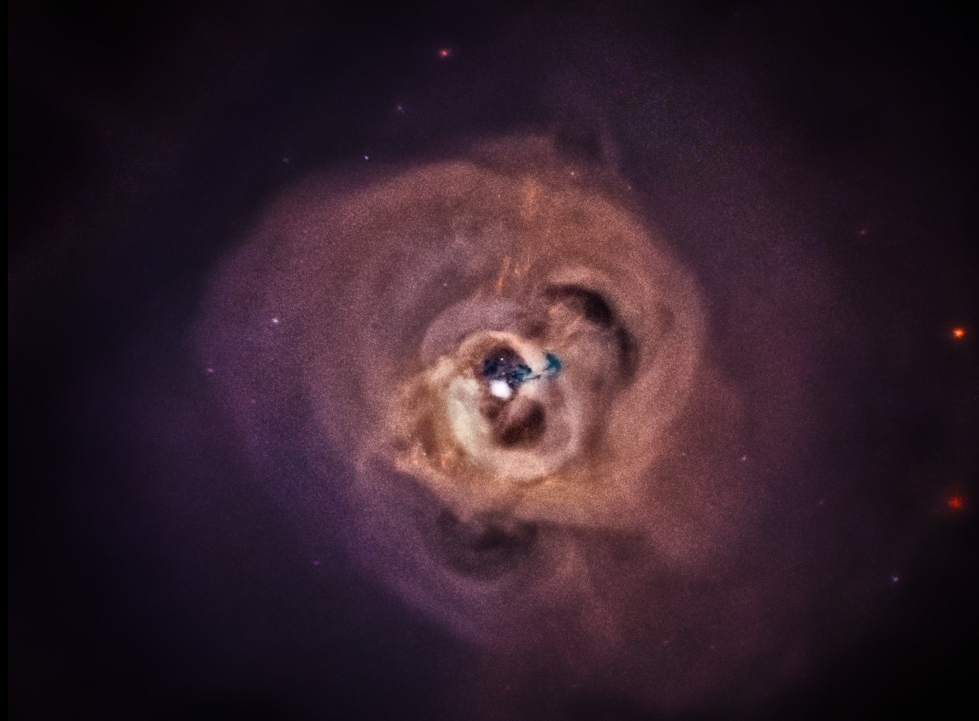


Searching for Dark Matter with X-ray lines



Perseus Cluster
(Chandra)



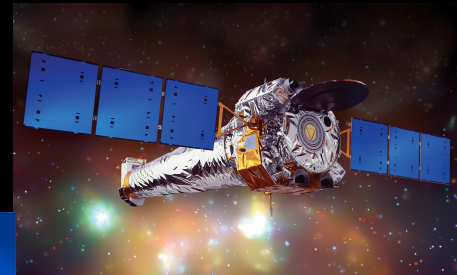
Kenny, Chun Yu Ng (吳震宇)
Weizmann Institute of Science



X-ray line Searches of Dark Matter

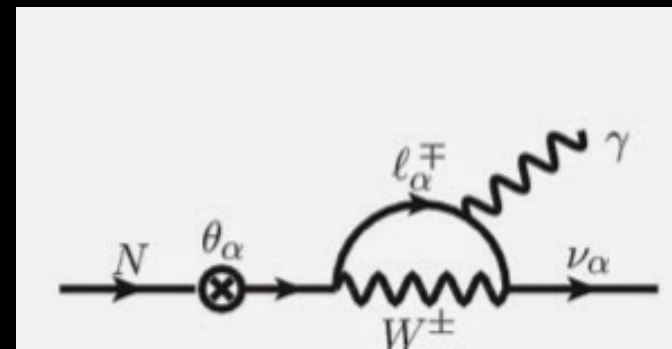
- Particle Dark matter Identification
 - Collider
 - Direct detection
 - Indirect detection
- Well Motivated Candidates
 - *Sterile Neutrino (keV)*
 - Axion-like Dark Matter
 - Gravitino
 - Exciting Dark Matter
 - ++++++
- Line signal (smoking gun signal)

Chandra (1999 -)



Suzaku (2005 - 2015)

XMM Newton (1999 -)

