

# *Search for Graviton in BESSIII: $J/\Psi \rightarrow \gamma + \text{graviton}$*

Wen Chen (IHEP)

*In collaboration with  
D. Bai (NJU), and Y. Jia (IHEP)*

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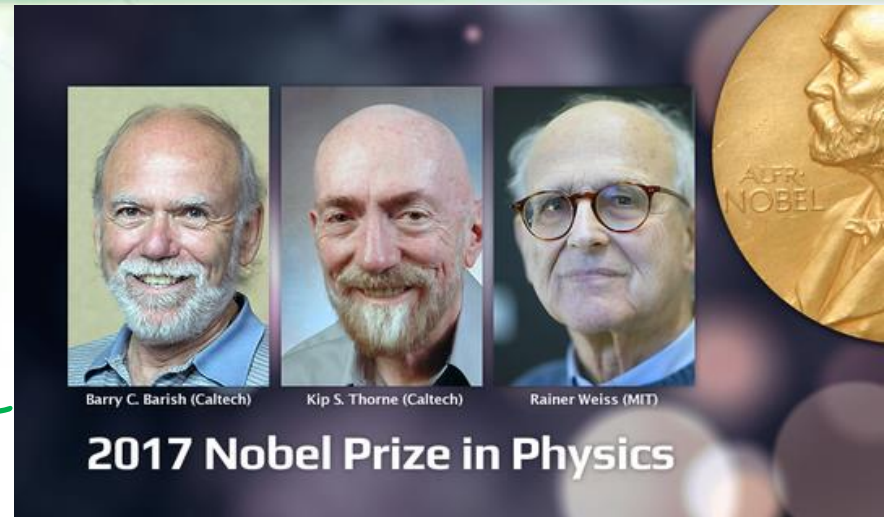
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# *Outline*

- *Motivation*
- *General Relativity as an Effective Field Theory*
- *Calculations*
- *Results and Phenomenology*
- *Conclusion*

# Motivation

- Gravitational waves from binary black hole mergers and the binary neutron star inspiral are recently observed directly on earth by LIGO Scientific and Virgo collaboration, which initiates a new era for fundamental physics and astronomy. (Abbott et al., PRL (2016), Abbott et al., PRL (2017))
  - Physical properties of graviton remain untouched.
  - A full theory of quantum gravity is far from available.
  - General relativity as an effective theory can be used in making predictions. (Donoghue, PRD (1994) .)
  - Various predictions based on perturbative quantum gravity are reported: QG corrections to Newton's law of universal gravitation, QG corrections to bending of light, etc. (Donoghue, PRL (1994), Bjerrum-Bohr, et al., PRL(2015))
- $$\theta_\gamma - \theta_\varphi = \frac{8(\text{bu}^\gamma - \text{bu}^\varphi)}{\pi} \frac{G^2 \hbar M}{b^3} \quad \text{bu}^\gamma - \text{bu}^\varphi = -\frac{43}{20}$$
- $J/\Psi \rightarrow \gamma + \text{graviton}$  provides a possible channel to study the quantum nature of graviton.



# General Relativity as an Effective Field Theory

background field method: DeWitt, PR (1967)

$$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu} \quad g^{\mu\nu} = \eta^{\mu\nu} - \kappa h^{\mu\nu} + \kappa^2 h^{\mu\lambda} h_{\lambda}^{\nu} + \dots$$

classical background

quantum fluctuation

fermion-gravity coupling:

$$\bar{\Psi} \gamma^{\mu} \partial_{\mu} \Psi \quad \longrightarrow \quad \bar{\Psi} \gamma^{\mu} D_{\mu} \Psi \quad D_{\mu} \Psi \equiv \partial_{\mu} \Psi + \frac{1}{2} \omega_{\mu ab} \sigma^{ab} \Psi,$$

spin connection:

$$\omega_{\mu}^{ab} = \frac{1}{2} (e^{[a\nu} \partial_{[\mu} e_{\nu]}^b) + e^{a\rho} e^{b\sigma} \partial_{[\sigma} e_{c\rho]} e_{\mu}^c$$

