

Belle II iTOP Optics: Design, Construction and Performance

Boqun Wang

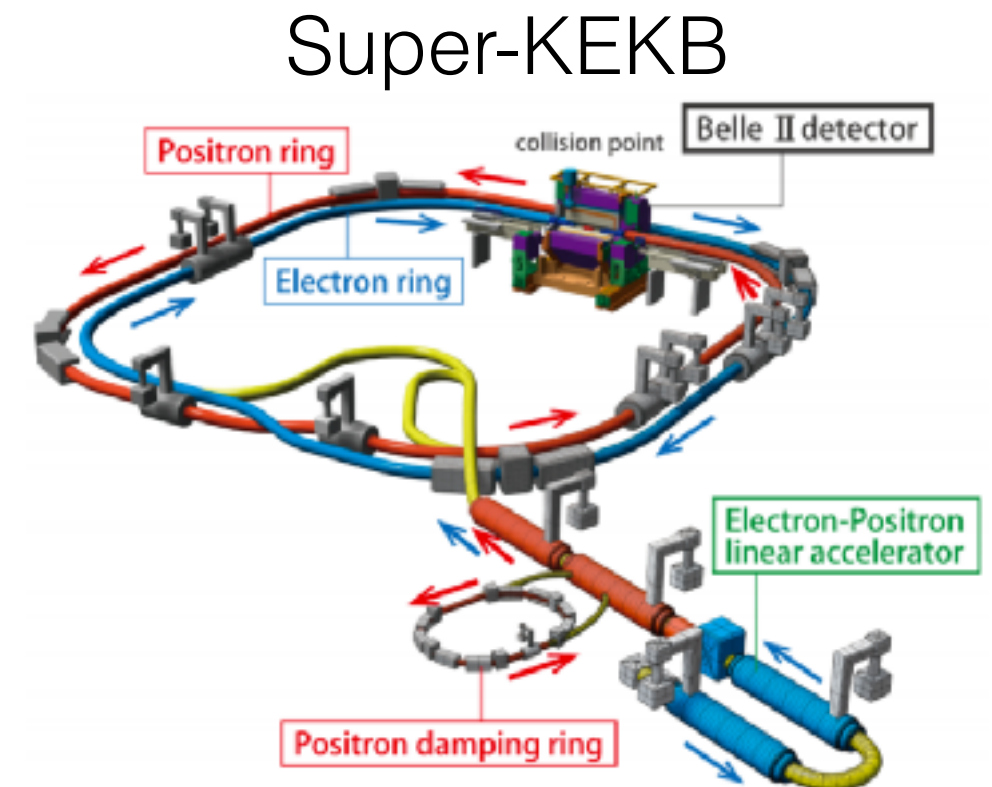
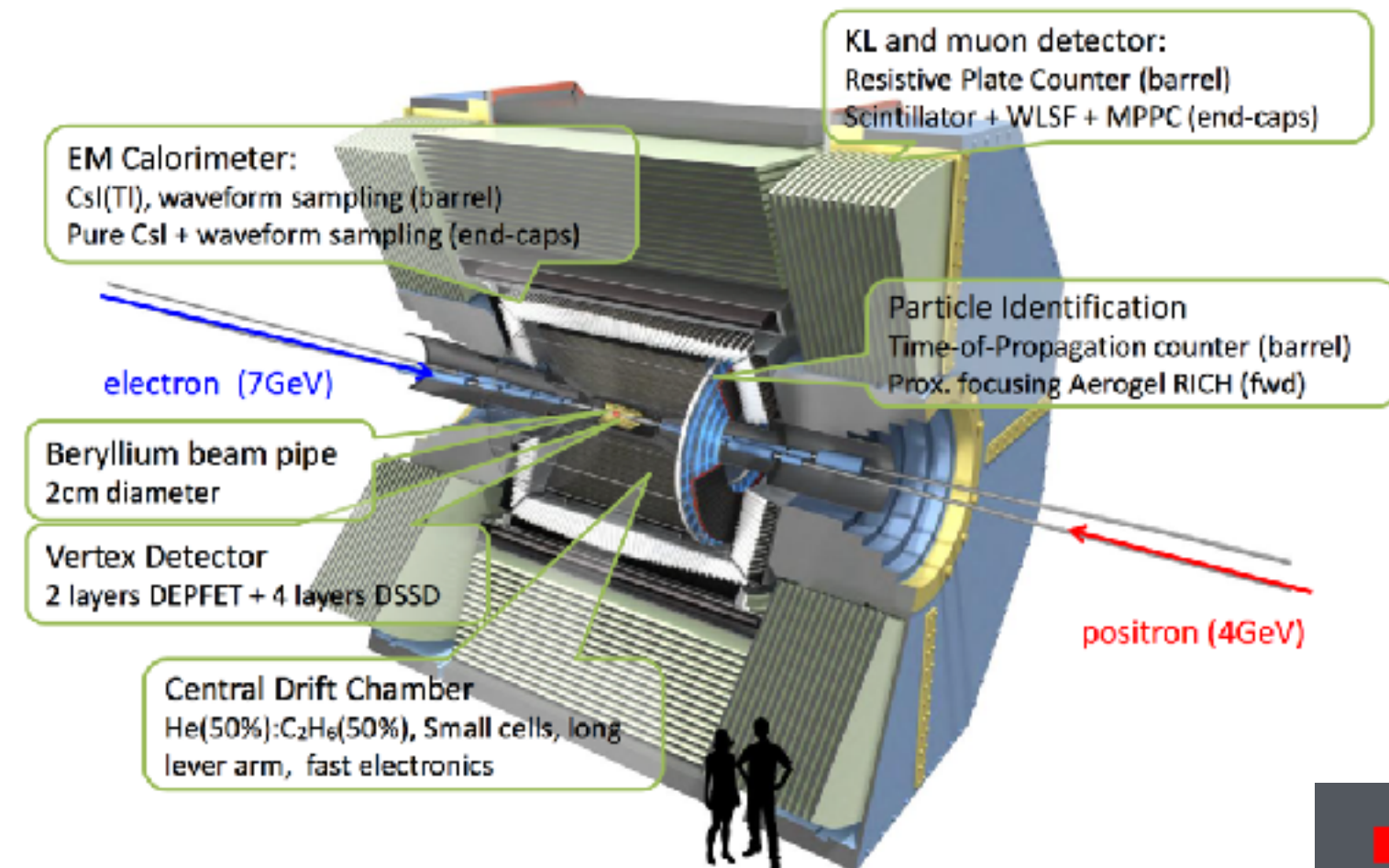
Saurabh Sandilya, Bilas Pal, Alan Schwartz
Department of Physics, University of Cincinnati

May 24, 2017

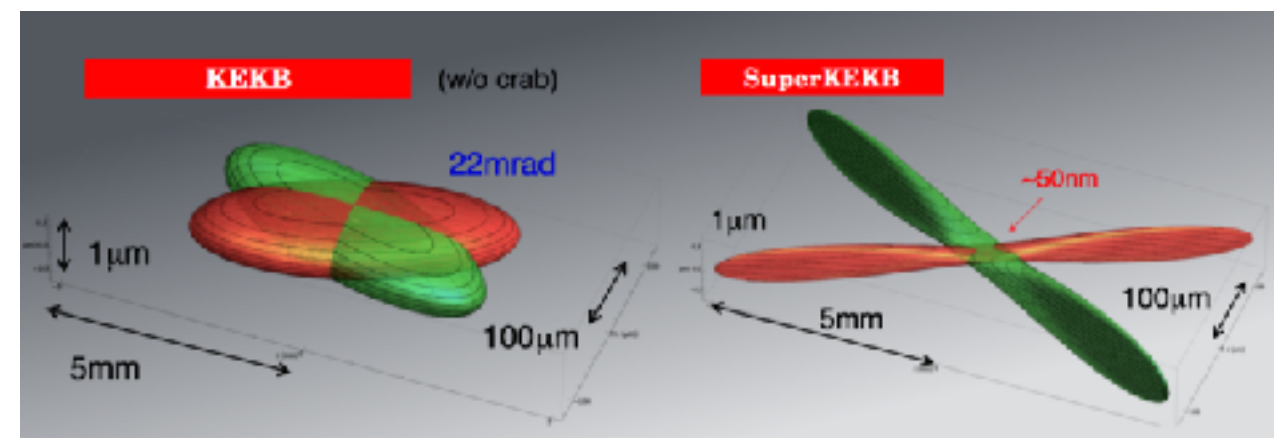
TIPP 2017, Beijing, China



Overview of Super-KEKB and Belle II

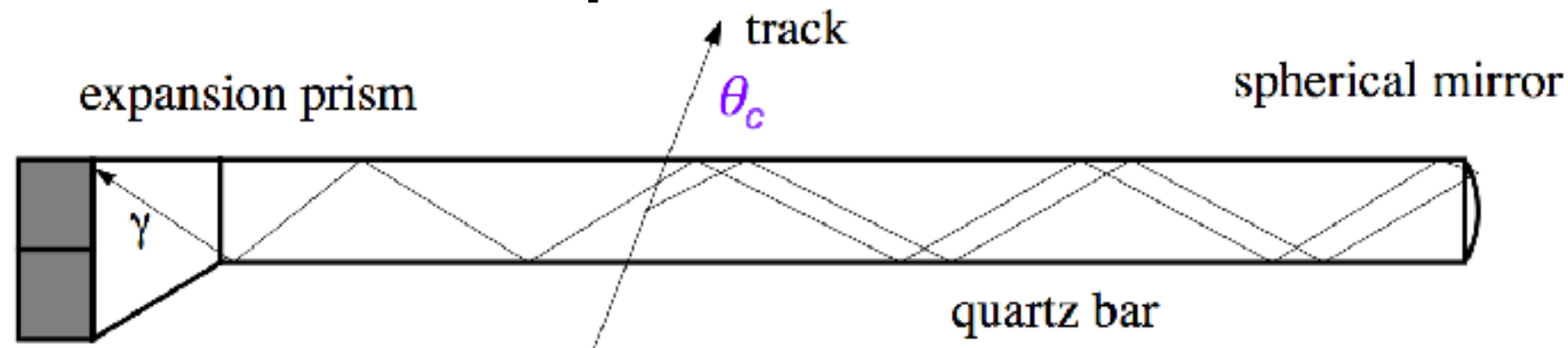


Target integrated luminosity: $\sim 50 \text{ ab}^{-1}$
 Peak luminosity: $8 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$ (by using nano-beam technology)
 Physics run: ~ 2018

