

Readout and precision calibration of square meter sized micromegas detectors using the Munich cosmic ray facility

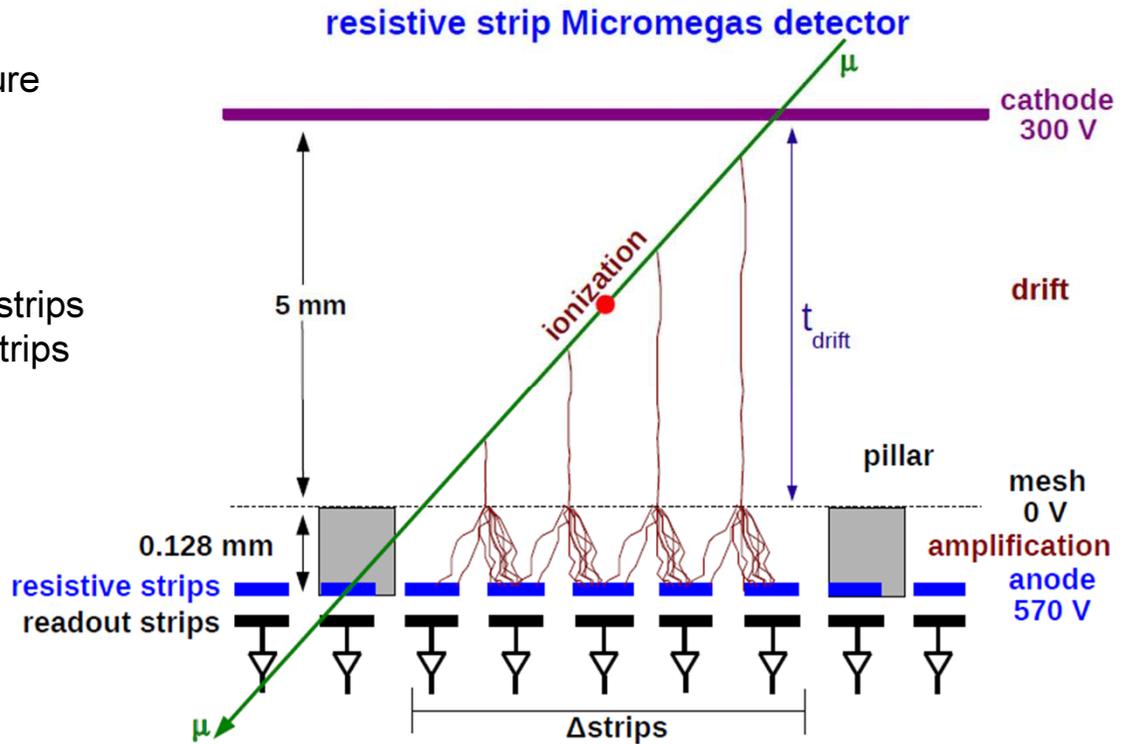
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Micromegas principles

- MICROMESH Gaseous Structure
- Electron drift region
- Amplification region
- Charge collection on resistive strips
- Charge detection on readout strips by capacitive coupling



Centroid method

$$x_{cen} = \frac{\sum_{strips} x_{strip} \cdot q_{strip}}{\sum_{strips} q_{strip}}$$

μ TPC method: angle reconstruction

$$\Rightarrow \Theta = \arctan\left(\frac{1}{slope_{fit}} \times \frac{pitch}{V_{drift}}\right)$$

$$slope_{fit} = \frac{t_{drift}}{\Delta strip}$$

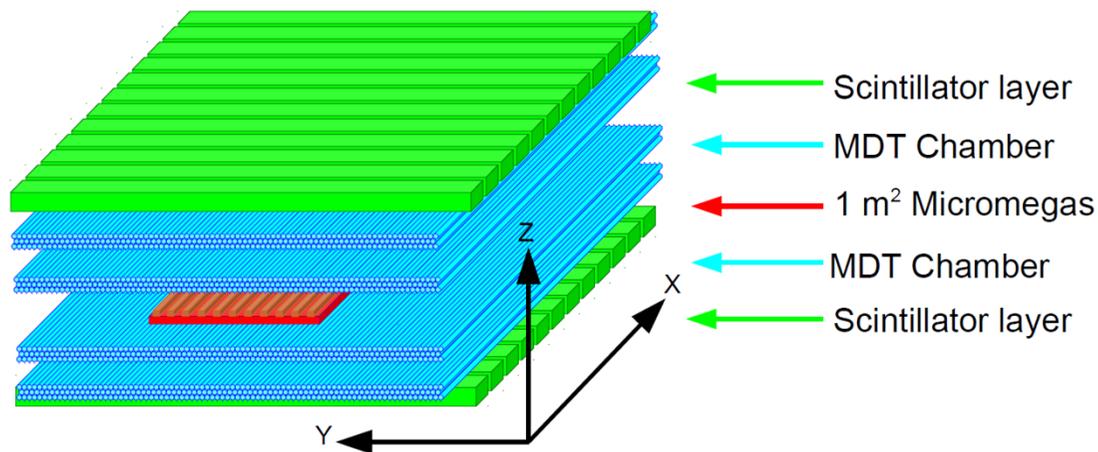
Large micromegas

- Small micromegas detectors up to 50 cm x 50 cm size are well understood and proven to work with high precision.
- Going to larger detectors requires precise assembly of multiple printed circuit boards in order to maintain the precision