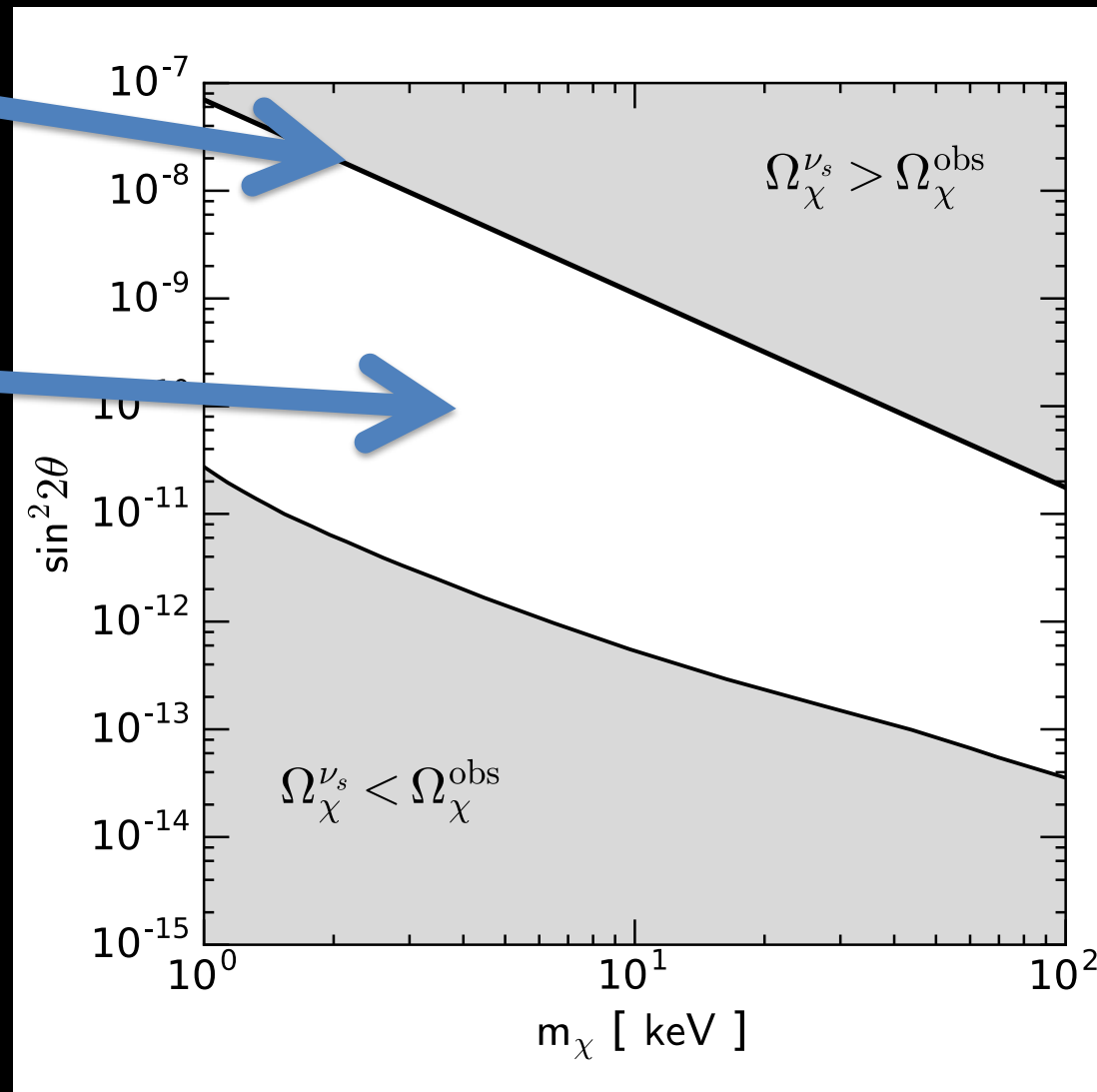


Sterile Neutrino Dark Matter Production

- Dodelson-Widrow 1994

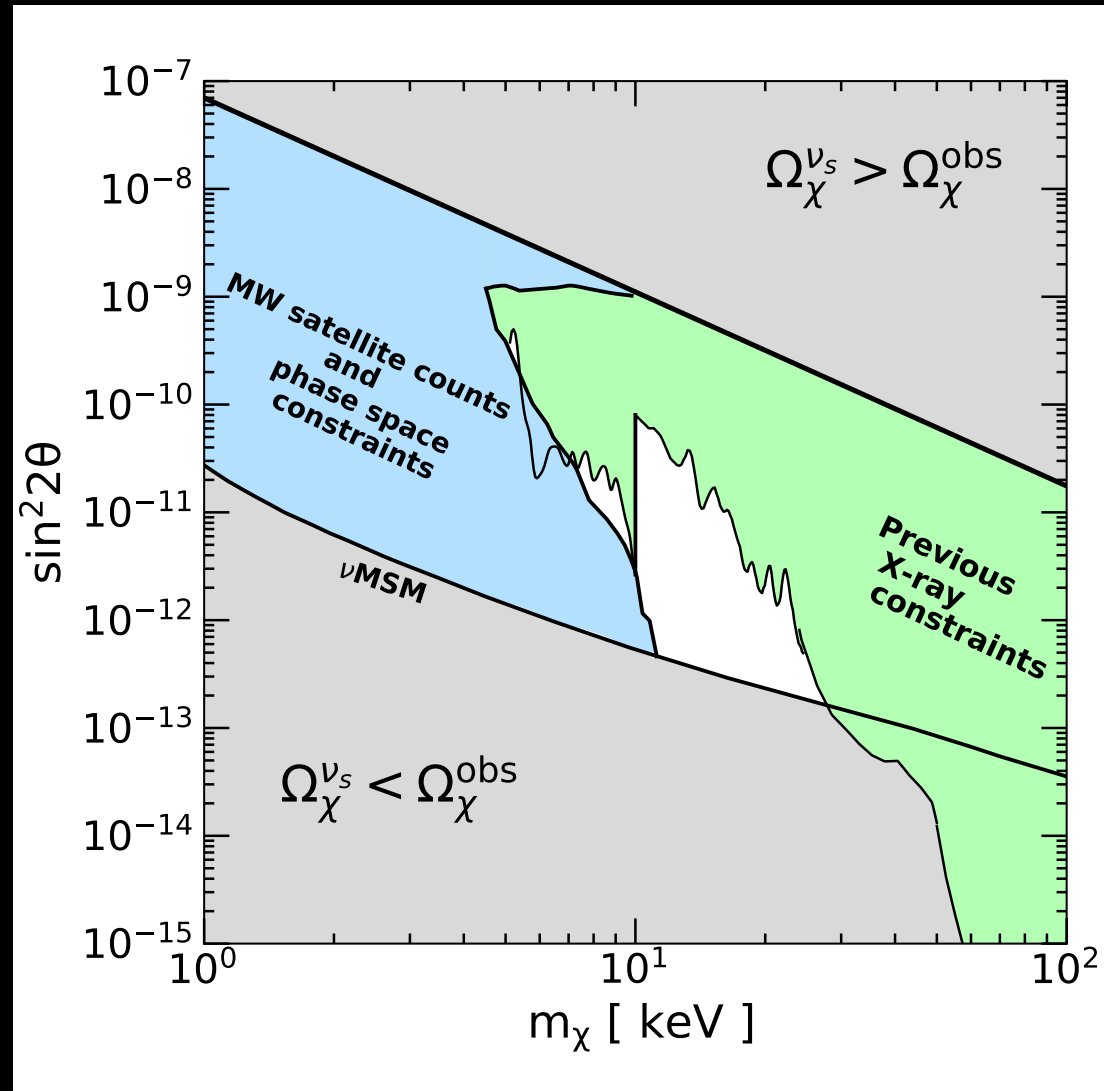
$$\Omega_4 h^2 \simeq 0.3 \frac{\sin^2 2\theta}{10^{-8}} \left(\frac{m_4}{10 \text{keV}} \right)^2$$

