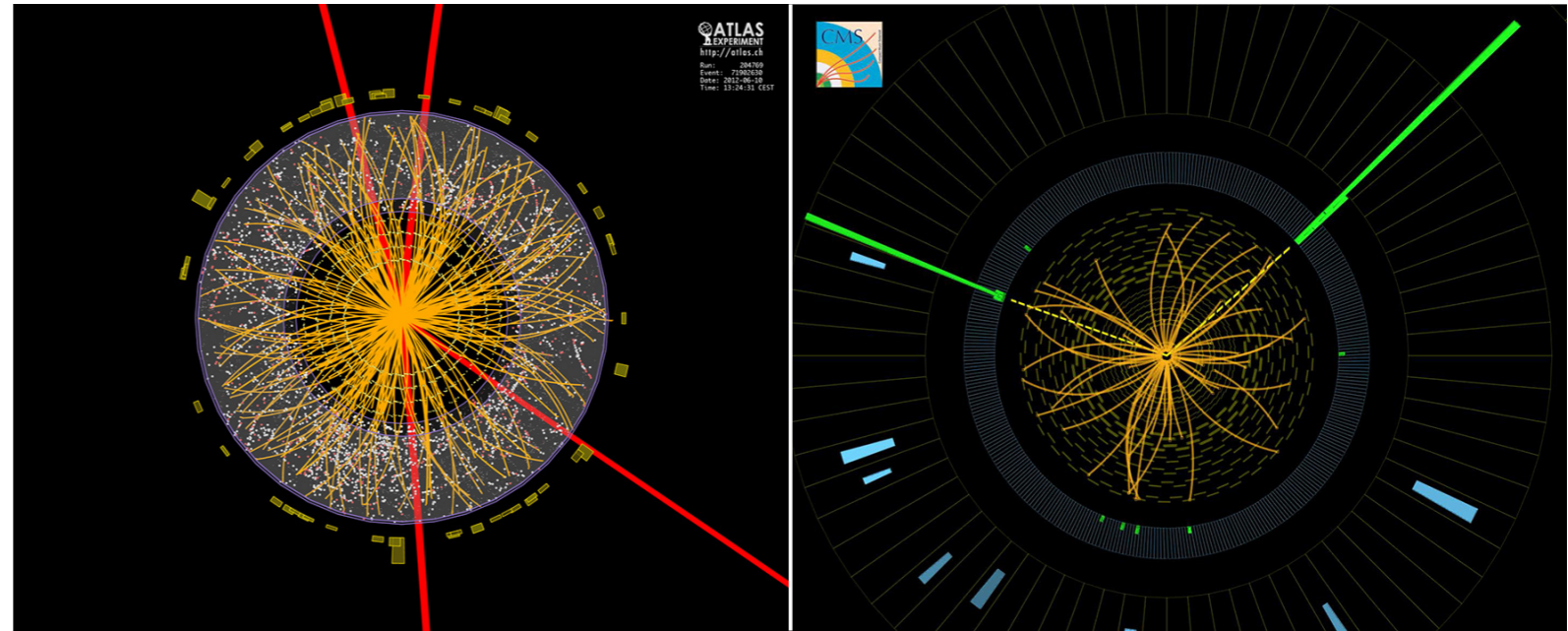


Higgs: what we have learned And next steps

LianTao Wang
University of Chicago

Lepton-Photon 2017, Guang Zhou Aug. 7 2017

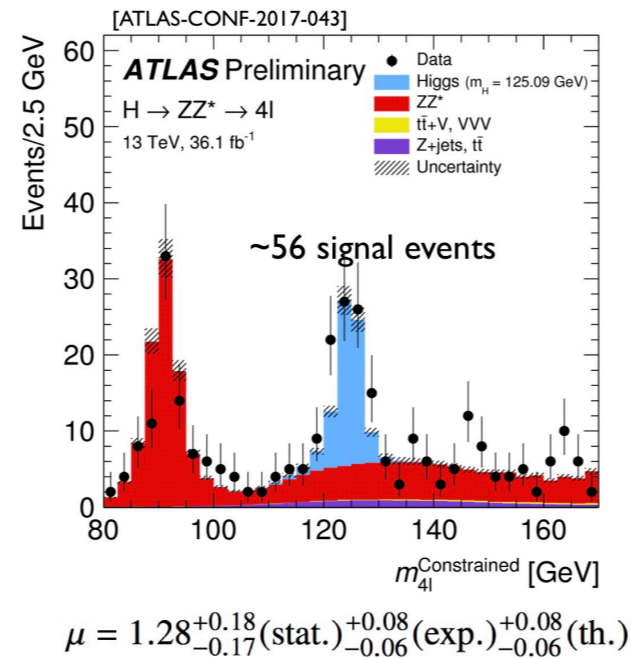
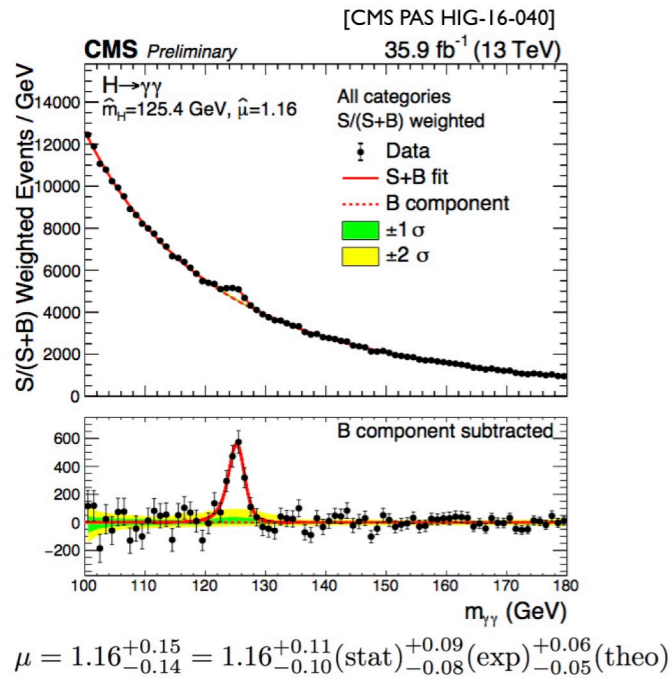
After the discovery



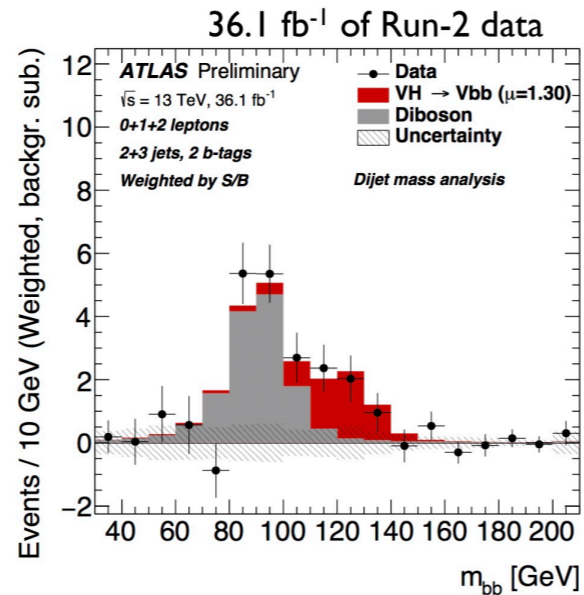
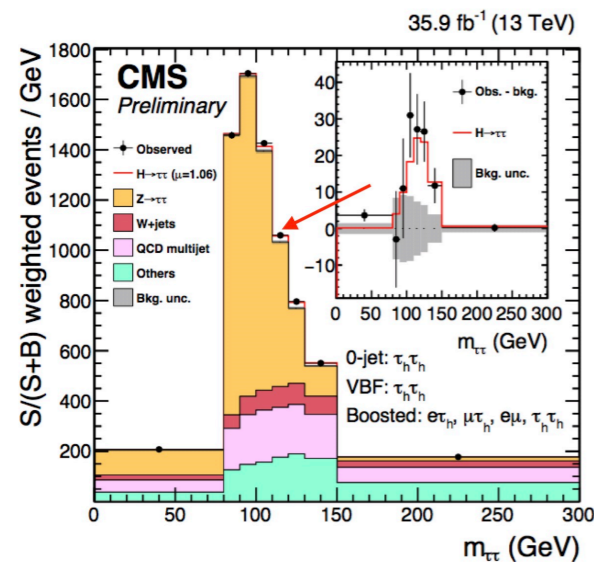
We have made significant progresses.

There is still a long way to go to understand the Higgs.
LHC can't finish the job, but it can do a lot.

Behaving like a Higgs boson



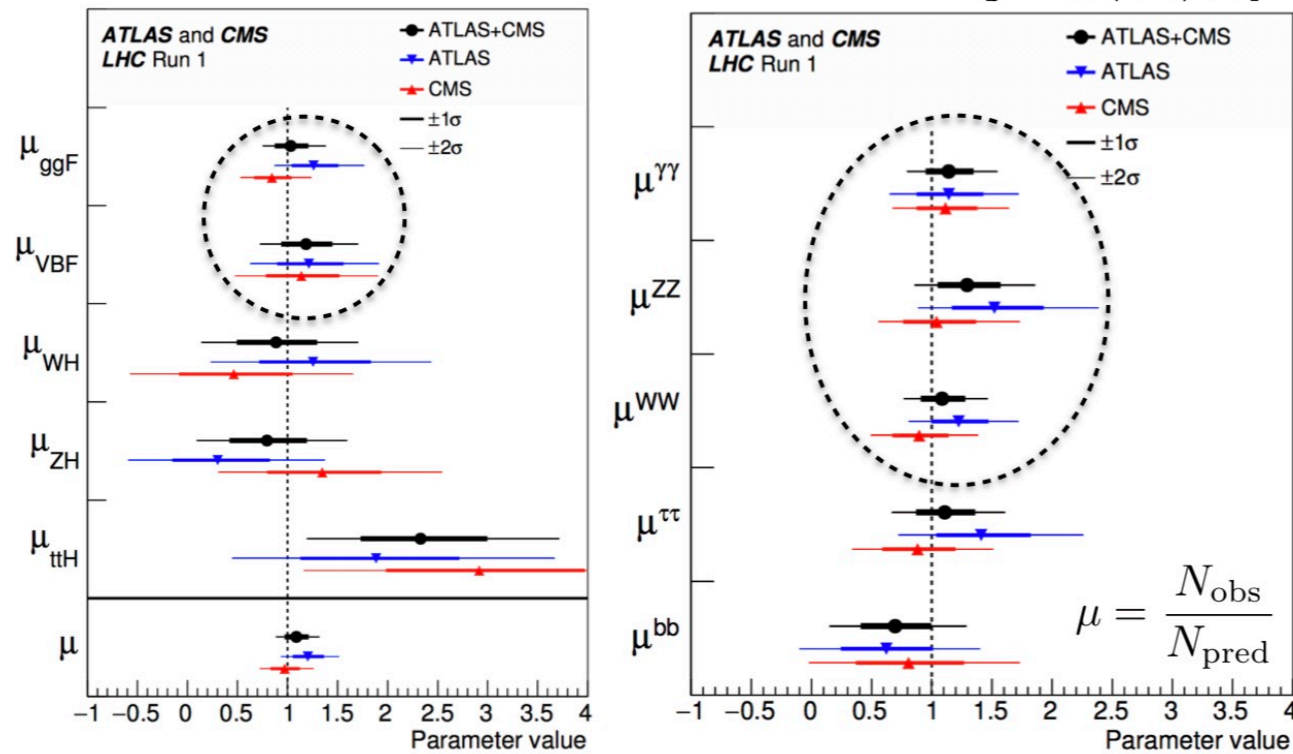
Higgs gauge boson coupling well established.



Started to see Higgs fermion coupling as well.