⇒ A calibration method is needed



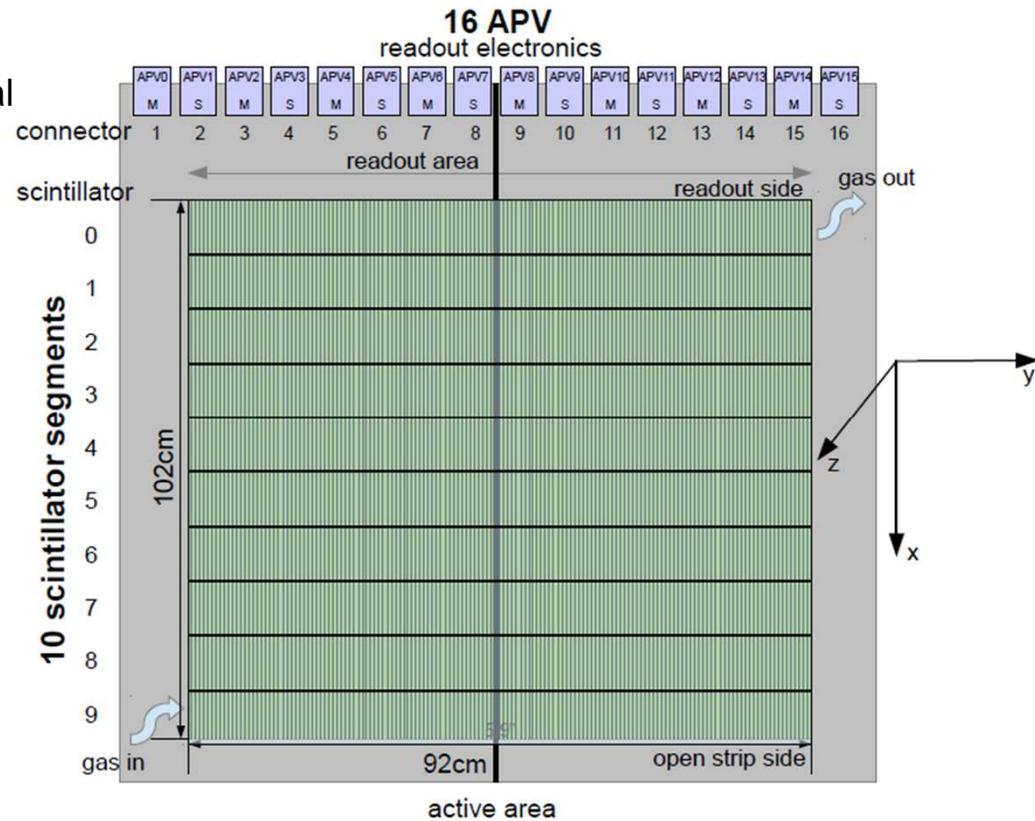
- Two Monitored Drift Tube chambers (MDT) => two reference tracks
- Two trigger scintillator Hodoscopes => second coordinate => segmentation of detector under test in 10 cm wide segments
- 34 cm iron absorber => $E_{\mu} > 600 \text{ MeV}$
- Active area of 9 m^2 , $\Theta \in [-30^{\circ}, 30^{\circ}]$



=> Investigation of the whole active area of 2-3 m² micromegas

1m² L1 micromegas detector

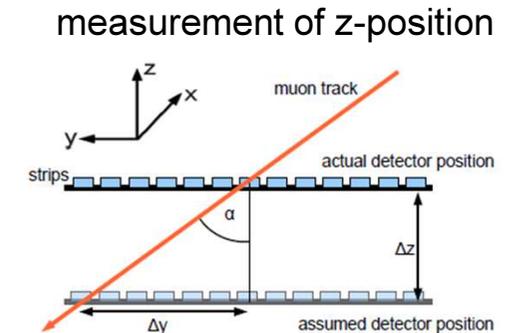
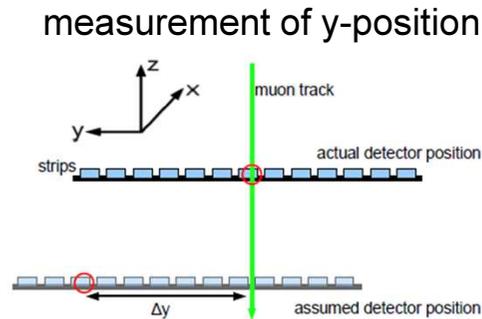
- Resistive strip technology
 - Active area: 0.92 x 1.02 m²
 - Two readout boards with 2048 strips total
 - Pitch: 0.45 mm
 - Ar:CO₂ 93:7 @ atm. Pressure
 - 16 APV25 frontend boards
57.6 mm wide (y - coordinate)
 - 10 scintillator segments
100 mm wide (x - coordinate)
- ⇒ Subdivision of detector in
16 APV * 10 scintillator
= 160 partitions



