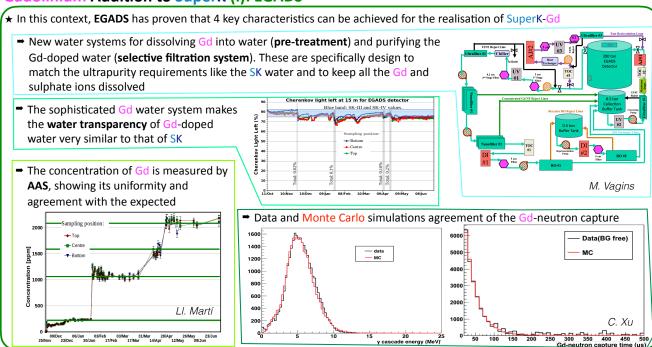


- $\star$  In this context, **EGADS** has proven that 4 key characteristics can be achieved for the realisation of SuperK-Gd
  - → New water systems for dissolving Gd into water (pre-treatment) and purifying the Gd-doped water (selective filtration system). These are specifically design to match the ultrapurity requirements like the SK water and to keep all the Gd and sulphate ions dissolved
  - → The sophisticated Gd water system makes the water transparency of Gd-doped water very similar to that of SK
  - → The concentration of Gd is measured by AAS, showing its uniformity and agreement with the expected





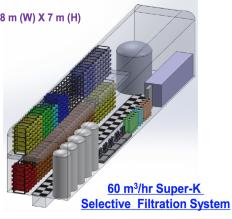




#### Gadolinium Addition to SuperK (II): SuperK-Gd water system

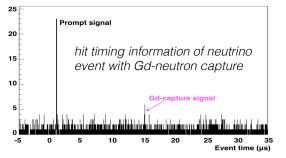
★ The works for scaling the **EGADS** water system to SK dimensions have alredy began designing the SuperK-Gd selective filtration system and the excavation of the new hall which will host it





★ Algorithm for detecting the Gd-neutron capture: a two-step process

- ★ Algorithm for detecting the Gd-neutron capture: a two-step process
  - Candidates are selected by scanning the number of hits clustered in the hit timing distribution after the prompt signal is triggered



- ★ Algorithm for detecting the Gd-neutron capture: a two-step process
  - Candidates are selected by scanning the number of hits clustered in the hit timing distribution after the prompt signal is triggered

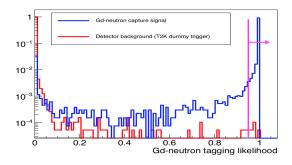
20
hit timing information of neutrino
event with Gd-neutron capture

10
Gd-capture signal

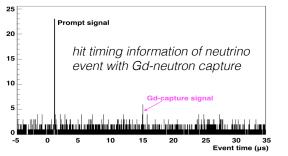
Gd-capture signal

5
Event time (µs)

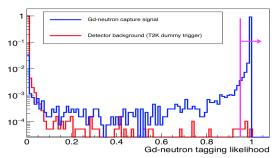
■ In order to discern Gd-neutron captures from the background candidates, a likelihood method is used



- ★ Algorithm for detecting the Gd-neutron capture: a two-step process
  - Candidates are selected by scanning the number of hits clustered in the hit timing distribution after the prompt signal is triggered



■ In order to discern Gd-neutron captures from the background candidates, a likelihood method is used



	Gd-neutron capture	Background
Candidate selection	99.3%	$2.72  \mathrm{event^{-1}}$
Likelihood selection	93.3%	0.009%
Total efficiency	92.7%	$2.4 \cdot 10^{-4} \text{ event}^{-1}$

★ As ~90% of neutrons are captured by Gd —> total Gd-neutron tagging efficiency: ~80%

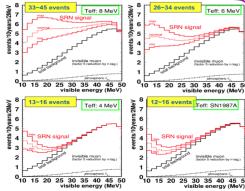
## **Gadolinium Impact on Physics**

- **★ Diffuse Supernova Neutrino Background (DSNB)** 
  - → SuperK-Gd will be able to first measure the antineutrinos coming from all the past supernovae explossions through their inverse β interaction in the detector, largely supressing the current spallation background