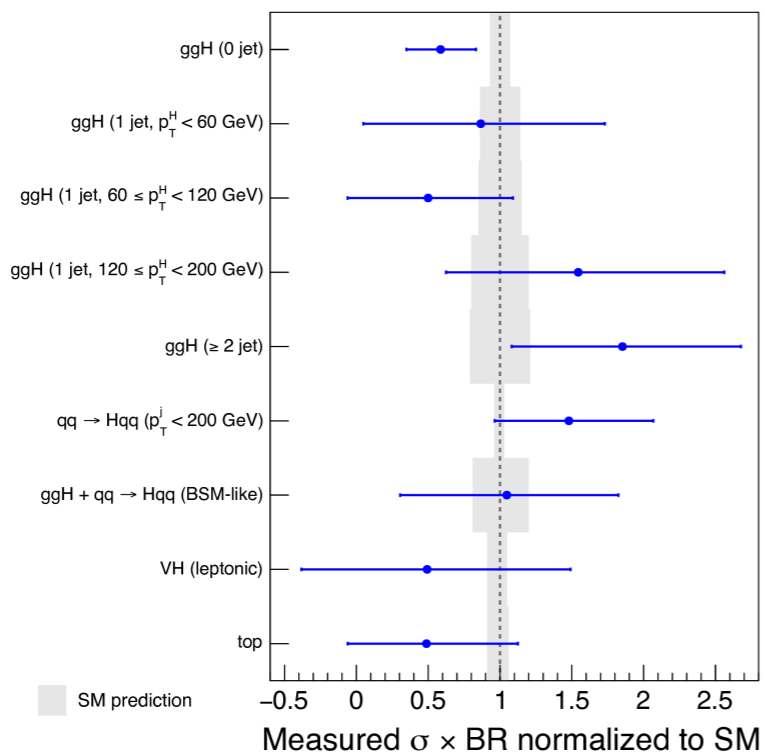
Roughly agree with Standard Model

[JHEP 08 (2016) 045]

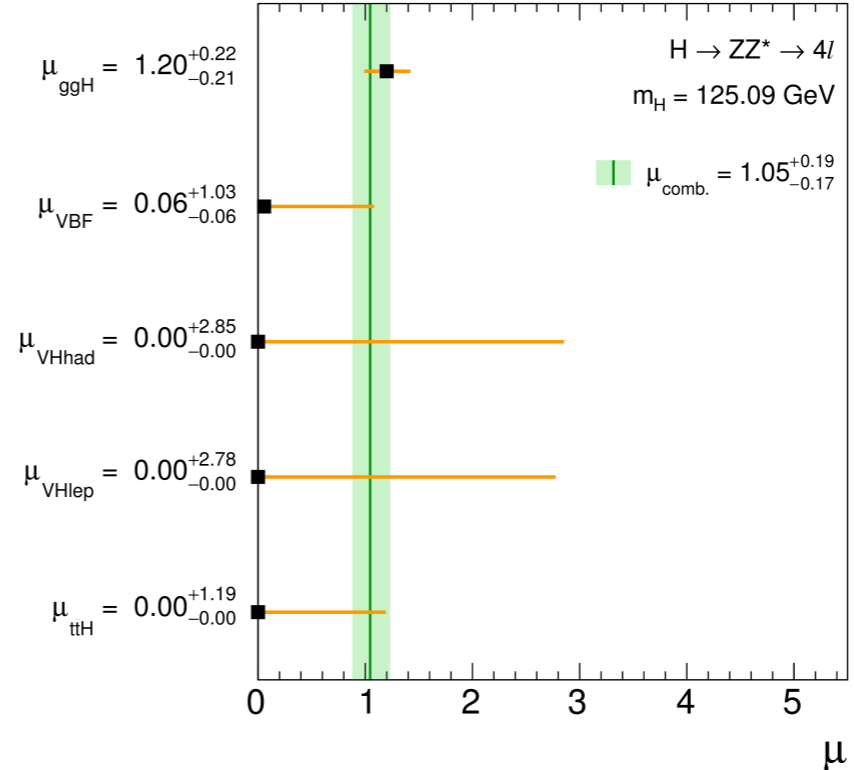


Agree to about
10-20%

ATLAS Preliminary $\sqrt{s}=13$ TeV, 36.1 fb^{-1}
 $H \rightarrow \gamma\gamma$, $m_H = 125.09$ GeV



CMS Preliminary 35.9 fb^{-1} (13 TeV)



Not entirely surprising

- In general, deviation induced by new physics is of the form

$$\delta \simeq c \frac{v^2}{M_{\text{NP}}^2}$$

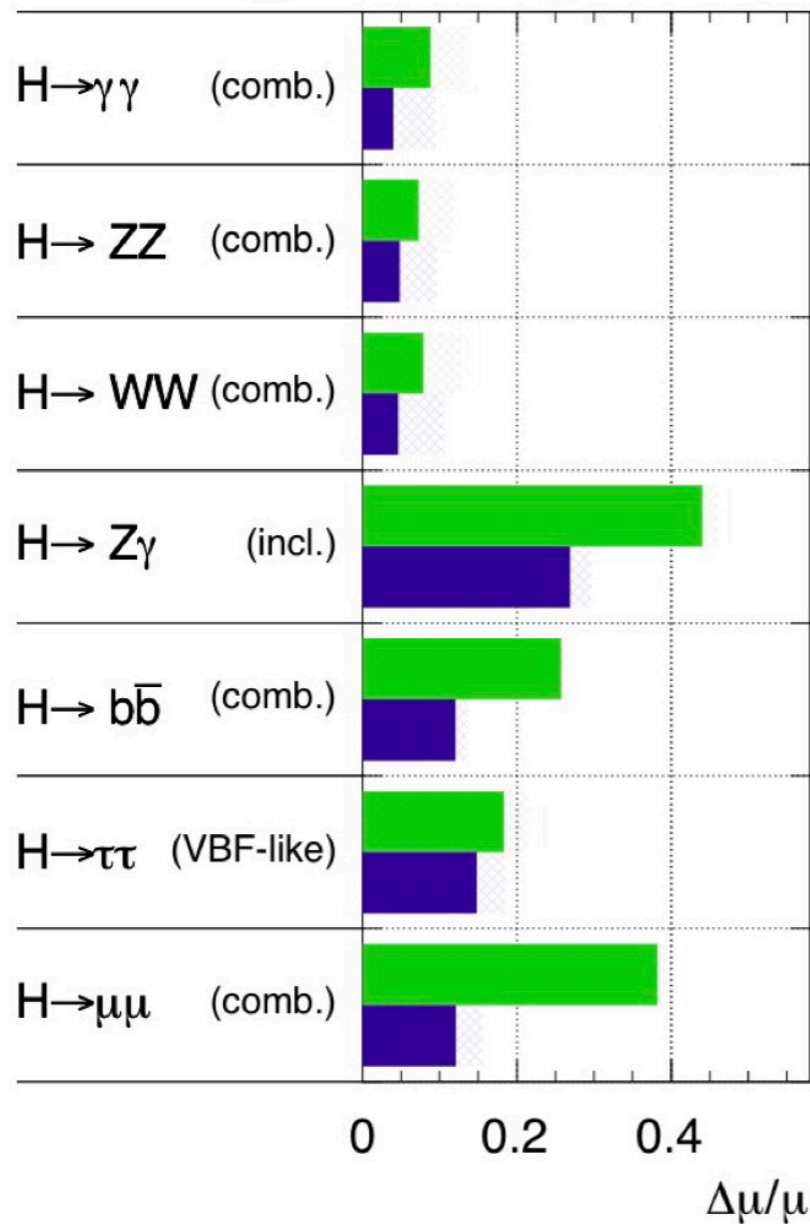
M_{NP} : mass of new physics
 c : $O(1)$ coefficient

- ▶ Current LHC precision: 10%
⇒ sensitive to $M_{\text{NP}} < 500\text{--}700$ GeV
- ▶ At the same time, direct searches constrain new physics below TeV already.
- ▶ **Unlikely to see $O(1)$ deviation.**

LHC entering precision measurement stage

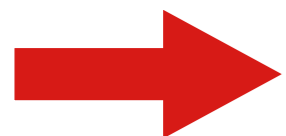
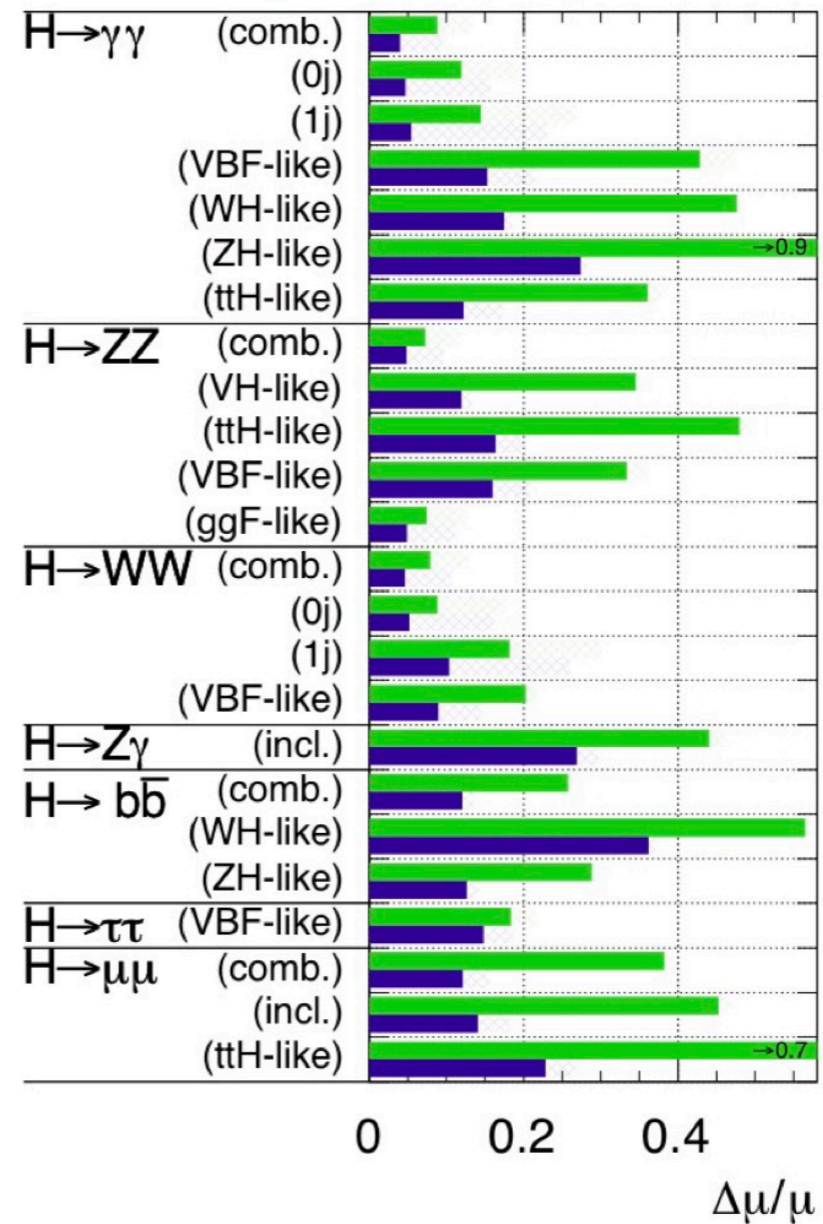
ATLAS Simulation Preliminary

$\sqrt{s} = 14$ TeV: $\int L dt = 300 \text{ fb}^{-1}$; $\int L dt = 3000 \text{ fb}^{-1}$



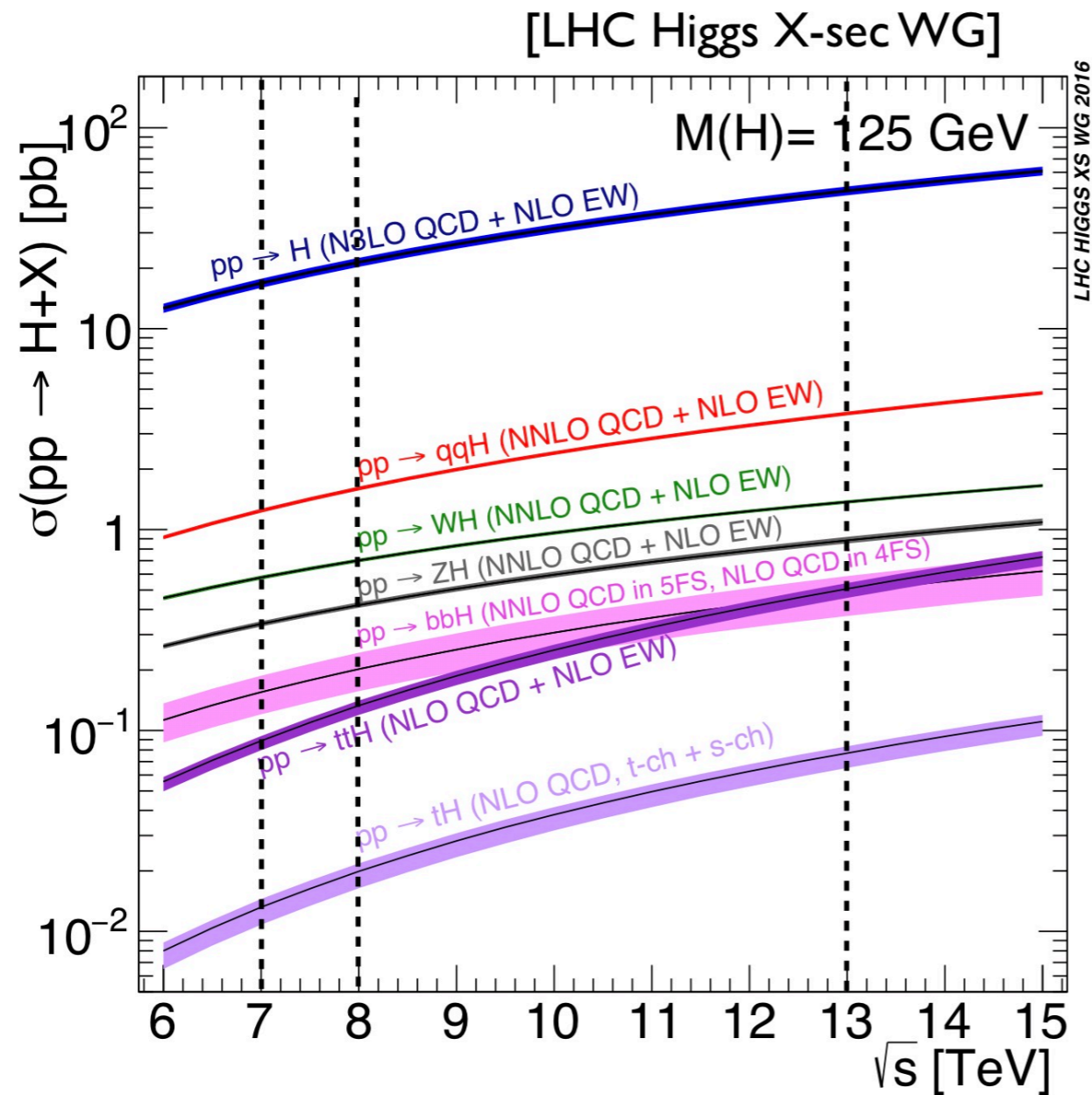
ATLAS Simulation Preliminary

$\sqrt{s} = 14$ TeV: $\int L dt = 300 \text{ fb}^{-1}$; $\int L dt = 3000 \text{ fb}^{-1}$