⇒ Calibration and alignment for each of the 160 partitions

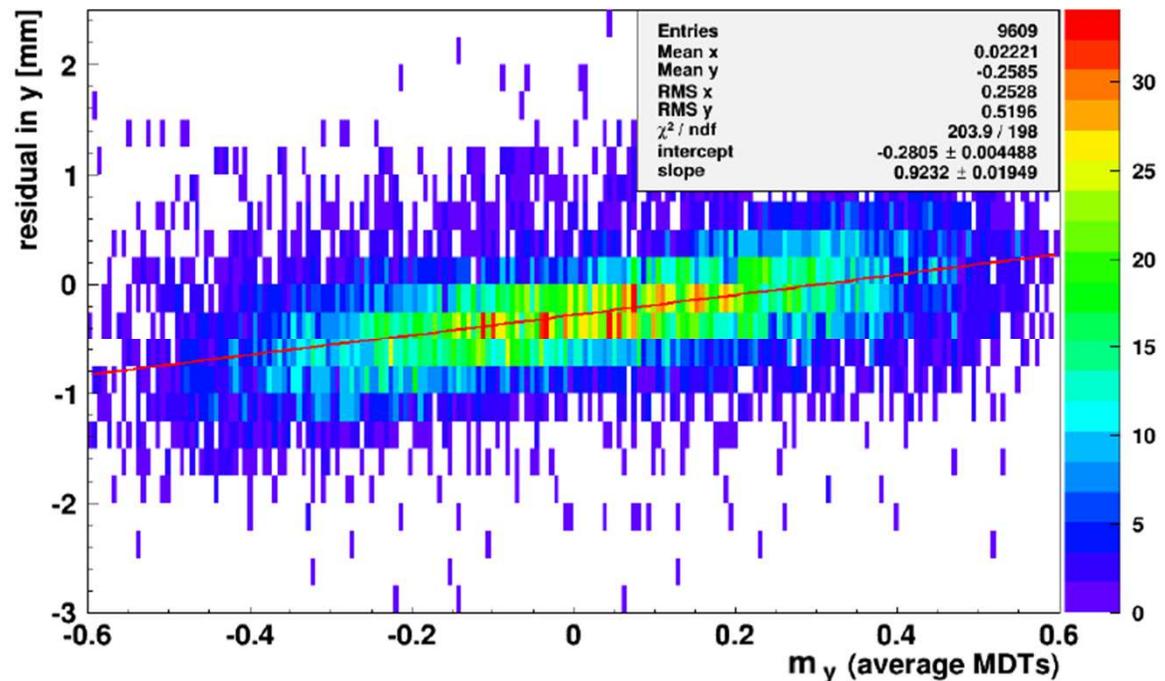
Calibration using reference tracks

- Measurement of y-position
(perpendicular tracks)
residual via centroid method
 $res = y_{\text{measured}} - y_{\text{predicted}}$
 $\Delta y = res$



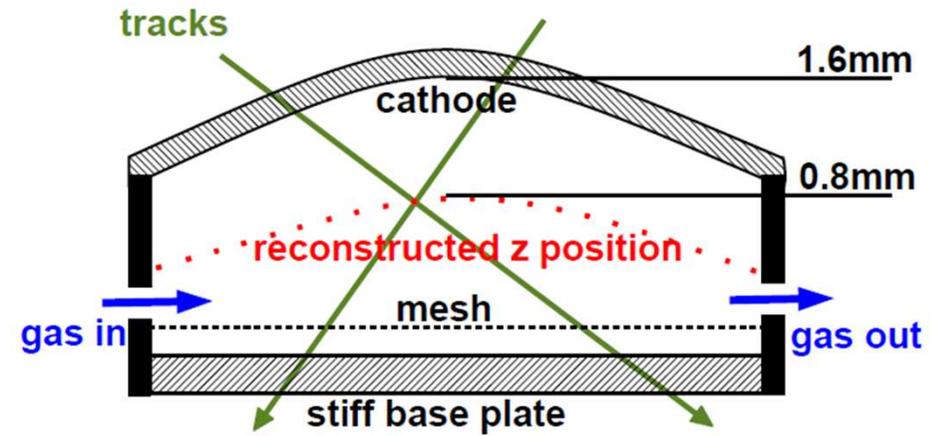
- Measurement of z-position
(inclined tracks)
 $\Delta z = res / \tan \alpha$
 $res = m_y * \Delta z$

- Fit with straight line
⇒ $\Delta z = \text{slope}$
 $\Delta y = \text{intercept}$

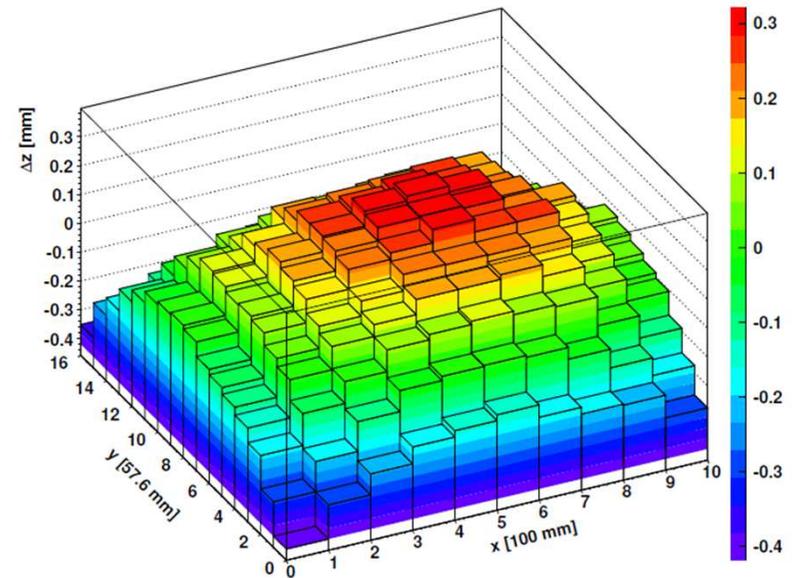
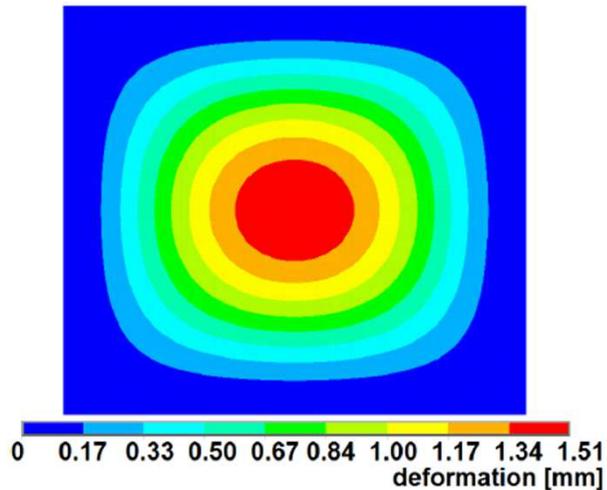


Chamber deformation in z

- Inclined muon tracks:
 - Drift gap deformation due to slight overpressure (10 mbar)
 - Maximum deviation from plane measured ~ 0.8 mm
 - \Rightarrow 1.6 mm cathode deformation due to stiff base plate
- Resolution better 0.1 mm