Signal: ~5 events/year/SK

→ The main remaining background is due to <sup>238</sup>U SF that might get into the detector as contaminant of the Gd salt

Background: ~0.11 events/year/SK per mBq/kg of Gd2(SO4)3



#### **★ Diffuse Supernova Neutrino Background (DSNB)**

SuperK-Gd will be able to first measure the antineutrinos coming from all the past supernovae explossions through their inverse β interaction in the detector, largely supressing the current spallation background

Signal: ~5 events/year/SK

→ The main remaining background is due to <sup>238</sup>U SF that might get into the detector as contaminant of the Gd salt

Background: ~0.11 events/year/SK per mBq/kg of Gd2(SO4)3

#### **★ Supernova Early Warning**

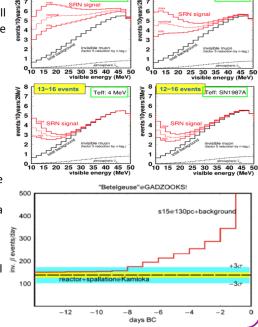
- → This refers to th Si burning phase previous to the core-collapse where ~1% of the total enegy is released in the form of low energy neutrinos
- → the number of events above 3 MeV assuming a very close supernova such as Betelgeuse (0.2 kpc away) in the 24h before the explossion is

#### Signal: 16.4 events/24h/SK

 The main remaining background is due to reactor neutrinos, assuming all Japanese reactors

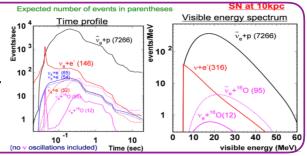
Background: ~30 events/day/SK

Significance: 3.4σ



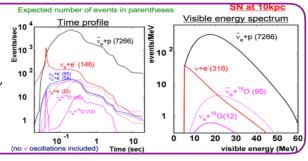
#### ★ Supernova Burst

- Supernova electron antineutrinos will be much better distinguished from neutrinos due to the Gd-neutron tagging
- → This measurement, will provide much information about early stages of the core-collapse process, its spectrum and time profile, yielding to more detailed picture of the whole core-collapse
- SuperK-Gd will detect thousends of events with negligible background where v and v fluxes independently will be extracted independently

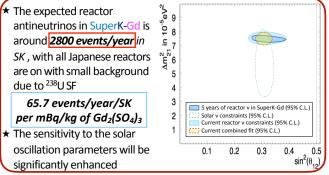


#### **★ Supernova Burst**

- Supernova electron antineutrinos will be much better distinguished from neutrinos due to the Gd-neutron tagging
- → This measurement, will provide much information about early stages of the core-collapse process, its spectrum and time profile, yielding to more detailed picture of the whole core-collapse
- SuperK-Gd will detect thousends of events with negligible background where v and v fluxes independently will be extracted independently

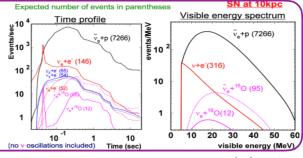


### Low Energy Neutrinos in SuperK-Gd (II): Reactor

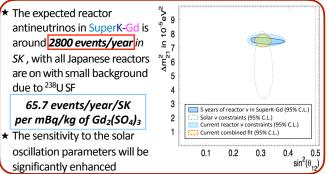


#### ★ Supernova Burst

- Supernova electron antineutrinos will be much better distinguished from neutrinos due to the Gd-neutron tagging
- → This measurement, will provide much information about early stages of the core-collapse process, its spectrum and time profile, yielding to more detailed picture of the whole core-collapse
- SuperK-Gd will detect thousends of events with negligible background where v and v fluxes independently will be extracted independently



#### Low Energy Neutrinos in SuperK-Gd (II): Reactor



#### Low Energy Neutrinos in SuperK-Gd (III): Solar

- ★ With the addition of Gd to SK water solar neutrinos may be affected due to the radioactive β decays from the contamination in the Gd salt
- ★ Low energy (from 3 MeV) events in SK after solar cuts are