4-5% on Higgs coupling, reach TeV new physics

LHC as a Higgs factory



> 3 million Higgses
at run 2 already

100 times more by the
end of HL-LHC

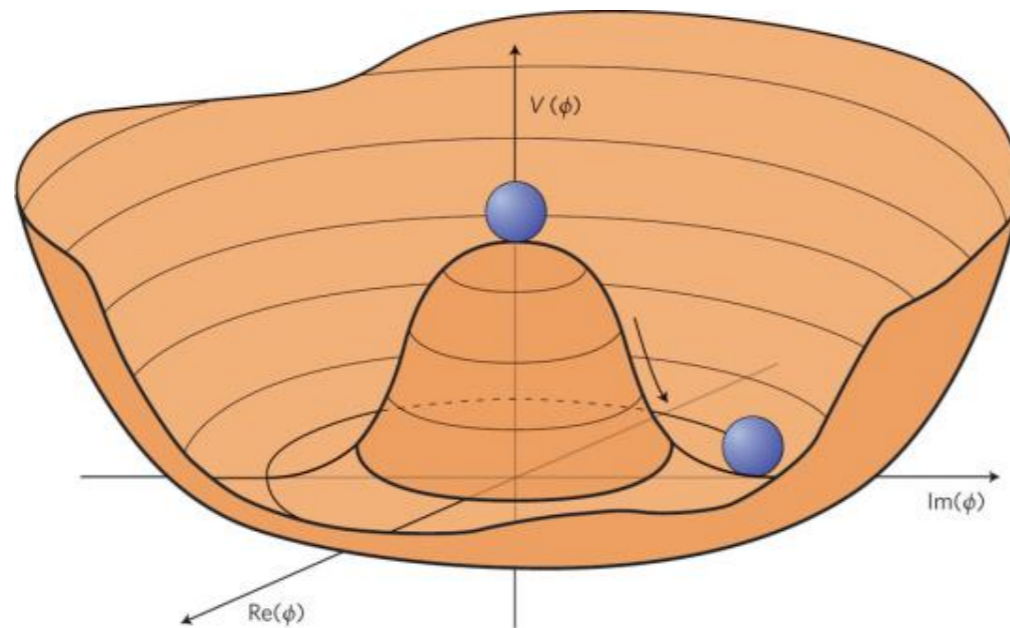
Great for clean yet very rare decay channels!

e.g. multiple leptons, displaced, etc.

Potentially 10^{-7} sensitivity on BR.

Questions to be addressed
by Higgs measurement

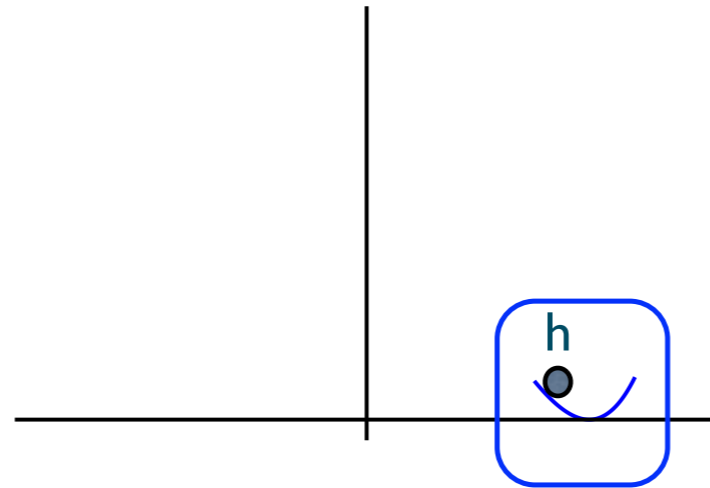
Mysteries of the electroweak scale.



$$V(h) = \frac{1}{2}\mu^2 h^2 + \frac{\lambda}{4}h^4$$

$$\langle h \rangle \equiv v \neq 0 \rightarrow m_W = g_W \frac{v}{2}$$

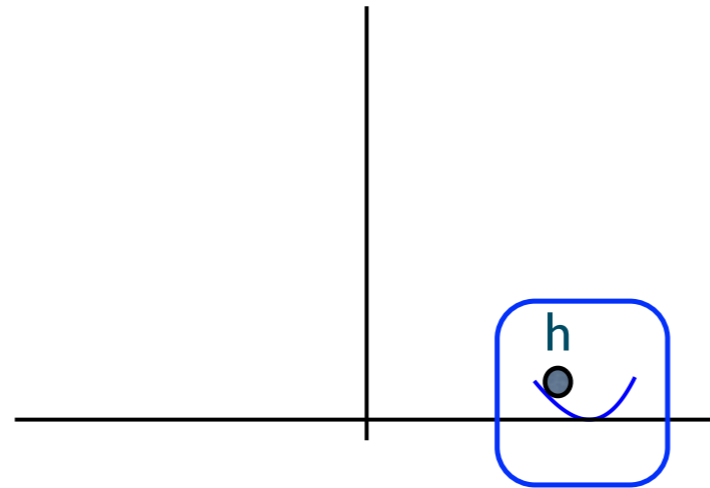
Mysteries of the electroweak scale.



What we know now

- How to predict/calculate Higgs mass?
- What does the rest of the Higgs potential look like? Nature of electroweak phase transition.

Mysteries of the electroweak scale.



What we know now

- How to predict/calculate Higgs mass?
- What does the rest of the Higgs potential look like? Nature of electroweak phase transition.

How to predict Higgs mass?

.....

The energy scale of new physics
responsible for EWSB

What is this energy scale?

$$M_{\text{Planck}} = 10^{19} \text{ GeV}, \dots?$$

If so, why is so different from 100 GeV?
The so called naturalness problem



Electroweak scale, 100 GeV.

$m_h, m_W \dots$

Naturalness of electroweak symmetry breaking



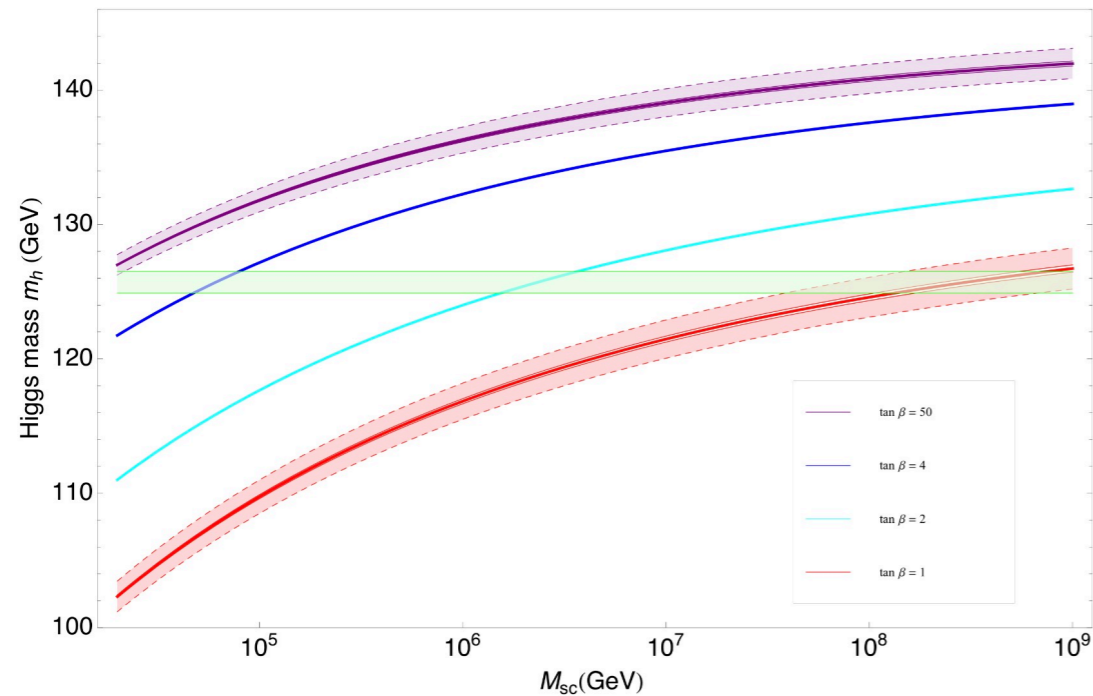
The energy scale of new physics
responsible for EWSB

TeV new physics.
Naturalness motivated
Many models, ideas.