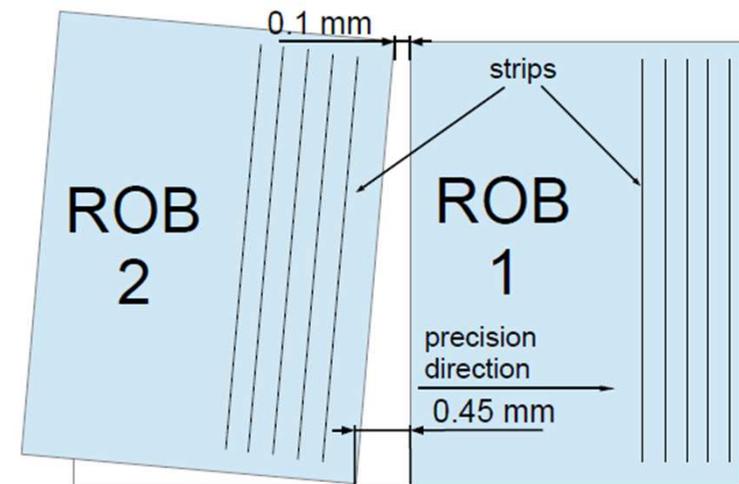
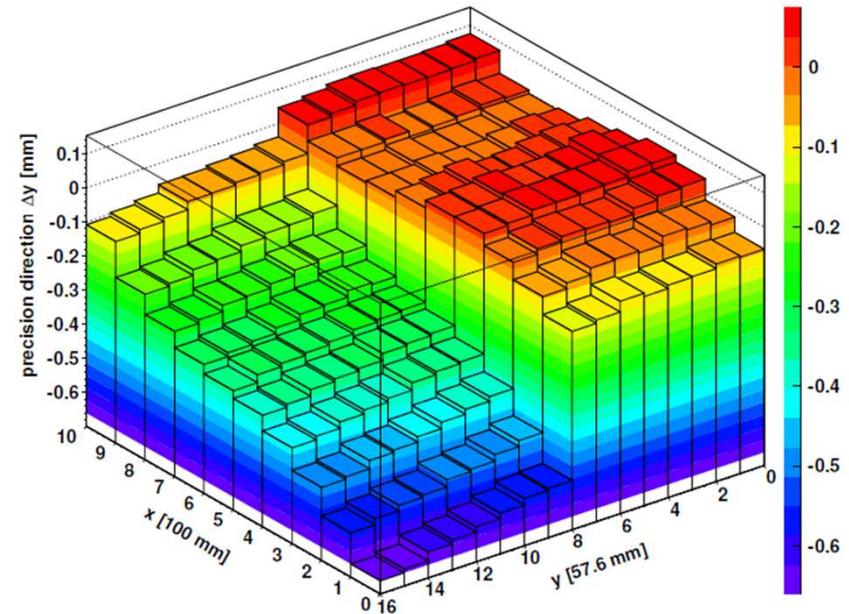
Finite elements simulation (ANSYS)



Readout board alignment

- L1 micromegas chamber:
 - 2 readout boards with resistive and readout strips
 - No dedicated alignment tooling used during the construction
 - Variation of gap size in between boards

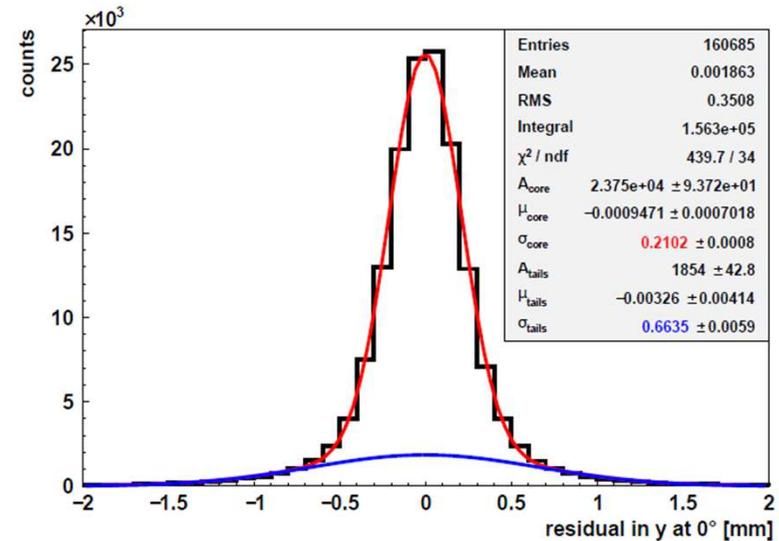
- Using perpendicular tracks:
 - Shift: $100\ \mu\text{m}$
 - Rotation: $350\ \mu\text{m}/\text{m}$
 - Determination of strip position within $15\ \mu\text{m}$