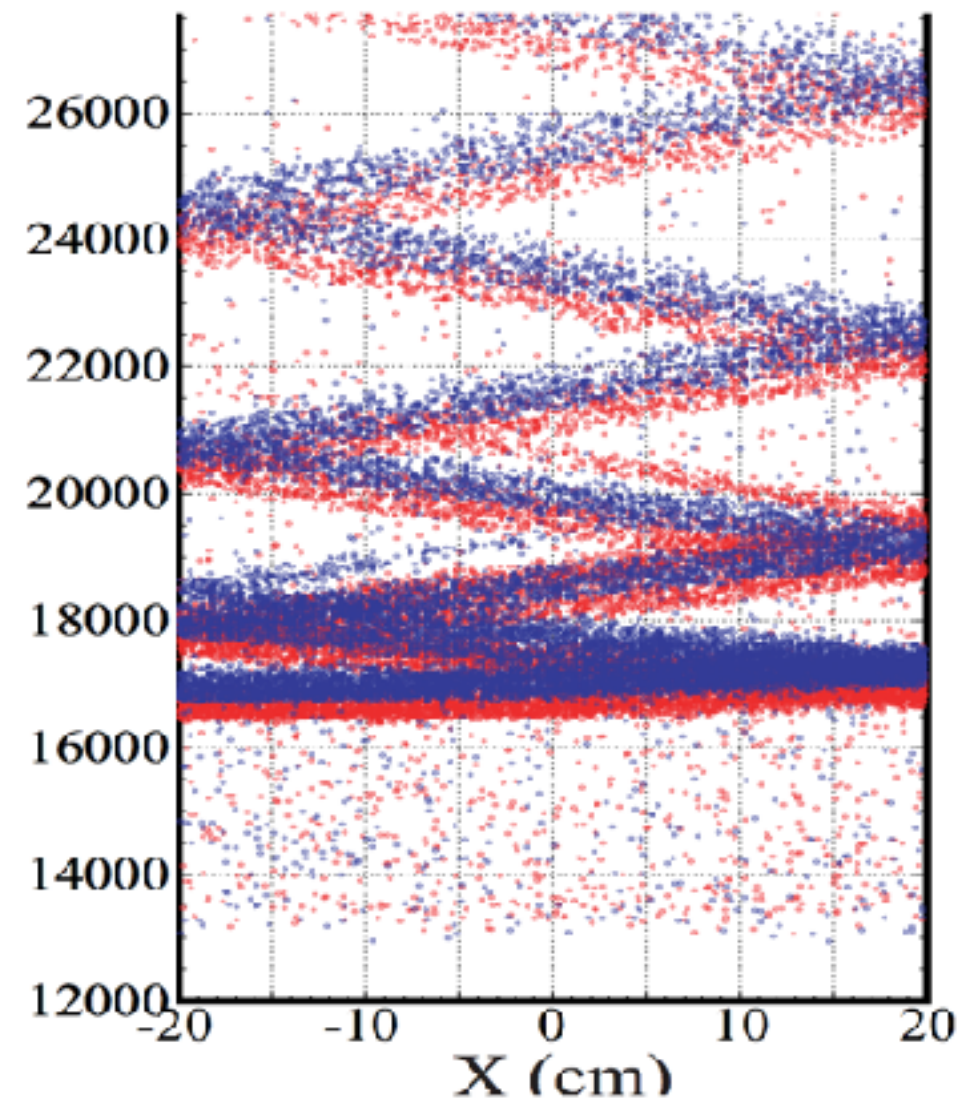
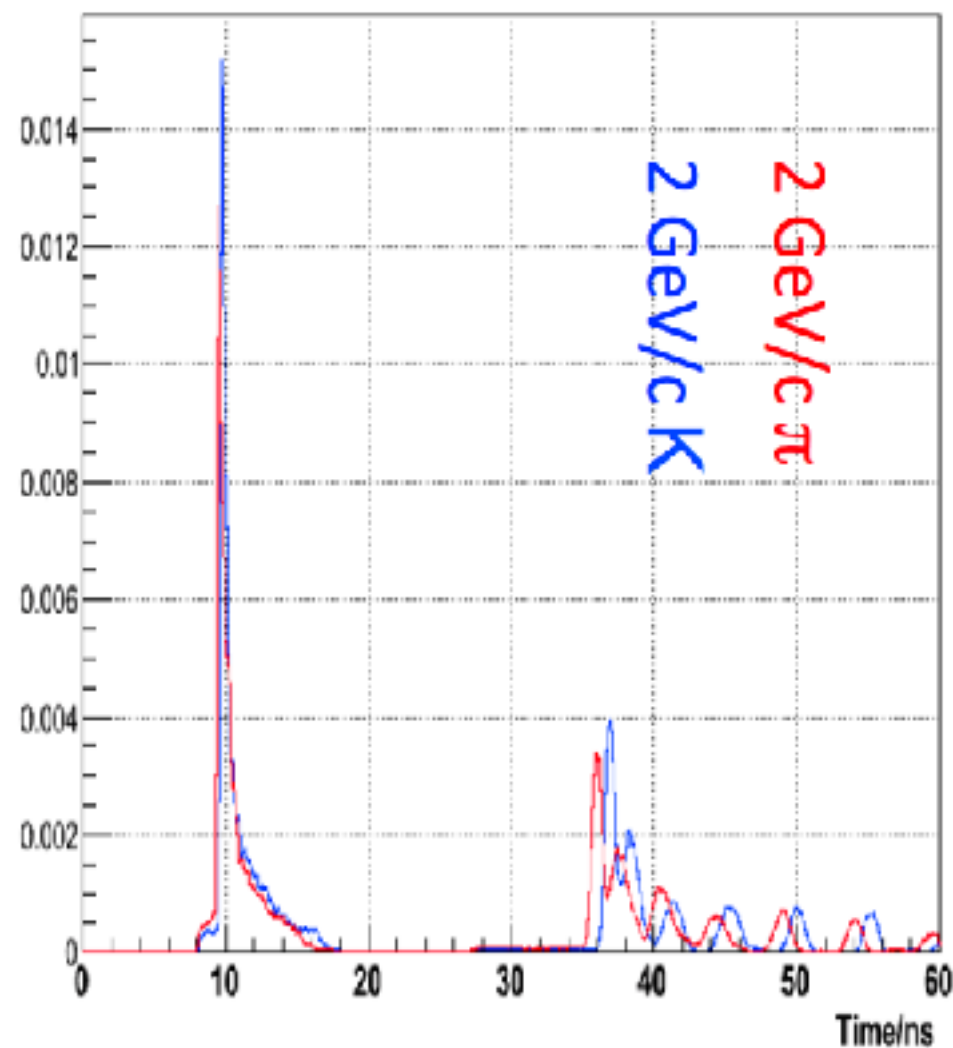


Nano-beam

Principles of iTOP Detector

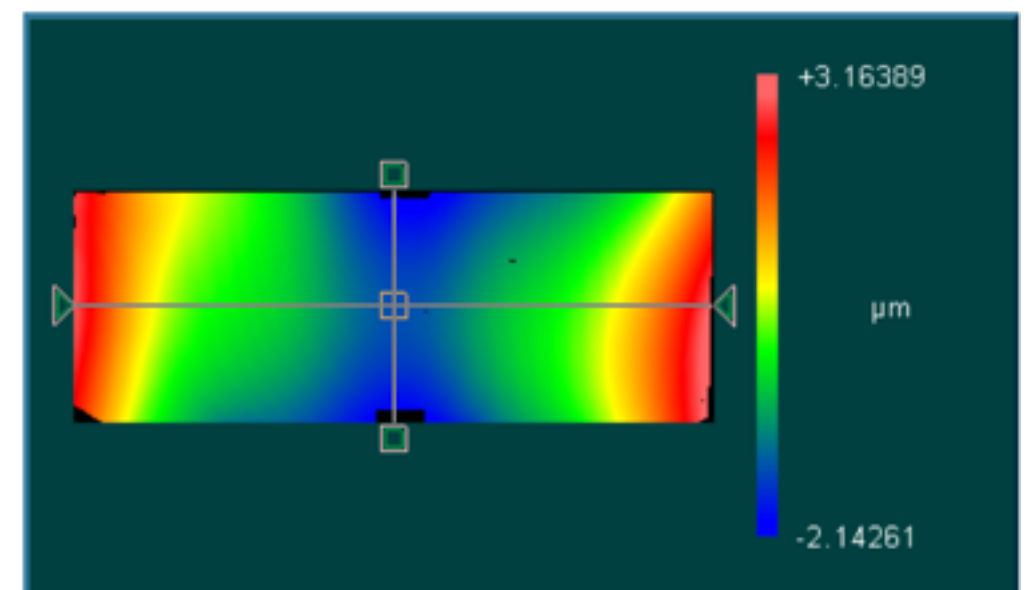
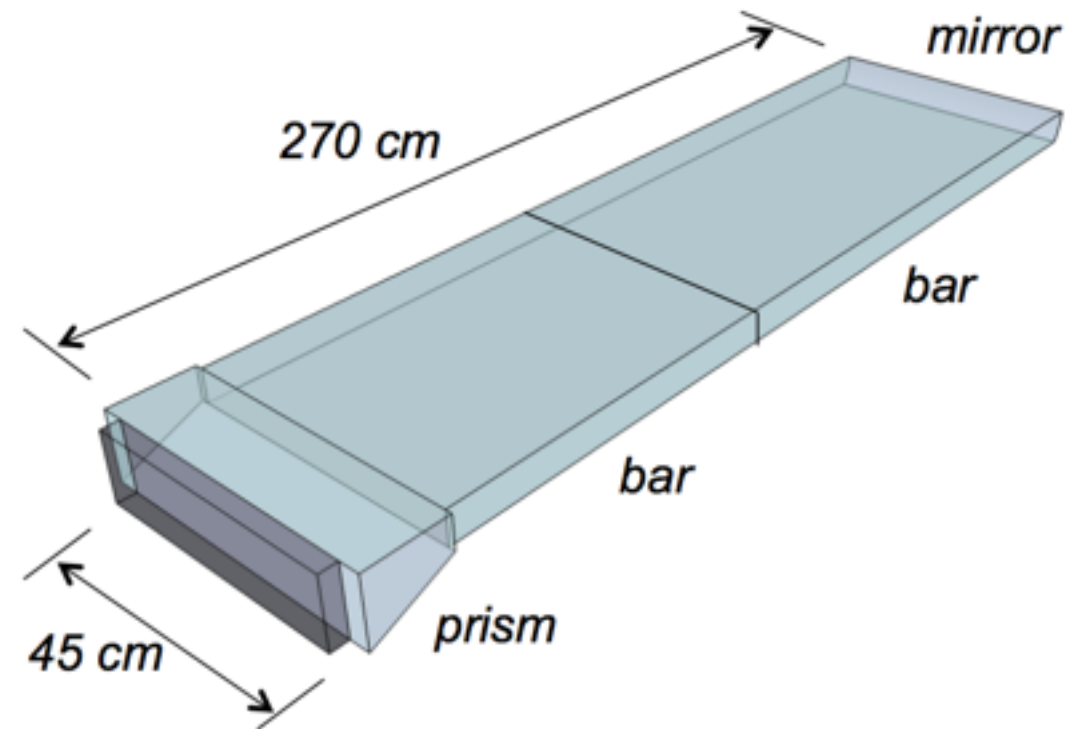


$\cos\theta_c = 1/n\beta$
 π and K have different θ_c
 Different hit positions and
 arrival times of photons



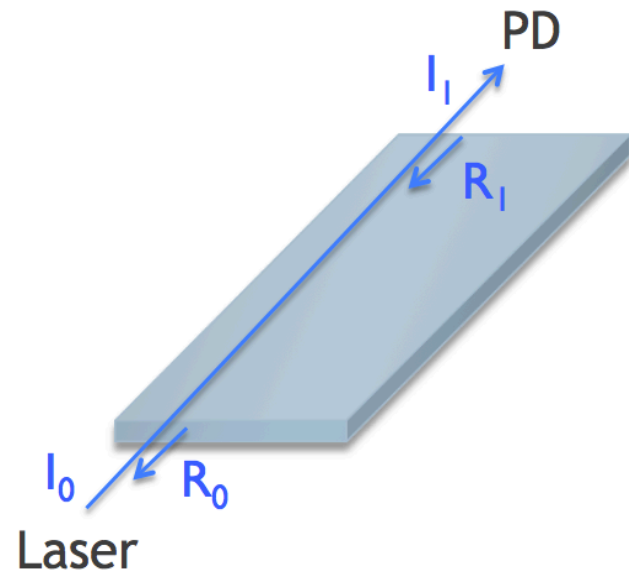
Procurement of Synthetic Fused Silica (Quartz)

- Two bars, one mirror and one prism per module.
- Totally 16 modules and 1 spare module.
- Acceptance test:
 - For 32+ bars: chip inspection, bulk transmittance, internal surface reflectance.
 - For 16+ mirrors: chip inspection, reflectivity, position of optical axis, focal point and focal length, spherical aberration, astigmatism
 - For 16+ prisms: chip inspection, transmission, angle of tilted surface.
- Surface flatness, surface roughness, parallelism, perpendicularity and chamfer specs were qualified by vender.



