

Introductory remarks

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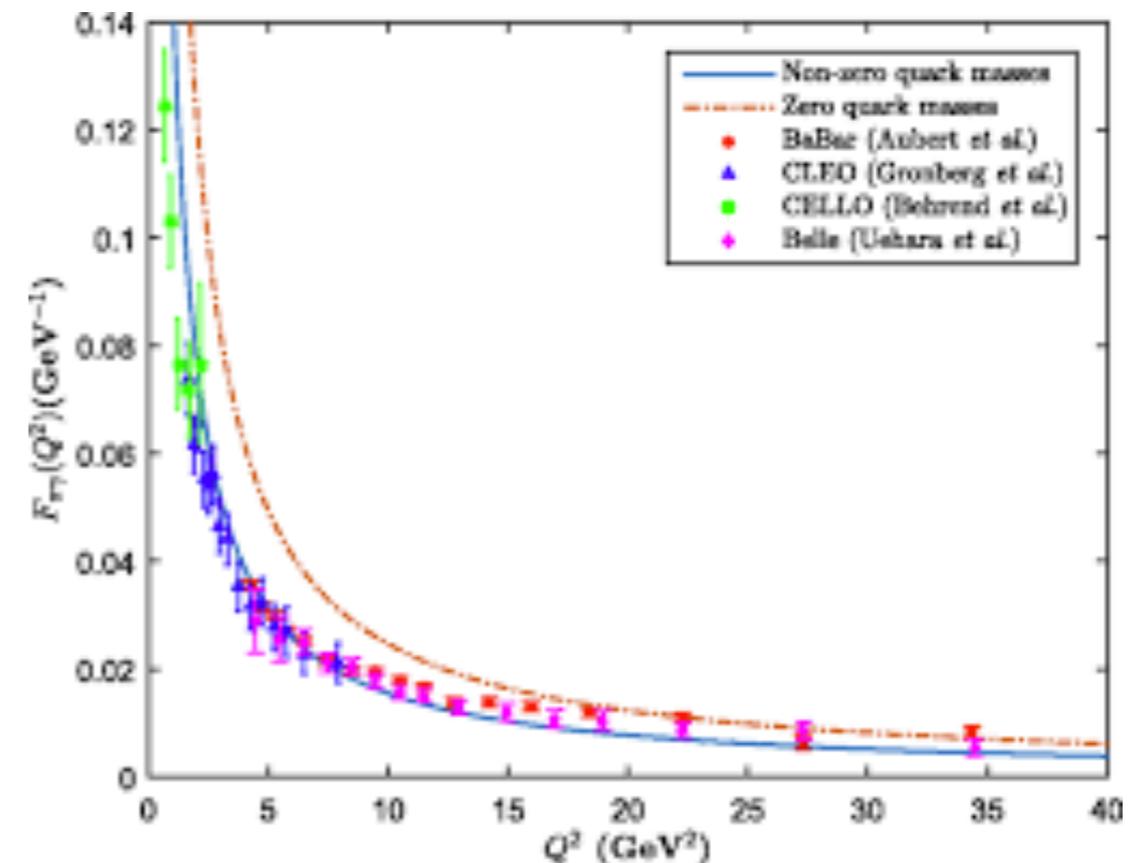
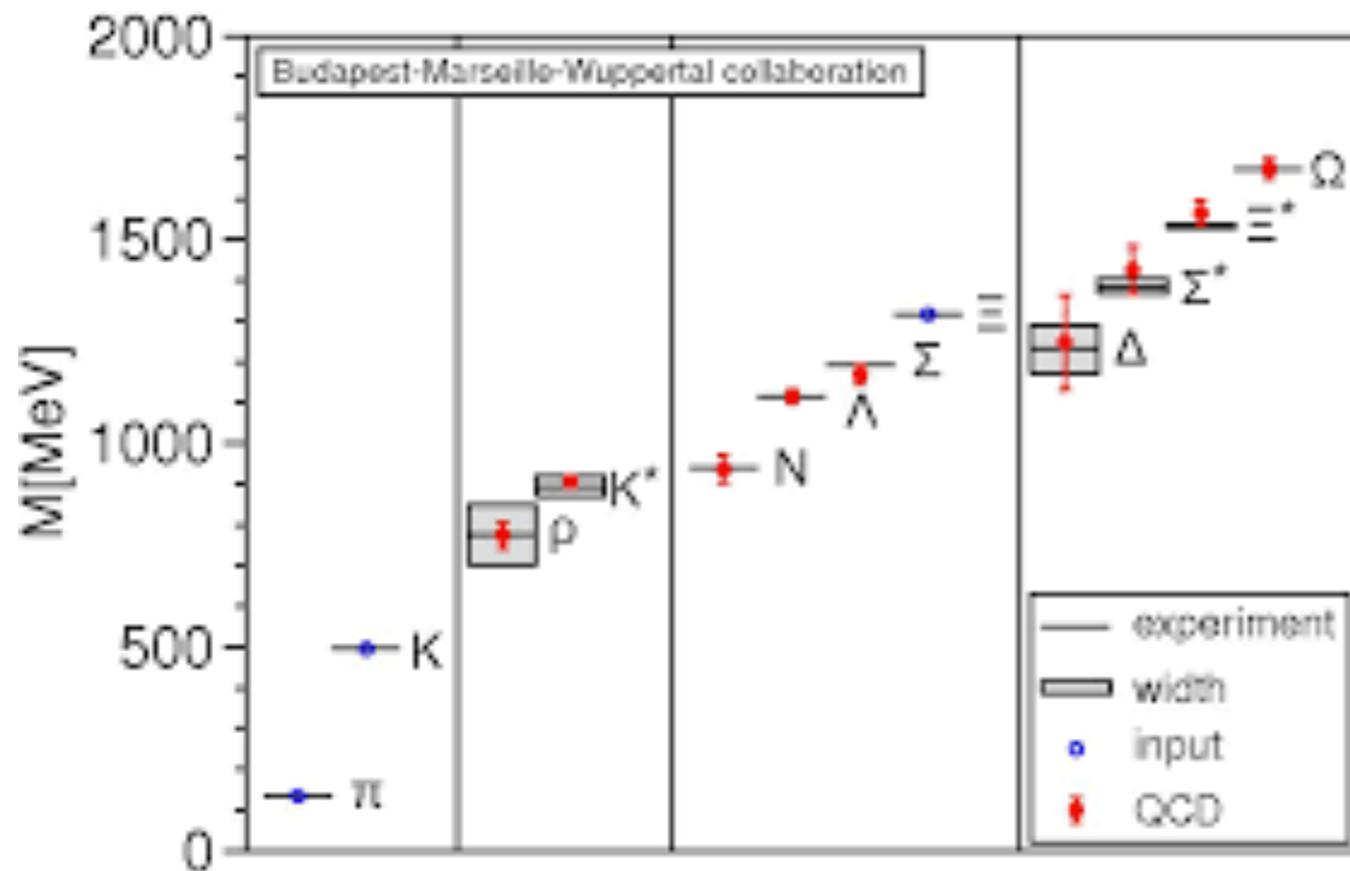
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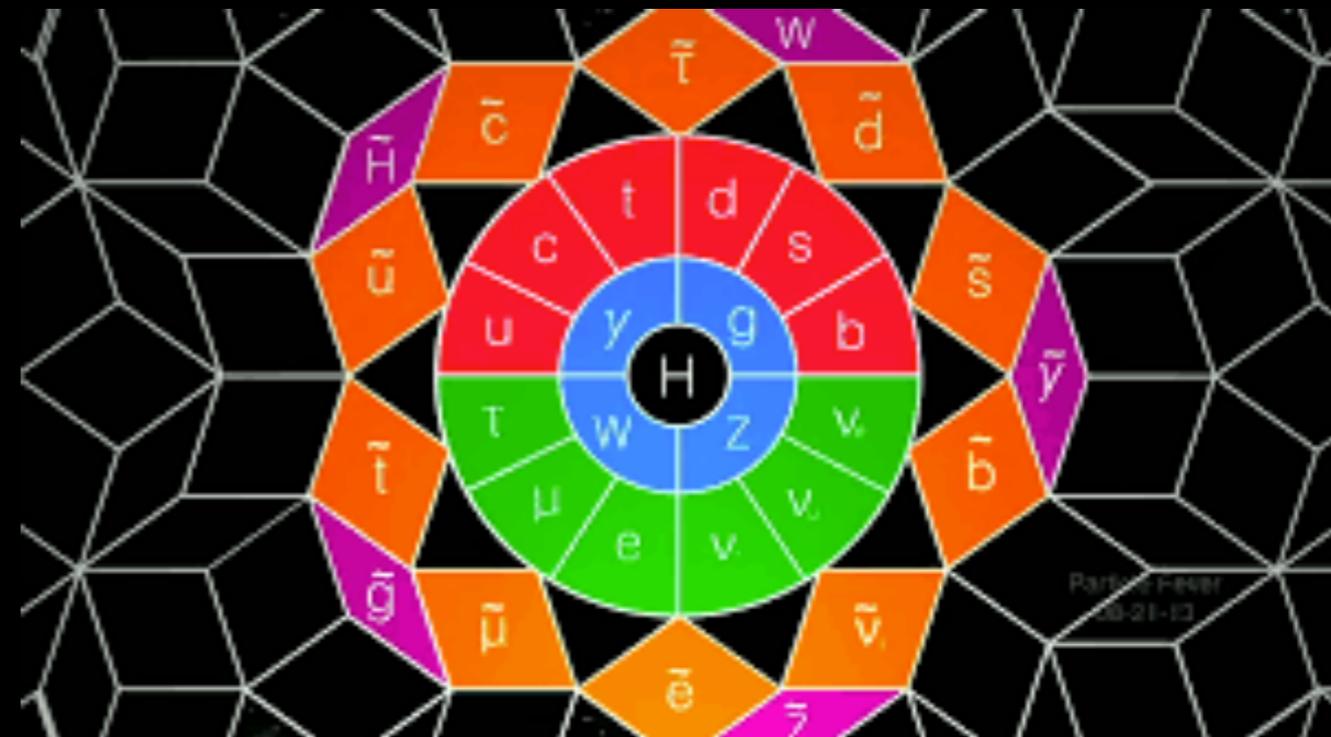
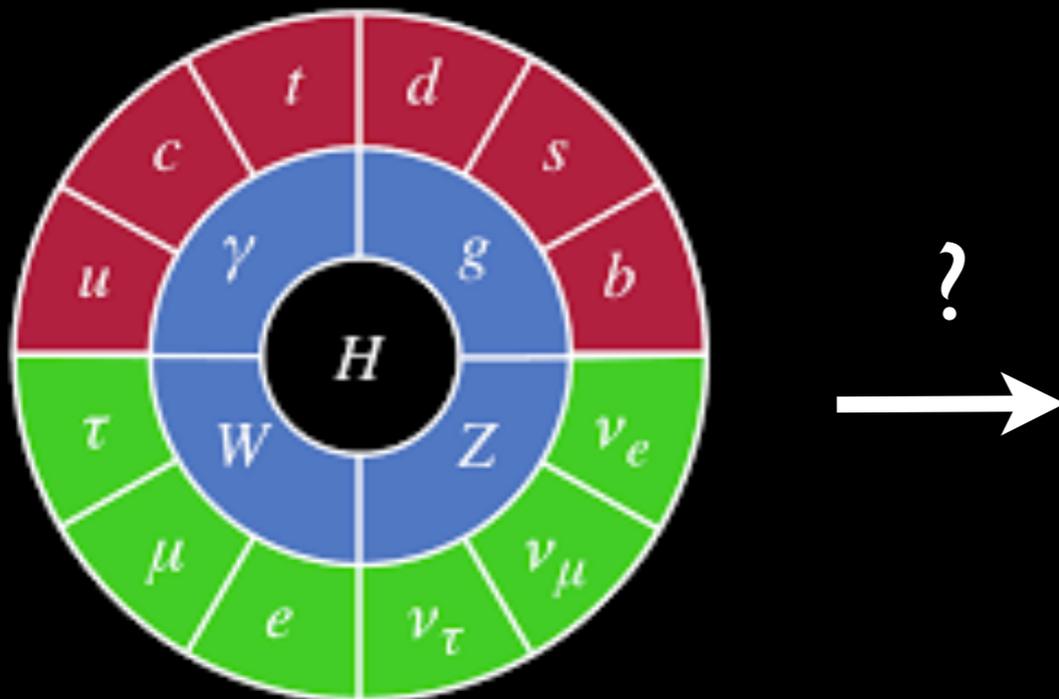
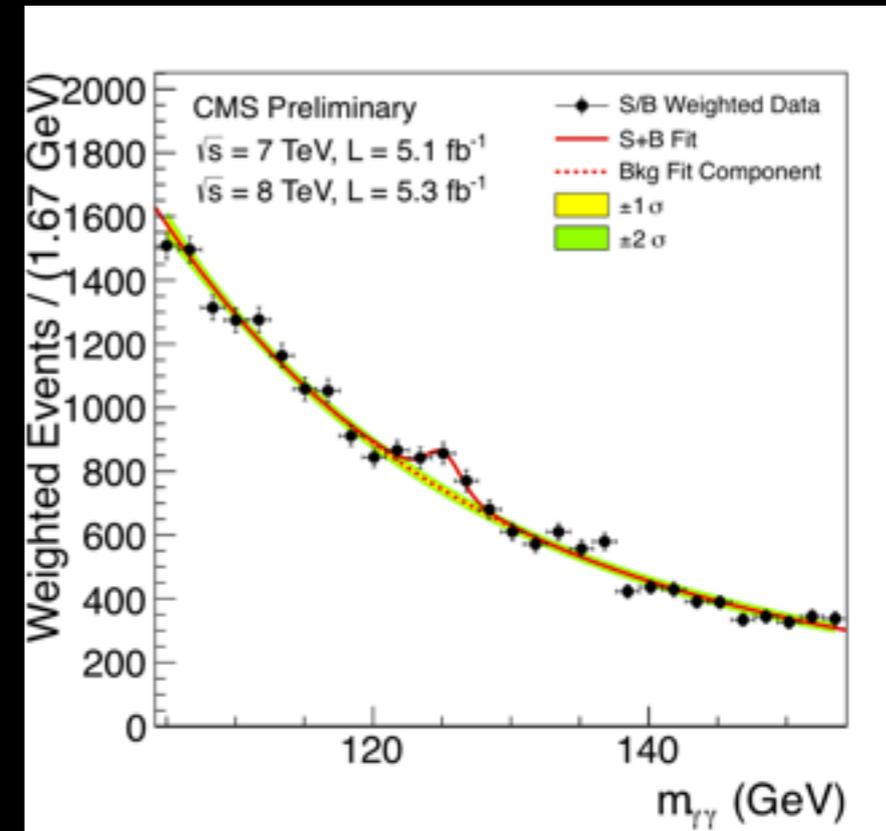
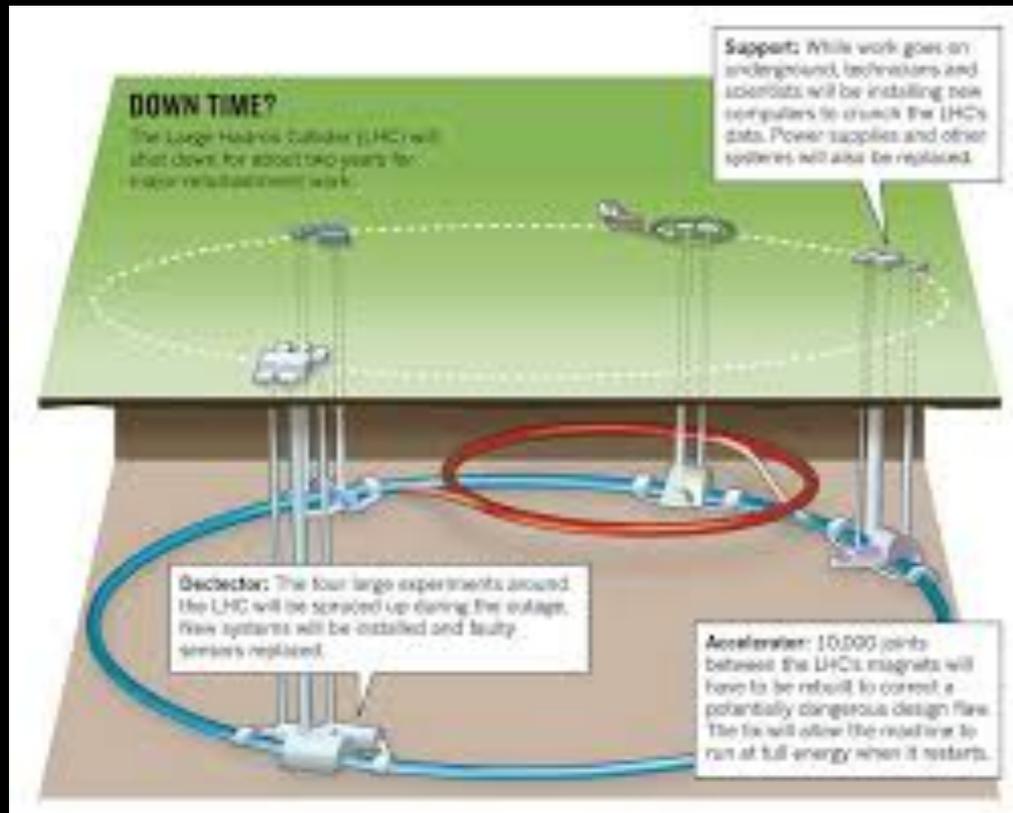
QCD is multi-facet

This school is about QCD - the theory of strong interactions. Strong interactions are interactions between hadrons; they determine hadron properties, such as masses, magnetic moments, lifetimes, form factors etc.

Since our focus will be the LHC physics, we will mostly talk about QCD in the context of quarks and gluons.



What is the physics beyond the Standard Model?



Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1-2 \tau$	2-10 jets/3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.85 TeV	$m(\tilde{q})=m(\tilde{g})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{q}	1.57 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(1^{\text{st}} \text{ gen. } \tilde{q})=m(2^{\text{nd}} \text{ gen. } \tilde{q})$	ATLAS-CONF-2017-022
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	3.2	\tilde{q}	608 GeV	$m(\tilde{q})-m(\tilde{\chi}_1^0) < 5$ GeV	1604.07773
