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Based the following work:

arXiv:1705.09500 (Phys. Rev. D97(2018)035018.)

arXiv:1708.04909 (Phys. Lett. B774 (2017) 217.)

arXiv: 1801.01266 (Phys. Rev. D97(2018)113003.)

Shanghai, China

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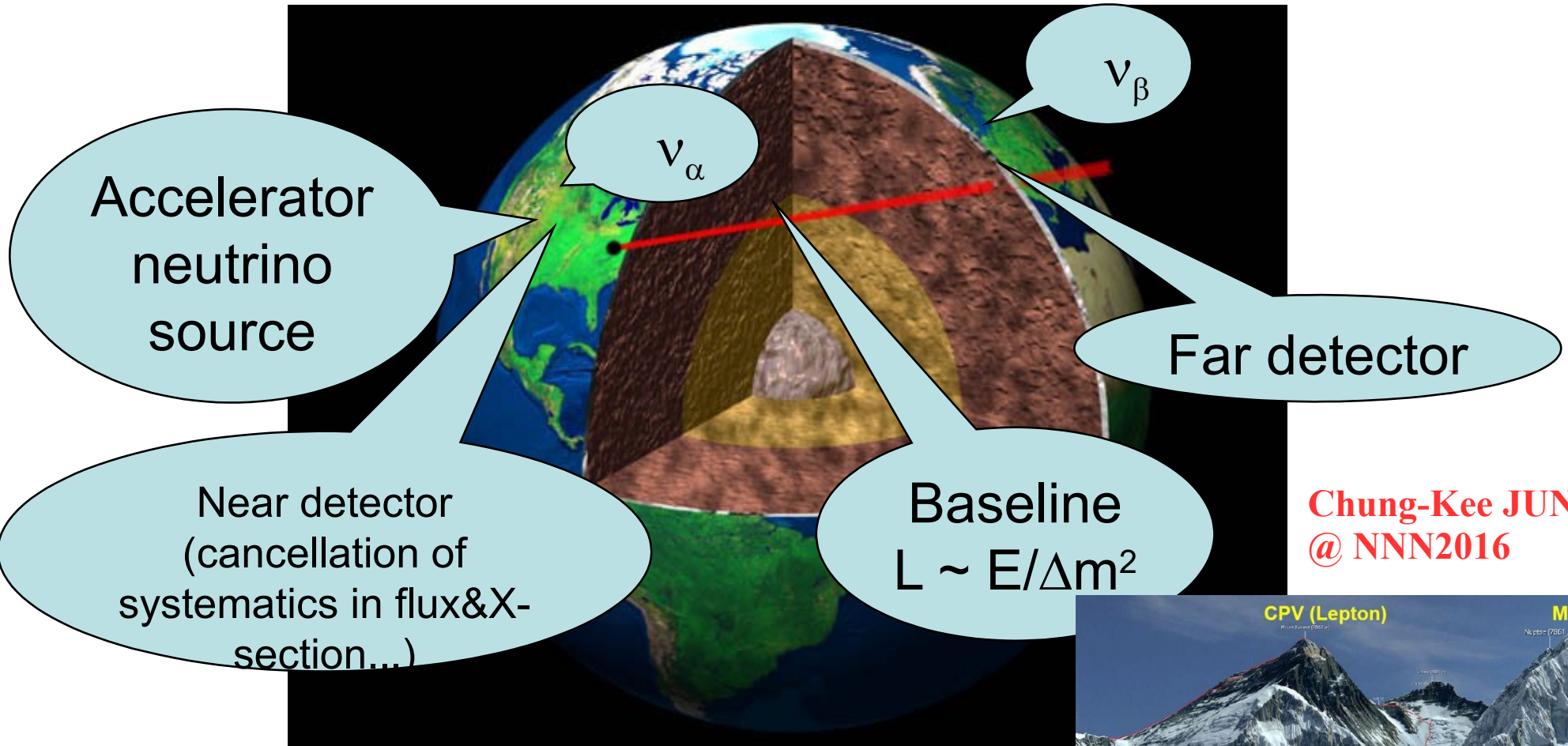
- **Motivations of accelerator neutrino experiments**
- **CC-NSIs at MOMENT (MuOn decay MEdium-baseline NeuTrino beam facility)**
- **Tests of non-unitarity violation with future's accelerator neutrino facilities**
- **Neutrino Activation Analysis with accelerator neutrinos**
- **Summary**



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# Simulation of accelerator neutrino oscillations



**Chung-Kee JUNG**  
@ NNN2016



## Neutrino physics:

- 1 Are there any **sterile** neutrinos in Nature?
- 2 The precise value of angles such as  $\theta_{13}$  and CP phase  $\delta \dots$
- 3 The mass hierarchy:  $\Delta m_{31}^2 > 0$  or  $\Delta m_{31}^2 < 0$ ?
- 4 Can one determine the matter density in a high precision by neutrino oscillation in matter?
- 5 The existence of Non-Standard Interactions?

# Survey of high-power accelerators around the world



- **High-power proton accelerators are scarce resources and very expensive to construct.**
- **Should benefit as more as possible research fields**
- **Hundred-kW beams mostly available now, energy range from 0.5 to 450 GeV**
- **MW beams:**
  - **two in 1-1.5 MW in operation (PSI, SNS)**
  - **one to reach the design goal 1-MW (J-PARC/RCS)**
  - **one 5 MW under construction (ESS)**
  - **one to start construction soon (CiADS, 2.5 MW)**
  - **two to upgrade: 2.4 MW (FNAL/PIP-II), 1.3 MW (J-PARC/MR)**



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# Overview of a Chinese proposed MOMENT

## (Muon-decay MEdium baseline NeuTrino beam facility)

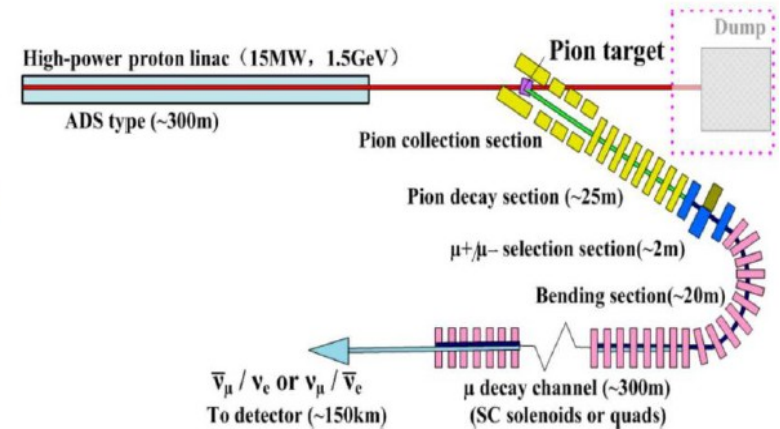
- **MOMENT**: the proposal is still in an early stage ; the details have not been completely fixed.