~200 events/day/SK

- ★ The radiopurity requirements for low energy solar neutrino analysis are
  - ⇒ <sup>228</sup>Th: <0.03 mBq/kg of Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub>
  - ⇒ <sup>226</sup>Ra: <0.51 mBq/kg of Gd<sub>2</sub>(SO<sub>4</sub>)<sub>3</sub>

#### Low Energy Neutrinos in SuperK-Gd (IV): Radioactive contamination

★ The Gd salt will be dissolved uniformly along the whole volume of SK and, therefore, special attention has to paid to its induced radioactive contamination as seen previously

★ Exhaustive radioactive maesuring capaings are being carried out at Canfranc Underground Lab. (Spain), Kamioka Observatory (Japan) and Boulby mine (UK)

	Measured radioactivity in $mBq/kg$ for all the $\mathrm{Gd}_2(\mathrm{SO}_4)_3$ batches									
Chain		$^{238}{ m U}$		$^{232}{ m Th}$		$^{235}{ m U}$		Others		
	Sub-Chain	$^{238}\mathbf{U}$	$^{226}$ Ra	$^{228}$ Ra	$^{228}\mathbf{Th}$	$^{235}\mathbf{U}$	$^{227}Ac$	$^{40}\mathbf{K}$	$^{138}$ La	$^{176}$ Lu
,	Company A	$51 \pm 21$	8 ± 1	$11 \pm 2$	28 ± 3	< 32	$214\pm10$	29 ± 5	8 ± 1	80 ± 8
	Company A (10/08)	< 33	$2.8 \pm 0.6$	$270 \pm 16$	86 ± 5	< 32	$1700 \pm 20$	$12\pm3$	<	$21\pm2$
	Company B	$292 \pm 6$	$74\pm2$	$1099 \pm 12$	$504 \pm 6$	< 112	$2956 \pm 30$	101 ± 10	$683 \pm 15$	566 ± 6
anies	Company C (13/02)	$74 \pm 28$	$13\pm1$	$205\pm6$	$127\pm3$	< 25	$1423 \pm 21$	$60 \pm 7$	$3\pm1$	$12\pm1$
Supplying Companies	Company B (13/03)	$242\pm 6$	$13\pm2$	$21\pm3$	$374 \pm 6$	< 25	$175\pm42$	$18\pm8$	$42\pm3$	8 ± 2
Supplyi	Company A (13/08)	$71\pm20$	8 ± 1	$6\pm1$	$159\pm3$	< 32	$295\pm10$	$3\pm 2$	$5\pm1$	30 ± 1
	Company D (13/07a)	$47\pm26$	$5\pm1$	$14 \pm 2$	$13\pm1$	< 12	< 6	$2\pm2$	< 1	$1.6\pm0.3$
	Company D (13/07b)	$73\pm27$	6 ± 1	$3\pm1$	$411\pm5$	< 30	< 18	$8\pm4$	< 2	< 2
	Company A	< 76	< 1.4	$2\pm1$	$29\pm2$	< 1.8	190 ± 6	< 5	$23\pm1$	$2.5\pm0.6$
	Company E	< 34	< 0.8	< 1.1	$2.0\pm0.5$	< 0.6	$11 \pm 4$	< 3	< 0.6	$2.9\pm0.2$
	Company F	< 139	< 2.1	$2.8 \pm 1.9$	$1.8 \pm 0.9$	< 2.4	< 10	< 14	< 1.9	< 1.6
	Company F (16/04)	< 20	< 0.64	< 0.67	$0.5 \pm 0.2$	< 0.7	< 2.3	< 1.6	< 0.3	< 0.4

L. Labarga, J. Pérez

#### Low Energy Neutrinos in SuperK-Gd (IV): Radioactive contamination

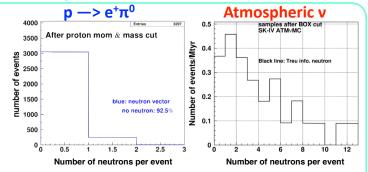
- ★ The Gd salt will be dissolved uniformly along the whole volume of SK and, therefore, special attention has to paid to its induced radioactive contamination as seen previously
- ★ Exhaustive radioactive maesuring capaings are being carried out at Canfranc Underground Lab. (Spain), Kamioka Observatory (Japan) and Boulby mine (UK)
- ★ Current gadolinium sulphates batches have lower radioactive contamination due to cooperation with suplying companies
- \* At EGADS, big efforts are being done to remove the remaining most relevant contaminants for these measurements
  - → <u>AmberJet</u> removes **U** more than a factor 100
  - Resin <u>AJ1020Gd</u> is being developed for removing **Ra**
  - → Several methods, such pH shock, are being studied for **Th** removal