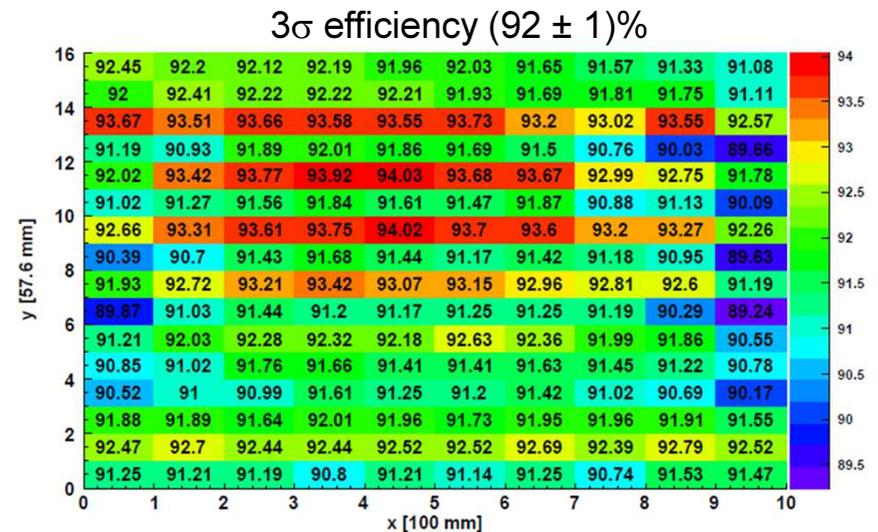
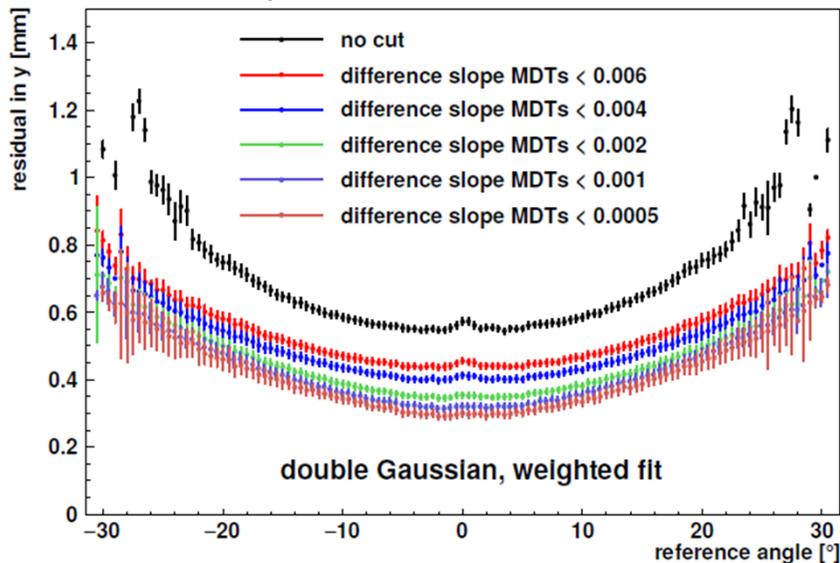
Spatial resolution

- Determination of spatial resolution:
 - Residual with centroid method
 - Fitted with double Gaussian function
 - Sigmas weighted by integral

- Spatial resolution limited by:
 - Multiple scattering of cosmic muons
 - Extrapolation of MDT tracks

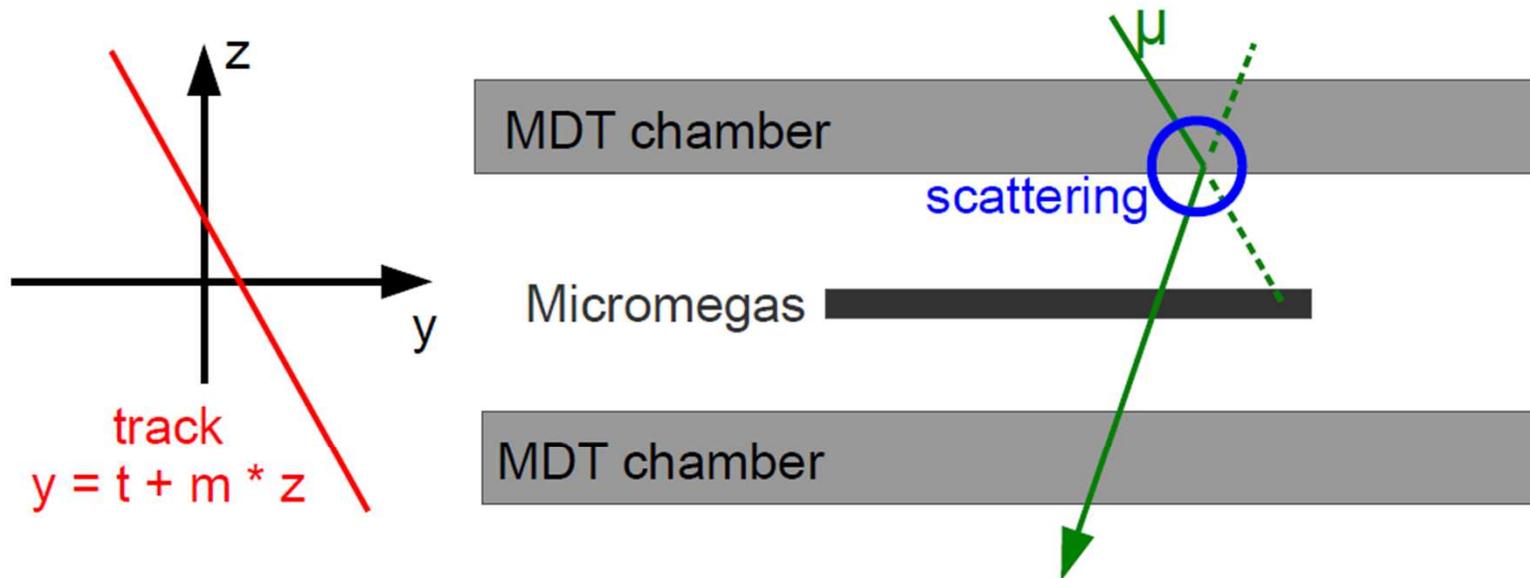


Cut on slope difference of reference detectors



Cosmic muon scattering

- Two track segments with different slope
⇒ reconstruct the deflection point in space