- **Peak energy**: 200 MeV  
Neutrino energy range: 100MeV—800MeV

- The lower beam energy at  $\sim 300$  MeV:  
free from  $\pi^0$  background

- **Baseline**:  $L=150$  km

In the MOMENT: the neutrino flux peak at low energies  
 require a very massive detector to compensate  
 the low interaction cross section

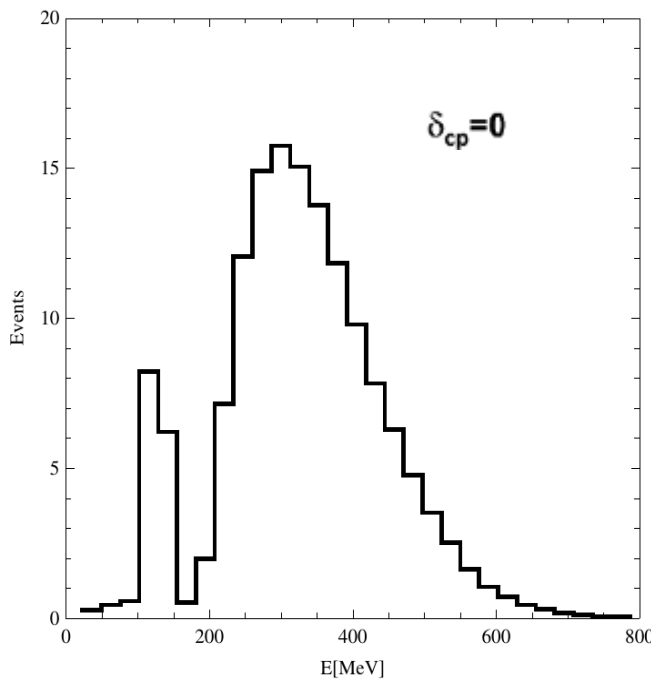


$$N_\nu(E) \sim \Phi_\nu(E) \times \sigma_\nu(E) \times \text{target}$$

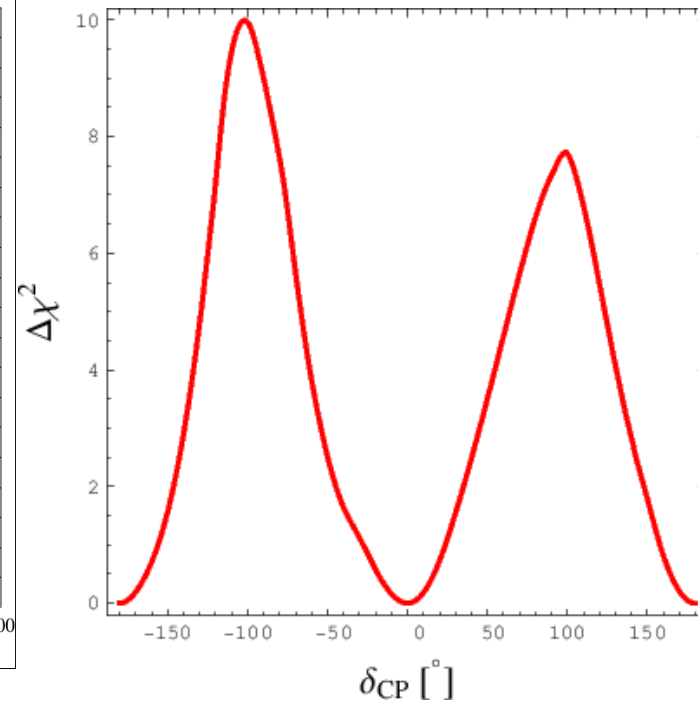
$\nu$  flux (# neutrinos) depends on your  $\nu$  source **make this large!**  
 detector (# targets) **make this large!**  
 $\nu$  cross section tiny ( $\sim 10^{-38}$  cm<sup>2</sup>)  
 $\sigma_{\nu, \text{tot}} \sim E_\nu$  **go to higher energies**

# Updates of CPV sensitivity

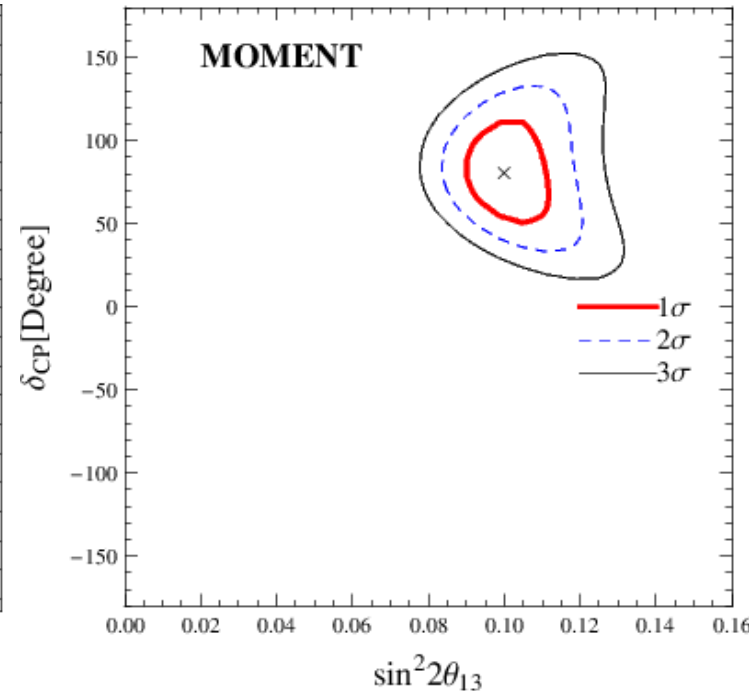
- Neutrino fluxes and detector info inherited from Miao He& Jiashu Lu
- Loss of CPV by a factor of 2 after including both systematic and statistic uncertainties
- All backgrounds highly suppressed, especially atmospheric bckgs!



Detected neutrino spectra



Discovery of CPV



Precision measurements

- **First physics study performed by Pilar, Matthias and Eriquer in arXiv:1511.02859**
- **NC-NSIs in matter considered by Pouya and Yasaman in arXiv: 1602.07099**



# Links between NSIs and neutrino oscillations

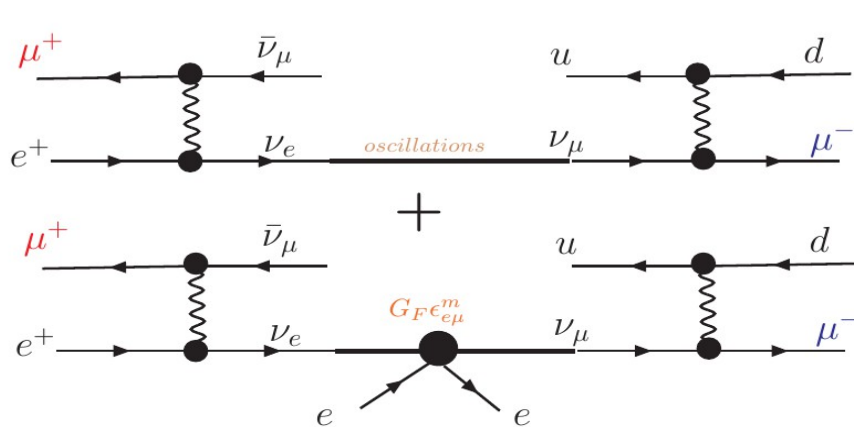
**New physics beyond SM: new particles, new couplings, new phenomenon...**