L. Labarga, J. Pérez

• •	industry .										
					tivity in m				batches		
Chain		$^{238}U$		$^{232}\mathrm{Th}$		<sup>235</sup> U		Others			
		Sub-Chain	$^{238}U$	$^{226}$ Ra	$^{228}$ Ra	$^{228}\mathbf{Th}$	$^{235}\mathbf{U}$	$^{227}Ac$	$^{40}\mathbf{K}$	$^{138}$ La	$^{176}$ Lu
		Company A	$51 \pm 21$	8 ± 1	$11 \pm 2$	$28\pm3$	< 32	$214\pm10$	29 ± 5	8 ± 1	80 ± 8
		Company A	< 33	$2.8\pm0.6$	$270\pm16$	$86 \pm 5$	< 32	$1700\pm20$	$12\pm3$	<	$21\pm2$
		Company B	$292\pm 6$	$74\pm2$	$1099\pm12$	$504 \pm 6$	< 112	$2956 \pm 30$	$101 \pm 10$	$683 \pm 15$	$566\pm6$
	anies	Company C (13/02)	$74\pm28$	$13\pm1$	$205\pm 6$	$127\pm3$	< 25	$1423\pm21$	60 ± 7	$3\pm1$	$12\pm1$
	ng Comp	Company B (13/03)	$242\pm 6$	$13\pm 2$	$21\pm3$	$374\pm6$	< 25	$175\pm42$	18 ± 8	$42\pm3$	$8\pm2$
	Supplyi	Company A (13/08)	$71 \pm 20$	8 ± 1	$6\pm1$	$159\pm3$	< 32	$295\pm10$	$3\pm 2$	$5\pm1$	$30\pm1$
		Company D (13/07a)	$47\pm26$	$5\pm1$	$14\pm2$	$13\pm1$	< 12	< 6	$2\pm2$	< 1	$1.6\pm0.3$
		Company D (13/07b)	$73\pm27$	$6\pm1$	$3\pm1$	$411\pm5$	< 30	< 18	$8\pm4$	< 2	< 2
		Company A (14/12)	< 76	< 1.4	$2\pm1$	$29\pm2$	< 1.8	$190\pm 6$	< 5	$23\pm1$	$2.5\pm0.6$
		Company E	< 34	< 0.8	< 1.1	$2.0\pm0.5$	< 0.6	$11\pm4$	< 3	< 0.6	$2.9 \pm 0.2$
		Company F (15/12)	< 139	< 2.1	$2.8 \pm 1.9$	$1.8\pm0.9$	< 2.4	< 10	< 14	< 1.9	< 1.6
		Company F	< 20	< 0.64	< 0.67	$0.5 \pm 0.2$	< 0.7	< 2.3	< 1.6	< 0.3	< 0.4

#### High Energy Physics in SuperK-Gd (I): Proton decay

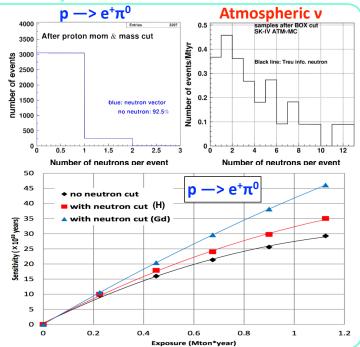
- ★ Neutron tagging provides an improvement on the sensitivity of proton decay searches
- \* As illustration, in the  $p \rightarrow e^+\pi^0$  decay mode the neutron multiplicity of proton decay is significantly different from the neutron multiplicity distribution of atmopsheric neutrinos after the proton decay cuts are applied