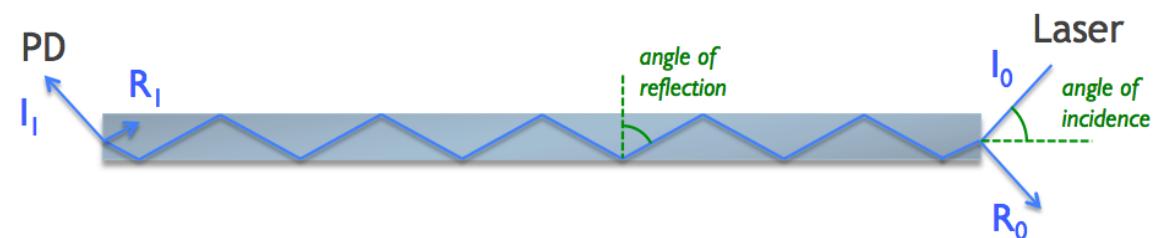
Interferograms from metrology report

QA: Quartz Bar



Bulk Transmission

$$I_0(1 - R_0) \tau (1 - R_1) = I_1$$



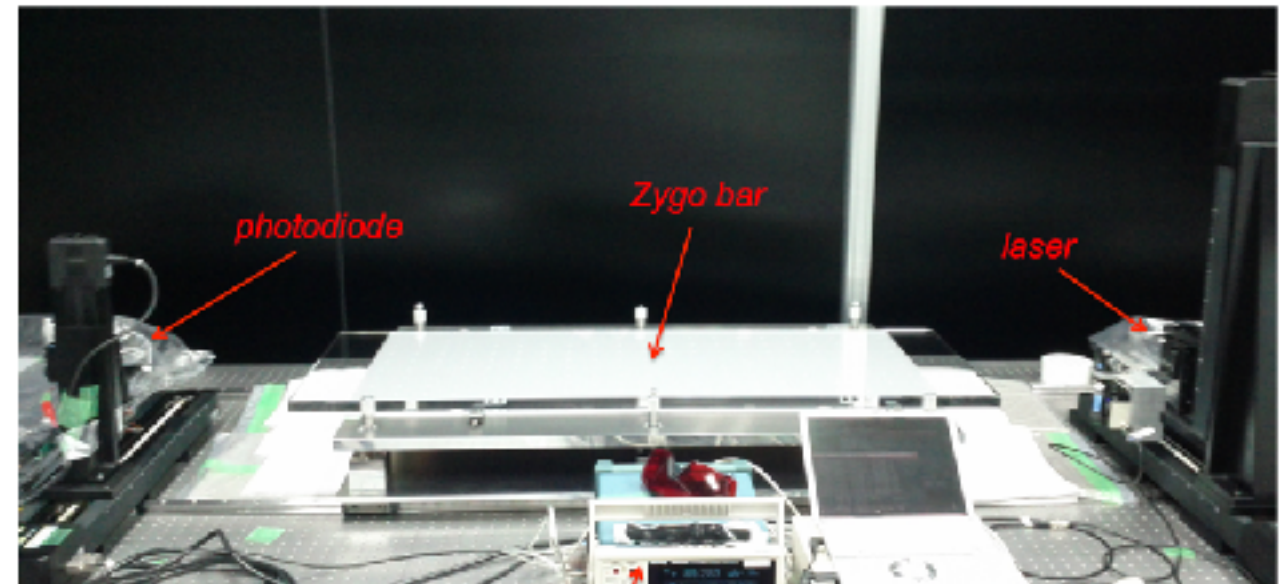
Internal Reflectivity

$$(I_1 - R_1) = (I_0 - R_0) \cdot \alpha^N \cdot \exp\left(-\frac{L}{\Lambda} \cdot \sqrt{1 + (Nh/L)^2}\right)$$

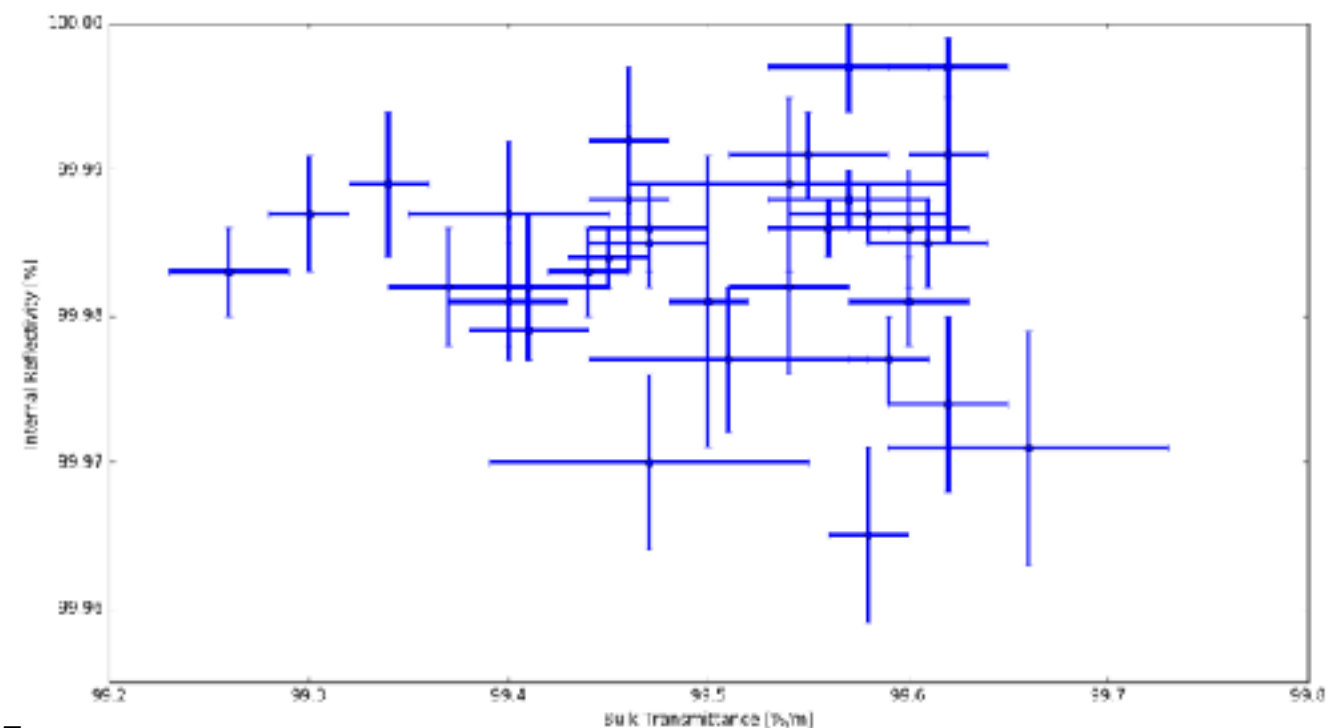
Requirement:

Bulk Transmittance: > 98.5 %/m

Internal Reflectivity: > 99.9 %

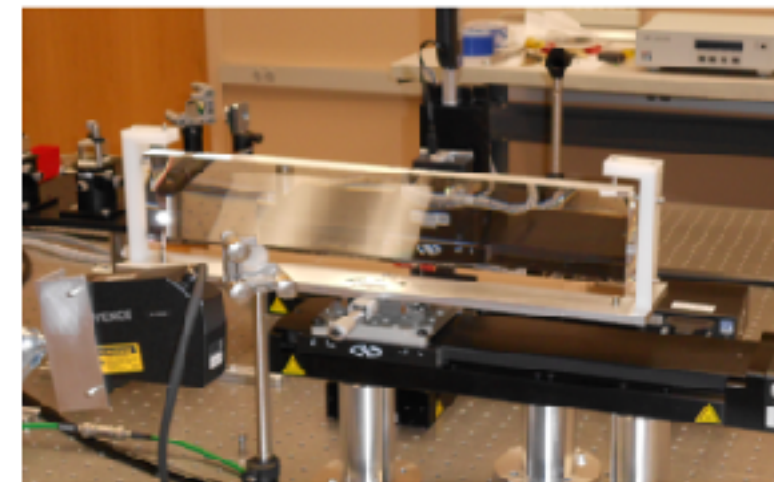
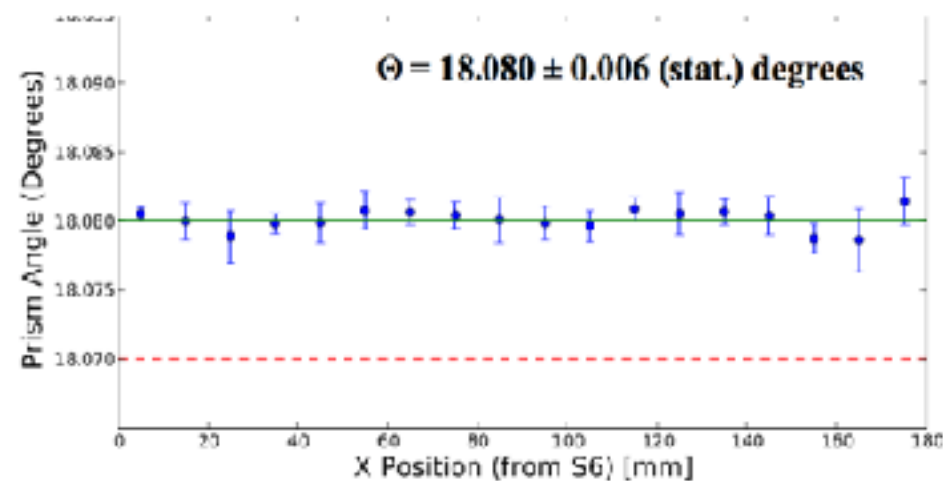
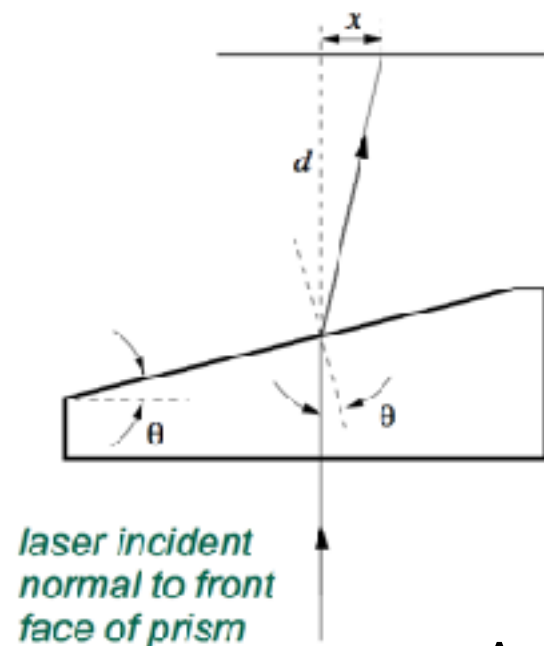
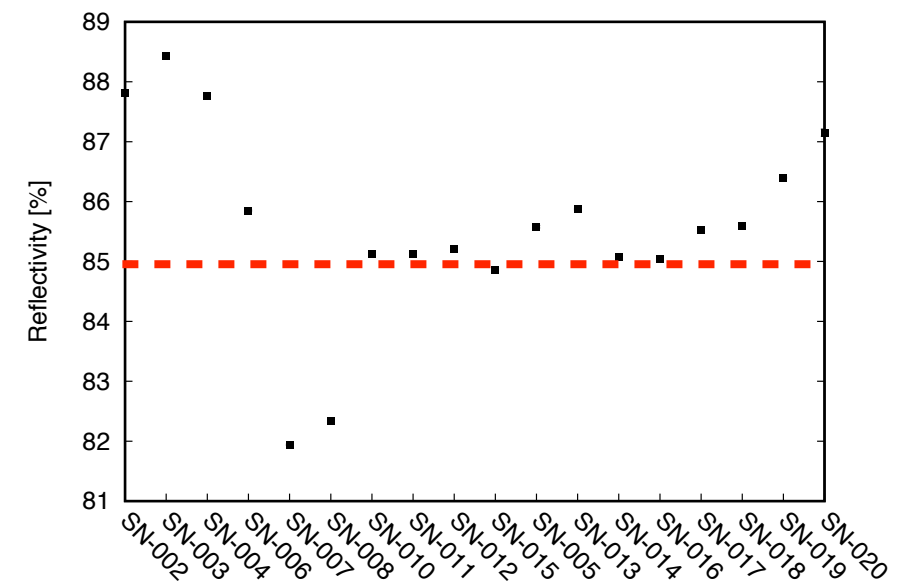
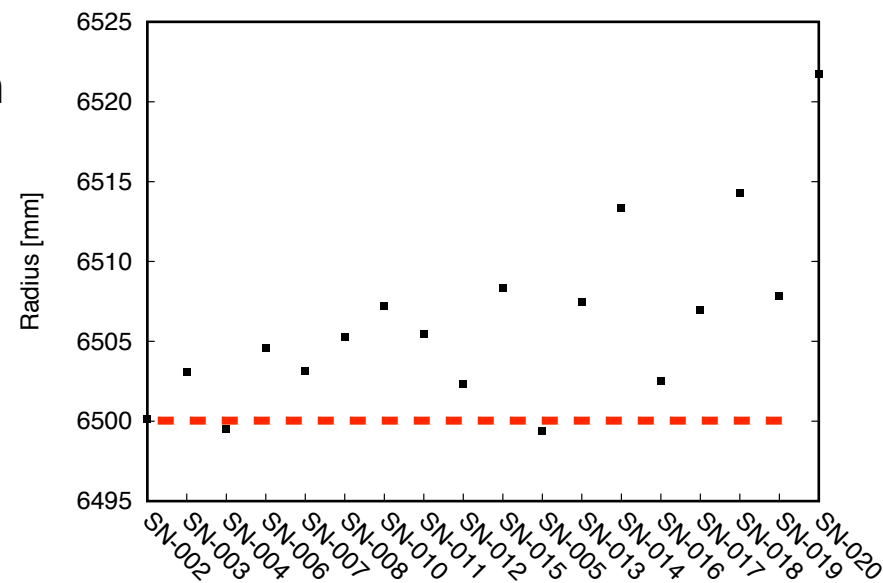
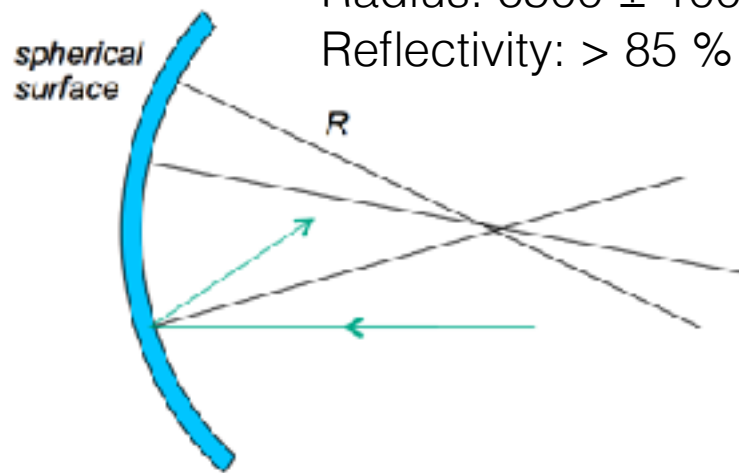


control system
(software runs as Python scripts)