- Shi-Fuller (1999)
 - MSW effect due to primordial lepton asymmetry
- ν MSM
 - Asaka, Blanchet, Shaposhnikov (2005)
 - Dark Matter
 - Neutrino mass
 - Leptogenesis
- Other production methods also proposed



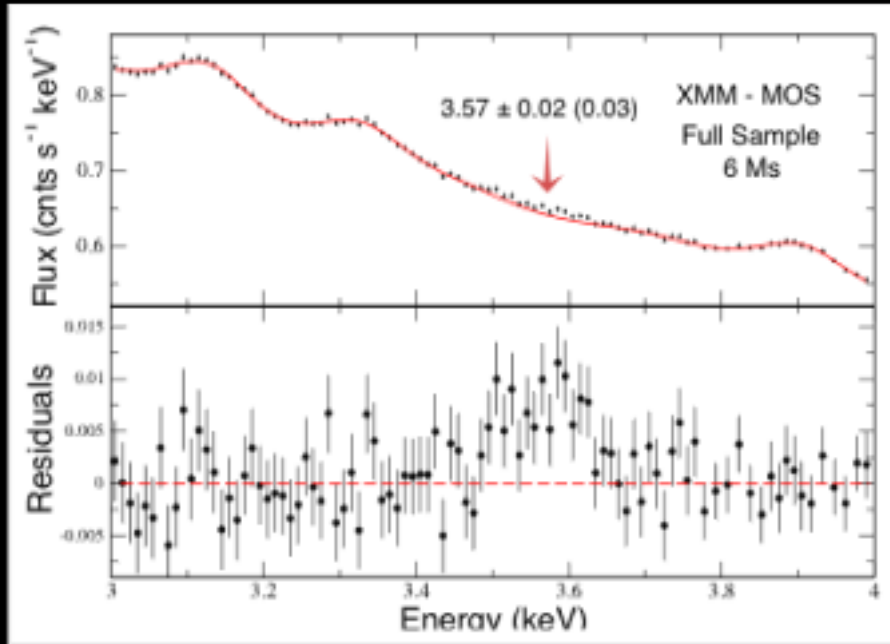
Constrained from all sides

- Warm dark matter candidate
 - Schneider 2016
 - Cherry, Horiuchi 2017
- May solve the Small Scale problem!
- X-rays searches
 - Chandra
 - NuSTAR Bullet Cluster
 - ***Fermi GBM (KCYN 2015)***
 - Integral
 - ..



3.5 keV line excess!

- Bulbul et al (2014)

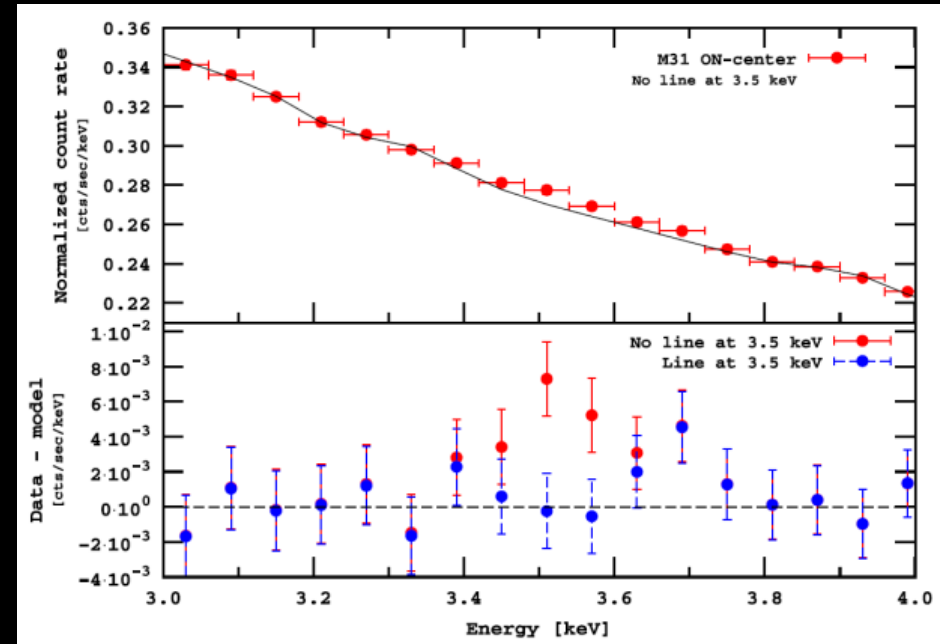


Stacked 73 clusters XMM-MOS ($4-5\sigma$)

Also

Chandra Perseus 2.5σ and 3.4σ

- Boyarsky et al (2014)