#### High Energy Physics in SuperK-Gd (I): Proton decay

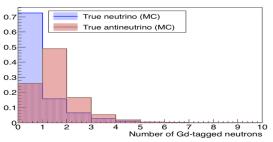
- ★ Neutron tagging provides an improvement on the sensitivity of proton decay searches
- \* As illustration, in the  $p \rightarrow e^+\pi^0$  decay mode the neutron multiplicity of proton decay is significantly different from the neutron multiplicity distribution of atmopsheric neutrinos after the proton decay cuts are applied

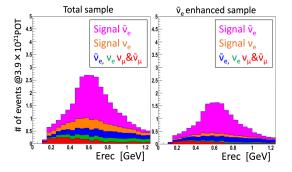
- ★ For the sensitivity study, 80% neutron tagging efficiency is assumed as discussed previously
  - → For this case, <u>92.5%</u> of proton decays produce no neutrons in their final state
  - → The remaining background for 10 years of observation in SK is 0.58, whereas for SuperK-Gd is only 0.098 events



#### High Energy Physics in SuperK-Gd (II): T2K Long baseline neutrinos

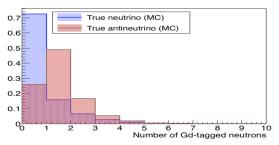
- $\star$  Long baseline neutrinos can also benefit from the  $v-\overline{v}$  separation induced by the 80% Gd-neutron tagging
- → For the separation a likelihood distribution is built using the number of neutrons and the scattering angle
- → This way, neutrino and antineutrino purity of samples is enhanced



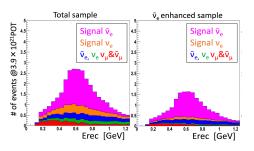


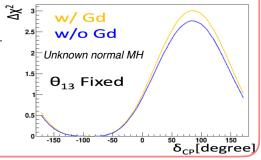
#### High Energy Physics in SuperK-Gd (II): T2K Long baseline neutrinos

- ★ Long baseline neutrinos can also benefit from the v-v separation induced by the 80% Gd-neutron tagging
  - → For the separation a likelihood distribution is built using the number of neutrons and the scattering angle
  - → This way, neutrino and antineutrino purity of samples is enhanced



- With the T2K analysis tools the official sensitivity to the <u>CP phase</u> is compared with that using the neutrino-antineutrino separation from Gdneutron tagging
- The sensitivity curves are done assuming the PDG values,  $\sin^2\theta_{23}$ =0.5 and 3.9·10<sup>21</sup>POT

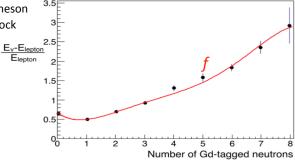




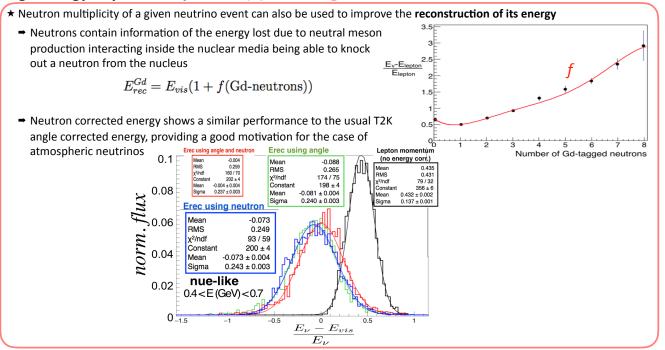
#### High Energy Physics in SuperK-Gd (II): T2K Long baseline neutrinos

- ★ Neutron multiplicity of a given neutrino event can also be used to improve the reconstruction of its energy
  - Neutrons contain information of the energy lost due to neutral meson production interacting inside the nuclear media being able to knock out a neutron from the nucleus