QA: Mirror & Prism

Specification
Radius: 6500 ± 100 mm
Reflectivity: $> 85\%$

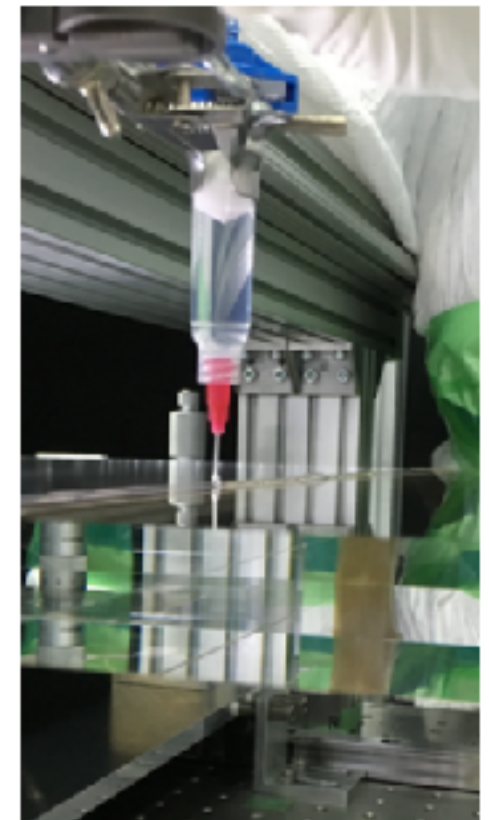
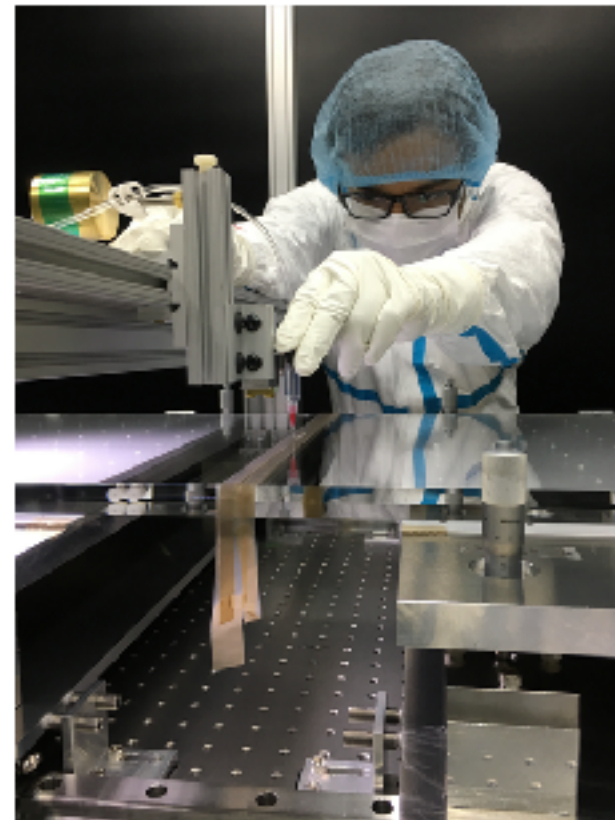
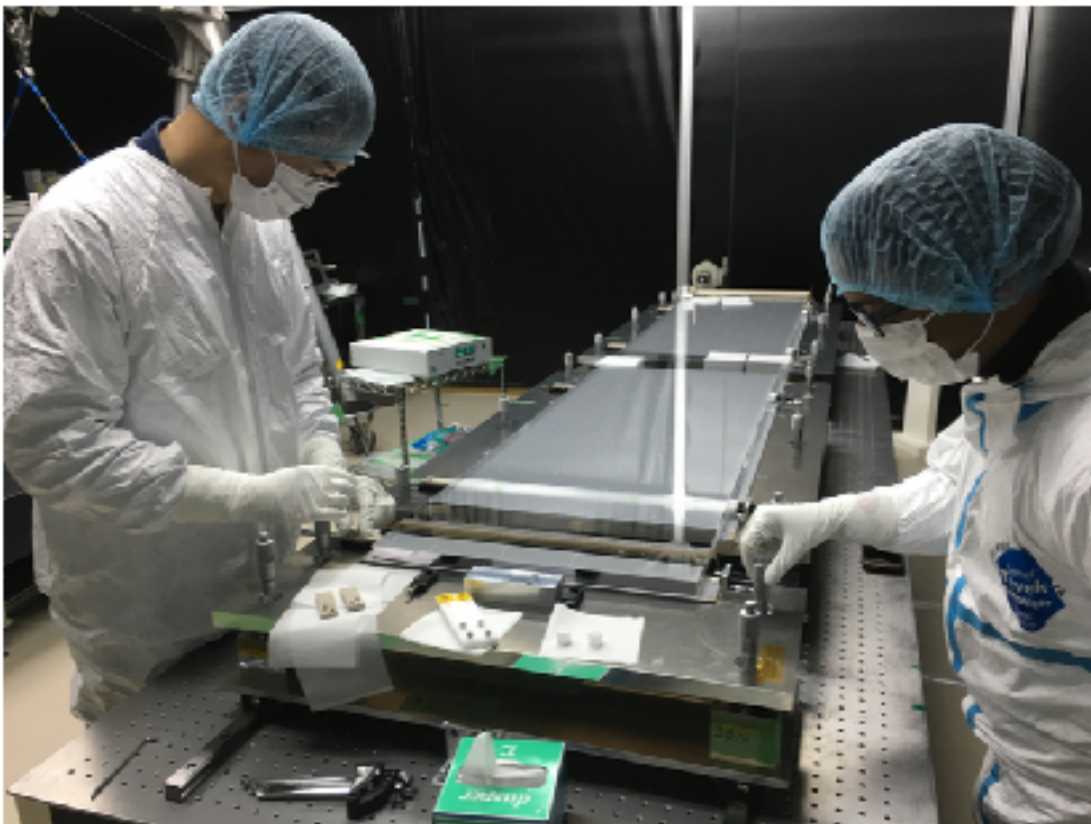
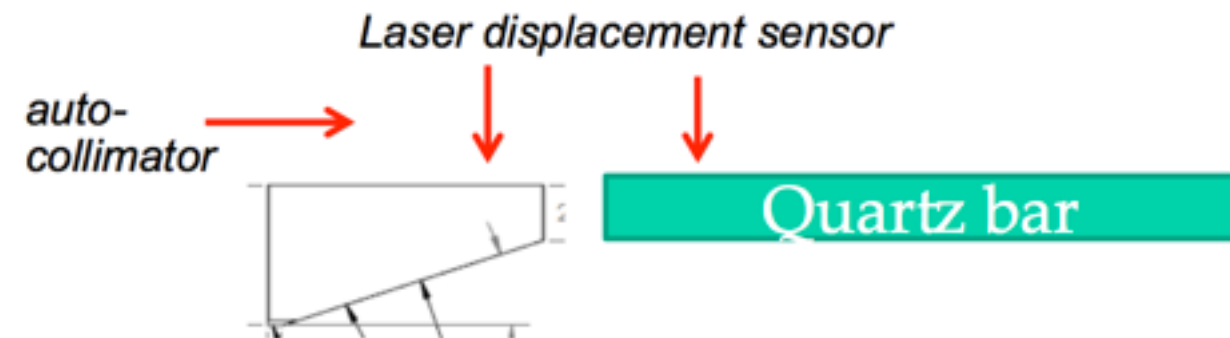


Angle of tilted face: 18.07 ± 0.04 deg (± 144 arcsec)

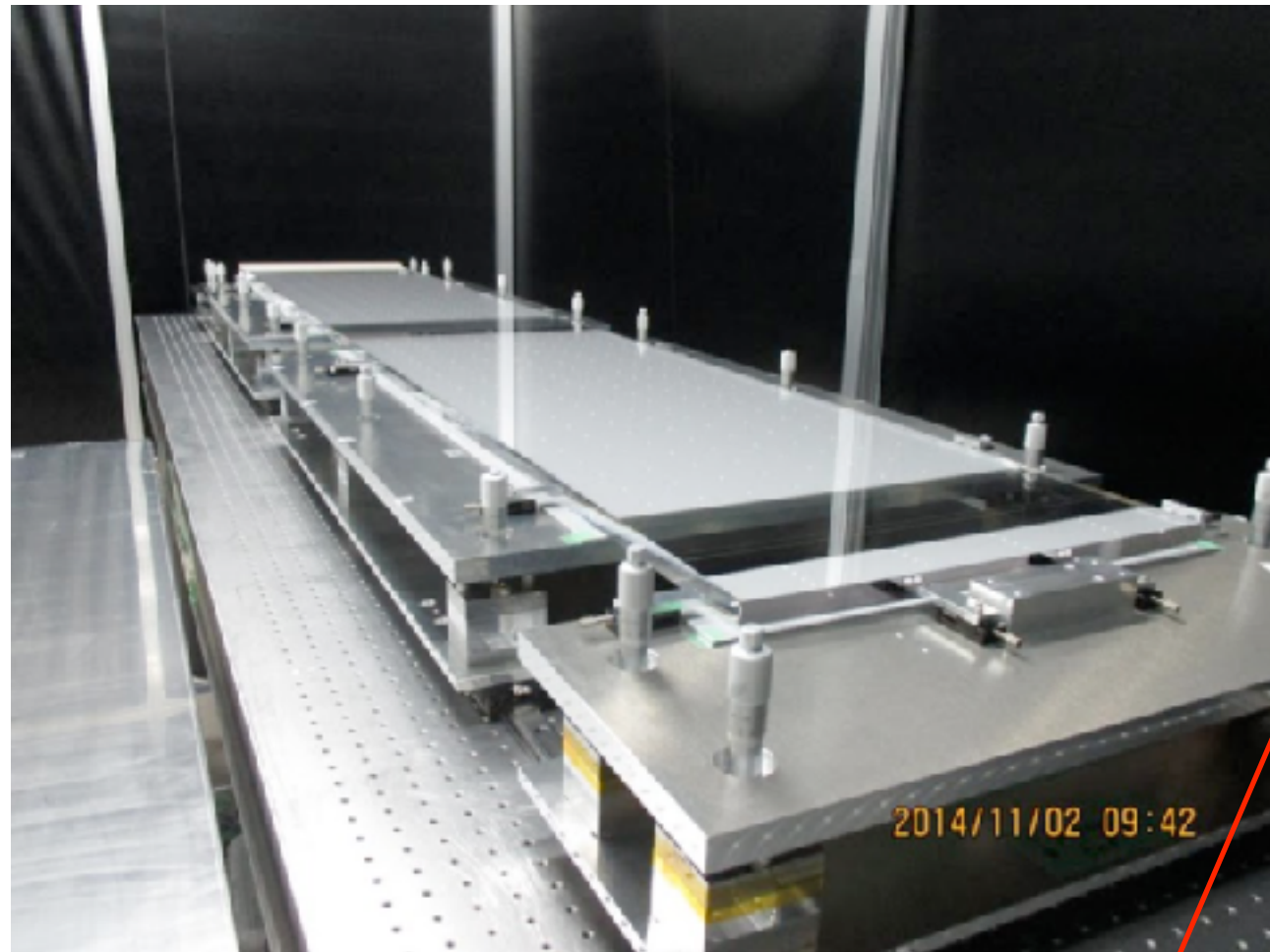
Alignment & Gluing

Procedure:

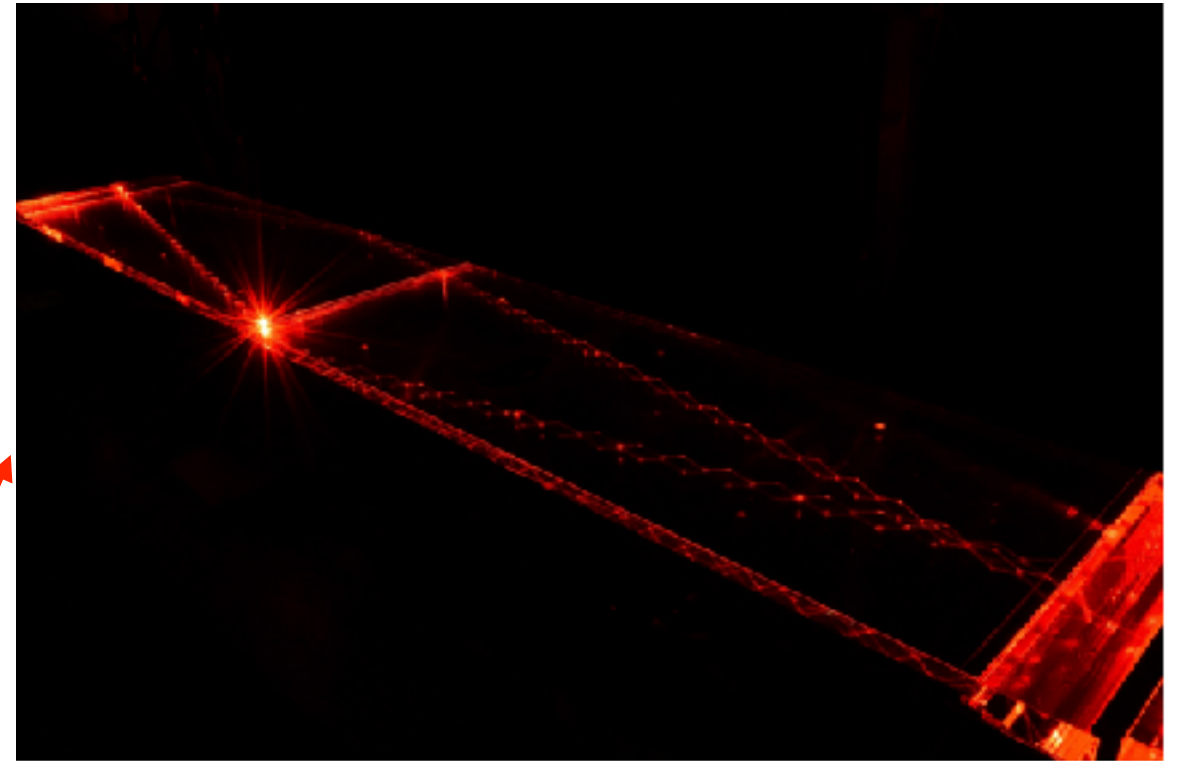
- adjust surfaces positions using laser displacement sensor and micrometers
- adjust surfaces angles using autocollimator and micrometers
- insert shims, tape joint and repeat steps 1, 2
- apply epoxy (EPOTEK 301-2) to joint



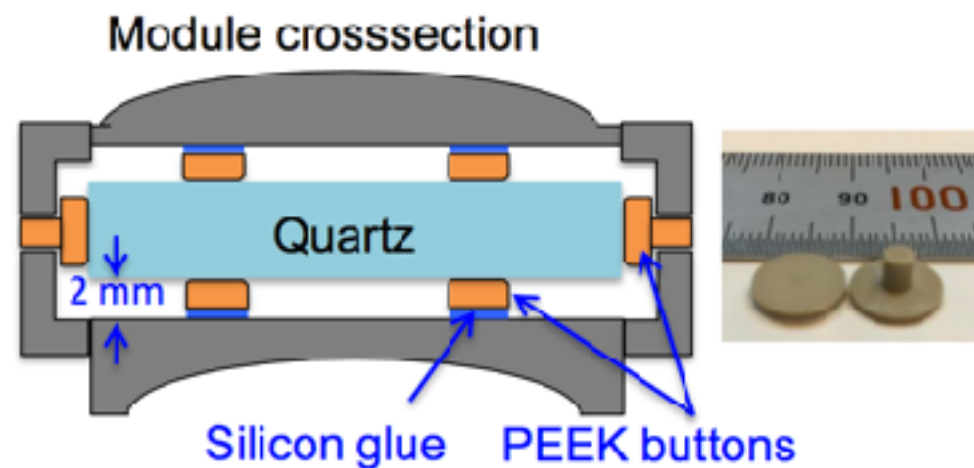
After Gluing



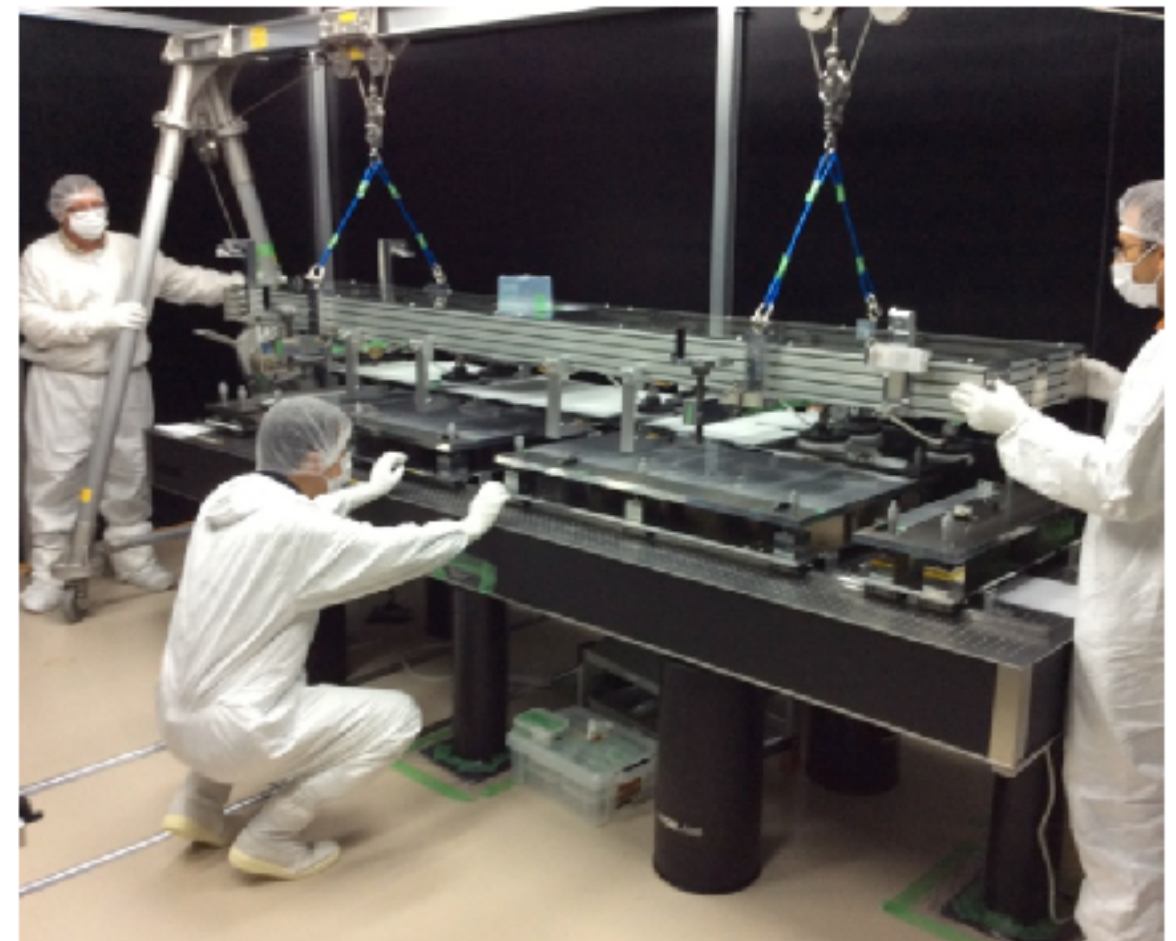
Long time exposure with laser
input from prism end
Laser scattering on the surface
and inside the bar



QBB Assembly



Use vacuum based lifting jig to move glued optics to QBB assembly table

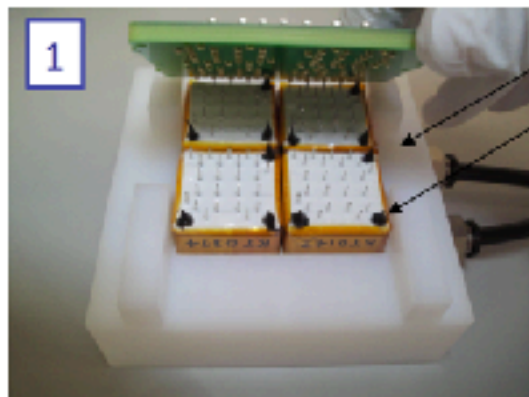
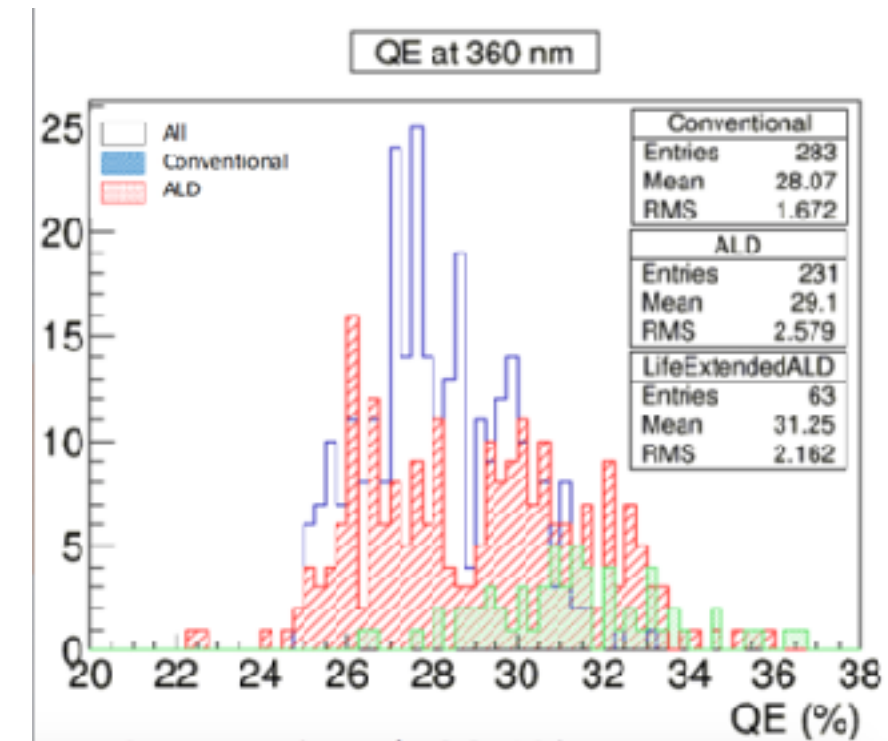


MCP-PMT

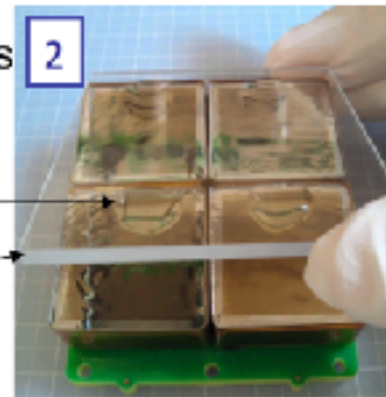


4 x 4 anodes
27.6 x 27.6 mm²
23.0 x 23.0 mm² active

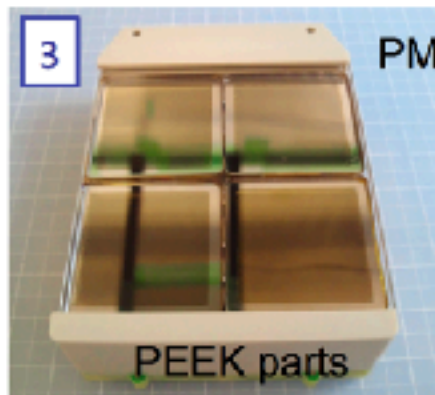
QE requirement:
> 24% at peak λ
> 28% average