XMM-Newton M31

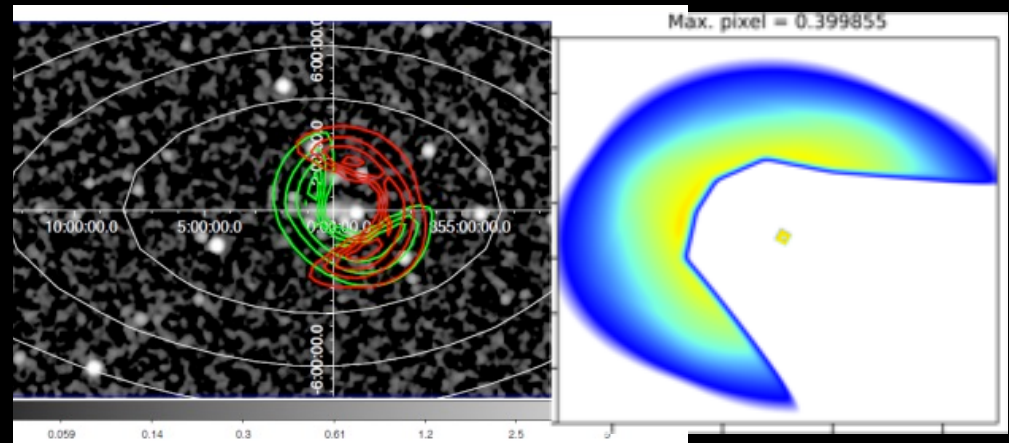
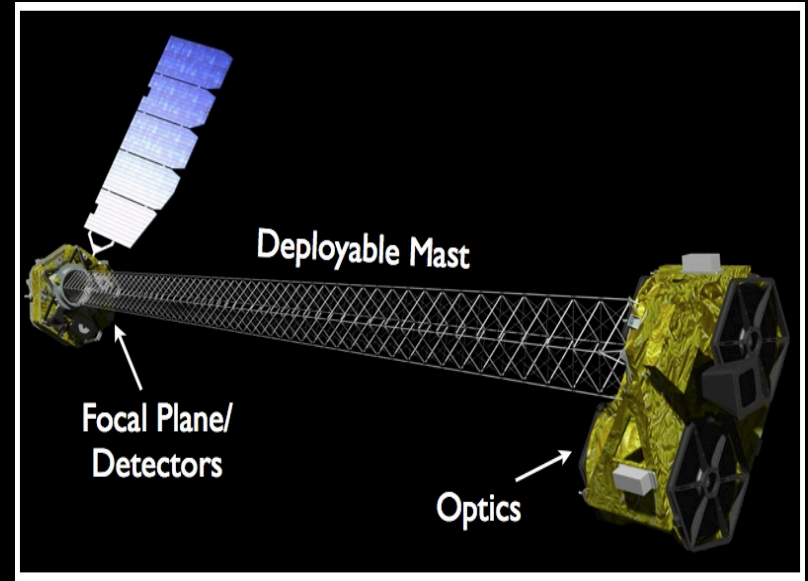
Many Follow-up detections and non-detections! But not ruled out!
Nature not clear!

NuSTAR



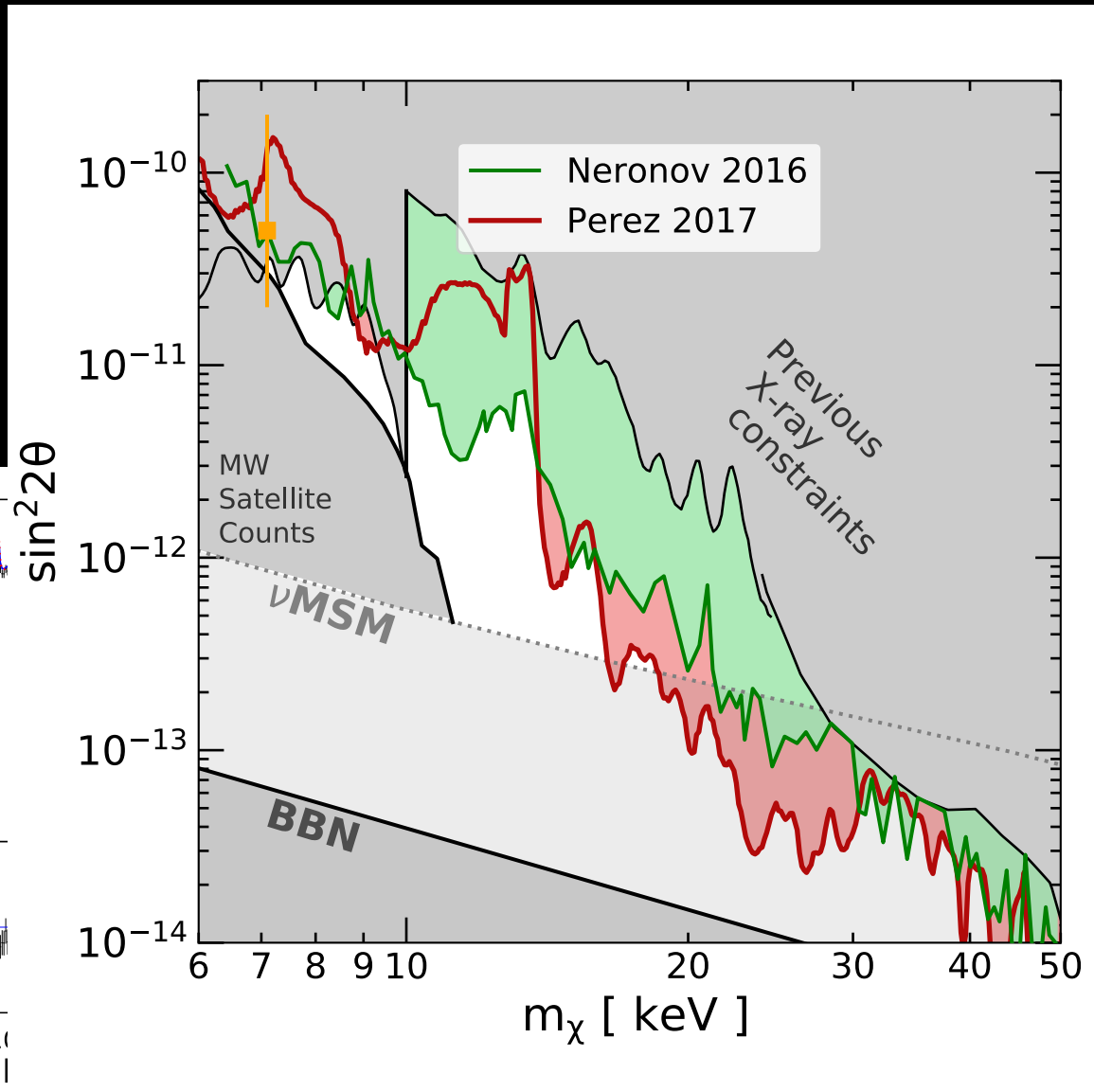
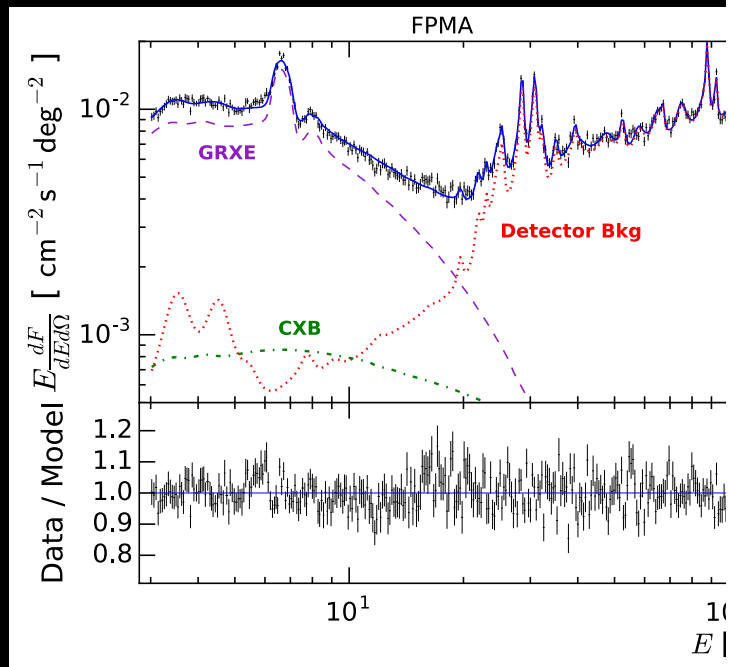
- **Nuclear Spectroscopic Telescope Array**