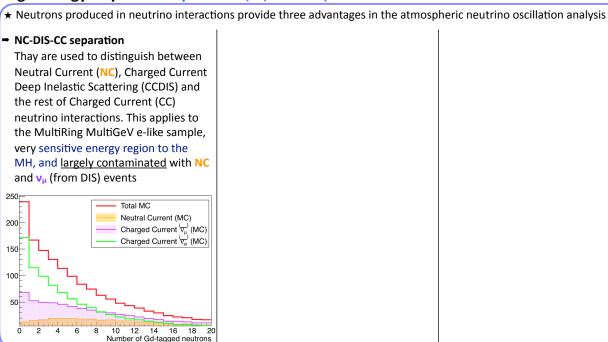
$$E_{rec}^{Gd} = E_{vis}(1 + f(\text{Gd-neutrons}))$$



## High Energy Physics in SuperK-Gd (II): T2K Long baseline neutrinos



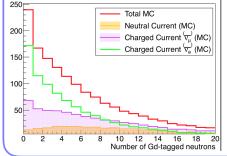
# High Energy Physics in SuperK-Gd (III): Atmospheric neutrinos ★ Neutrons produced in neutrino interactions provide three advantages in the atmospheric neutrino oscillation analysis



★ Neutrons produced in neutrino interactions provide three advantages in the atmospheric neutrino oscillation analysis

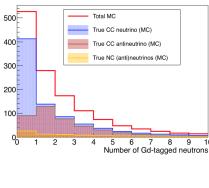
#### → NC-DIS-CC separation

Thay are used to distinguish between Neutral Current (NC), Charged Current Deep Inelastic Scattering (CCDIS) and the rest of Charged Current (CC) neutrino interactions. This applies to the MultiRing MultiGeV e-like sample, very sensitive energy region to the MH, and <u>largely contaminated</u> with NC and v<sub>u</sub> (from DIS) events



#### ⇒ v-v̄ separation

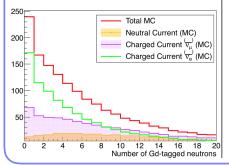
This is done for all the atmospheric neutrino samples. Here, the fact that antineutrinos tend to produce more neutrons in the final state than neutrinos in charged current interactions



\* Neutrons produced in neutrino interactions provide three advantages in the atmospheric neutrino oscillation analysis

#### ■ NC-DIS-CC separation

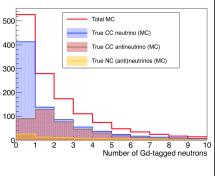
Thay are used to distinguish between Neutral Current (NC), Charged Current Deep Inelastic Scattering (CCDIS) and the rest of Charged Current (CC) neutrino interactions. This applies to the MultiRing MultiGeV e-like sample, very sensitive energy region to the MH, and largely contaminated with NC and  $v_{\mu}$  (from DIS) events



#### ⇒ ν-ν̄ separation

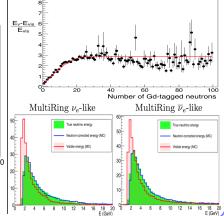
This is done for all the atmospheric neutrino samples.

Here, the fact that antineutrinos tend to produce more neutrons in the final state than neutrinos in charged current interactions

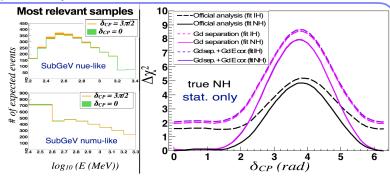


#### Neutron energy corrections

It is observed that neutron multiplicity provides information about the fraction of the neutrino energy invisible to the detector, due to its relation with the neutral hadron production  $(\pi, \eta, \kappa...)$ . inside the nuclear media



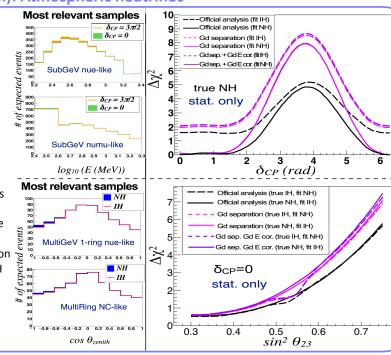
\* The sensitivity to the <u>CP violation phase</u> is largely improved due to the high efficiency of the  $v-\overline{v}$  separation in the SubGeV samples Adjacent plots show the spectra of two of the most sensitive samples to the  $\delta_{CP}$  and its  $\chi 2$  distribution for rejecting  $\delta_{CP}=0$ 