Vacuum chuck to align the PMT faces
RTV silicon rubber to hold the PMTs



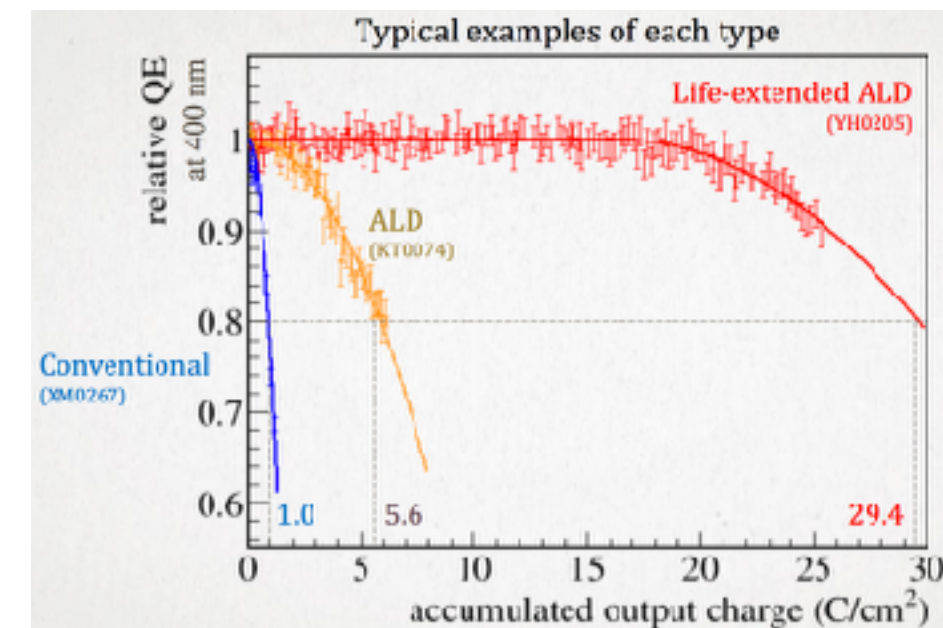
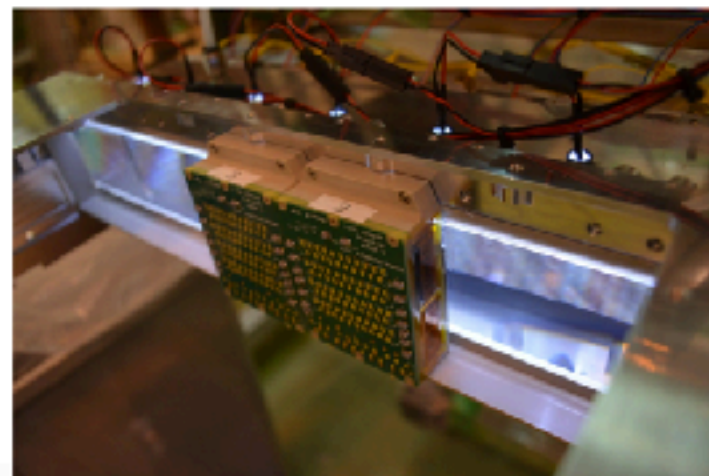
Silicon rubber TSE3032 (before curing) to be filled between the PMTs and the wavelength cut filter



PMT module completed

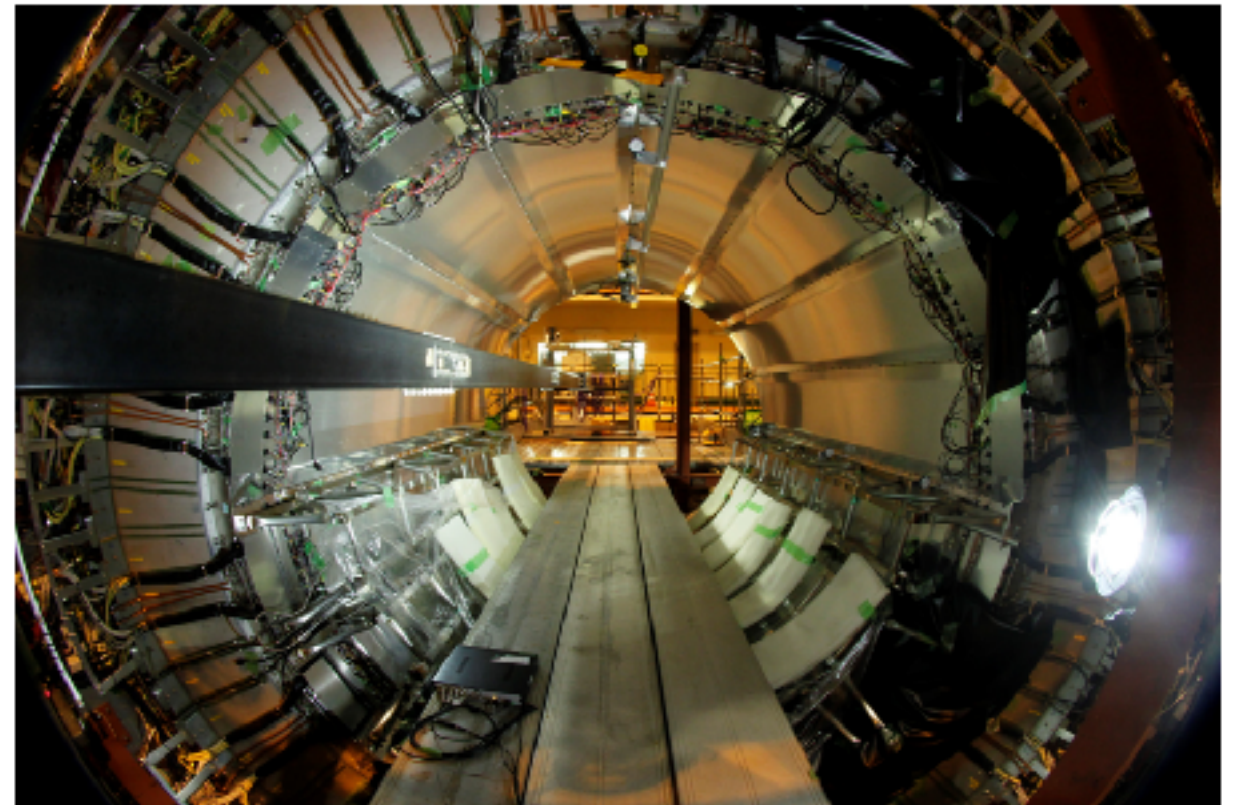
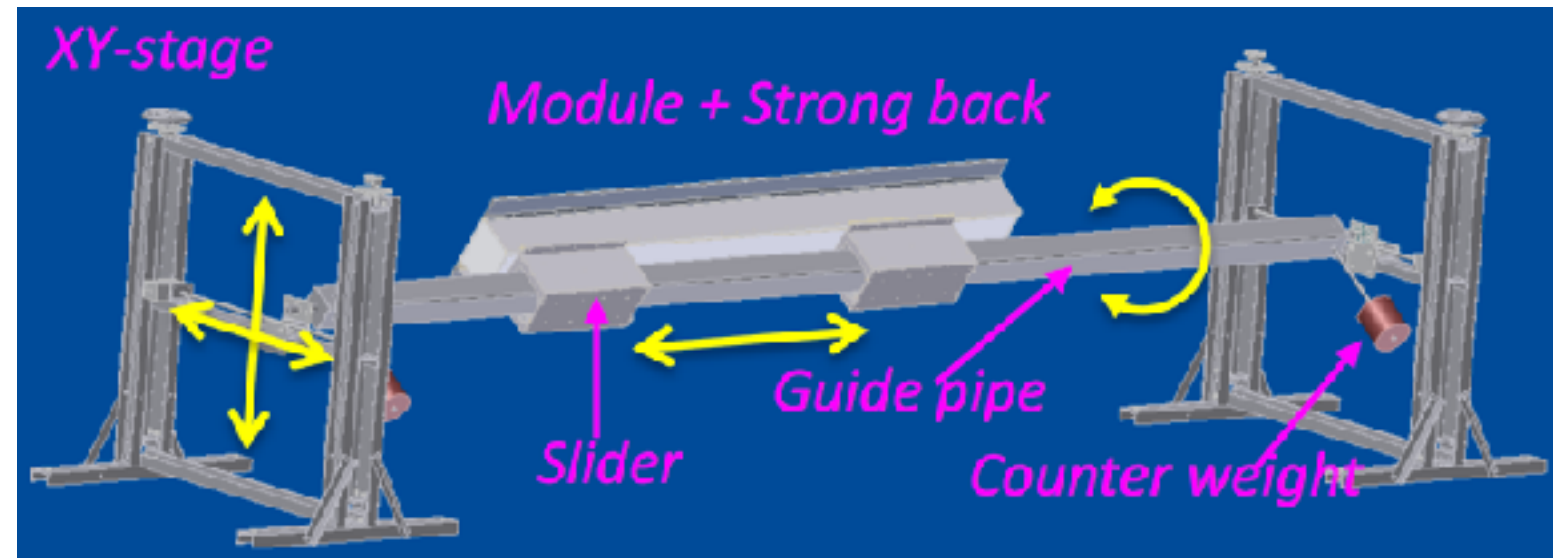
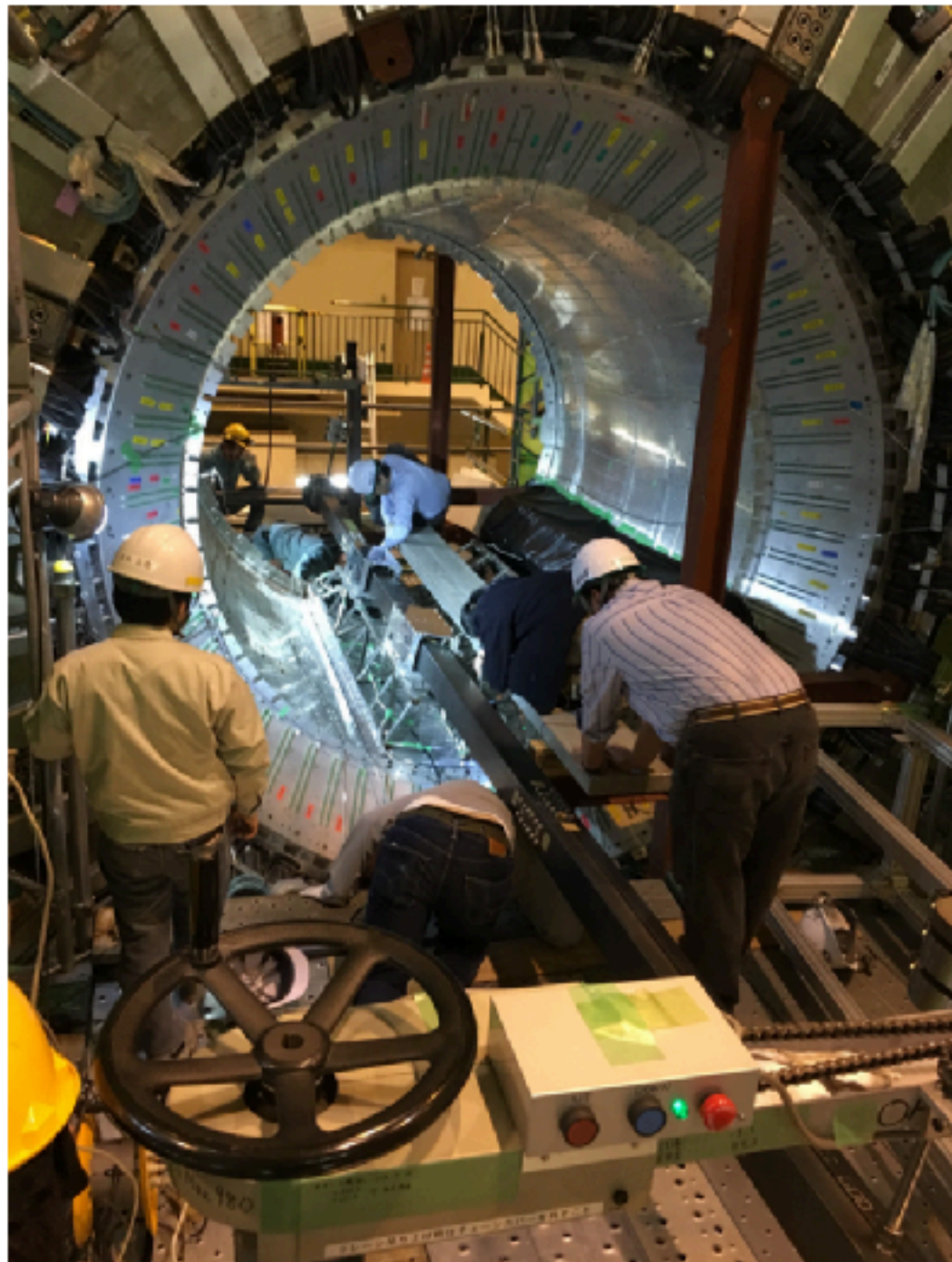
PEEK parts