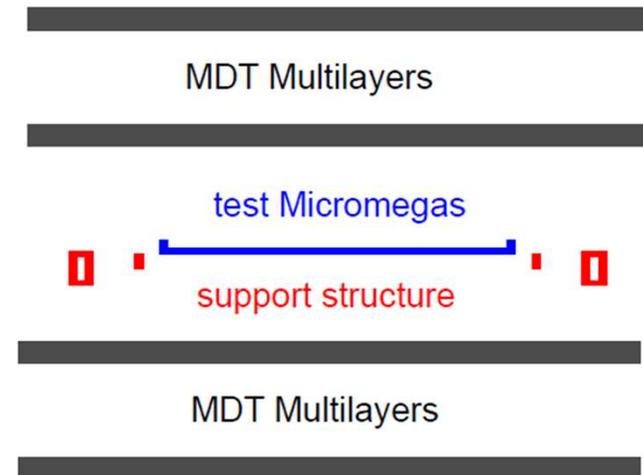
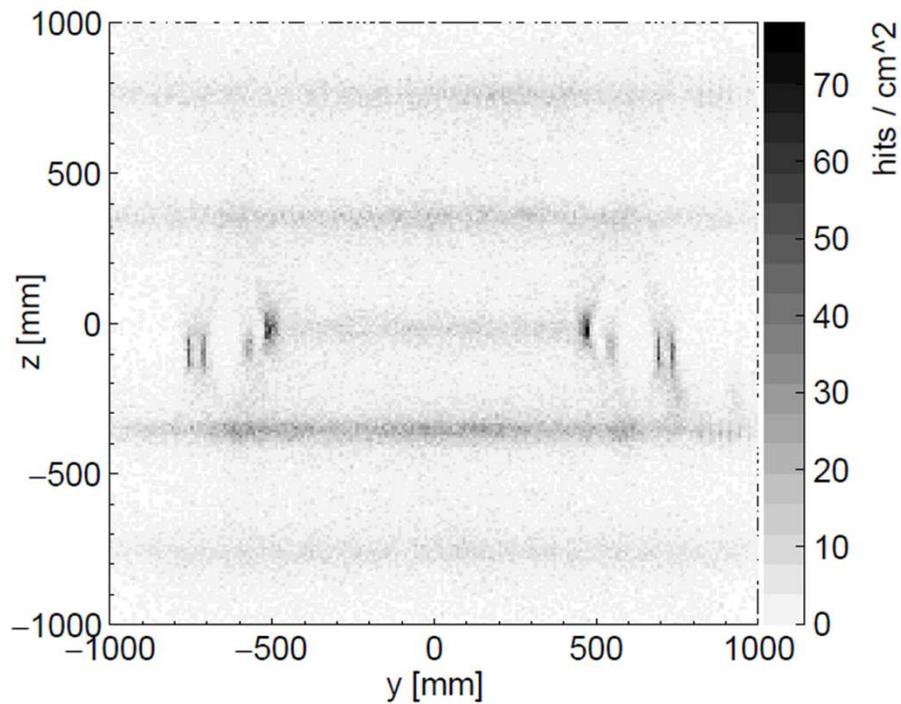
$$\Rightarrow z_{\text{scattered}} = \frac{t_{\text{track1}} - t_{\text{track2}}}{m_{\text{track2}} - m_{\text{track1}}}$$

$$\Rightarrow y_{\text{scattered}} = t_{\text{track1}/2} + m_{\text{track1}/2} \cdot z_{\text{scattered}}$$

Muon tomography in the y-z-plane

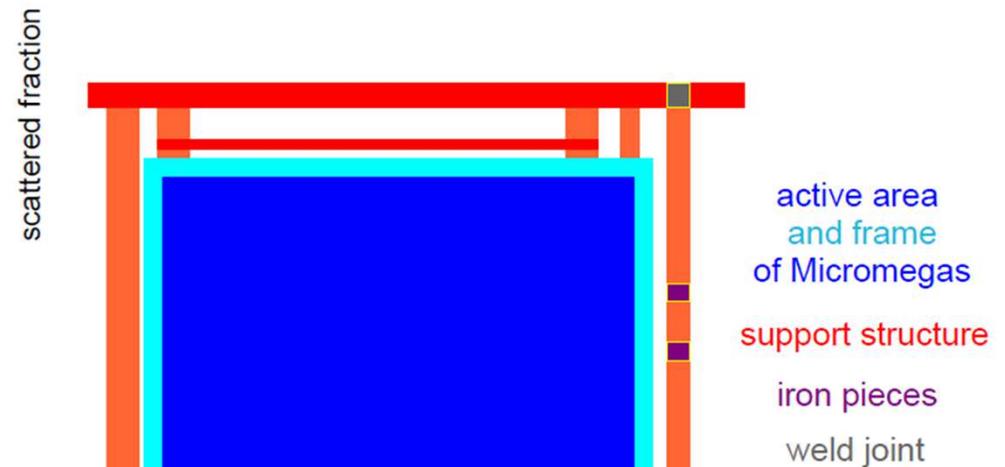
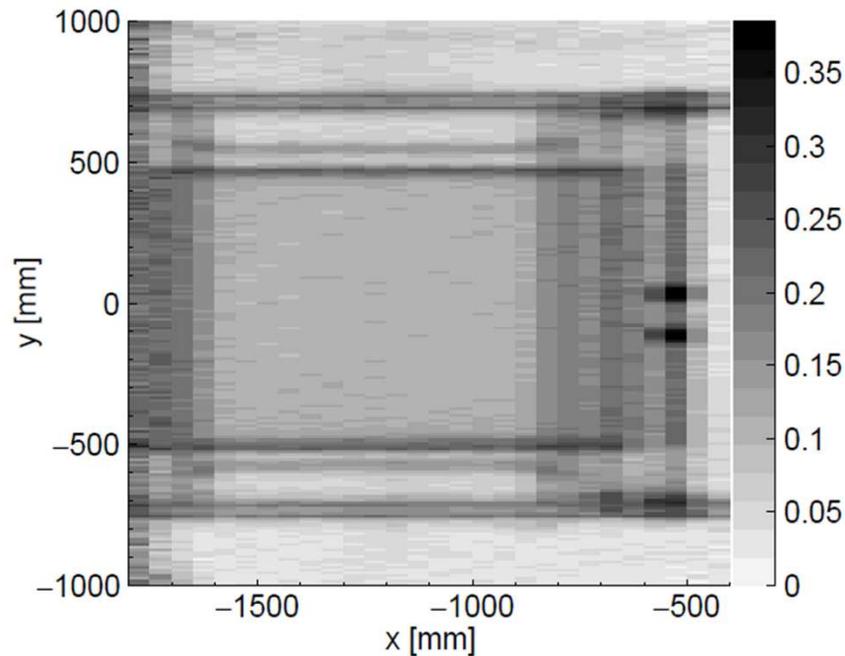
- Reconstruction of track deflection point in the y-z-plane
- Track slope difference $> 0,03$ selected

