

Vorticity in Heavy-Ion Collisions

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- 1 Vorticity in Hydrodynamics
 - Non-relativistic case
 - Relativistic case
- 2 Setup of the numerical simulations
 - Definition of the velocity field
 - ω_{1y} VS. ω_{2y}

Non-relativistic case

The vorticity (pseudo) vector is defined by,

$$\boldsymbol{\omega}_1(\mathbf{x}, t) = \nabla \times \mathbf{v}$$

The Euler equation for ideal fluid,

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} = -\frac{1}{\rho} \text{grad} p$$

It comes from the equation of motion $\rho \frac{d\mathbf{v}}{dt} = -\text{grad} p$,

- The change during dt in the velocity at a point fixed in space,

$$\frac{\partial \mathbf{v}}{\partial t} dt$$

- The difference between the velocities (at the same instant) at two points $d\mathbf{r}$ apart,

$$dx \frac{\partial \mathbf{v}}{\partial x} + dy \frac{\partial \mathbf{v}}{\partial y} + dz \frac{\partial \mathbf{v}}{\partial z} = (d\mathbf{r} \cdot \nabla) \mathbf{v}$$

Thus $d\mathbf{v} = \frac{\partial \mathbf{v}}{\partial t} dt + (d\mathbf{r} \cdot \nabla) \mathbf{v}$ dividing both sides by dt .

(1) Circulation conservation

$$\frac{d}{dt} \oint \mathbf{v} \cdot d\mathbf{x} = 0$$

(2) Helicity (pseudoscalar)field

$$h_f(\mathbf{x}, t) = \mathbf{v} \cdot \boldsymbol{\omega}_1 \rightarrow \text{helicity density}$$

$$\mathcal{H}_f = \int d^3\mathbf{x} h_f \rightarrow \text{total helicity} \rightarrow \text{conserved}$$

$$\frac{\partial \boldsymbol{\omega}_1}{\partial t} = \nabla \times (\mathbf{v} \times \boldsymbol{\omega}_1)$$

Relativistic case

A natural extension of the definition to relativistic fluid is,

$$\omega^\mu = \epsilon^{\mu\nu\rho\sigma} u_\nu \partial_\rho u_\sigma = \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} u_\nu \omega_{\rho\sigma}$$
$$\omega_{\mu\nu} = \partial_\mu u_\nu - \partial_\nu u_\mu \quad (\text{kinematic vorticity tensor})$$

It is worth writing down the components of ω_2^μ

- Spatial components

$$\omega_2 = \gamma^2 \omega_1 + \gamma^2 \mathbf{v} \times \partial_t \mathbf{v}$$

- time component

$$\omega_2^\mu = \gamma^2 \mathbf{v} \cdot \omega_1 = \mathbf{v} \cdot \omega_2$$

Thus, in the non-relativistic limit, $\omega_2^0 \rightarrow (h_f, \omega_1)$

In order to maintain the circulation conservation and helicity conservation, another definition of vorticity tensor has been introduced as,

$$\Omega_{\mu\nu} = \partial_\mu(Tu_\nu) - \partial_\nu(Tu_\mu)$$

The relativistic Euler equation for ideal fluid,

$$(\varepsilon + P) \frac{d}{d\tau} u^\mu = \nabla^\mu P$$

can be deduced to

$$\frac{d}{d\tau}(Tu^\mu) = \partial^\mu T$$

Thus, the relativistic circulation conservation,

$$\frac{d}{d\tau} \oint Tu_\mu dx^\mu = \oint \partial_\mu T dx^\mu = 0$$

The vorticity(pseudo) vector corresponding to $\Omega_{\mu\nu}$

$$\omega^\mu = \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} T u_\nu \Omega_{\rho\sigma} = T^2 \omega_2^\mu$$

Therefore, the integral of Ω^0 over space is conserved and we can identify Ω^0 as the conserved helicity density

If the fluid carries a conserved charge, one can define the vorticity tensor as

$$\begin{aligned}\tilde{\Omega}_{\mu\nu} &= \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} \partial_\rho (\beta u_\sigma) \\ \omega^\mu &= \tilde{\Omega}^{\mu\nu} u_\nu\end{aligned}$$

Setup of the numerical simulations

Definition of the velocity field

Velocity:

- Use program of Hydro to generate a velocity field
- Use a non-boost-invariant solution of relativistic hydrodynamics in 1+3 dimensions

$$u_t = -\frac{1}{t^2 - x_{\perp}^2} \left(\frac{t^2 (t^2 - x_{\perp}^2 + L^2 + z^2)}{\sqrt{(t^2 - x_{\perp}^2 - L^2 - z^2)^2 + 4L^2 (t^2 - x_{\perp}^2)}} - \alpha x_{\perp}^2 \right)$$
$$\vec{u}_{\perp} = \frac{t \vec{x}_{\perp}}{t^2 - x_{\perp}^2} \left(\frac{t^2 - x_{\perp}^2 + L^2 + z^2}{\sqrt{(t^2 - x_{\perp}^2 - L^2 - z^2)^2 + 4L^2 (t^2 - x_{\perp}^2)}} - \alpha \right)$$
$$u_z = \frac{2zt}{\sqrt{(t^2 - x_{\perp}^2 - L^2 - z^2)^2 + 4L^2 (t^2 - x_{\perp}^2)}}$$

Velocity Field

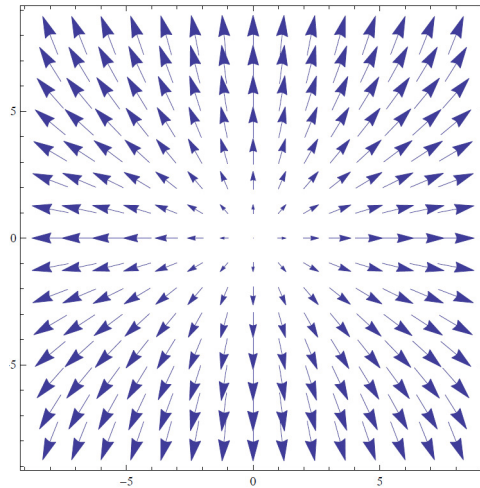


Figure: (v_z, v_x) in xoz plane

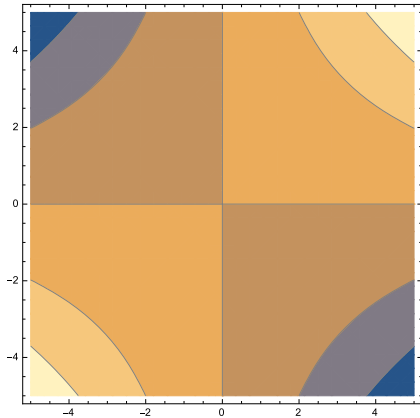


Figure: ω_y in xoz plane

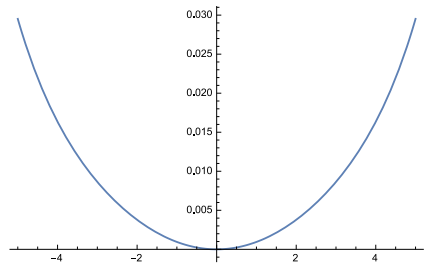


Figure: ω_y when $x=z$



THANK YOU