2 PMT modules mounted to prism with a "cookie" (+oil):



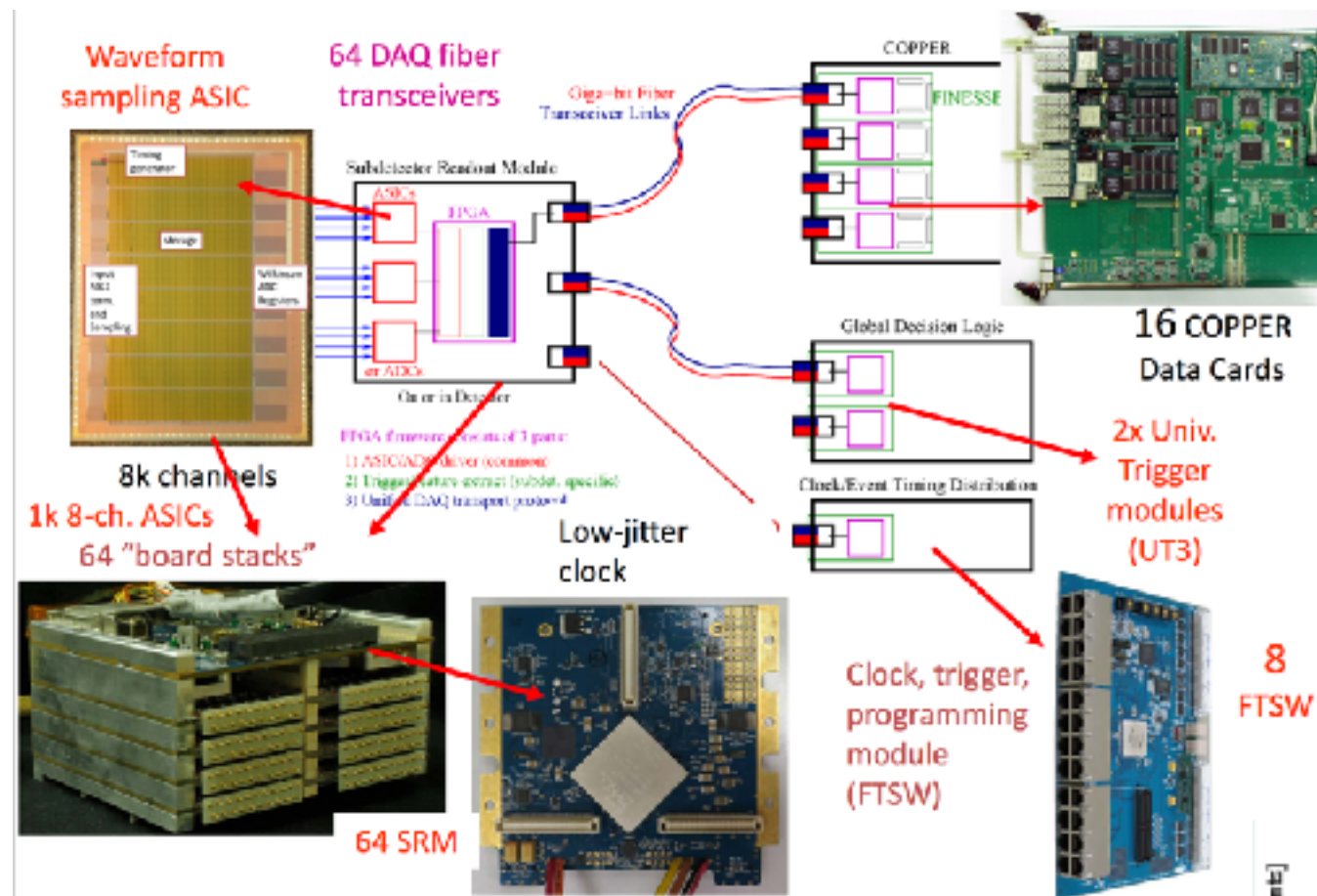
Life time extension
See Kodai Matsuoka's talk

Installation

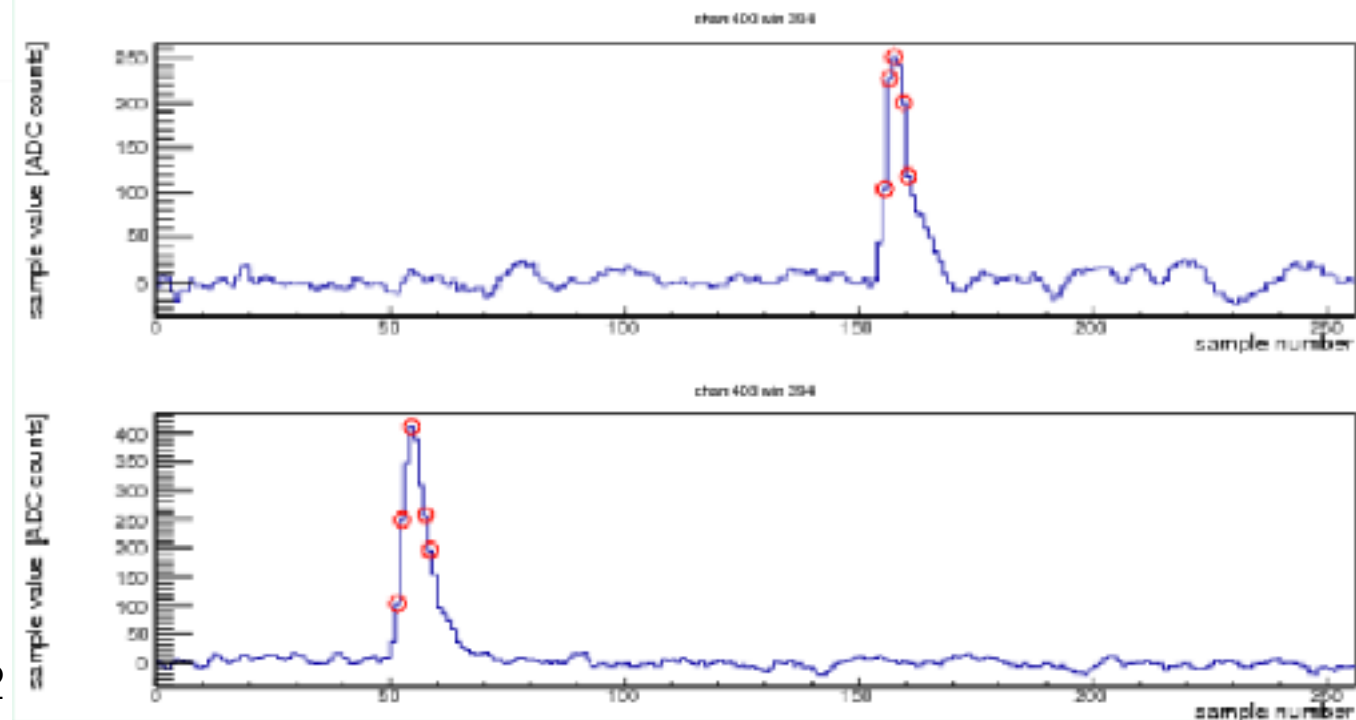
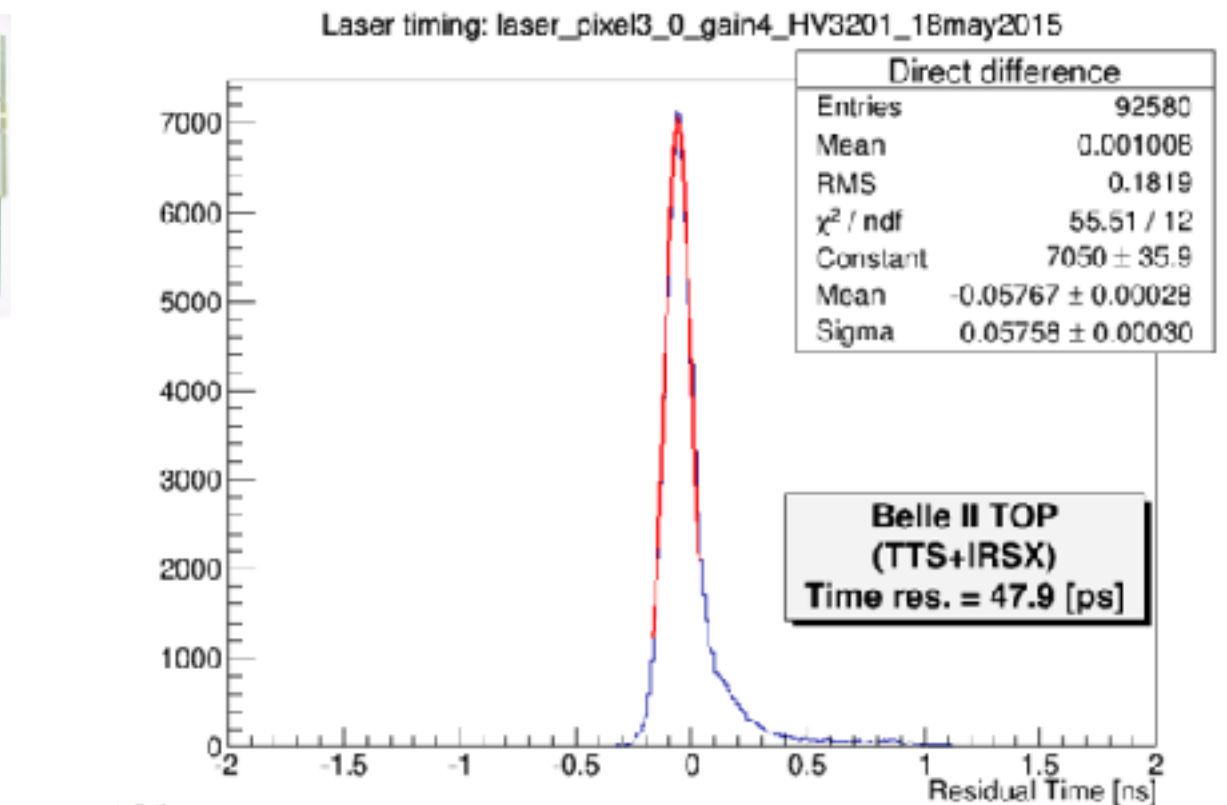


All module installed by May 2016

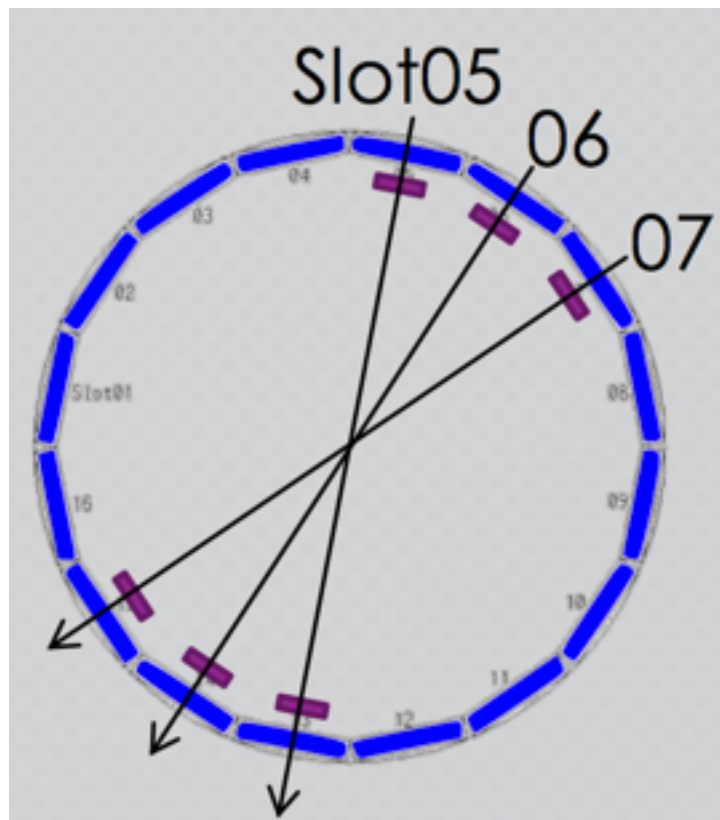
Readout System



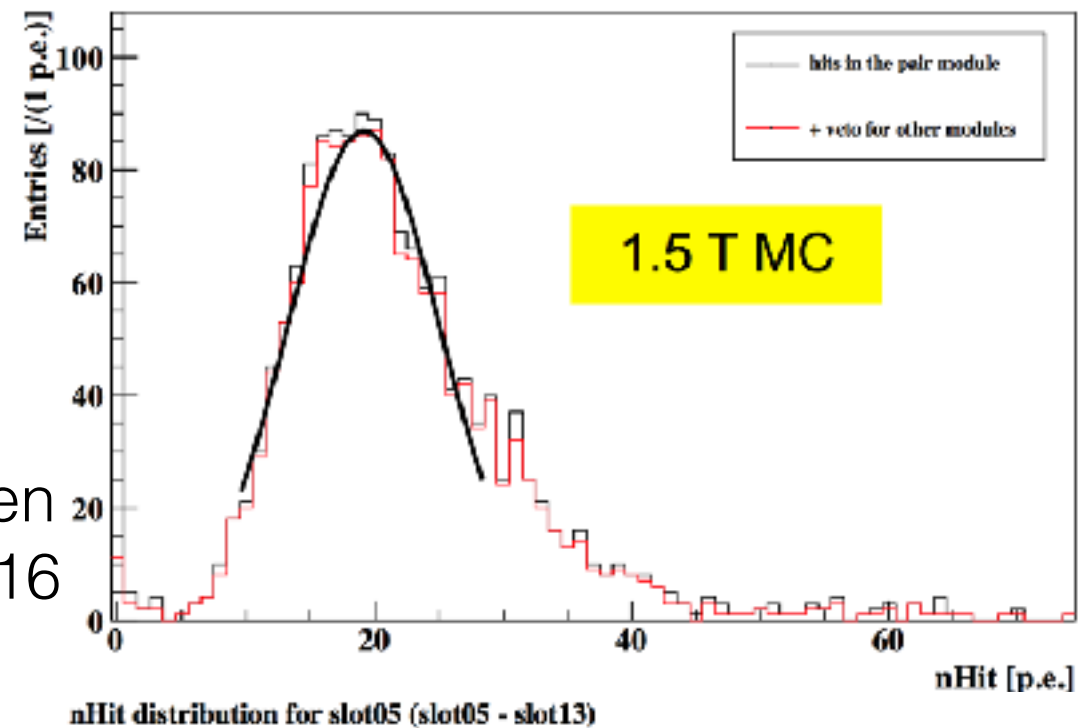
Full Waveform
vs
Region of Interest & Feature Extraction