⇒ 2D image of the detectors and support structure



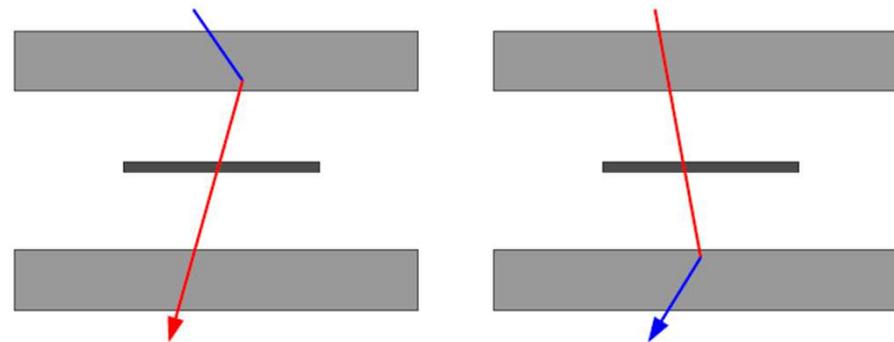
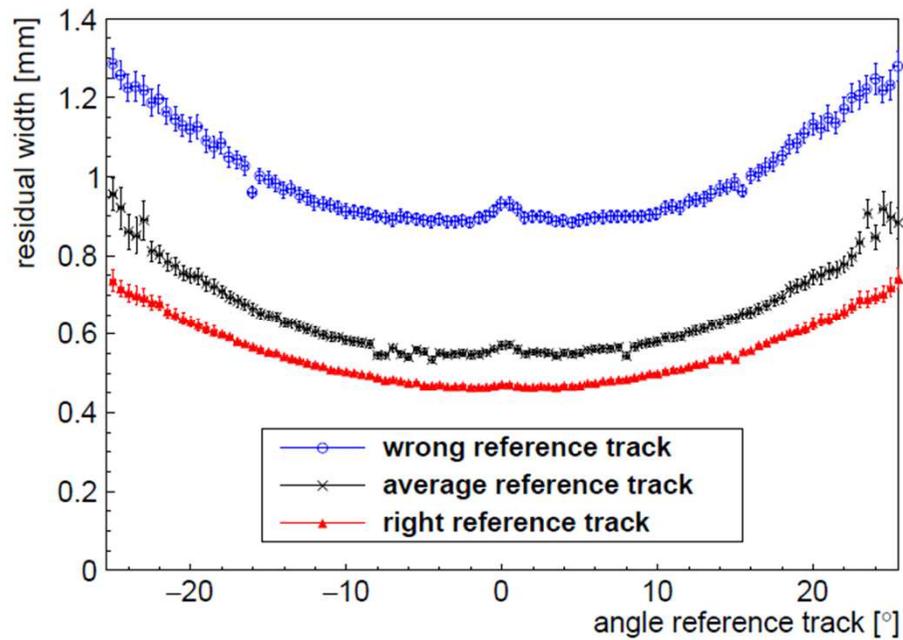
Muon tomography in the x-y-plane

- Coarse segmentation of the scintillator hodoscope along the x-axis
⇒ reconstruction of the scattering in the x-y-plane
- Only a slice in z is considered



Influence on spatial resolution

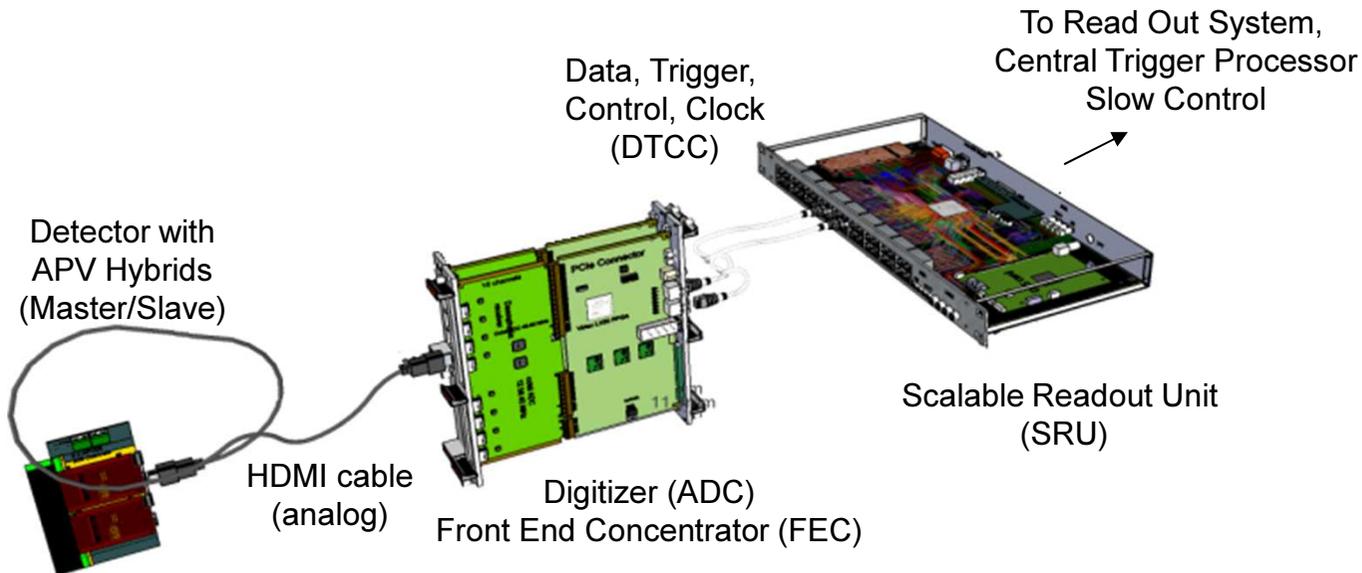
- Scattering point known
⇒ select just the track segments through the detector under test
- Better residual width distribution without loss of statistics