- Flavor violating interactions with neutrinos:

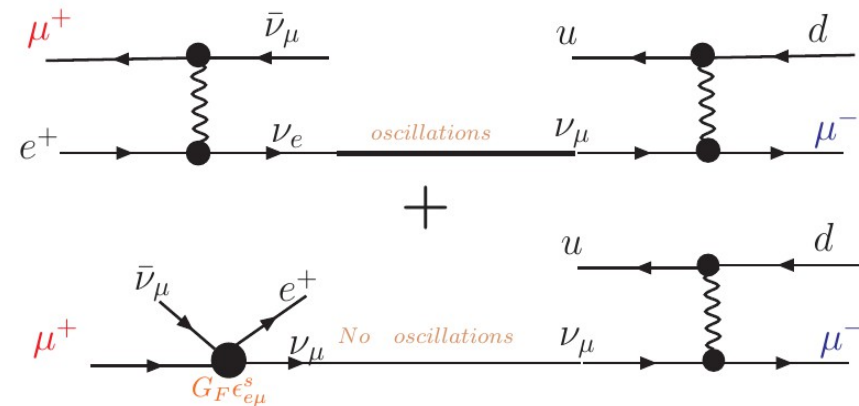
$$\nu_\alpha f \rightarrow \nu_\beta f, l_\alpha^- \rightarrow \nu_\beta e^- \bar{\nu}_e \dots$$

- 4-fermion vertices:

$$L_{\text{eff}} = 2\sqrt{2}G_F(\epsilon^{L/R})_{\beta\delta}^{\alpha\gamma}(\bar{\nu}^\beta\gamma^\rho P_L\nu_\alpha)(\bar{\ell}^\delta\gamma^\rho P_{L/R}\ell_\gamma)$$

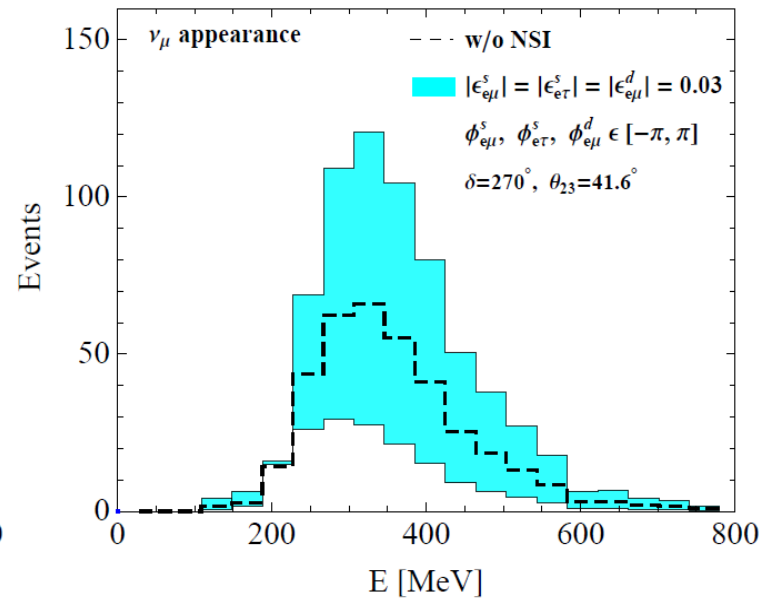
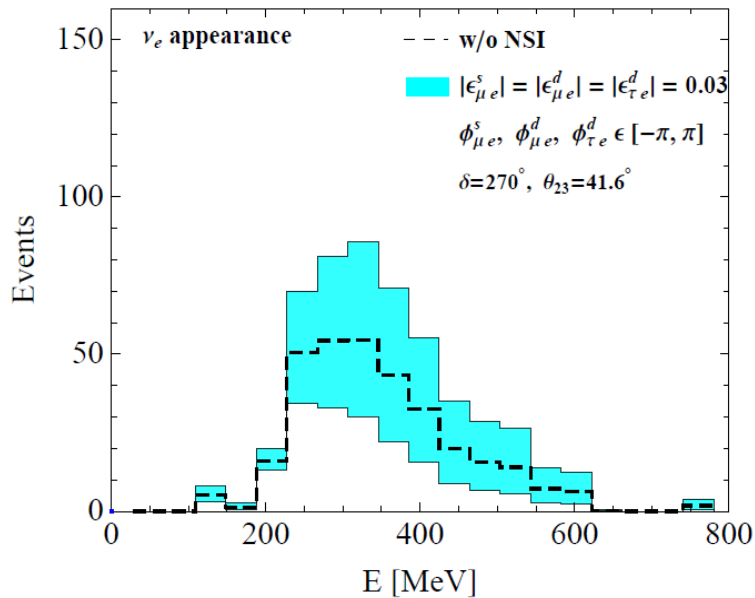
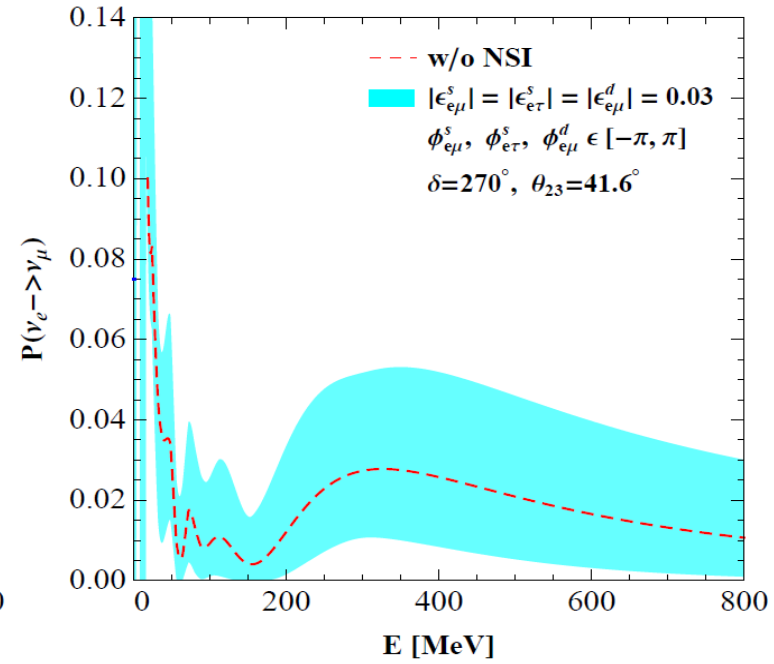
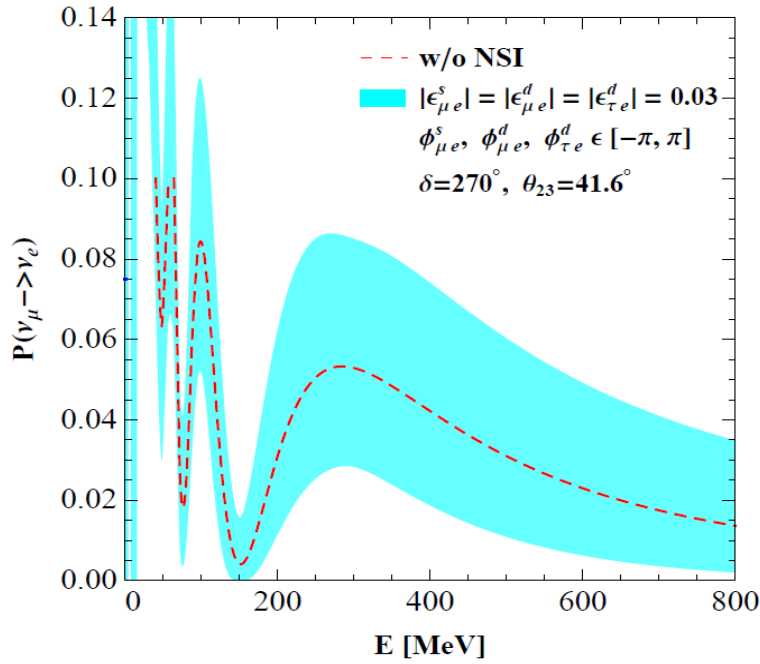