★ The sensitivity to the <u>CP violation phase</u> is largely improved due to the high efficiency of the v- $\overline{v}$  separation in the SubGeV samples Adjacent plots show the spectra of two of the most sensitive samples to the  $\delta_{CP}$  and its  $\chi 2$  distribution for rejecting  $\delta_{CP}$ =0

★ The sensitivity to the <u>neutrino mass hierarchy</u> is improved due to the NC-DIS-CC and v-v separation in the MultiGeV and MultiRing e-like samples

Adjacent plots show the zentih angle distribution of two of the most sensitive samples to the MH and its  $\chi 2$  distribution as function of  $\theta_{23}$ 



- ★ <u>Gd addition to water-Cherenkov detectors enhance their capabilities</u> through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very <u>important effects in low and high</u> energy processes

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - → Low Energy Physics:
    - DSNB will be first measured within few years time given the current radioactive contamination levels

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - → Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from <u>supernovae</u> within our galaxy will be measured
    - ✓ If a SN is close enough (10<sup>-1</sup> kpc), the <u>Si burning phase</u> of the star can alsobe measured

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - → Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from supernovae within our galaxy will be measured
    - ✓ If a SN is close enough (10<sup>-1</sup> kpc), the Si burning phase of the star can alsobe measured
    - Reactor antineutrinos can also be mesured, enabling to constraint the solar oscillation parameters

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - → Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from supernovae within our galaxy will be measured
    - ✓ If a SN is close enough (10<sup>-1</sup> kpc), the Si burning phase of the star can alsobe measured
    - √ Reactor antineutrinos can also be mesured, enabling to constraint the solar oscillation parameters
    - ✓ <u>Low energy solar neutrino</u> analysis will be still alive due to the great works limiting the radioactive impurities of Gd salt

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - **→** Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from supernovae within our galaxy will be measured
    - ✓ If a SN is close enough (10<sup>-1</sup> kpc), the Si burning phase of the star can alsobe measured
    - ▼ Reactor antineutrinos can also be mesured, enabling to constraint the solar oscillation parameters.
    - Low energy solar neutrino analysis will be still alive due to the great works limiting the radioactive impurities
      of Gd salt
  - → High Energy Physics:
    - ✓ <u>Proton decay</u> sensitivity is significantly enhanced due to neutron veto

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - → Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from supernovae within our galaxy will be measured
    - ✓ If a SN is close enough (10<sup>-1</sup> kpc), the Si burning phase of the star can alsobe measured
    - √ Reactor antineutrinos can also be mesured, enabling to constraint the solar oscillation parameters
    - Low energy solar neutrino analysis will be still alive due to the great works limiting the radioactive impurities
      of Gd salt
  - → High Energy Physics:
    - ✓ Proton decay sensitivity is significantly enhanced due to neutron veto
    - ✓ T2K will also benefit of Gd-neutron tagging enhancing the sensitivity to the CP phase

- ★ Gd addition to water-Cherenkov detectors enhance their capabilities through the efficient detection of neutrons produced from neutrino interations and proton decays
- ★ SuperK-Gd project is in a very advanced status thanks, mainly, to the excellent works at EGADS demostrator
- ★ The main physics outcomes of SuperK-Gd have been reviewed showing very important effects in low and high energy processes
  - **→** Low Energy Physics:
    - ✓ DSNB will be first measured within few years time given the current radioactive contamination levels
    - ✓ Neutrino and antineutrino spectra from supernovae within our galaxy will be measured
    - $\checkmark$  If a SN is close enough (10<sup>-1</sup> kpc), the Si burning phase of the star can alsobe measured
    - ▼ Reactor antineutrinos can also be mesured, enabling to constraint the solar oscillation parameters.
    - Low energy solar neutrino analysis will be still alive due to the great works limiting the radioactive impurities
      of Gd salt
  - → High Energy Physics:
    - ✓ Proton decay sensitivity is significantly enhanced due to neutron veto
    - ✓ T2K will also benefit of Gd-neutron tagging enhancing the sensitivity to the CP phase
    - ✓ Several advantages will be achieved in <u>atmospheric neutrino</u> analysis proving better sensitivity to the neutrino MH and the CP phase