$$g_{\mu\nu} = e_{\mu a} e_{\nu}^a$$

$$e_{\mu a} = \bar{e}_{\mu a} + \kappa c_{\mu a}$$



$$c_{\{\mu\nu\}} \approx h_{\nu\nu} + \dots$$

## Effective Lagrangian:

$$\mathcal{L} = \mathcal{L}_{VK} + \mathcal{L}_{GGH} + \mathcal{L}_{AAH} + \mathcal{L}_{FK} + \mathcal{L}_{FFAH} + \mathcal{L}_{FFGH} + \mathcal{L}_{FFH} + \mathcal{L}_{HK} + \dots$$

$$\mathcal{L}_{VK} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}G_{\mu\nu}^a G^{a\mu\nu},$$

$$\mathcal{L}_{GGH} = \kappa \left\{ -\frac{1}{4}h_{\alpha}^{\alpha} (\partial_{\mu}G_{\nu}^a \partial^{\mu}G^{a\nu} - \partial_{\nu}G_{\mu}^a \partial^{\mu}G^{a\nu}) \right. \\ \left. + \frac{1}{2}h_{\rho\sigma} (\partial_{\alpha}G^{a\sigma} \partial^{\alpha}G^{a\rho} + \partial^{\sigma}G_{\alpha}^a \partial^{\rho}G^{a\alpha} - \partial_{\alpha}G^{a\sigma} \partial^{\rho}G^{a\alpha} - \partial^{\sigma}G_{\alpha}^a \partial^{\alpha}G^{a\rho}) \right\}$$

$$\mathcal{L}_{AAH} = \kappa \left\{ -\frac{1}{4}h_{\alpha}^{\alpha} (\partial_{\mu}A_{\nu} \partial^{\mu}A^{\nu} - \partial_{\nu}A_{\mu} \partial^{\mu}A^{\nu}) \right. \\ \left. + \frac{1}{2}h_{\rho\sigma} (\partial_{\alpha}A^{\sigma} \partial^{\alpha}A^{\rho} + \partial^{\sigma}A_{\alpha} \partial^{\rho}A^{\alpha} - \partial_{\alpha}A^{\sigma} \partial^{\rho}A^{\alpha} - \partial^{\sigma}A_{\alpha} \partial^{\alpha}A^{\rho}) \right\},$$

$$\mathcal{L}_{FK} = i\bar{\psi}\not{D}\psi - m\bar{\psi}\psi,$$

$$\mathcal{L}_{FFAH} = \kappa \left\{ -\frac{1}{3}e\bar{\psi}\gamma^{\mu}\psi (h_{\mu\nu} - h_{\alpha}^{\alpha}\eta_{\mu\nu}) A^{\nu} \right\},$$

$$\mathcal{L}_{FFGH} = \kappa \left\{ -\frac{1}{3}e\bar{\psi}\gamma^{\mu}T^a\psi (h_{\mu\nu} - h_{\alpha}^{\alpha}\eta_{\mu\nu}) G^{a\nu} \right\},$$

$$\mathcal{L}_{FFH} = \kappa \left\{ \frac{1}{2}h_{\alpha}^{\alpha} (i\bar{\psi}\gamma^{\mu}\partial_{\mu}\psi - m\bar{\psi}\psi) - \frac{i}{2}\bar{\psi}\gamma^{\nu}\partial^{\mu}\psi h_{\mu\nu} \right. \\ \left. + \frac{i}{2}\bar{\psi} \left( \frac{1}{2}\gamma^{\rho}\eta^{\mu\nu} - \frac{1}{2}\gamma^{\nu}\eta^{\mu\rho} \right) \psi \partial_{\rho}h_{\mu\nu} \right\},$$



# Calculations

*NRQCD factorization:*

$$\mathcal{M}(J/\Psi \rightarrow \gamma h) = (\mathcal{A}_\mu + \langle v^2 \rangle \mathcal{B}_\mu) \langle 0 | \chi^\dagger \sigma^\mu \psi | J/\Psi \rangle = \epsilon_\psi^\mu (\mathcal{A}_\mu + \langle v^2 \rangle \mathcal{B}_\mu) \sqrt{\langle \mathcal{O} \rangle_\psi} \equiv \mathcal{C} \sqrt{\langle \mathcal{O} \rangle_\psi},$$