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.02 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0 \rightarrow q\tilde{q}W^\pm\tilde{\chi}_1^\pm$	0	2-6 jets	Yes	36.1	\tilde{g}	2.01 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}_1^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\nu\nu)\tilde{\chi}_1^0$	3 e, μ	4 jets	-	36.1	\tilde{g}	1.825 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2017-030
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0	7-11 jets	Yes	36.1	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2017-033
	GMSB ($\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	\tilde{g}	2.0 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	1607.05979
	GGM (bino NLSP)	2 γ	-	Yes	3.2	\tilde{g}	1.65 TeV	$c\tau(\text{NLSP}) < 0.1$ mm	1606.09150
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 950$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu < 0$	1507.05493
GGM (higgsino-bino NLSP)	γ	2 jets	Yes	13.3	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) > 680$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu > 0$	ATLAS-CONF-2016-066	
GGM (higgsino NLSP)	2 e, μ (Z)	2 jets	Yes	20.3	\tilde{g}	900 GeV	$m(\text{NLSP}) > 430$ GeV	1503.03290	
Gravitino LSP	0	mono-jet	Yes	20.3	$F^{1/2}$ scale	865 GeV	$m(\tilde{G}) > 1.8 \times 10^{-1} e, m(\tilde{g})=m(\tilde{q})=1.5$ TeV	1502.01518	