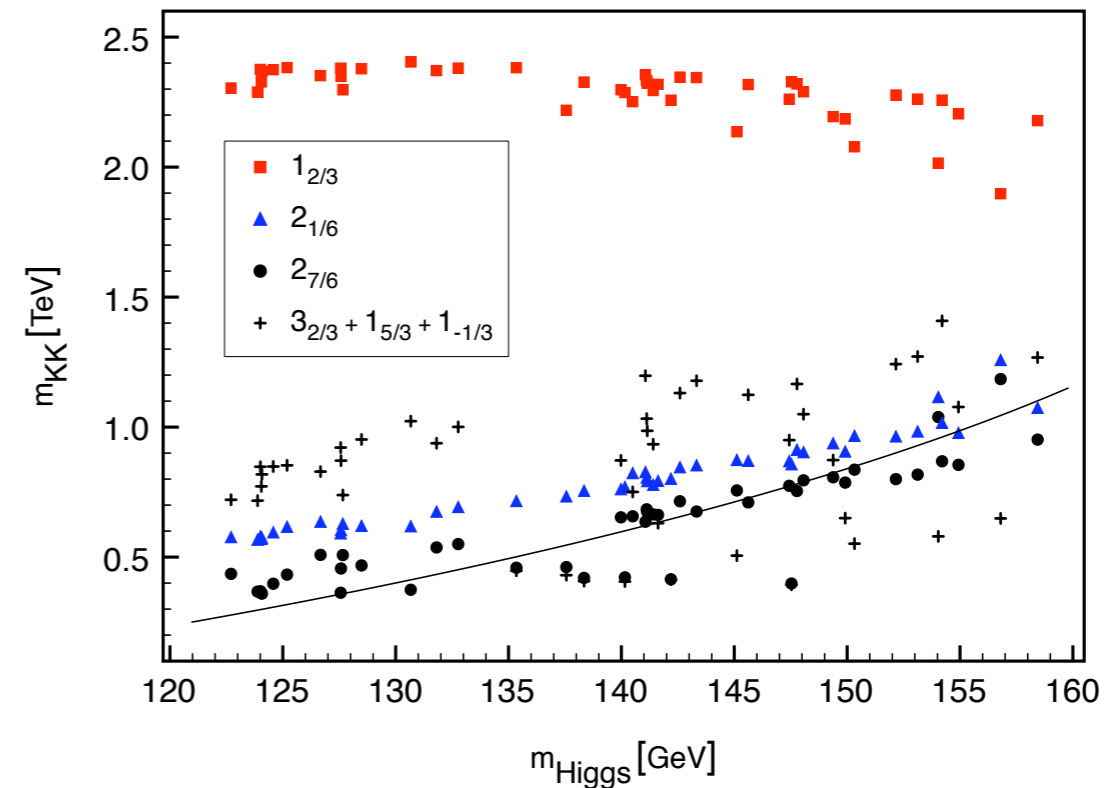


Electroweak scale, 100 GeV.
 $m_h, m_W \dots$

A confusing picture



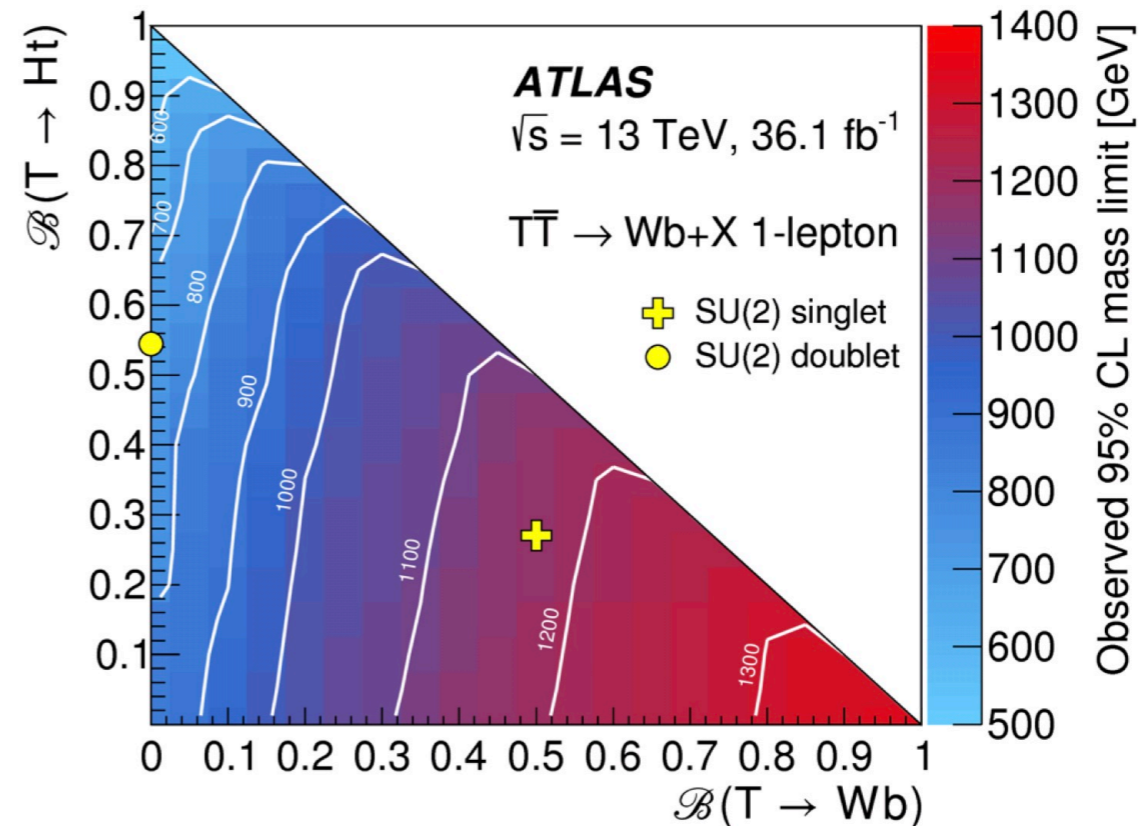
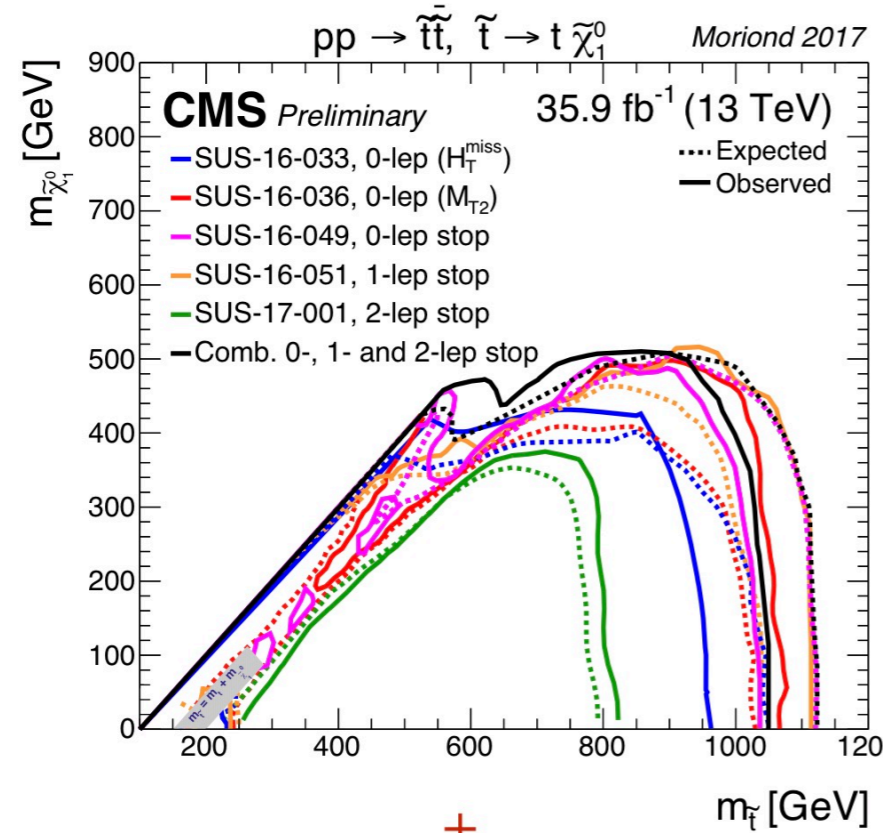
Stop too heavy to be natural



Composite top partner too light, excluded

Such conclusions too simplistic, "work around" available.
A bit uncomfortable, yes. Not time to give up just yet.

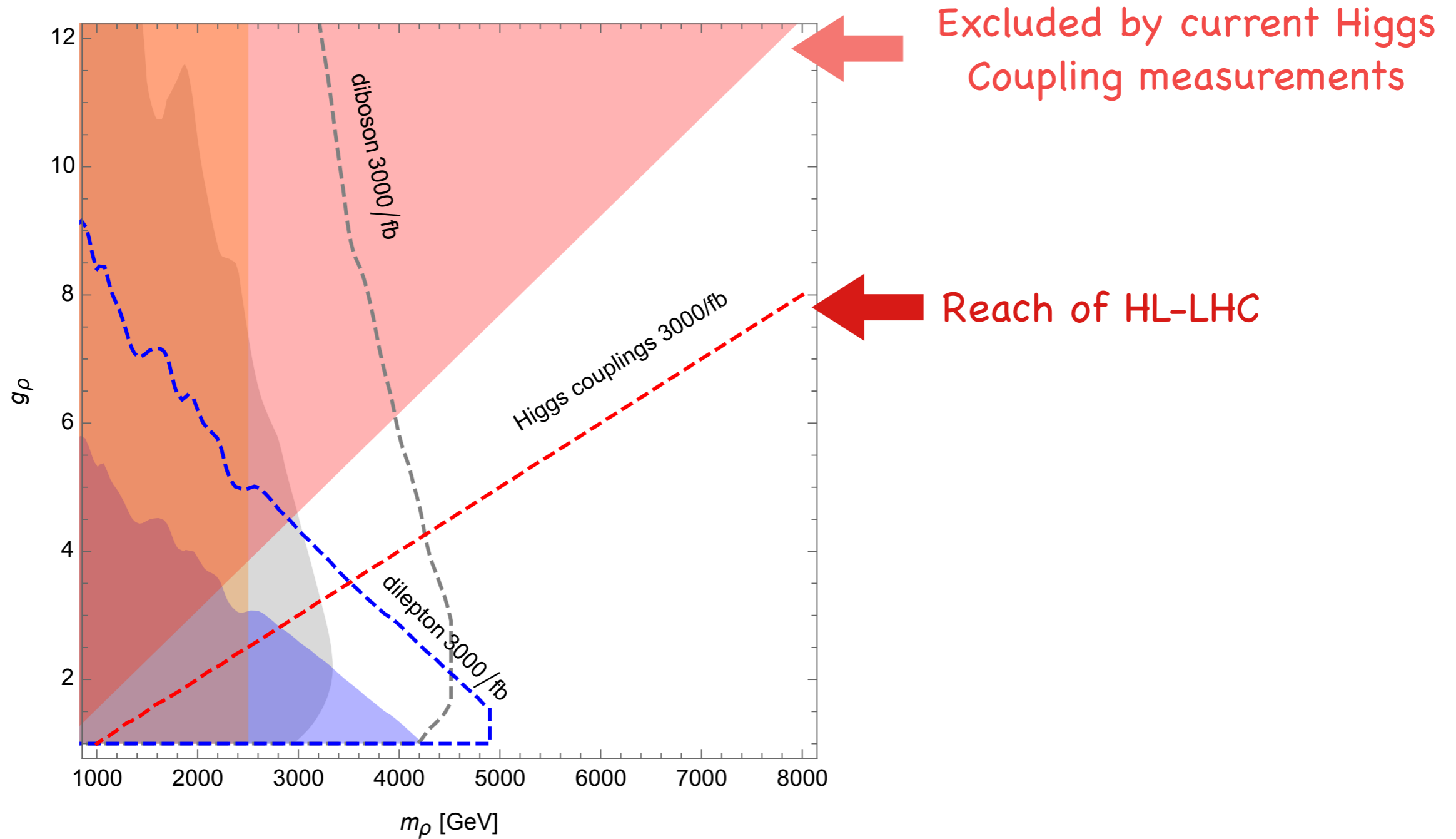
Direct searches



LHC will keep searching for such new particles

Future colliders, FCC-hh/SPPC, can continue the quest.

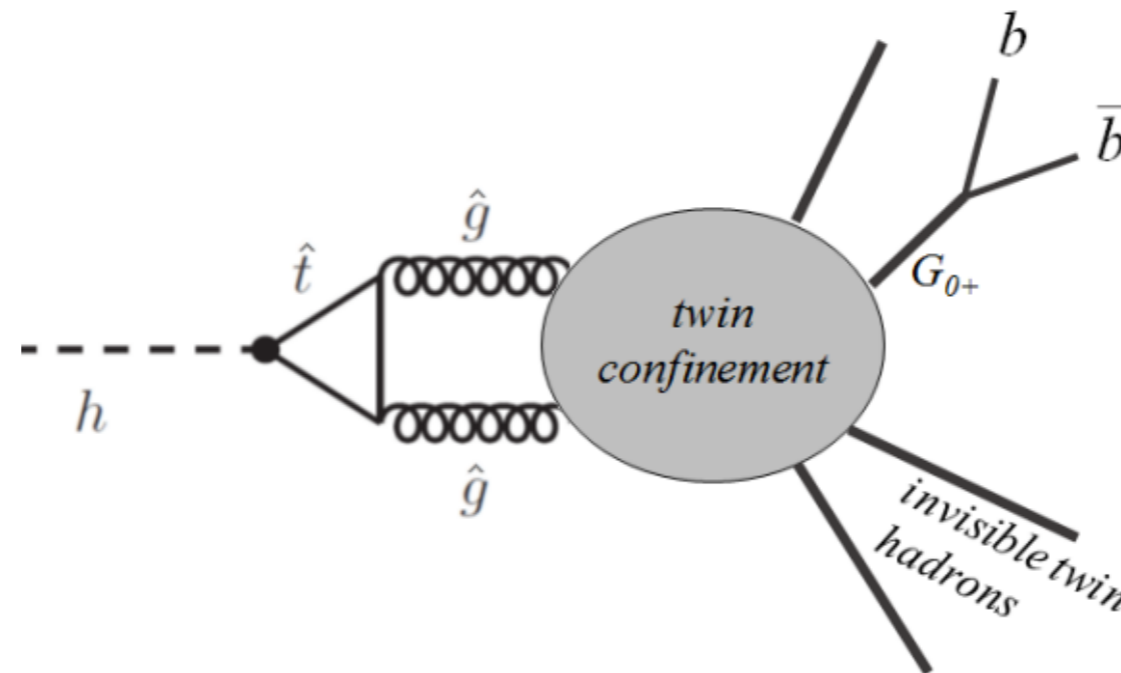
Higgs coupling vs direct search



Stealthy top partner. "twin"