- Neronov, Malyshev, Eckert [1607.07328]
 - Diffuse sky, MW halo
- Perez, *KCYN*, Beacom, Hersh, Horiuchi, Krivonos [1609.00667]
 - Galactic Center
- **Zero bounce photons**
 - **10X exposure!**



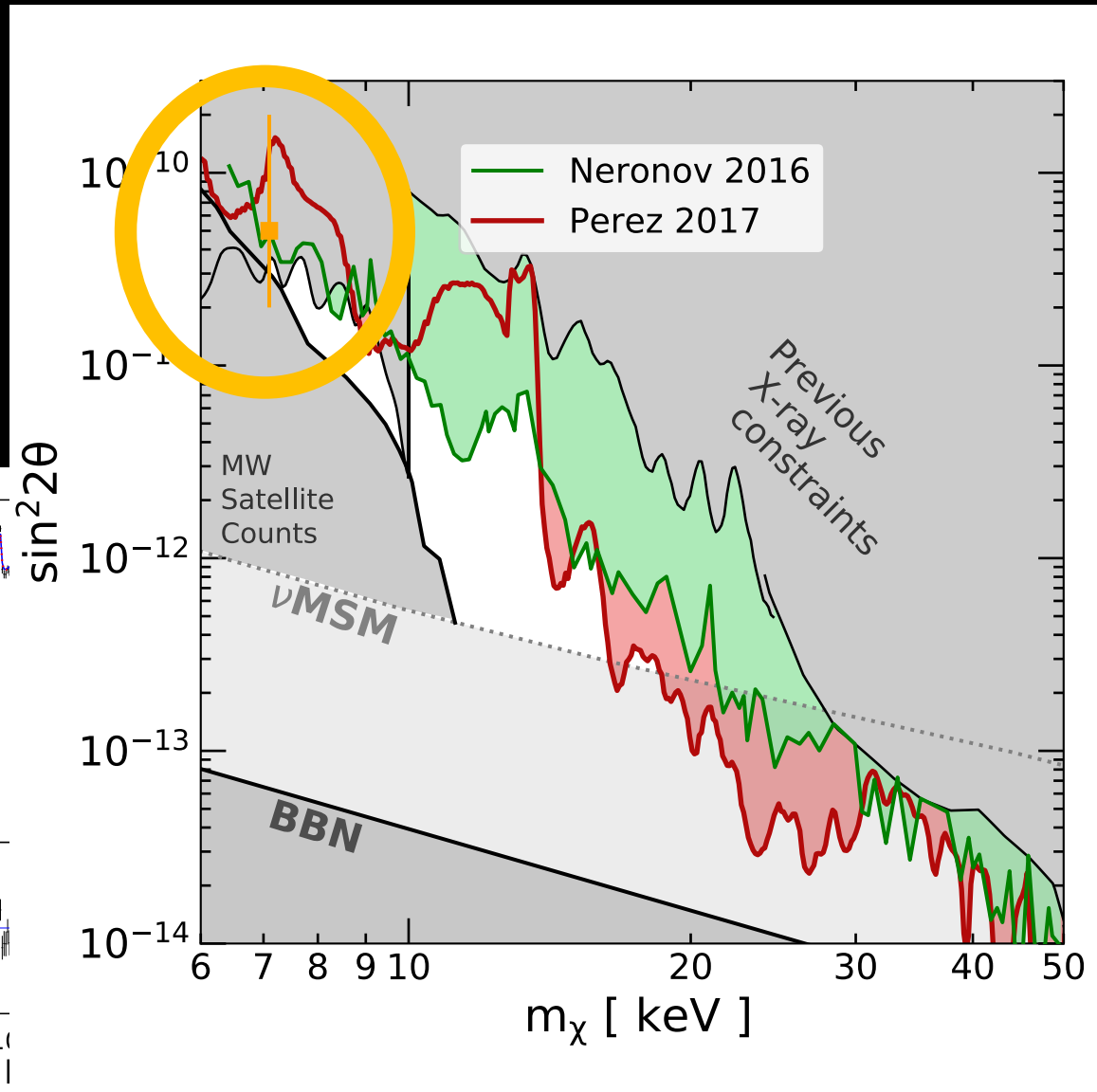
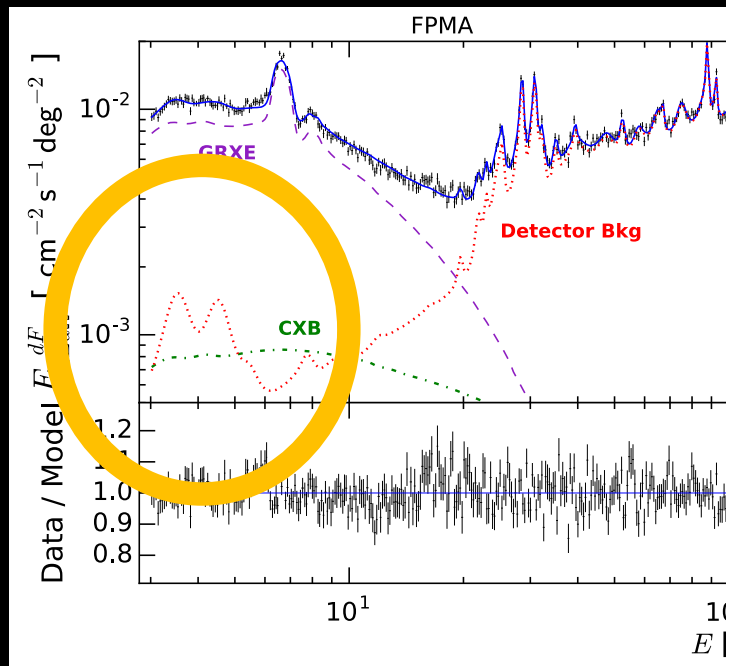
NuSTAR constraints