NSI happens to neutrino propagation in matter

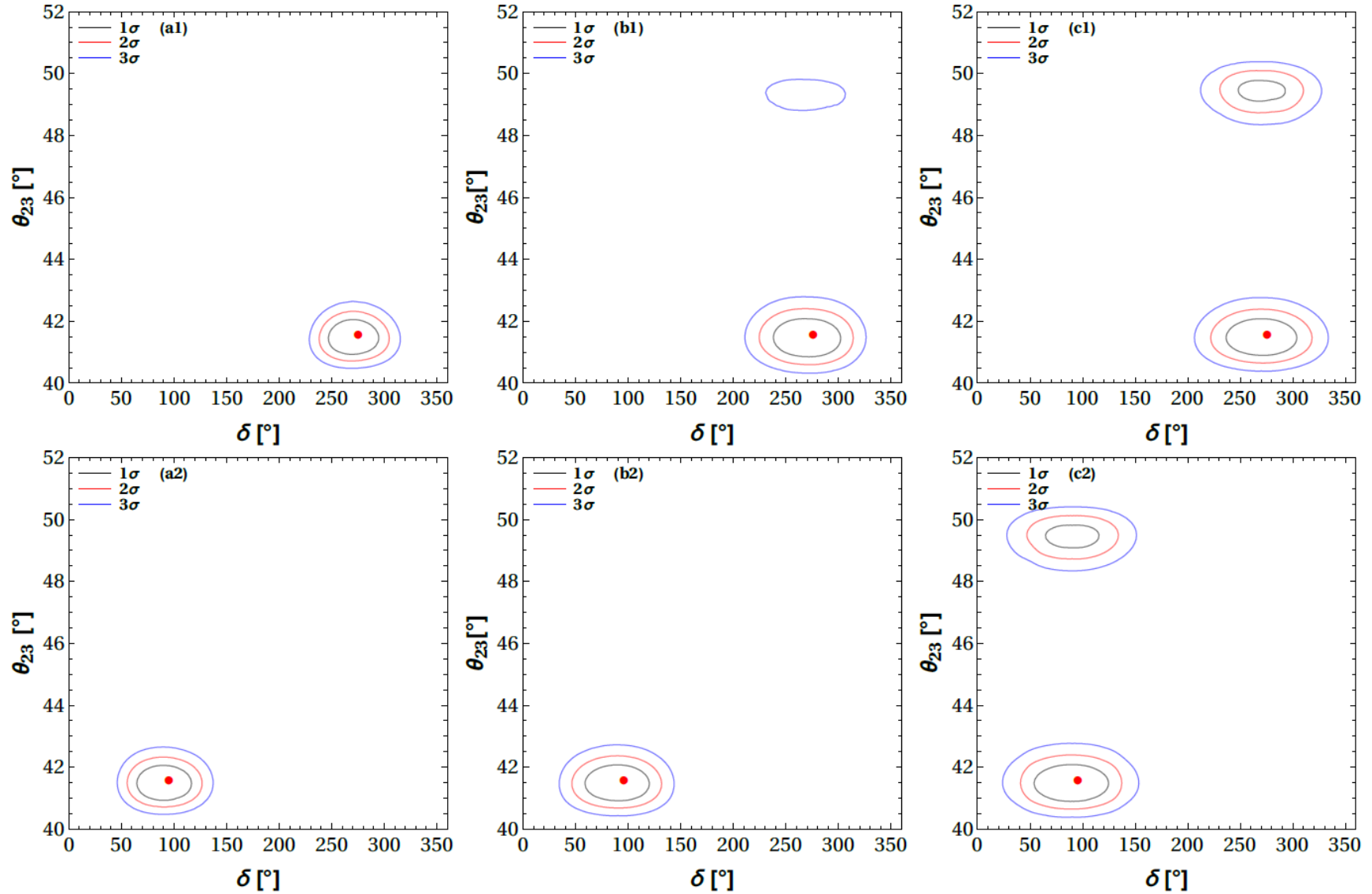


NSI at neutrino productions

# Numerical tests of oscillation probabilities and events at MOMENT



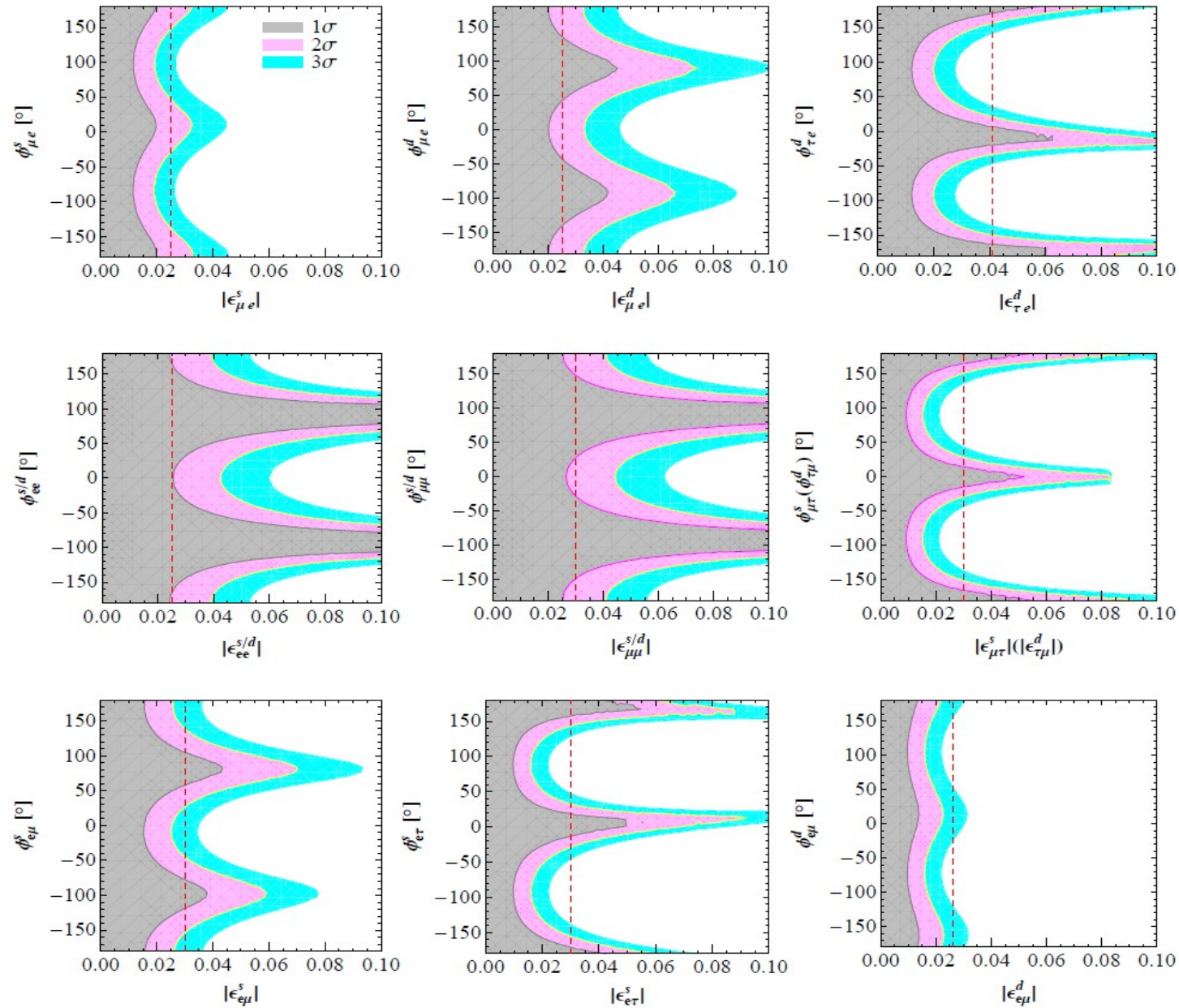
# Impacts on precision measurements by CC-NSIs



**Degeneracy shows up after an introduction of CC-NSIs at some parameter space.**

# Constraints of CC-NSIs with a far detector at MOMENT

- Colorful regions are allowed after running a far detector at MOMENT.
- The e-mu sector of NSI are the best constrained.
- Almost all NSI-induced CP phases change the exclusion limits severely except the e-mu sector.