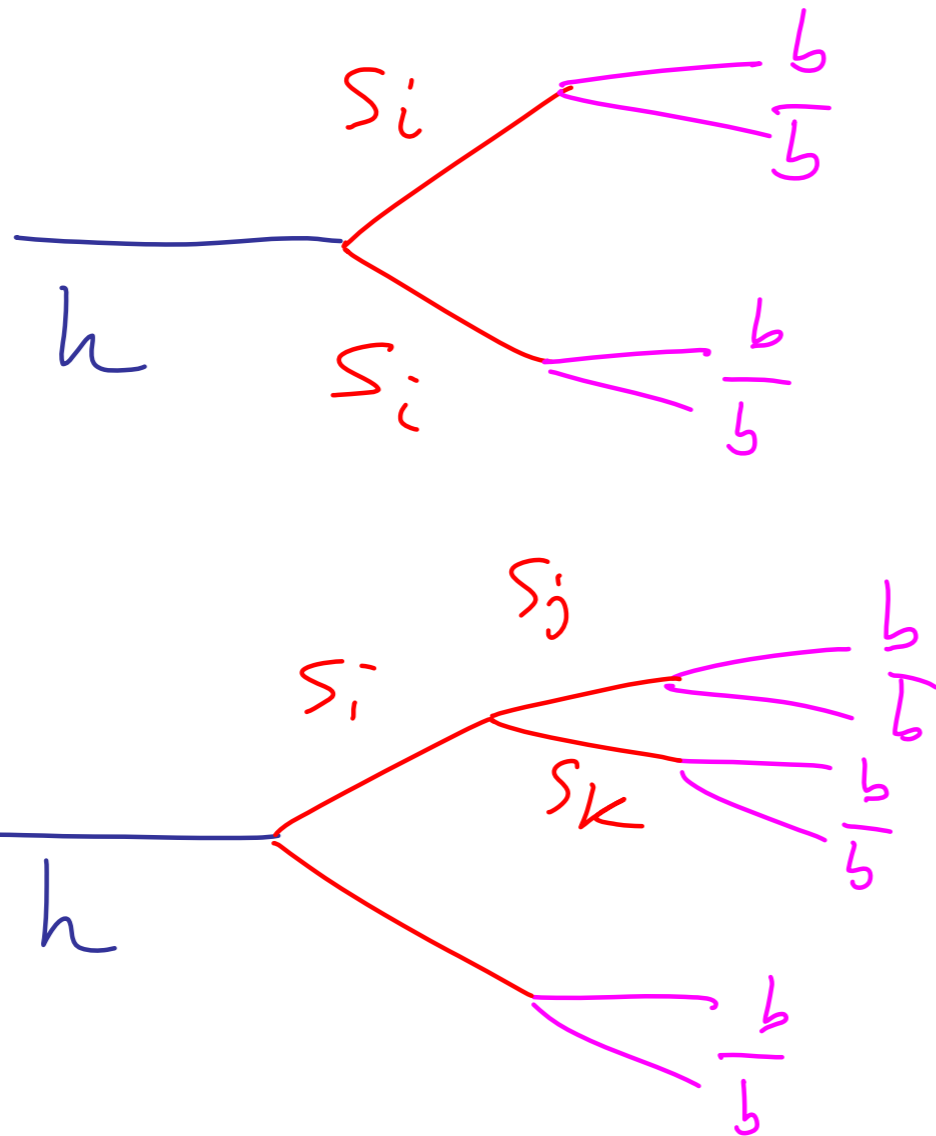
Chacko, Goh, Harnik

Craig, Katz, Strassler, Sundrum

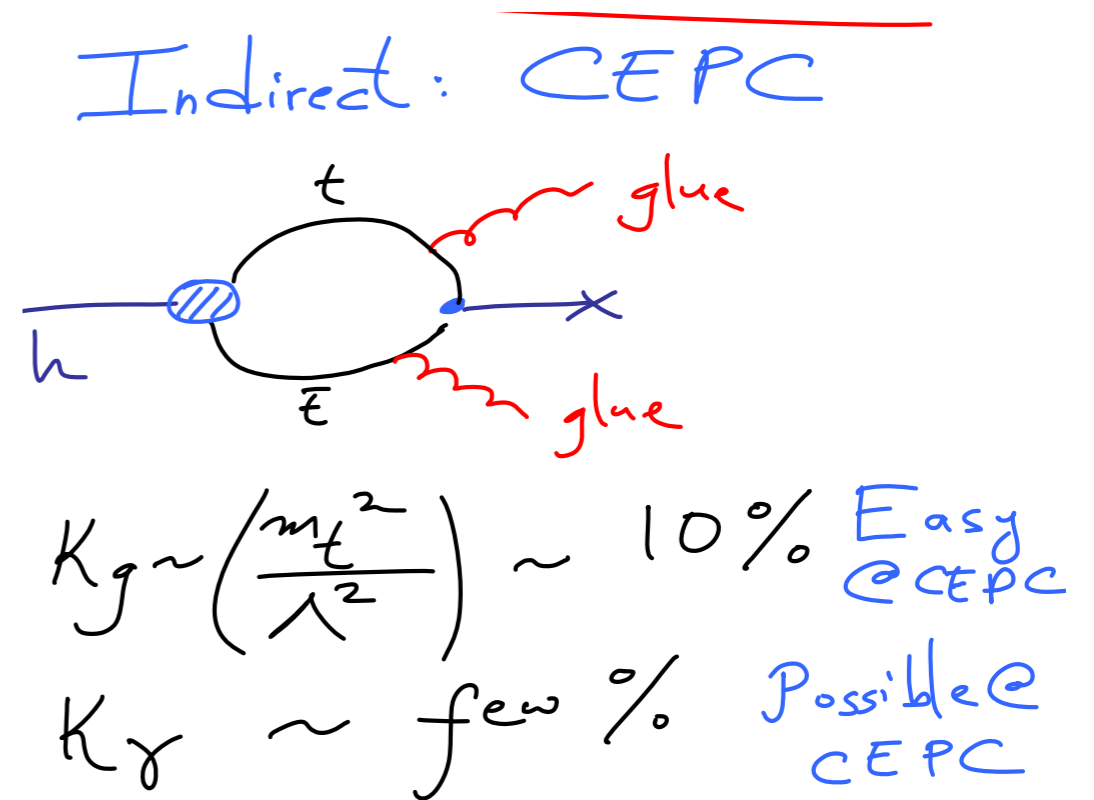


- Top partner not colored. Higgs decay through hidden world and back.
- Lead to Higgs rare decays.

More exotic ideas



Low scale landscape
Higgs rare decay.



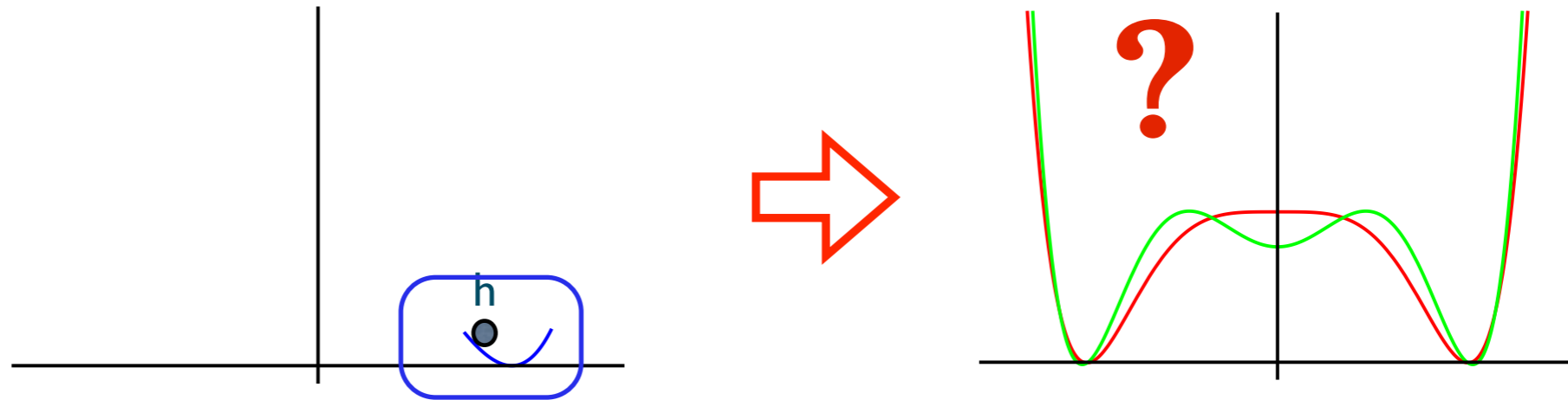
“fat” Higgs
Higgs coupling

Can't hide from the Higgs.

Bottom line on naturalness

- It is the most pressing question of EWSB.
 - ▶ How should we predict the Higgs mass?
- We have ideas, but maybe not the right one.
 - No confirmation of any of the proposed models.
 - Confusion is good for physics. Challenging the foundation of our understanding of Quantum Field Theory.
 - Need experimental guidance.
- Fortunately, with Higgs, we know where to look.
 - Clue to any possible way to address naturalness problem must show up in Higgs coupling measurement.

Nature of EW phase transition



What we know from LHC
LHC upgrades won't go much further

“wiggles” in Higgs potential

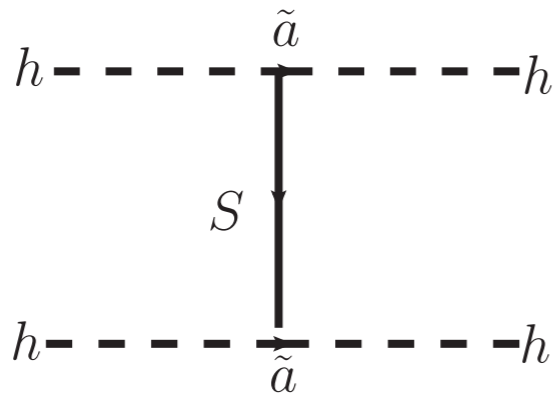
Big difference in triple Higgs coupling

Triple Higgs coupling measurement

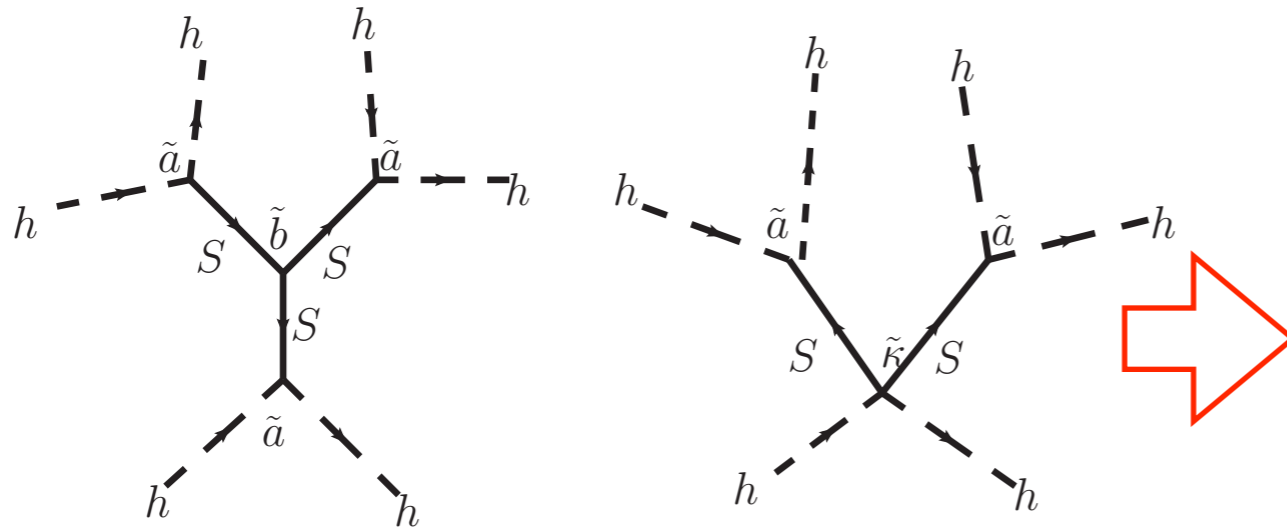
- Very difficult at HL-LHC: “order 1”
- 100 TeV pp collider or 1 TeV ILC can reach about 10%.
- However, if new physics modifies electroweak phase transition, it will also generate corrections to other Higgs couplings.
 - ▶ e.g. Generating sizable deviations in Higgs-Z coupling.

For example

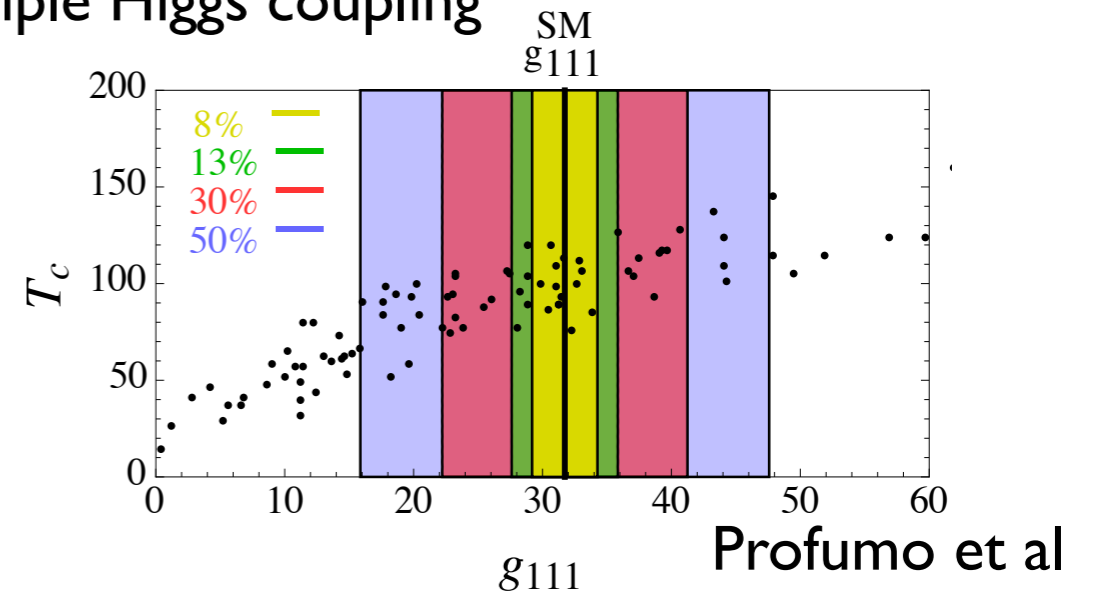
$$m^2 h^\dagger h + \tilde{\lambda} (h^\dagger h)^2 + m_S^2 S^2 + \tilde{a} S h^\dagger h + \tilde{b} S^3 + \tilde{\kappa} S^2 h^\dagger h + \tilde{h} S^4$$