3 rd gen. \tilde{g} med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	36.1	\tilde{g}	1.92 TeV	$m(\tilde{\chi}_1^0) < 600$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	36.1	\tilde{g}	1.97 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^+$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 300$ GeV	1407.0600
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	36.1	\tilde{b}_1	950 GeV	$m(\tilde{\chi}_1^0) < 420$ GeV	ATLAS-CONF-2017-038
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 e, μ (SS)	1 b	Yes	36.1	\tilde{b}_1	275-700 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_1^0) + 100$ GeV	ATLAS-CONF-2017-030
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	0-2 e, μ	1-2 b	Yes	4.7/13.3	\tilde{t}_1	117-170 GeV, 200-720 GeV	$m(\tilde{\chi}_1^\pm) = 2m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0)=55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 e, μ	0-2 jets/1-2 b	Yes	20.3/36.1	\tilde{t}_1	90-198 GeV, 205-950 GeV	$m(\tilde{\chi}_1^0)=1$ GeV	1506.08616, ATLAS-CONF-2017-020
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	3.2	\tilde{t}_1	90-323 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)=5$ GeV	1604.07773
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-600 GeV	$m(\tilde{\chi}_1^0) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	36.1	\tilde{t}_2	200-790 GeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2017-019
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 e, μ	4 b	Yes	36.1	\tilde{t}_2	320-880 GeV	$m(\tilde{\chi}_1^0)=0$ GeV	ATLAS-CONF-2017-019
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	36.1	$\tilde{\ell}$	320-400 GeV	$m(\tilde{\chi}_1^0)=0$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp, \tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\tilde{\ell}\bar{\nu})$	2 e, μ	0	Yes	36.1	$\tilde{\chi}_1^\pm$	720 GeV	$m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp/\tilde{\chi}_2^0, \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}\nu(\tilde{\tau}\bar{\nu}), \tilde{\chi}_2^0 \rightarrow \tilde{\tau}\tau(\nu\bar{\nu})$	2 τ	0	Yes	36.1	$\tilde{\chi}_1^\pm$	720 GeV	$m(\tilde{\chi}_1^0)=0, m(\tilde{\tau}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L(\bar{\nu}\nu), \tilde{\ell}\tilde{\nu}\tilde{\ell}_L(\bar{\nu}\nu)$	3 e, μ	0	Yes	36.1	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	1.16 TeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0Z\tilde{\chi}_1^0$	2-3 e, μ	0-2 jets	Yes	36.1	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	580 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \tilde{\ell}$ decoupled	