- Limits from other sectors are not as good as those from the e-mu sector of NSI.



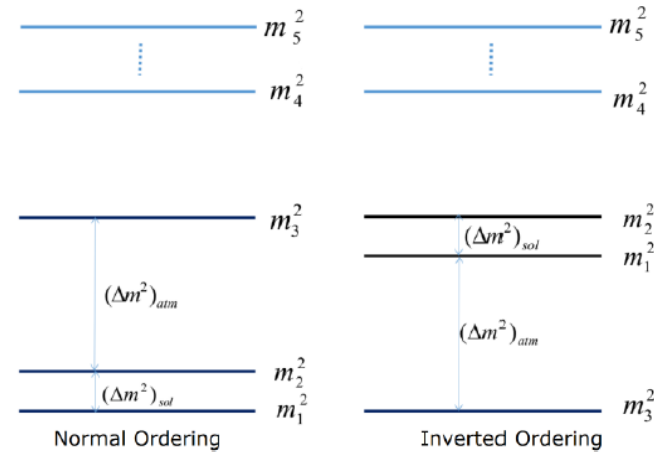


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# Tests of unitarity violation

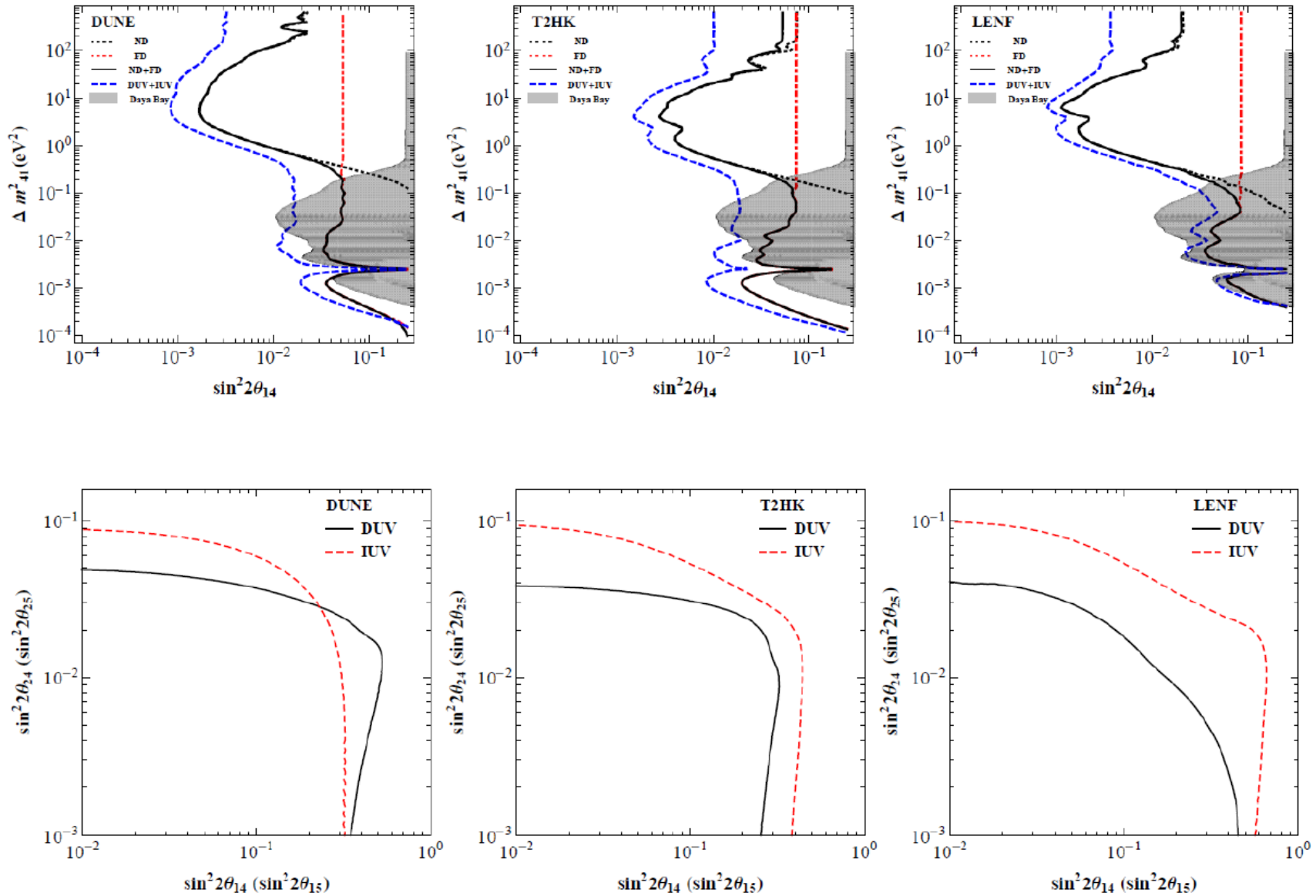
- Light sterile neutrino anomaly (eV scale)
- Heavy sterile neutrinos from see-saw model (GeV scale)
- Dark matter candidate (keV scale)
- IUUV (indirect unitary violation) by heavy sterile neutrinos
- DUV (direct unitary violation) by light sterile neutrinos: oscillation with active ones