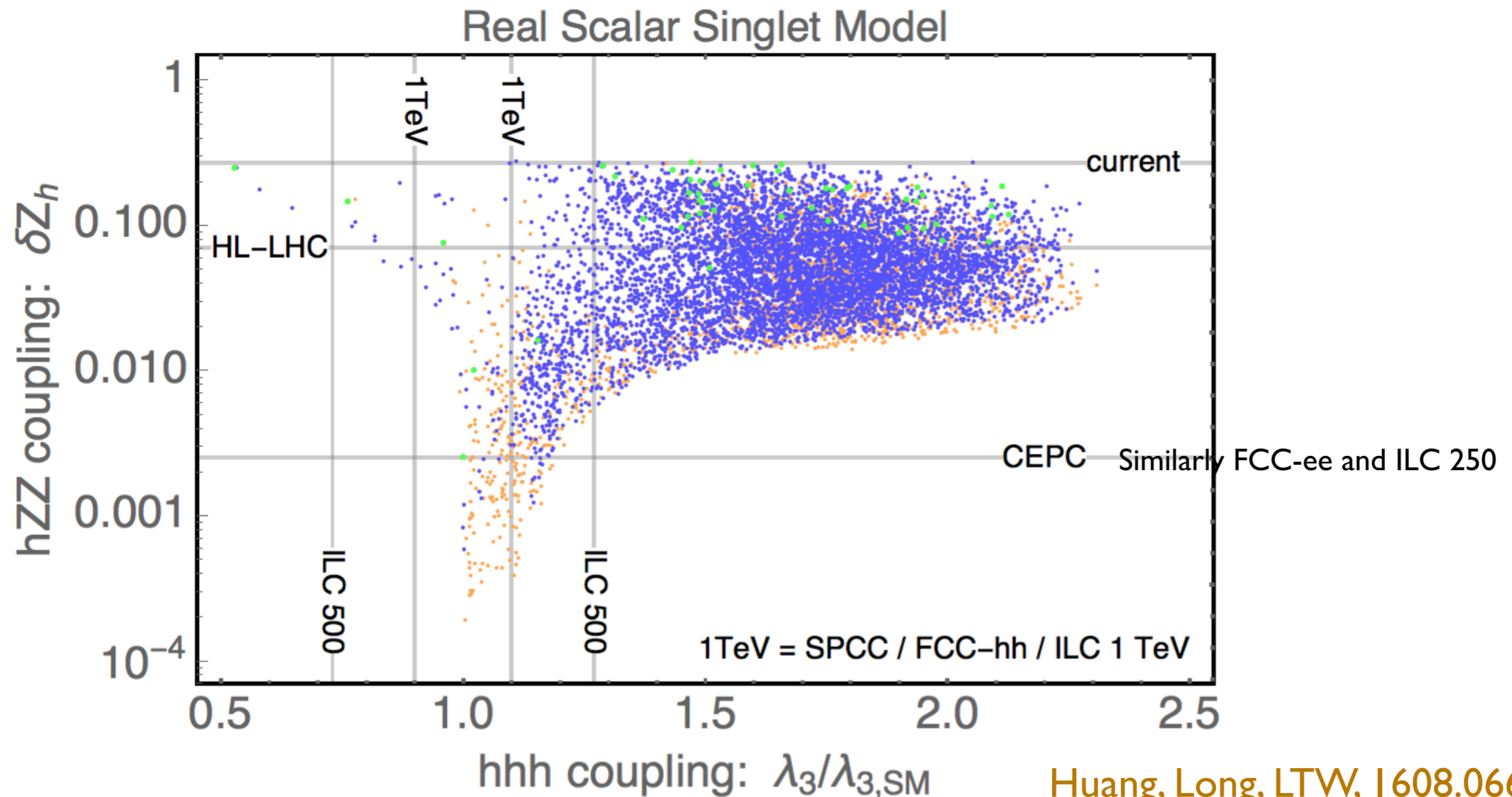
shift in h-Z coupling



triple Higgs coupling

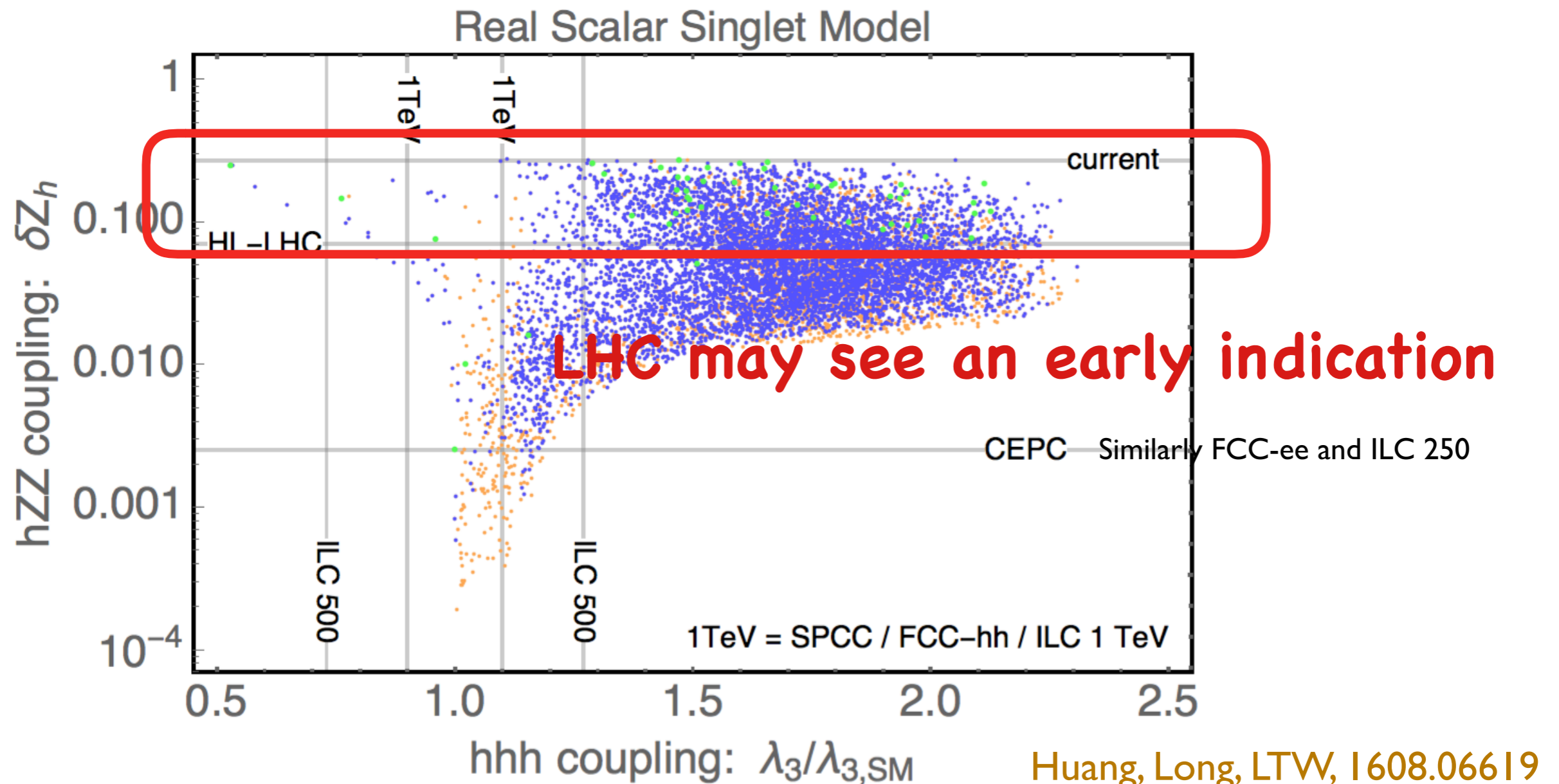


Probing nature of EW phase transition



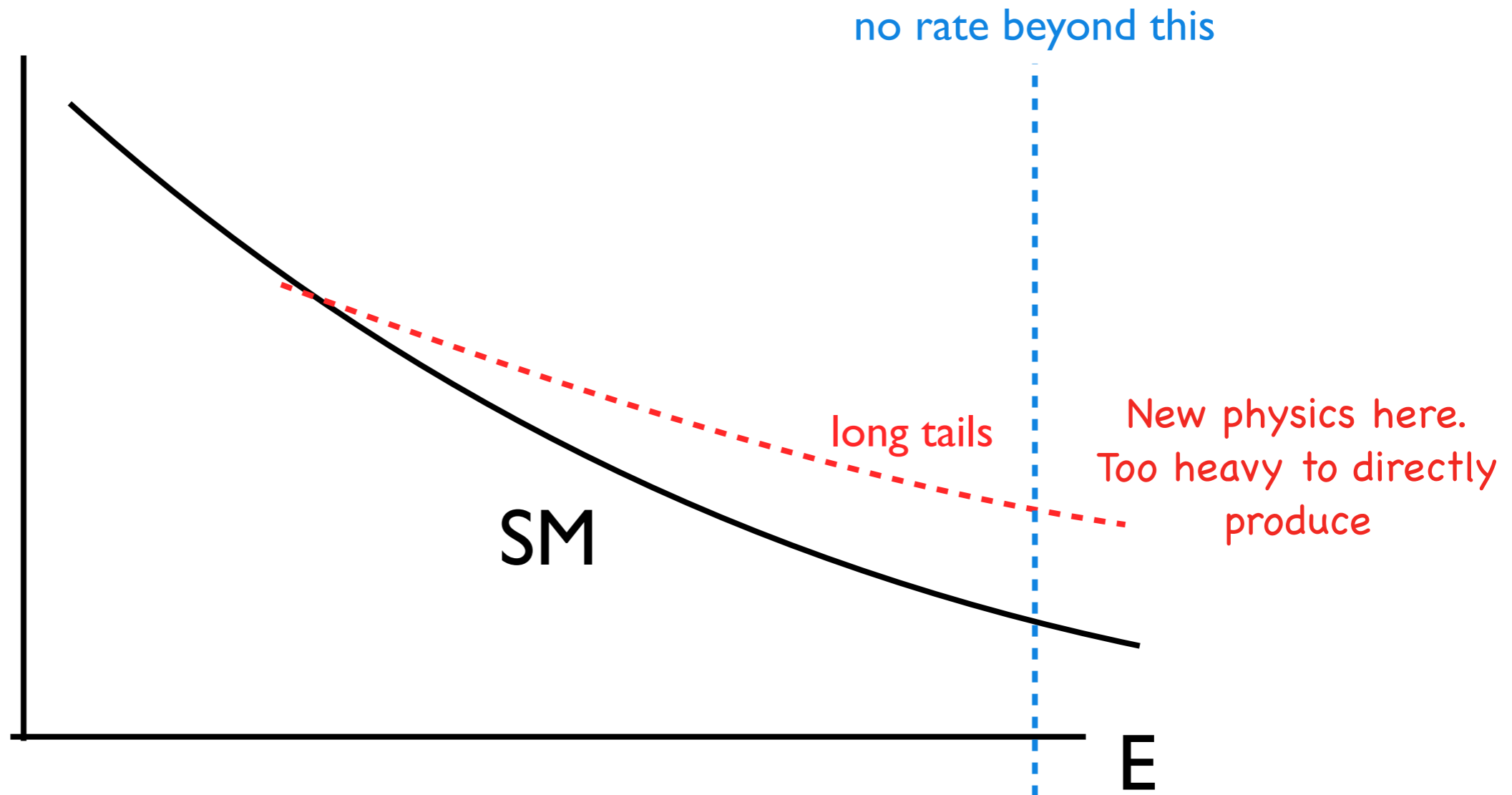
Orange = first order phase transition, $v(T_c)/T_c > 0$
Blue = “strongly” first order phase transition, $v(T_c)/T_c > 1.3$
Green = very strongly 1PT, could detect GWs at eLISA

Probing nature of EW phase transition



Beyond Higgs coupling
measurements.

Broad features with di-boson.



- Closely related to electroweak symmetry breaking
- Difficult. Systematics important.

Effect captured by:

$$\begin{aligned}
 \mathcal{O}_W &= \frac{ig}{2} \left(H^\dagger \sigma^a \overleftrightarrow{D}^\mu H \right) D^\nu W_{\mu\nu}^a, & \mathcal{O}_B &= \frac{ig'}{2} \left(H^\dagger \overleftrightarrow{D}^\mu H \right) \partial^\nu B_{\mu\nu} \\
 \mathcal{O}_{HW} &= ig (D^\mu H)^\dagger \sigma^a (D^\nu H) W_{\mu\nu}^a, & \mathcal{O}_{HB} &= ig' (D^\mu H)^\dagger (D^\nu H) B_{\mu\nu} \\
 \mathcal{O}_{3W} &= \frac{1}{3!} g \epsilon_{abc} W_\mu^{a\nu} W_{\nu\rho}^b W^{c\rho\mu}, & \mathcal{O}_T &= \frac{g^2}{2} (H^\dagger \overleftrightarrow{D}^\mu H) (H^\dagger \overleftrightarrow{D}_\mu H) H \\
 \mathcal{O}_R^u &= ig^2 \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \bar{u}_R \gamma^\mu u_R, & \mathcal{O}_R^d &= ig^2 \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \bar{d}_R \gamma^\mu d_R \\
 \mathcal{O}_L^q &= ig^2 \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \bar{Q}_L \gamma^\mu Q_L, & \mathcal{O}_L^{(3)q} &= ig^2 \left(H^\dagger \sigma^a \overleftrightarrow{D}_\mu H \right) \bar{Q}_L \sigma^a \gamma^\mu Q_L
 \end{aligned}$$

dim 6

$$\begin{aligned}
 {}_8\mathcal{O}_{TWW} &= g^2 \mathcal{T}_f^{\mu\nu} W_{\mu\rho}^a W_\nu^{a\rho} & {}_8\mathcal{O}_{TBB} &= g'^2 \mathcal{T}_f^{\mu\nu} B_{\mu\rho} B_\nu^\rho \\
 {}_8\mathcal{O}_{TWB} &= gg' \mathcal{T}_f^{a\mu\nu} W_{\mu\rho}^a B_\nu^\rho, & {}_8\mathcal{O}_{TH} &= g^2 \mathcal{T}_f^{\mu\nu} D_\mu H^\dagger D_\nu H \\
 {}_8\mathcal{O}_{TH}^{(3)} &= g^2 \mathcal{T}_f^{a\mu\nu} D_\mu H^\dagger \sigma^a D_\nu H
 \end{aligned}$$

dim 8

$$\mathcal{T}_f^{\mu\nu} = \frac{i}{4} \bar{\psi} (\gamma^\mu \overleftrightarrow{D}^\nu + \gamma^\nu \overleftrightarrow{D}^\mu) \psi \qquad \mathcal{T}_f^{a,\mu\nu} = \frac{i}{4} \bar{\psi} (\gamma^\mu \overleftrightarrow{D}^\nu + \gamma^\nu \overleftrightarrow{D}^\mu) \sigma^a \psi$$

Precision measurement at the LHC possible?