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0h\tilde{\chi}_1^0$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	270 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \tilde{\ell}$ decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_R\ell$	4 e, μ	0	Yes	20.3	$\tilde{\chi}_2^0$	635 GeV	$m(\tilde{\chi}_2^0)=m(\tilde{\chi}_3^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_2^0)+m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	1 $e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	115-370 GeV	$c\tau < 1$ mm	1507.05493
	GGM (bino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	2 γ	-	Yes	20.3	\tilde{W}	590 GeV	$c\tau < 1$ mm	1507.05493
	Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	36.1	$\tilde{\chi}_1^\pm$	430 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm)=0.2$ ns
Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp$ prod., long-lived $\tilde{\chi}_1^\pm$		dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^\pm$	495 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm) < 15$ ns	1506.05332
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9	\tilde{g}	850 GeV	$m(\tilde{\chi}_1^0)=100$ GeV, $10 \mu\text{s} < \tau(\tilde{g}) < 1000$ s	1310.6584
Stable \tilde{g} R-hadron		trk	-	-	3.2	\tilde{g}	1.58 TeV		1606.05129
Metastable \tilde{g} R-hadron		dE/dx trk	-	-	3.2	\tilde{g}	1.57 TeV	$m(\tilde{\chi}_1^0)=100$ GeV, $\tau > 10$ ns	1604.04520