$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$

- Simplifying the mixing matrix to deal with DUV and IUUV, Phys. Lett., B718:1447-1453, 2013
- Perturbation study of oscillation probabilities for DUV and IUUV, Phys. Rev., D93(3):033008

# Exclusion limits on mixing parameters with non-unitarity



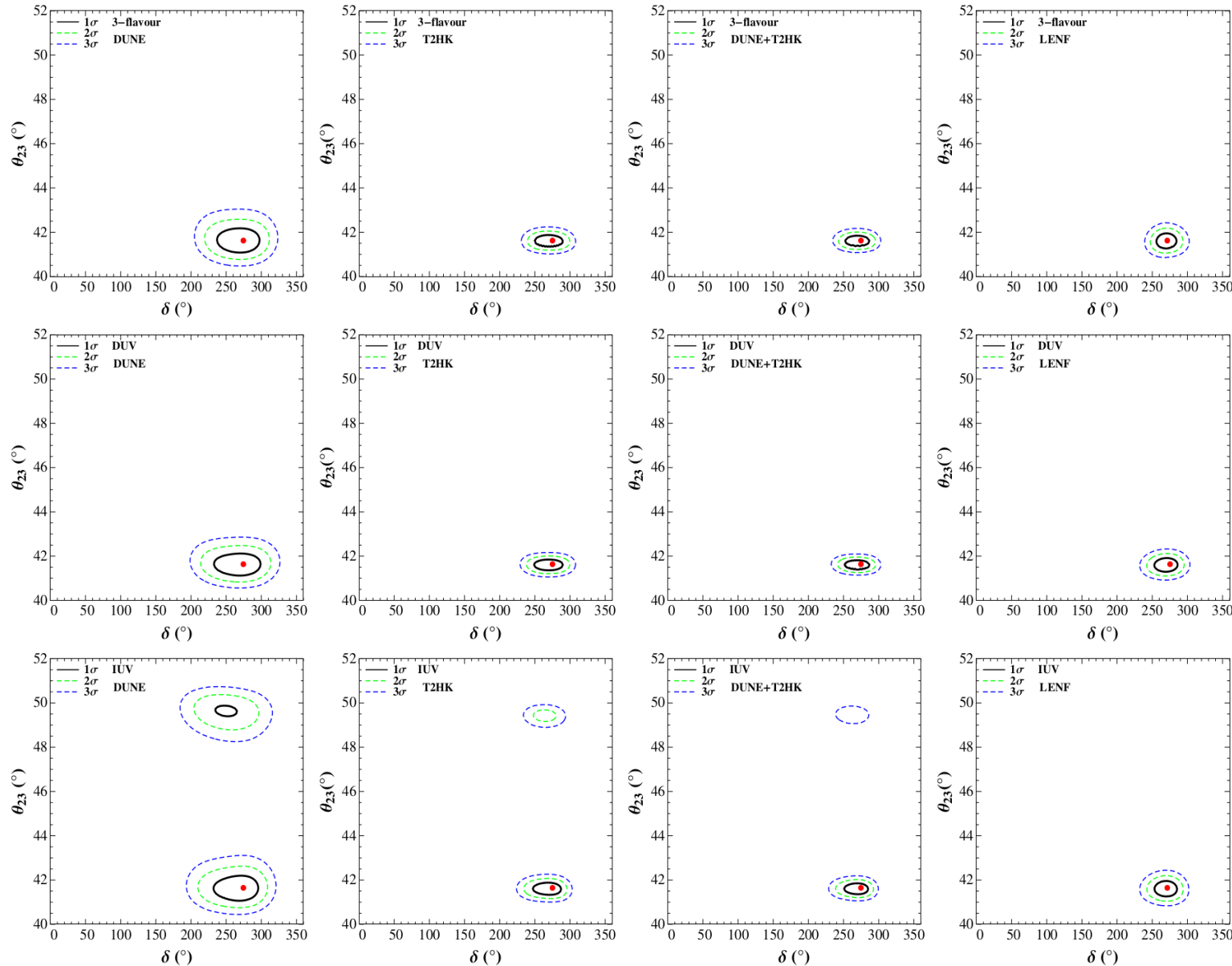
The limits to new parameters induced by the DUV and IUV effects

# Impacts on precision measurements

- IUV can only induce rate correlations to the three neutrino oscillation, but DUV contributes both rate and spectrum signatures to the experimental measurements.

- The DUV generally does not cause degeneracies for  $\theta_{23}$ .

- The IUV effects would cause degeneracies for  $\theta_{23}$  in DUNE and T2HK. Thus we can turn to the most powerful experiment LBNF to solve this problem;







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# Neutrino-nucleus coherent scatterings

Science

REPORTS

Cite as: D. Akimov *et al.*, *Science*  
10.1126/science.aao0990 (2017).