LEP precision tests probe NP about $\Lambda \sim 2 \text{ TeV}$

$$\frac{\delta\sigma}{\sigma_{\text{SM}}} \sim \frac{m_W^2}{\Lambda^2} \sim 2 \times 10^{-3}$$

At LHC

Signal-SM interference

Without interference

$E \sim \text{TeV}$

$$\frac{\delta\sigma}{\sigma_{\text{SM}}} \sim \frac{E^2}{\Lambda^2} \sim 0.25$$

$$\frac{\delta\sigma}{\sigma_{\text{SM}}} \sim \frac{E^4}{\Lambda^4} \sim 0.05$$

LHC has potential.

Both interference and energy growing behavior crucial

Helicity structure at LHC

$$f_L \bar{f}_R \rightarrow W^+ W^-$$

(h_{W^+}, h_{W^-})	SM	\mathcal{O}_W	\mathcal{O}_{HW}	\mathcal{O}_{HB}	\mathcal{O}_B	\mathcal{O}_{3W}	\mathcal{O}_{TWW}
(\pm, \mp)	1	0	0	0	0	0	$\frac{E^4}{\Lambda^4}$
$(0, 0)$	1	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	0	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$
$(0, \pm), (\pm, 0)$	$\frac{m_W}{E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^4}{\Lambda^4} \frac{m_W}{E}$
(\pm, \pm)	$\frac{m_W^2}{E^2}$	$\frac{m_W^2}{\Lambda^2}$	$\frac{m_W^2}{\Lambda^2}$	0	0	$\frac{E^2}{\Lambda^2}$	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$

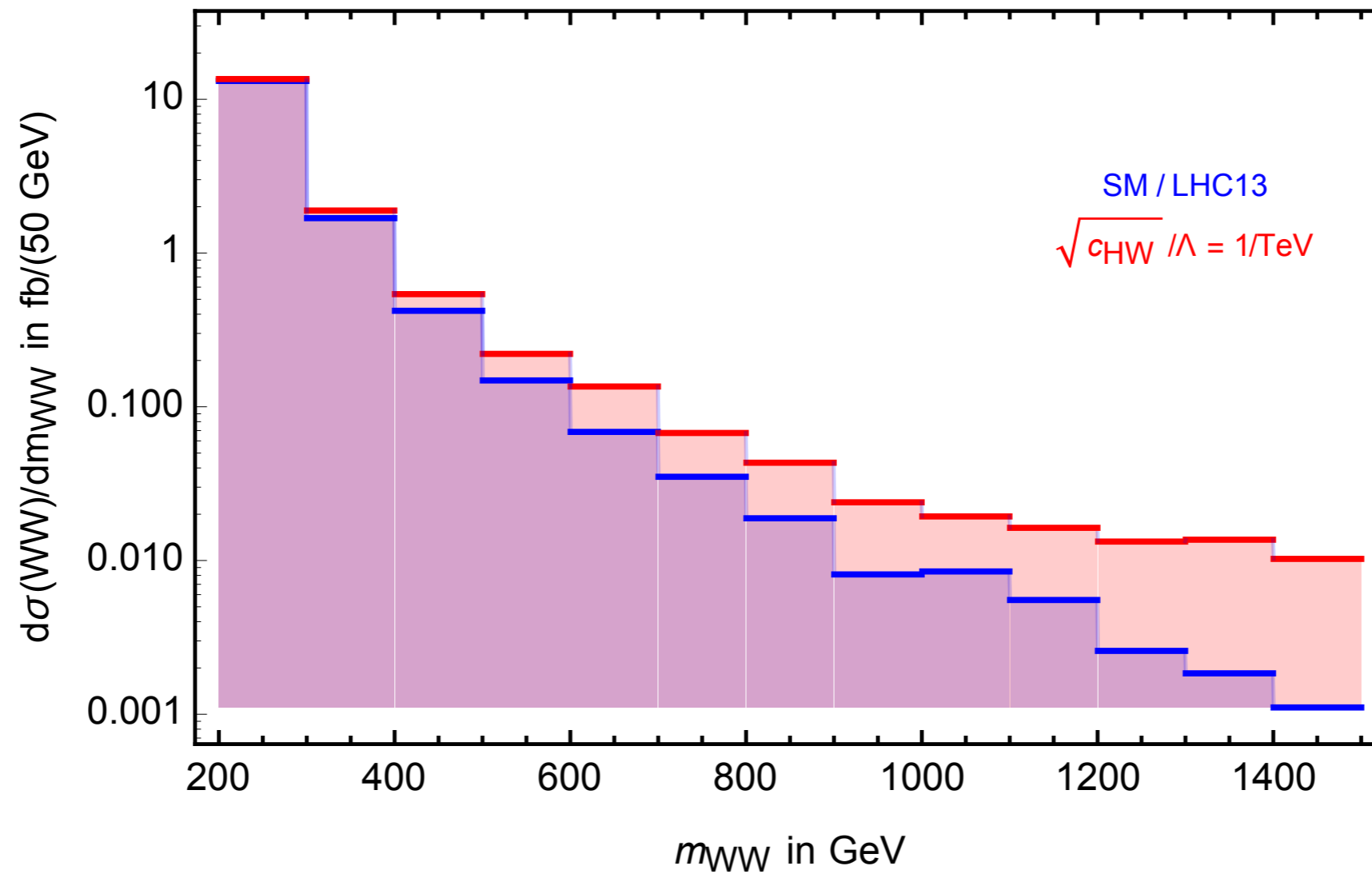
$$f_R \bar{f}_L \rightarrow W^+ W^-$$

 growing with energy

(h_{W^+}, h_{W^-})	SM	\mathcal{O}_W	\mathcal{O}_{HW}	\mathcal{O}_{HB}	\mathcal{O}_B	\mathcal{O}_{3W}	\mathcal{O}_{TWW}
(\pm, \mp)	0	0	0	0	0	0	$\frac{E^4}{\Lambda^4}$
$(0, 0)$	1	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	0	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$
$(0, \pm), (\pm, 0)$	$\frac{m_W}{E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{m_W^2 m_W}{\Lambda^2 E}$	$\frac{E^4}{\Lambda^4} \frac{m_W}{E}$
(\pm, \pm)	$\frac{m_W^2}{E^2}$	$\frac{m_W^2}{\Lambda^2}$	$\frac{m_W^2}{\Lambda^2}$	0	0	$\frac{m_W^2}{\Lambda^2}$	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$

- Whether interference or not depends on polarization of WW. Polarization differentiation can be crucial.
- Need large SM piece to interfere with. Longitudinal (0,0) most promising.

Growing with energy



Where to look?

“tail” parameterized by $\frac{\mathcal{O}}{\Lambda^d}$

$$\sigma_{\text{signal}} \propto \frac{1}{E^n} \left(\frac{E}{\Lambda} \right)^d \quad \sigma_{\text{SM}} \propto \frac{1}{E^n}$$

$\Lambda \approx$ scale of NP

E: energy bin of the measurement

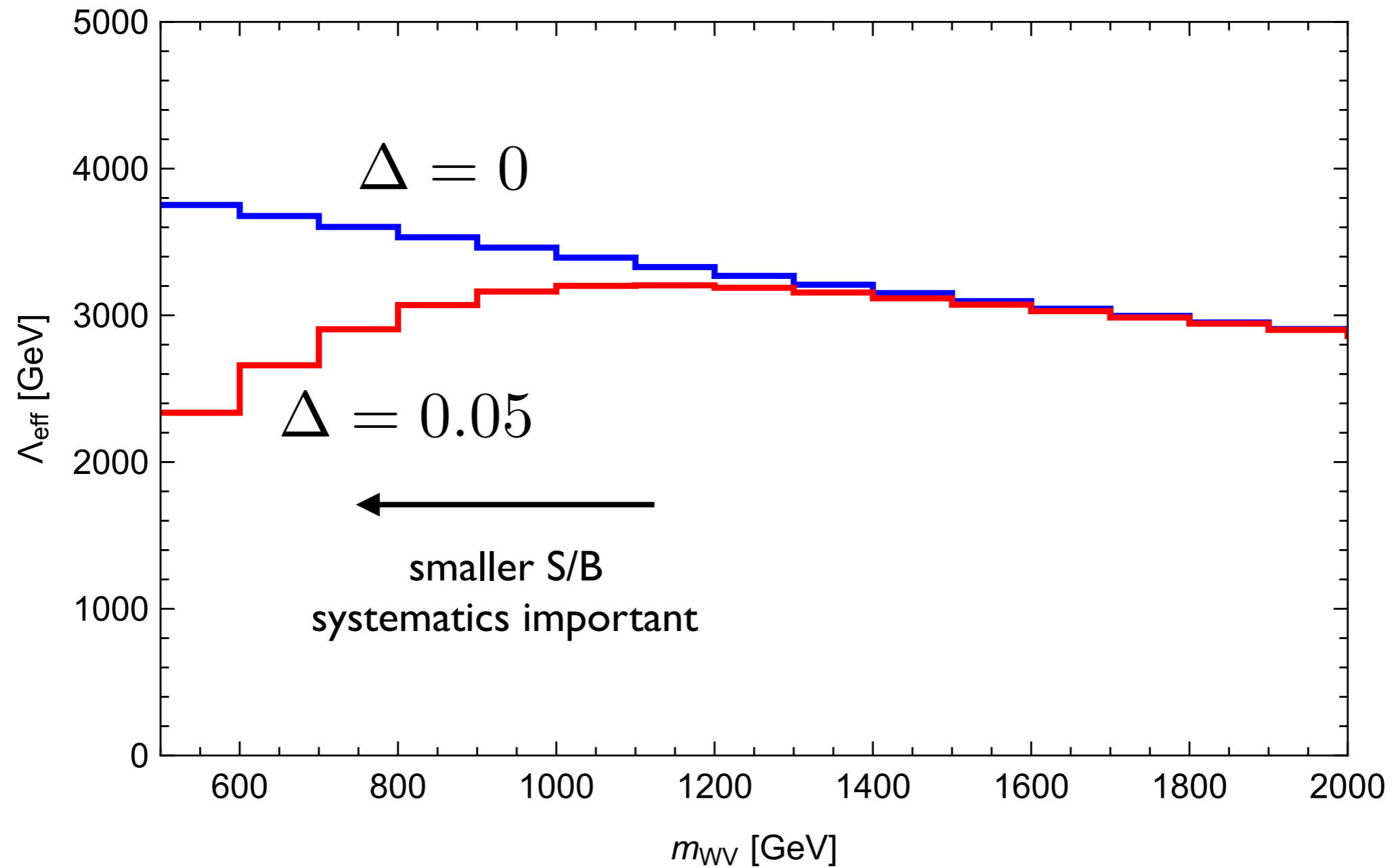
n: 5-8 falling parton luminosity

$$\frac{S}{\sqrt{B}} \sim \sqrt{\frac{\mathcal{L}}{E^n}} \left(\frac{E}{\Lambda} \right)^d$$

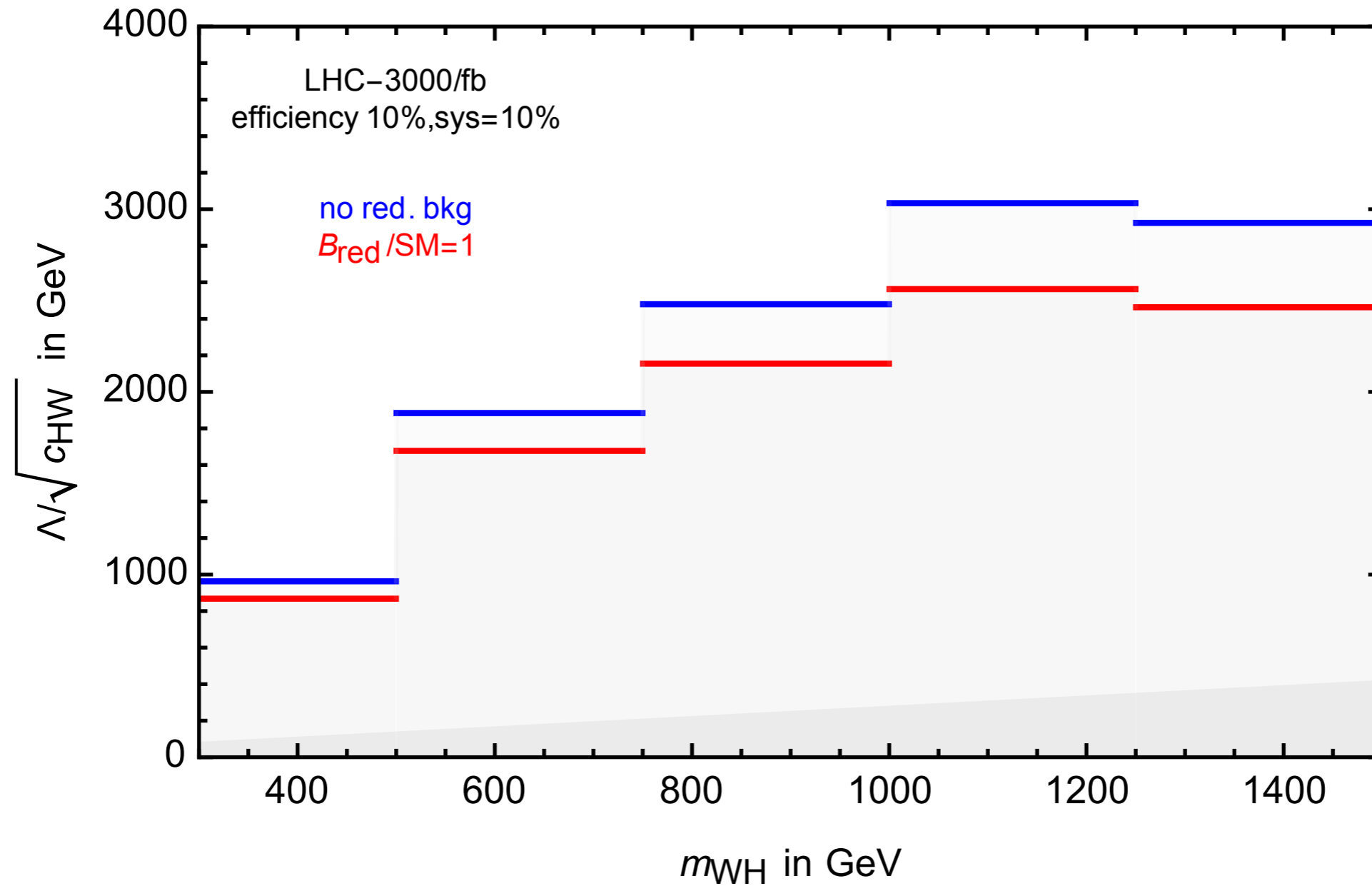
\mathcal{L} = integrated luminosity

- For small d, lower E with higher reach. (e.g. dim 6, d=2)
 - ▶ Limited by systematics.