- 0.5 Ms Milky Way observation
- Neronov et al
– 7.5 Ms blank sky



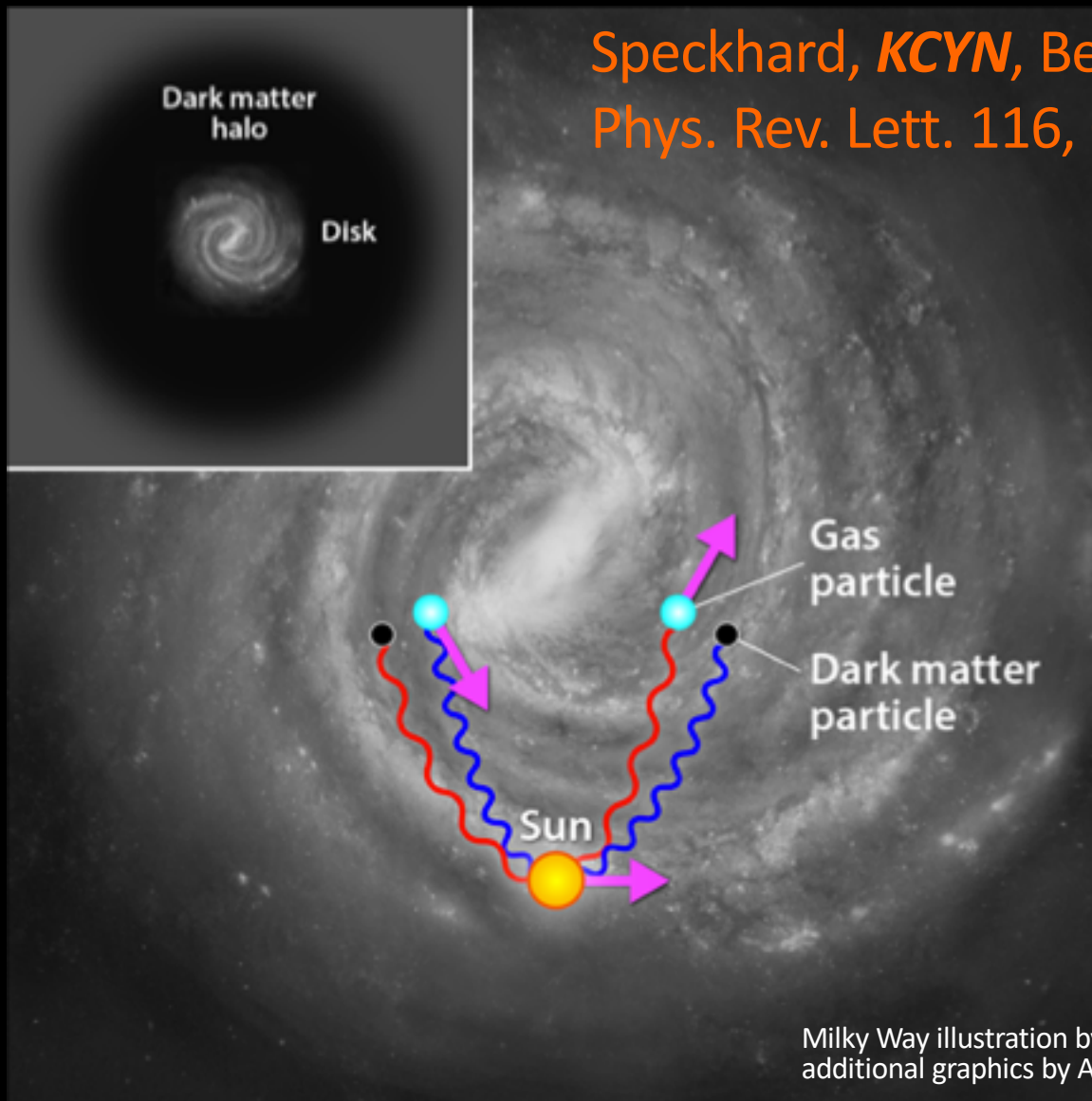
3.5 keV in NuSTAR?