## Observation of coherent elastic neutrino-nucleus scattering

D. Akimov,<sup>1,2</sup> J. B. Albert,<sup>3</sup> P. An,<sup>4</sup> C. Awe,<sup>4,5</sup> P. S. Barbeau,<sup>4,5</sup> B. Becker,<sup>6</sup> V. Belov,<sup>1,2</sup> A. Brown,<sup>4,7</sup> A. Bolozdynya,<sup>2</sup> B. Cabrera-Palmer,<sup>8</sup> M. Cervantes,<sup>5</sup> J. I. Collar,<sup>9\*</sup> R. J. Cooper,<sup>10</sup> R. L. Cooper,<sup>11,12</sup> C. Cuesta,<sup>13†</sup> D. J. Dean,<sup>14</sup> J. A. Detwiler,<sup>13</sup> A. Eberhardt,<sup>13</sup> Y. Efremenko,<sup>6,14</sup> S. R. Elliott,<sup>12</sup> E. M. Erkela,<sup>13</sup> L. Fabris,<sup>14</sup> M. Febraro,<sup>14</sup> N. E. Fields,<sup>9‡</sup> W. Fox,<sup>3</sup> Z. Fu,<sup>13</sup> A. Galindo-Uribarri,<sup>14</sup> M. P. Green,<sup>4,14,15</sup> M. Hai,<sup>9§</sup> M. R. Heath,<sup>3</sup> S. Hedges,<sup>4,5</sup> D. Hornback,<sup>14</sup> T. W. Hossbach,<sup>16</sup> E. B. Iverson,<sup>14</sup> L. J. Kaufman,<sup>3||</sup> S. Ki,<sup>4,5</sup> S. R. Klein,<sup>10</sup> A. Khromov,<sup>2</sup> A. Konovalov,<sup>1,2,17</sup> M. Kremer,<sup>4</sup> A. Kumpan,<sup>2</sup> C. Leadbetter,<sup>4</sup> L. Li,<sup>4,5</sup> W. Lu,<sup>14</sup> K. Mann,<sup>4,15</sup> D. M. Markoff,<sup>4,7</sup> K. Miller,<sup>4,5</sup> H. Moreno,<sup>11</sup> P. E. Mueller,<sup>14</sup> J. Newby,<sup>14</sup> J. L. Orrell,<sup>16</sup> C. T. Overman,<sup>16</sup> D. S. Parno,<sup>13¶</sup> S. Penttila,<sup>14</sup> G. Perumpilly,<sup>9</sup> H. Ray,<sup>18</sup> J. Raybern,<sup>5</sup> D. Reyna,<sup>8</sup> G. C. Rich,<sup>4,14,19</sup> D. Rimal,<sup>18</sup> D. Rudik,<sup>1,2</sup> K. Scholberg,<sup>5</sup> B. J. Scholz,<sup>9</sup> G. Sinev,<sup>5</sup> W. M. Snow,<sup>3</sup> V. Sosnovtsev,<sup>2</sup> A. Shakirov,<sup>2</sup> S. Suchyta,<sup>10</sup> B. Suh,<sup>4,5,14</sup> R. Tayloe,<sup>3</sup> R. T. Thornton,<sup>3</sup> I. Tolstukhin,<sup>3</sup> J. Vanderwerp,<sup>3</sup> R. L. Varner,<sup>14</sup> C. J. Virtue,<sup>20</sup> Z. Wan,<sup>4</sup> J. Yoo,<sup>21</sup> C.-H. Yu,<sup>14</sup> A. Zawada,<sup>4</sup> J. Zettlemoyer,<sup>3</sup> A. M. Zderic,<sup>13</sup> COHERENT Collaboration#

- Progress of low-threshold DM detectors made it come true.
- What else can we do with CEvNS?

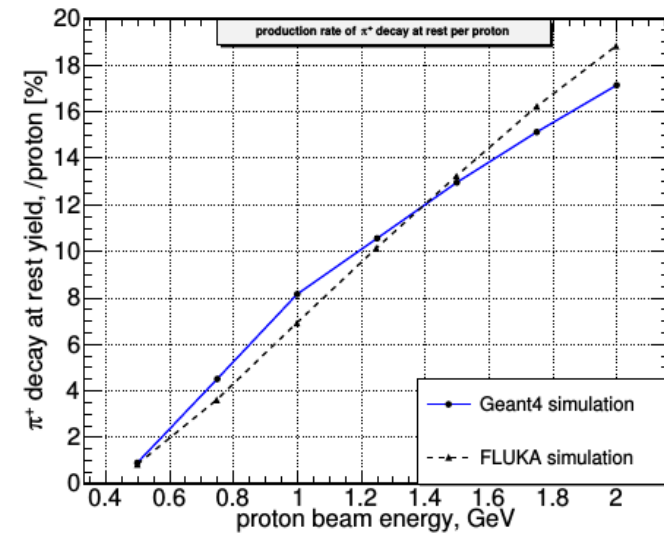
# Neutrino Activation Analysis

$$\frac{d\sigma(E_\nu, E_r)}{dE_r} = \frac{G_F^2 [N - (1 - 4 \sin^2 \theta_w) Z]^2 F^2(Q^2) M^2}{4\pi} \times \frac{1}{M} \left(1 - \frac{E_r}{E_{max}}\right)$$

- **CEvNS is proportional to the number of neutrons in the nucleus.**
- **Nuclear effects are factorized in the form factor:  
a transformation of the density distribution**

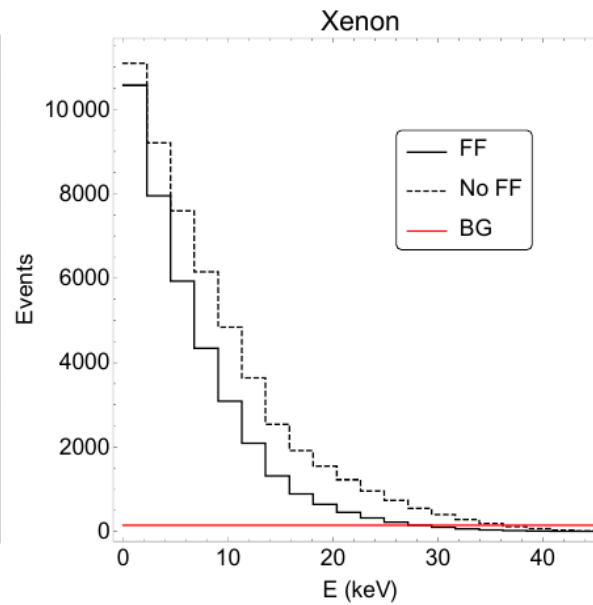
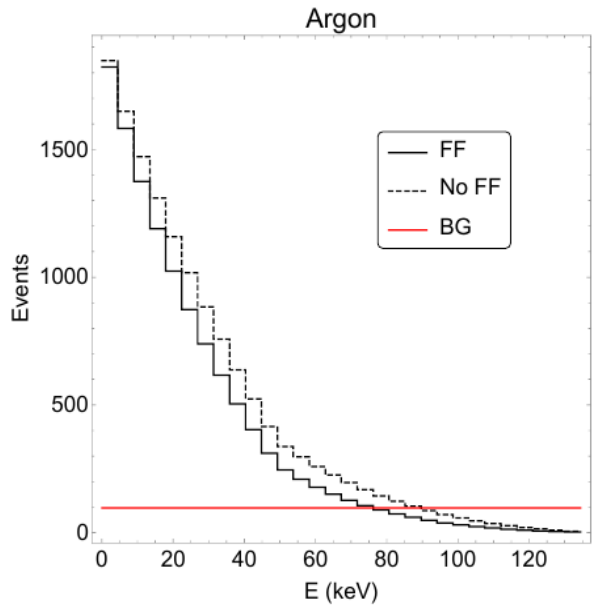
$$F(Q^2) = \frac{1}{Q_w} \int [\rho_n(r) - (1 - 4 \sin^2 \theta_w) \rho_p(r)] \frac{\sin(Qr)}{Qr} r^2 dr$$