*perturbative matching:*

$$\begin{aligned} \mathcal{M}(c\bar{c} \rightarrow \gamma h) = & \sqrt{2N_c} \epsilon_{h\mu\nu}^* \left( C^{(a)} \epsilon_\gamma^{*\mu} \epsilon_\psi^\nu + C^{(b)} P \cdot \epsilon_\gamma^* k \cdot \epsilon_\psi P^\mu P^\nu + C^{(c)} P \cdot \epsilon_\gamma^* P^\mu \epsilon_\psi^\nu \right. \\ & \left. + C^{(d)} k \cdot \epsilon_\psi P^\mu \epsilon_\gamma^{*\nu} + C^{(e)} \epsilon_\gamma^* \cdot \epsilon_\psi P^\mu P^\nu \right), \end{aligned}$$

*Ward identities leads to:*

$$C^{(c)} = \frac{2C^{(a)}}{M^2}, \quad C^{(b)} = -\frac{4C^{(a)}}{M^2}, \quad C^{(d)} = \frac{2C^{(a)}}{M^2}, \quad C^{(e)} = 0.$$

$$\mathcal{C} = C^{(a)} \epsilon_{h\mu\nu}^* \left( \epsilon_\gamma^{*\mu} \epsilon_\psi^\nu - \frac{4}{M^2} P \cdot \epsilon_\gamma^* k \cdot \epsilon_\psi P^\mu P^\nu - \frac{2}{M^2} P \cdot \epsilon_\gamma^* P^\mu \epsilon_\psi^\nu + \frac{2}{M^2} k \cdot \epsilon_\psi P^\mu \epsilon_\gamma^{*\nu} \right).$$

*helicity amplitude: ( $\mathcal{C}_{\lambda_\psi \lambda_\gamma \lambda_h}$ )*

$$\mathcal{C}_{1\pm 1\pm 2} = \mathcal{C}_{-1\mp 1\mp 2} = \frac{1}{2} C^{(a)} (-1 \pm \cos \theta);$$