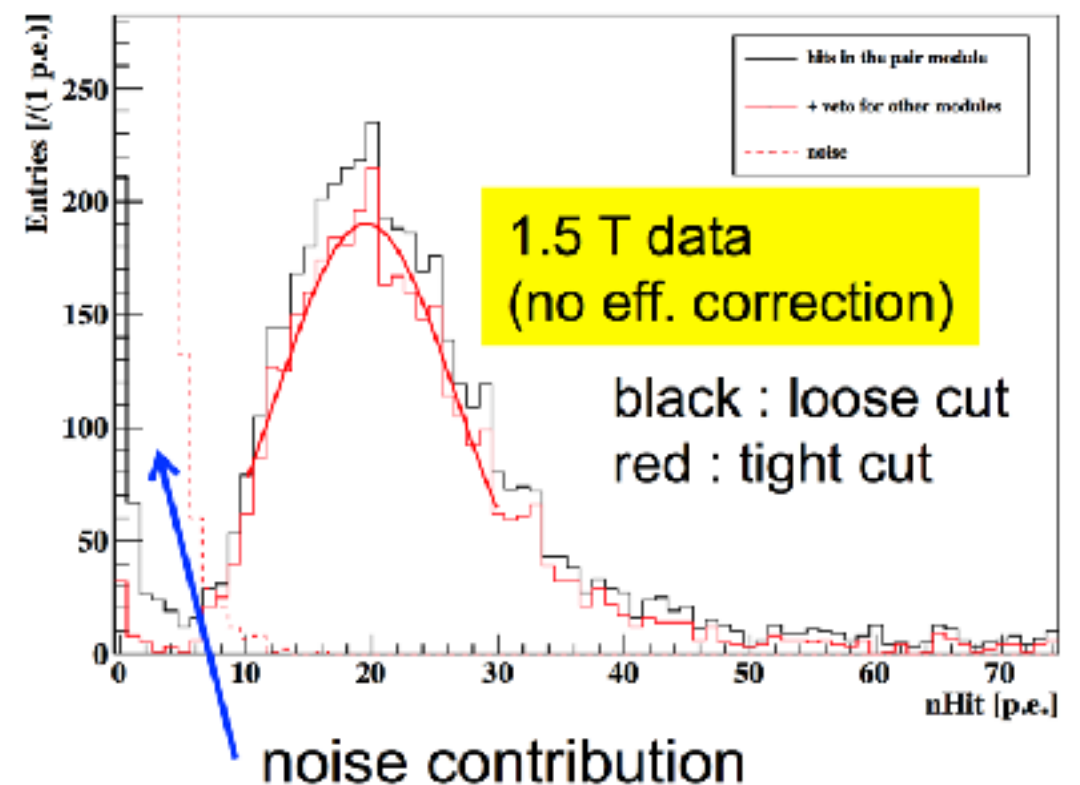
Cosmic Ray Test



Cosmic ray data taken during Sep - Oct, 2016



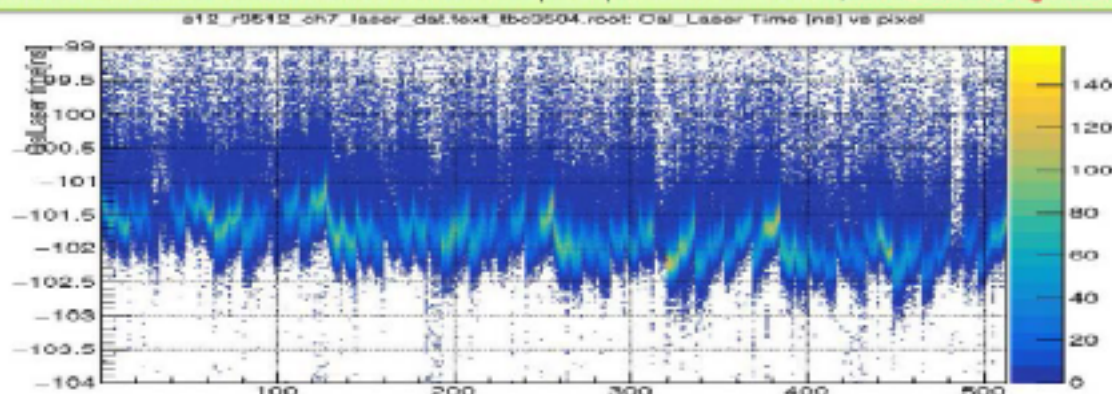
Scintillator paddles as trigger



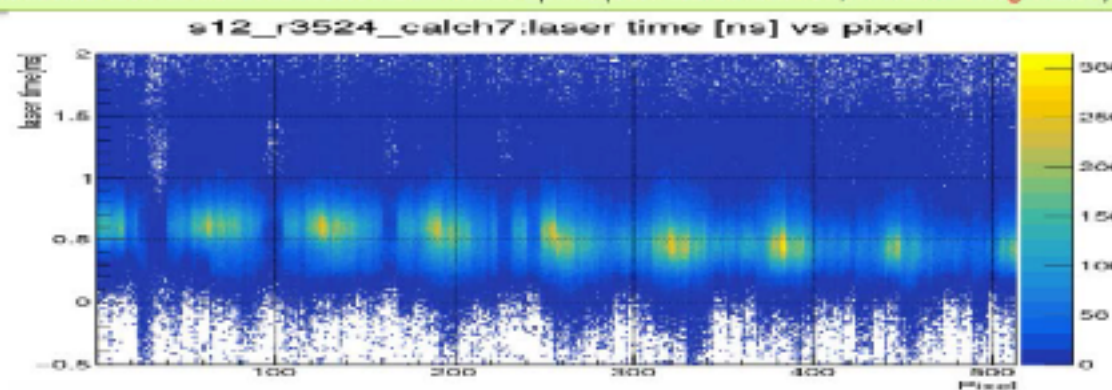
More Ongoing Work

Channel by Channel Time Alignment

DATA slot12-r3512: Laser time as a function of pixel (after TB correction, before time alignment)

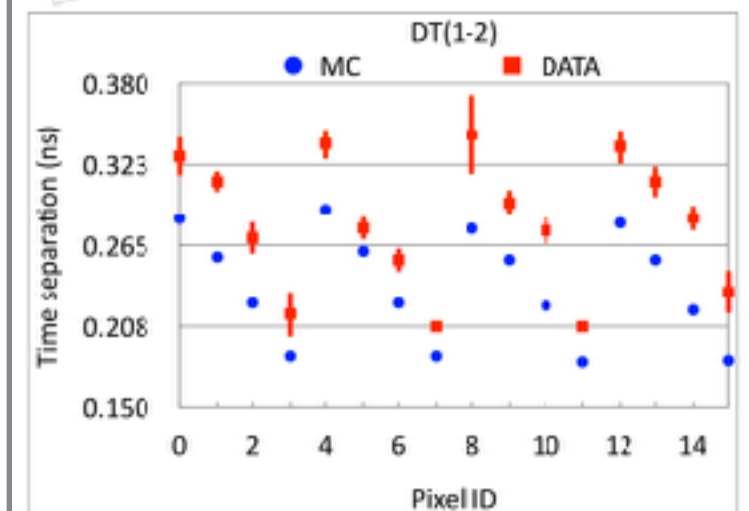


DATA slot12-r3512: Laser time as a function of pixel (after TB correction, after time alignment)



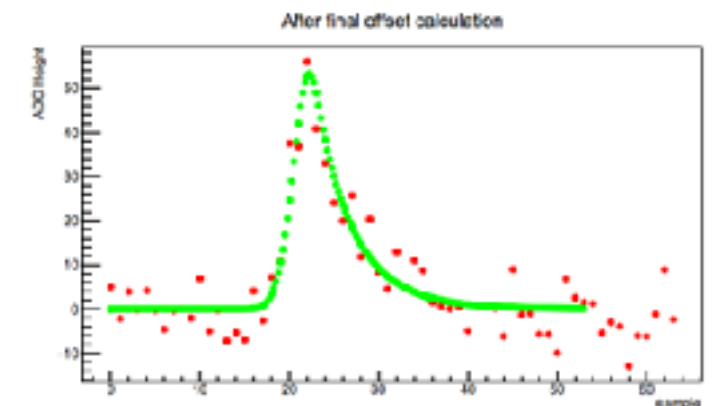
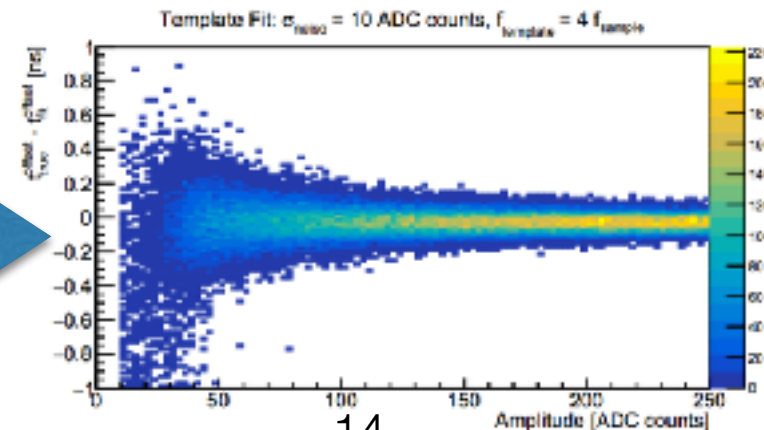
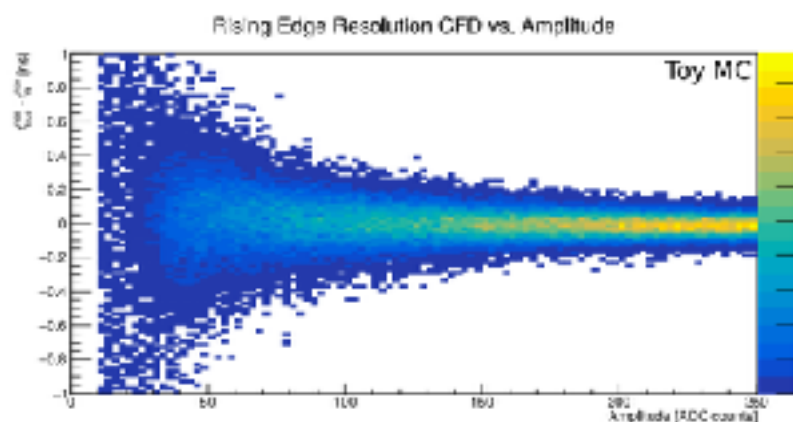
NOTE: Very Different Time Scales!

Understanding laser data



More from Prof. Gary Varner's talk

Improve low PMT gain: use template fit to waveform data



Summary

- iTOP is a new type of ring-image Cherenkov detector for Belle II. It provides particle identification ability in the barrel region.
- The construction of all modules took ~18 months. They're all installed to Belle II and cabled, ready for testing.
- The cosmic ray data and laser data have been taken and the analysis of these data is ongoing.
- Belle II will take physics data with all sub-detectors installed in late 2018.