- **Lots of proton accelerators around the world.**
- **Use CEvNS to measure the nuclear structure while it is complementary to CC-scatterings?**
- **Which kind of detector can do the job?**
- **What are requirements to measure the nuclear structure?**

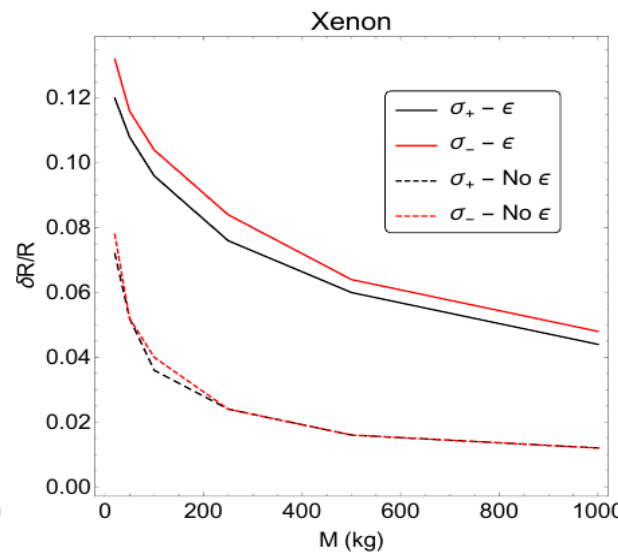
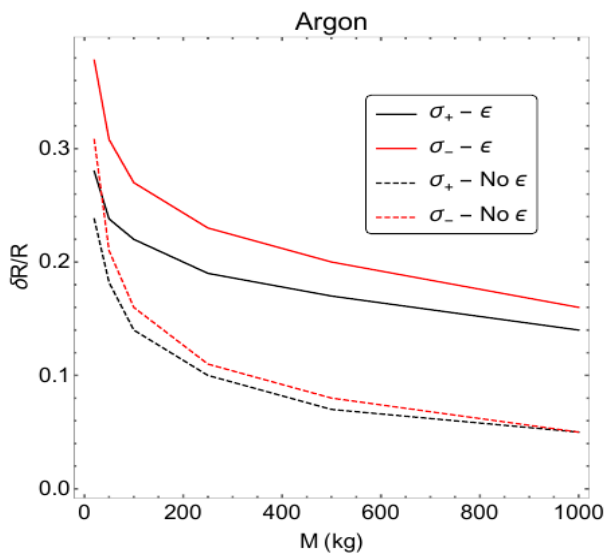


# LAr and LXe TPC

- Learn from DM detection experiments: LAr and LXe TPC.



- Threshold is the key
- Beam-related backgrounds: timing structures
- Cosmic-induced backgrounds: passive and active vetos



- A ton-scale detector reaches the sub-percent precision of the neutron radius in the nucleus.
- LXe TPC is doing better given the same fiducial mass.
- Good to distinguish nuclear physics models.



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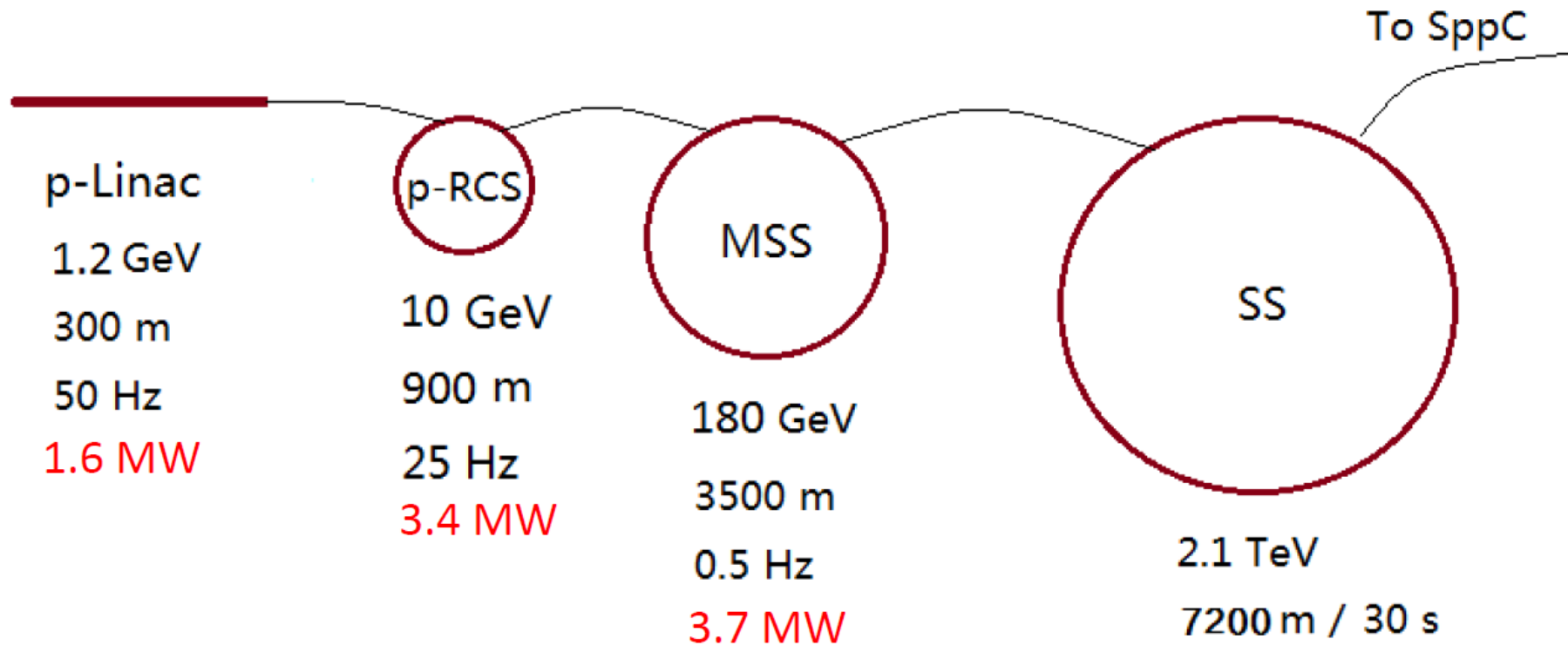


# Summary

- **Lots of physics to be done with accelerator neutrinos.**
- **CC-NSI study at MOMENT.**
- **Probe of the direct and indirect unitarity violations at future accelerator neutrino facilities.**
- **Neutrino activation analysis to probe the nuclear structure.**
- **Welcome to work together on new physics searches.**

**Thanks for your attention!**

# SPPC proton driver for neutrino physics



Very powerful injector beams to support rich physics programs including neutrino physics

- Three proton beams in MW level: 1.2 GeV, 10 GeV, 180 GeV