- 3.5 keV line in the NuSTAR background model
- Under Investigations



Dark Matter **Velocity** Spectroscopy

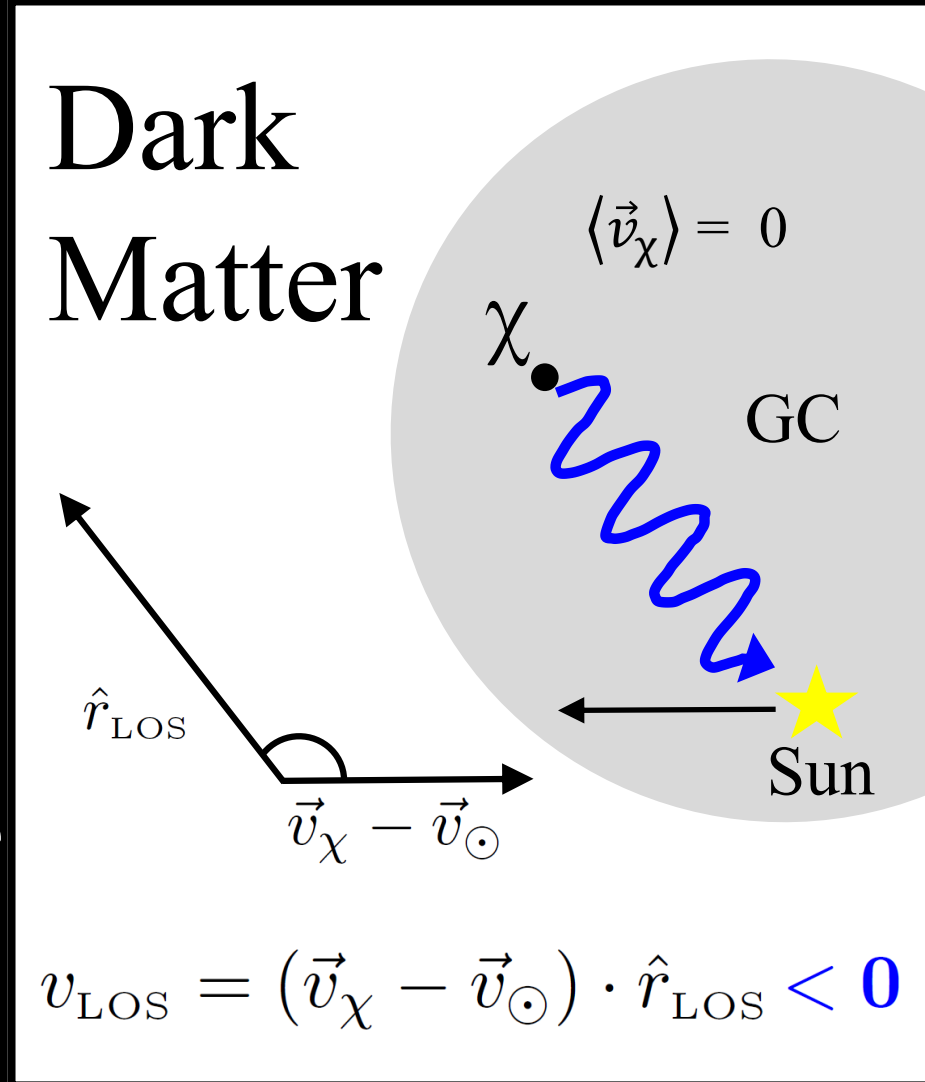
Speckhard, *KCYN*, Beacom, Laha
Phys. Rev. Lett. 116, 031301



Milky Way illustration by Nick Risinger (CC:BY);
additional graphics by APS/Alan Stonebraker

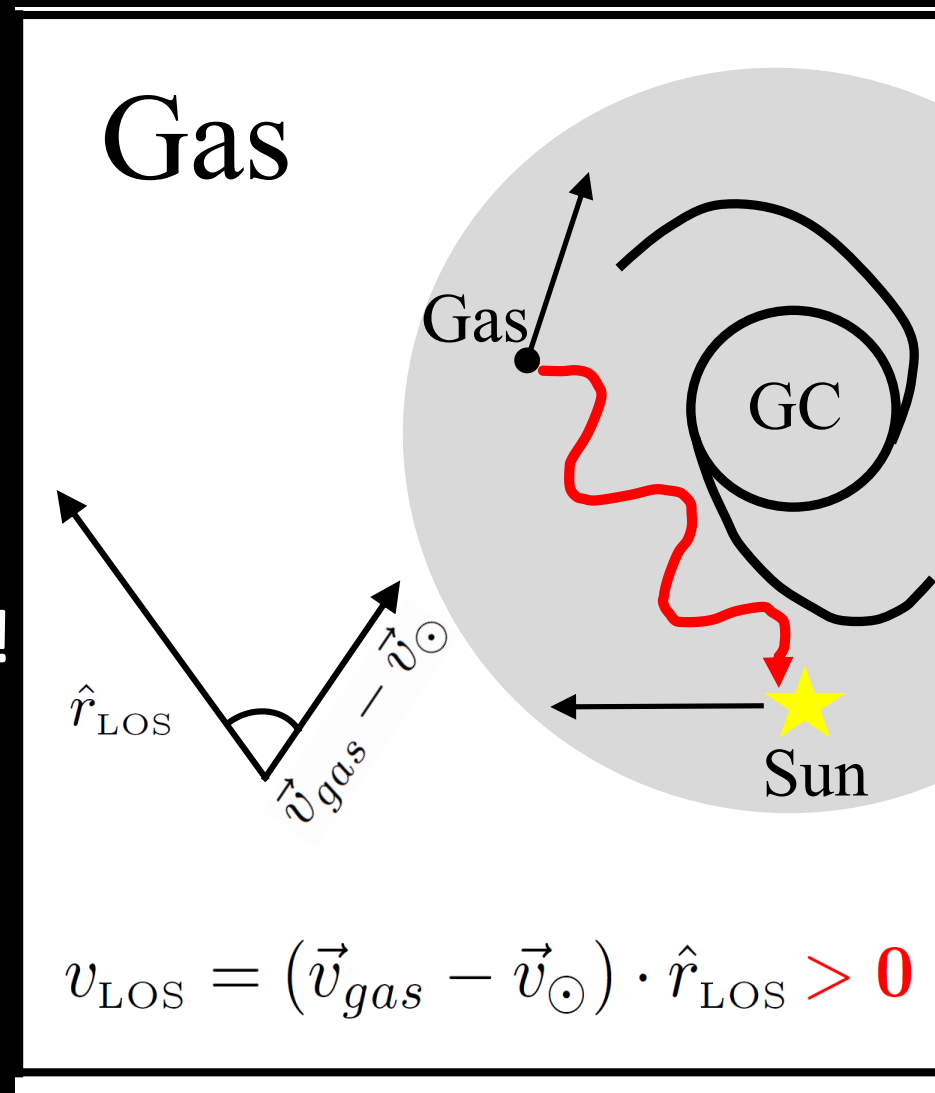
Milky Way dark matter (signal)