Readout synchronization

- MDT readout:
 - similar fashion as in the ATLAS detector
 - Clock and trigger distribution via optical TTC fiber

- Micromegas readout:
 - RD51 Scalable Readout System (SRS) with APV25 hybrids
 - Multiple FEC cards for larger detectors (>2048 channels)
 - ⇒ Including the optional Scalable Readout Unit (SRU) for synchronization

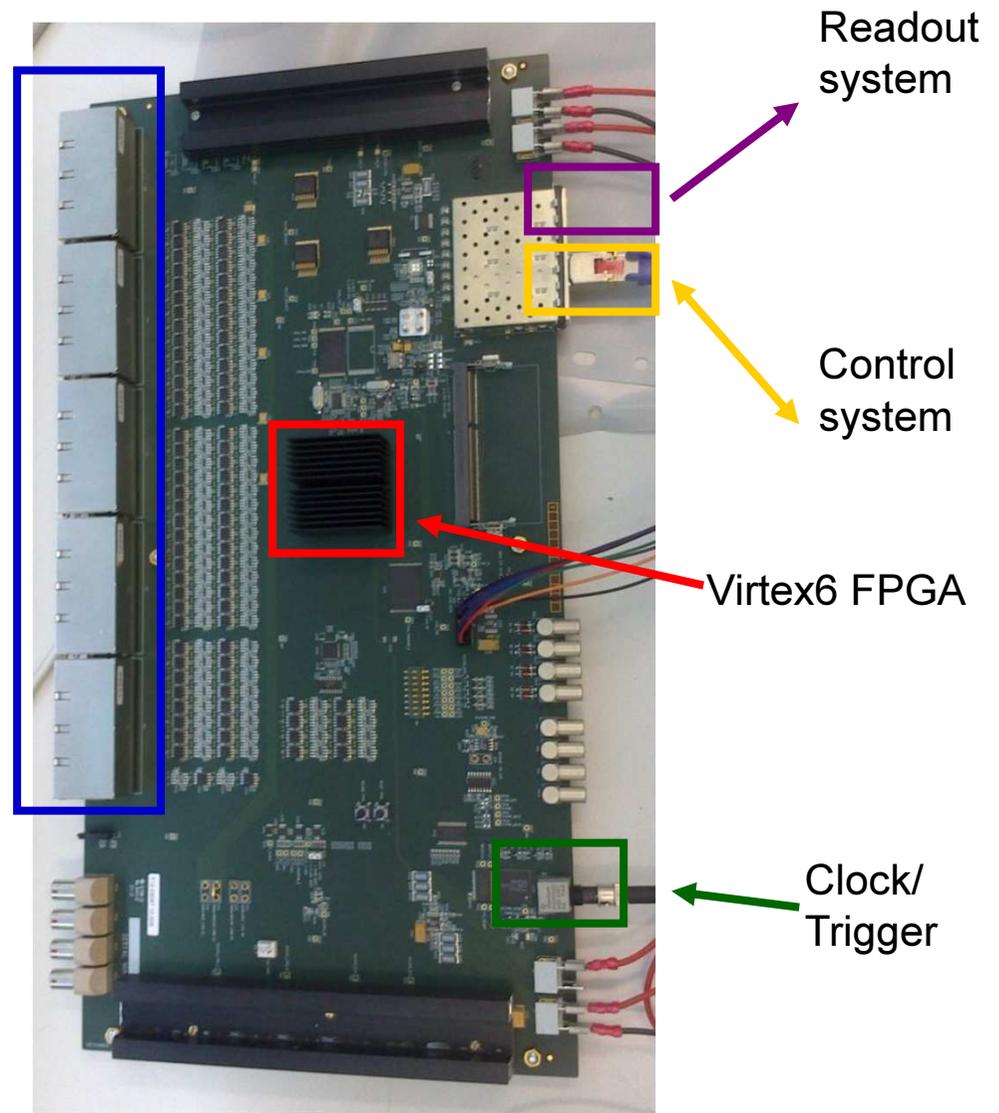


S. Martoiu, H. Muller, A. Tarazona, J. Toledo: Development of the scalable readout system for micro-pattern gas detectors and other applications, Journal of Instrumentation, Vol. 8 C03015 (2013)

Scalable Readout Unit (SRU)

Event building and
Interface to CRF infrastructure:

- Detector synchronization and data collection (**DTCC**, A. Martinez)
- Event building, data formatting and **ROD Firmware** (A. Zibell)
(correlation MDT \Leftrightarrow Micromegas)
- Data transmission to PC (ROS) via **SLINK** (M. della Volpe et al.)
- Slow control, data preview via **Ethernet**:
configuration data: Register settings on Frontend chip, SRS, SRU, ...
- Reception and distribution of triggers, synchronization (**TTC**)



- Calibration of the whole active area of large micromegas detectors in a cosmic ray facility with accuracy better than 30 μm
- Measurements include:
 - Homogeneity of efficiency and charge distribution
 - Determination of z position with 100 μm resolution
 - Calibration of readout strip position every 10 cm with an accuracy of 15 μm
- L1 micromegas (1 m^2)
 - Shift and rotation between readout boards calibrated
 - 1.6 mm deformation of the drift gap due to 10 mbar overpressure
 - Spatial resolution limited by multiple scattering, also in the MDTs
- Muon tomography with the system is possible with enough statistics
- Micromegas readout using APV25 chips and SRS electronics
- Synchronization in between multiple micromegas FEC cards and MDTs using a SRS SRU unit with custom firmware

128 preamplifier channels → Analogue pipeline buffer → Selected columns output