$$\mathcal{C}_{0\pm 1\pm 2} = \mp \frac{1}{\sqrt{2}} C^{(a)} \sin \theta.$$

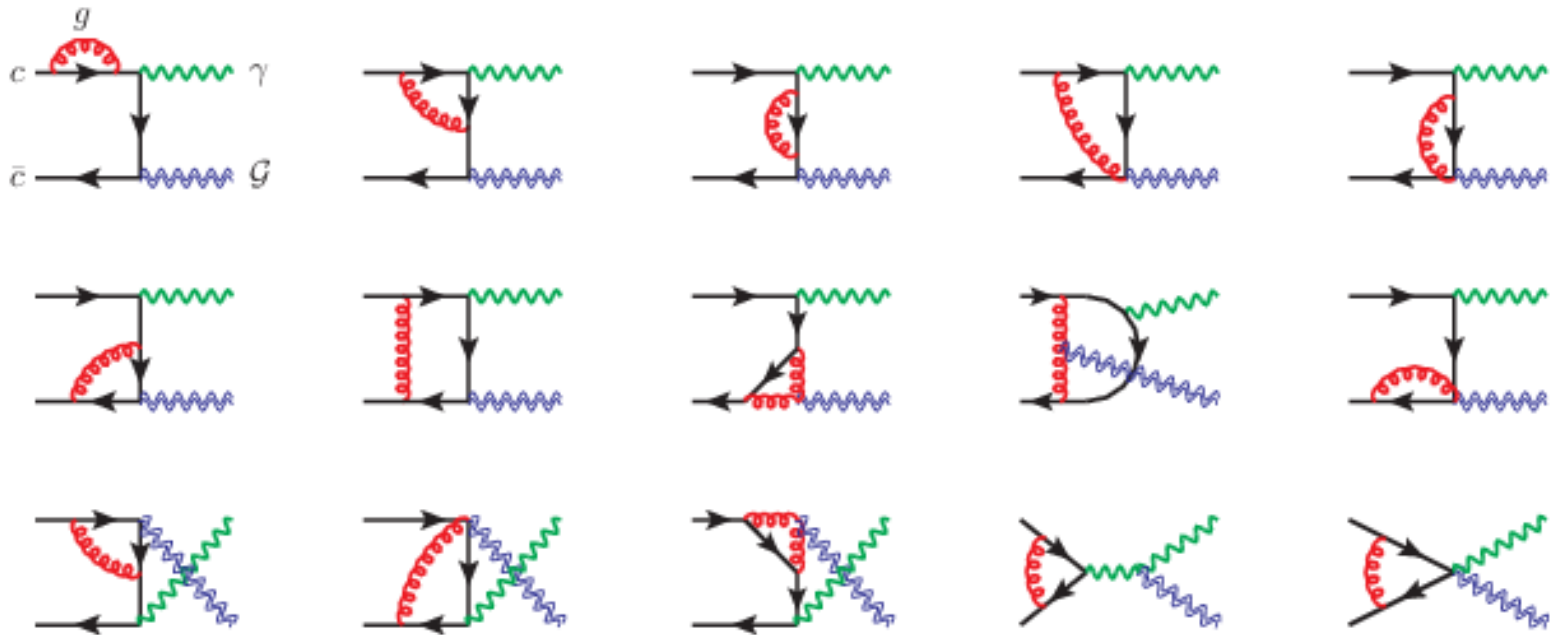
*decay rate:*

$$\Gamma(J/\Psi \rightarrow \gamma h) = \int d\Omega \frac{\pi}{4} \sum_{\text{pol}} |\mathcal{M}(J/\Psi \rightarrow \gamma h)|^2 = \frac{2\pi^2}{3} |C^{(a)}|^2 \langle \mathcal{O} \rangle_\psi$$

# Results



representative diagrams at leading order in  $\alpha_s$



representative diagrams at next-to-leading order in  $\alpha_s$

*analytic results:*  $C^{(a)} = \left( -\frac{4 \log 2 - 1}{8\pi} C_F \alpha_s + \frac{1}{6} \langle v^2 \rangle \right) Q_c e \kappa,$

*numerical inputs: (PDG and Bodwin et al., PRD (2008))*

$$\begin{aligned} \alpha &= 7.30 \times 10^{-3}, & \alpha_s(m) &= 0.31, \\ \kappa &= 8.21 \times 10^{-19} \text{ GeV}^{-1}, & M &= 3.096 \text{ GeV} \\ \Gamma_{J/\Psi} &= 9.29 \times 10^{-5} \text{ GeV}, \end{aligned}$$

$$\langle \mathcal{O} (^3S_1) \rangle_\psi = 0.440 \text{ GeV}^3, \quad \langle v \rangle_\psi^2 = 0.225.$$

*numerical results:*

$$\Gamma(J/\Psi \rightarrow \gamma h) = 5.5 \times 10^{-42} \text{ GeV}$$

$$\text{Br}(J/\psi \rightarrow \gamma + \text{graviton}) = 5.9 \times 10^{-38}$$



## massive gravity

$$\tilde{C}_{1\pm 1\pm 2} = \tilde{C}_{-1\mp 1\mp 2} = \frac{e\kappa Q_c m_h^2}{4m_\psi^2} (1 \mp \cos \theta);$$

$$\tilde{C}_{0\pm 1\pm 2} = \pm \frac{e\kappa Q_c m_h^2}{2\sqrt{2}m_\psi^2} \sin \theta;$$

$$\tilde{C}_{1\pm 1\pm 1} = -\tilde{C}_{-1\mp 1\mp 1} = -\frac{e\kappa Q_c m_h}{4m_\psi} \sin \theta;$$

$$\tilde{C}_{0\pm 1\pm 1} = -\frac{e\kappa Q_c m_h}{2\sqrt{2}m_\psi} \cos \theta;$$

$$\tilde{C}_{1\pm 10} = \tilde{C}_{-1\mp 10} = \frac{e\kappa Q_c}{4\sqrt{6}} (1 \pm \cos \theta);$$

$$\tilde{C}_{0\pm 10} = \mp \frac{e\kappa Q_c}{4\sqrt{3}} \sin \theta,$$

*vDVZ discontinuity*



$$\Gamma(J/\Psi \rightarrow \gamma \tilde{h}) = \frac{\pi^2}{36} e^2 \kappa^2 Q_c^2 \langle \mathcal{O} \rangle_\psi = 3.3 \times 10^{-39} \text{ GeV}$$

# Conclusion

*As an application of perturbative quantum gravity and NRQCD, we calculated the NLO QCD correction and relativistic correction to the process  $J/\Psi \rightarrow \gamma + \text{graviton}$ . The branch ratio is about  $5.9 \times 10^{-38}$ . We also generalized the calculations to the case of massive gravity. It is found that the longitudinal polarizations are not decoupled at massless limit, which reveals the well-known vDVZ discontinuity.*

*Thanks for your attention!*