- Velocity of the Sun
 - (+)220km/s, +longitude
- Mean dark matter velocity ~ 0
- DM line
 - Blue shifted for +longitude



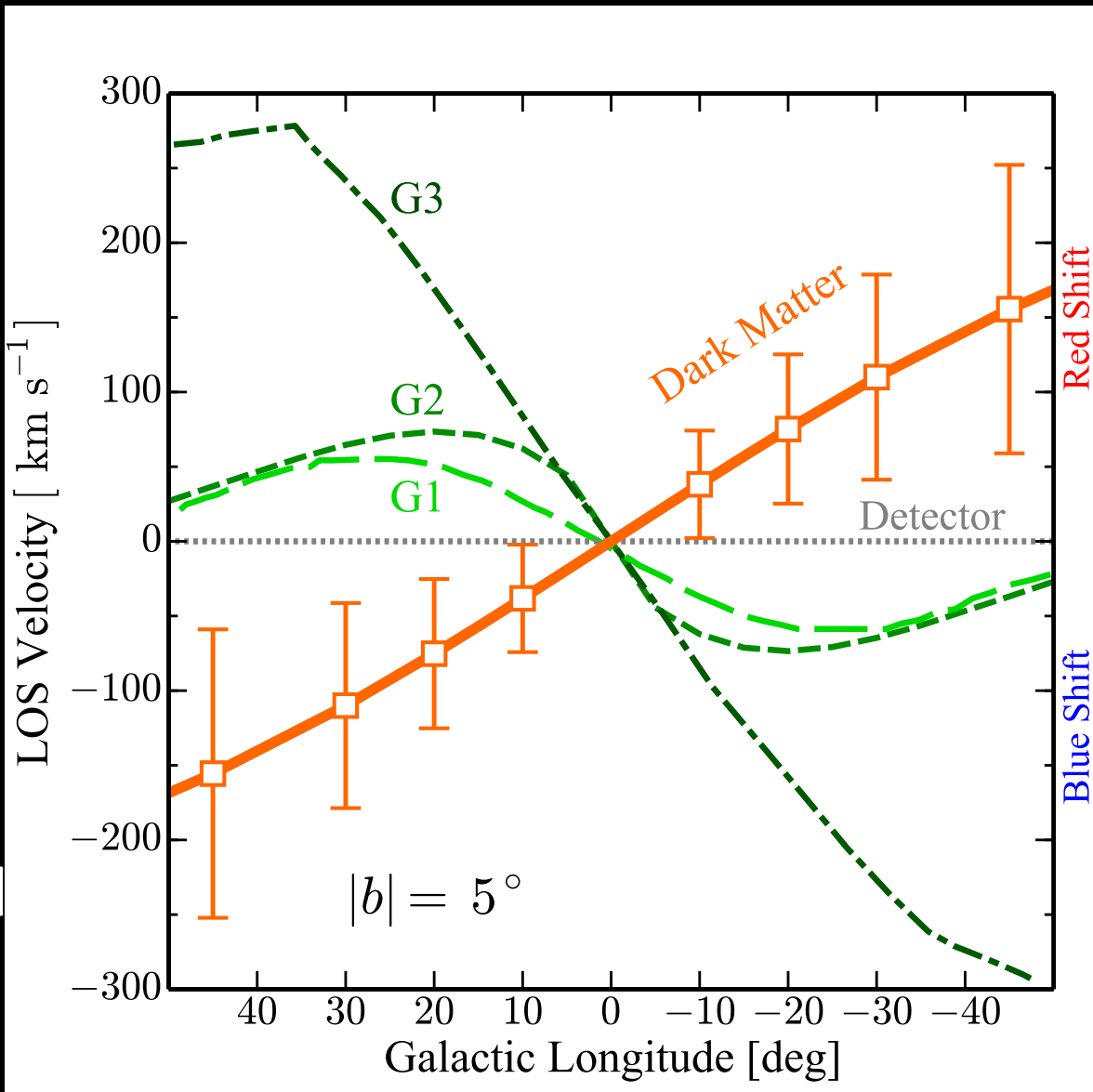
Milky Way Gas (Background)

- Gas and the Sun co-rotate in a disk
 - $V^2 \sim GM/r$
- Astro-physical line
 - **Red shifted** in + longitude!



DM – Astro Separation (MW)

- 0.1% energy resolution
- Clean separation
 - DM
 - Astro
 - Detector effect
- Two obs. $\rightarrow 3.6\sigma$
- Minimal theoretical uncertainty



DM Velocity Spectroscopy

- Extra handle for testing line-like signal
 - The “smoking gun” sometimes is not enough



- X-ray/ 3.5 keV line

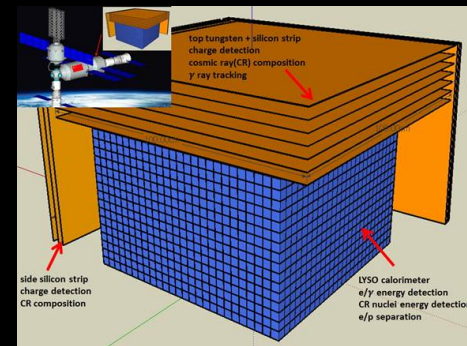
- ~~Astro-H (Hitomi)~~ ----> XRAM (2021)
- Micro-X/ sounding rockets

Figuroa-Feliciano+ [1506.05519]
Powell, Laha, *KCYN*, Abel [1611.02714]

- If DM decay/annihilation produces a line.

- HERD (GeV-TeV)
 - Photons and electrons
 - 2020?

- Dark astronomy/cosmology



Conclusion

- Searching for dark matter with X-rays
 - Production via mixing / ν MSM under test
 - NuSTAR (new results in 2018 and 2019)
 - NuSTAR 3.5 line under investigation
 - *HXMT?*
- Velocity Spectroscopy
 - New Hitomi (maybe 2021)
 - Micro-X

Thanks you!