The role of systematics



Projecting the reach: Wh channel



Can set interesting limit!

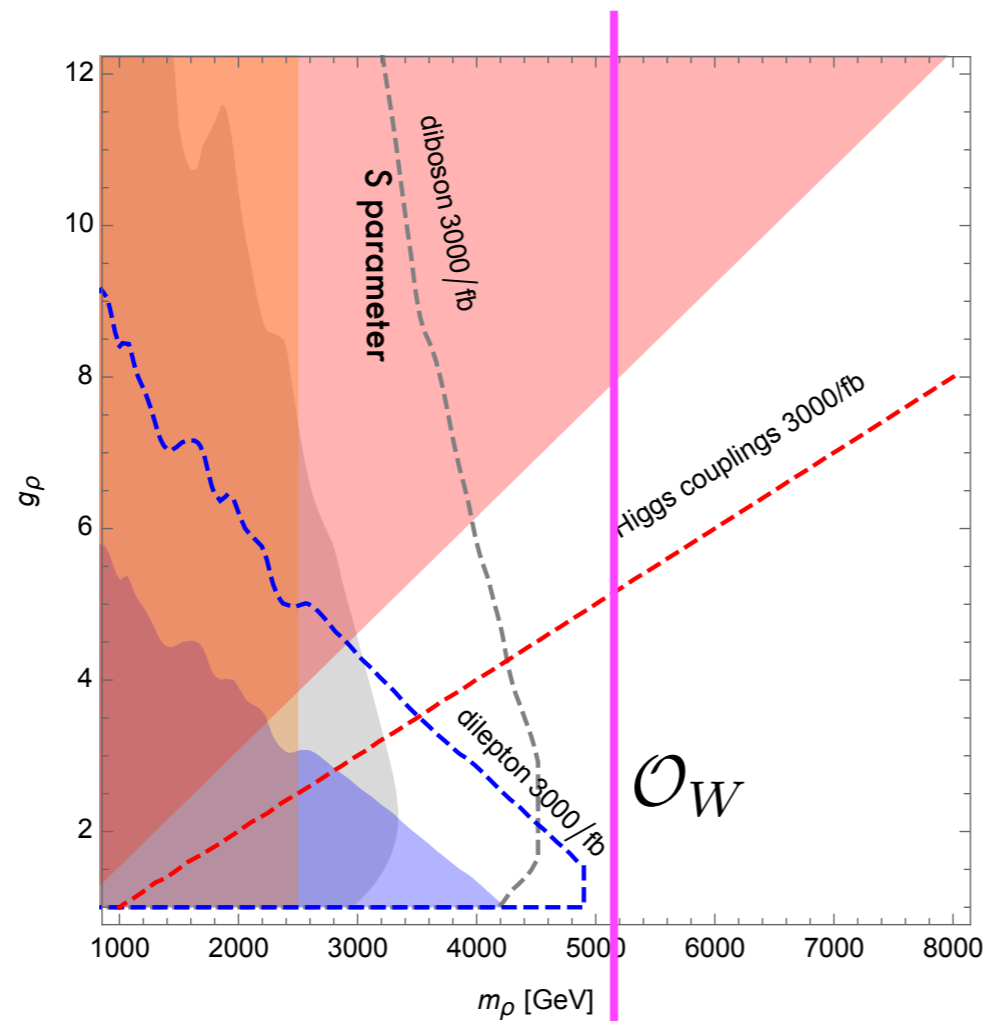
LHC benchmarks

Λ [TeV]	\mathcal{O}_W	\mathcal{O}_B	\mathcal{O}_{HW}	\mathcal{O}_{HB}	\mathcal{O}_{3W}
LEP	2.5	2.5	0.3	0.3	0.4
$WV(\ell + jets)$	4.8(1.9)	1.5(0.71)	4.8(1.9)	1.5(0.71)	1.2
$W^\pm h(\ell bb)$	(4.0,2.9,2.3)		(4.0,2.9,2.3)		
$W^\pm h(\ell + \ell\nu\nu)$	1.6		1.6		
$h \rightarrow Z\gamma$			1.7	1.7	

- ideal case, perfect pol tagging, no systematics
- tagging eff 50%, mis-tagging rate 10%, no systematics
- reducible bkg 0, 3, 10 times of the irreducible rate
- interference effect not important.

– Can beat LEP precision if some of these benchmarks can be reached.

Direct searches of composite resonance



Shaded areas:
current bounds

Most optimistic case can be competitive with direct narrow resonance searches.

The resonance may be broad, not covered by direct searches.

Conclusion

- Entering a precision era of the LHC, with Higgs the prime target.
- Understanding Higgs is a central question in particle physics.
 - ▶ Higgs mass, electroweak phase transition.
 - ▶ The current picture is confusing. Opportunity for big discoveries.
- LHC will lead the way, setting the stage for next steps.

Triple Higgs coupling at 100 TeV pp collider
30 ab⁻¹

$$\frac{\lambda}{\lambda_{\text{SM}}} \in \begin{cases} [0.891, 1.115] & \text{no background syst.} \\ [0.882, 1.126] & 25\% hh, 25\% hh + \text{jet} \\ [0.881, 1.128] & 25\% hh, 50\% hh + \text{jet} \end{cases}$$

Barr, Dolan, Englert, de Lima, Spannowsky

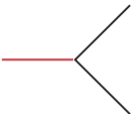
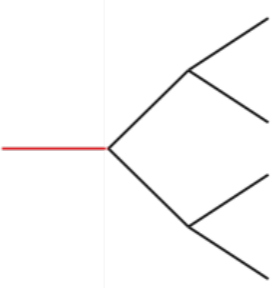
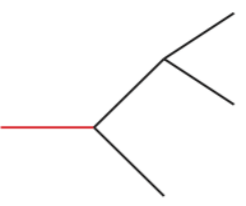
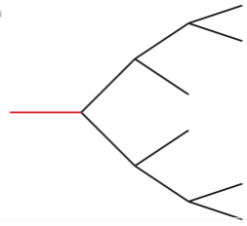
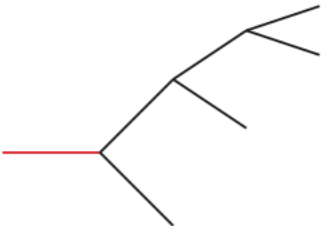
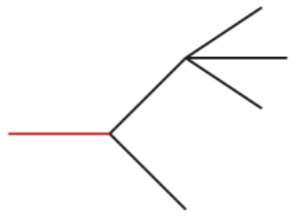
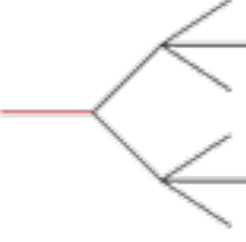
ILC 500: 27%
ILC ultimate, 1 TeV 5 ab⁻¹: 10%

But, there should be more

$$V(h) = \frac{m^2}{2}h^2 + \lambda h^4 + \frac{1}{\Lambda^2}h^6 + \dots$$

- 1st order EW phase transition means there is new physics close to the weak scale.
- Can be difficult to discover at the LHC.
 - ▶ Maybe only couple weakly to the Higgs.
- Will leave more signature in Higgs coupling.

Some possible channels

Decay Topologies	Decay mode \mathcal{F}_i	Decay Topologies	Decay mode \mathcal{F}_i
	$h \rightarrow \cancel{E}_T$		$h \rightarrow (b\bar{b})(b\bar{b})$
$h \rightarrow 2$	$h \rightarrow \gamma + \cancel{E}_T$		$h \rightarrow (b\bar{b})(\tau^+\tau^-)$
$h \rightarrow 2 \rightarrow 3$	$h \rightarrow (b\bar{b}) + \cancel{E}_T$		$h \rightarrow (b\bar{b})(\mu^+\mu^-)$
	$h \rightarrow (jj) + \cancel{E}_T$		$h \rightarrow (\tau^+\tau^-)(\tau^+\tau^-)$
	$h \rightarrow (\tau^+\tau^-) + \cancel{E}_T$		$h \rightarrow (\tau^+\tau^-)(\mu^+\mu^-)$
	$h \rightarrow (\gamma\gamma) + \cancel{E}_T$		$h \rightarrow (jj)(jj)$
	$h \rightarrow (\ell^+\ell^-) + \cancel{E}_T$		$h \rightarrow (jj)(\gamma\gamma)$
			$h \rightarrow (jj)(\mu^+\mu^-)$
$h \rightarrow 2 \rightarrow 3 \rightarrow 4$	$h \rightarrow (b\bar{b}) + \cancel{E}_T$		$h \rightarrow (\ell^+\ell^-)(\ell^+\ell^-)$
	$h \rightarrow (jj) + \cancel{E}_T$		$h \rightarrow (\ell^+\ell^-)(\mu^+\mu^-)$
	$h \rightarrow (\tau^+\tau^-) + \cancel{E}_T$		$h \rightarrow (\mu^+\mu^-)(\mu^+\mu^-)$
	$h \rightarrow (\gamma\gamma) + \cancel{E}_T$		$h \rightarrow (\gamma\gamma)(\gamma\gamma)$
	$h \rightarrow (\ell^+\ell^-) + \cancel{E}_T$		$h \rightarrow \gamma\gamma + \cancel{E}_T$
	$h \rightarrow (\mu^+\mu^-) + \cancel{E}_T$		$h \rightarrow \gamma\gamma + \cancel{E}_T$
$h \rightarrow 2 \rightarrow (1+3)$	$h \rightarrow b\bar{b} + \cancel{E}_T$	$h \rightarrow 2 \rightarrow 4 \rightarrow 6$	$h \rightarrow (\ell^+\ell^-)(\ell^+\ell^-) + \cancel{E}_T$
	$h \rightarrow jj + \cancel{E}_T$		$h \rightarrow (\ell^+\ell^-) + \cancel{E}_T + X$
	$h \rightarrow \tau^+\tau^- + \cancel{E}_T$		$h \rightarrow \ell^+\ell^-\ell^+\ell^- + \cancel{E}_T$
	$h \rightarrow \gamma\gamma + \cancel{E}_T$		$h \rightarrow \ell^+\ell^- + \cancel{E}_T + X$
	$h \rightarrow \ell^+\ell^- + \cancel{E}_T$		



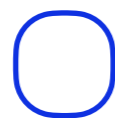
Helicity structure at LHC

$$f_L \bar{f}_R \rightarrow W^+ W^-$$

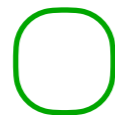
(h_{W^+}, h_{W^-})	SM	\mathcal{O}_W	\mathcal{O}_{HW}	\mathcal{O}_{HB}	\mathcal{O}_B	\mathcal{O}_{3W}	\mathcal{O}_{TWW}
(\pm, \mp)	1	0	0	0	0	0	$\frac{E^4}{\Lambda^4}$
$(0, 0)$	1	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	0	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$
$(0, \pm), (\pm, 0)$	$\frac{m_W}{E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^4}{\Lambda^4} \frac{m_W}{E}$
(\pm, \pm)	$\frac{m_W^2}{E^2}$	$\frac{m_W^2}{\Lambda^2}$	$\frac{m_W^2}{\Lambda^2}$	0	0	$\frac{E^2}{\Lambda^2}$	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$

$$f_R \bar{f}_L \rightarrow W^+ W^-$$

(h_{W^+}, h_{W^-})	SM	\mathcal{O}_W	\mathcal{O}_{HW}	\mathcal{O}_{HB}	\mathcal{O}_B	\mathcal{O}_{3W}	\mathcal{O}_{TWW}
(\pm, \mp)	0	0	0	0	0	0	$\frac{E^4}{\Lambda^4}$
$(0, 0)$	1	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	$\frac{E^2}{\Lambda^2}$	0	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$
$(0, \pm), (\pm, 0)$	$\frac{m_W}{E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{E^2 m_W}{\Lambda^2 E}$	$\frac{m_W^2 m_W}{\Lambda^2 E}$	$\frac{E^4}{\Lambda^4} \frac{m_W}{E}$
(\pm, \pm)	$\frac{m_W^2}{E^2}$	$\frac{m_W^2}{\Lambda^2}$	$\frac{m_W^2}{\Lambda^2}$	0	0	$\frac{m_W^2}{\Lambda^2}$	$\frac{E^4}{\Lambda^4} \frac{m_W^2}{E^2}$



growing with energy



SM piece is small. Interference does not grow with E.