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \mu)+\tau(e, \mu)$		1-2 μ	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$	1411.6795
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3	$\tilde{\chi}_1^0$	440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ee\nu/\mu\nu/\mu\mu\nu$		displ. $ee/\mu\mu/\mu\mu\nu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g})=1.3$ TeV	1504.05162
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$	displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480$ mm, $m(\tilde{g})=1.1$ TeV	1504.05162	
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu$	$e\mu, e\tau, \mu\tau$	-	-	3.2	$\tilde{\nu}_\tau$	1.9 TeV	$\lambda_{311}^2=0.11, \lambda_{132/133/233}^2=0.07$	1607.08079
	Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{\text{LSP}} < 1$ mm	1404.2500
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\nu, e\mu\nu, \mu\mu\nu$	4 e, μ	-	Yes	13.3	$\tilde{\chi}_1^\pm$	1.14 TeV	$m(\tilde{\chi}_1^0) > 400$ GeV, $\lambda_{12k} \neq 0$ ($k=1, 2$)	ATLAS-CONF-2016-075
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\mp, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\nu_e, e\tau\nu_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.08 TeV	$\text{BR}(t)=\text{BR}(b)=\text{BR}(c)=0\%$	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.55 TeV	$m(\tilde{\chi}_1^0)=800$ GeV	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	2.1 TeV	$m(\tilde{\chi}_1^0)=1$ TeV, $\lambda_{112} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	1.65 TeV	$m(\tilde{t}_1)=1$ TeV, $\lambda_{323} \neq 0$	ATLAS-CONF-2017-013
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	15.4	\tilde{t}_1	410 GeV, 450-510 GeV		ATLAS-CONF-2016-022, ATLAS-CONF-2016-084	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\ell}$	2 e, μ	2 b	-	36.1	\tilde{t}_1	0.4-1.45 TeV	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/\mu) > 20\%$	ATLAS-CONF-2017-036	
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	510 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹ 1 Mass scale [TeV]

In the midst of every crisis lies an opportunity

These results suggest a paradigm change:

indirect searches for new particles and interactions at hadron colliders will feature more and more prominently in the exploration of physics beyond the Standard Model.



CMS Experiment at the LHC, CERN

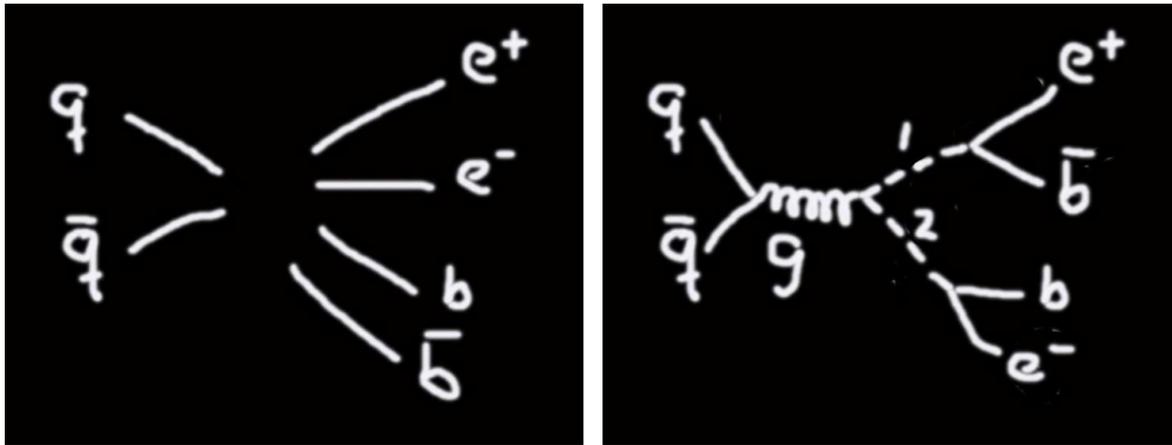
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The LHC was not envisaged as a precision machine, but it can be turned into one, provided that theory (QCD) can keep up.

The new paradigm of discovery

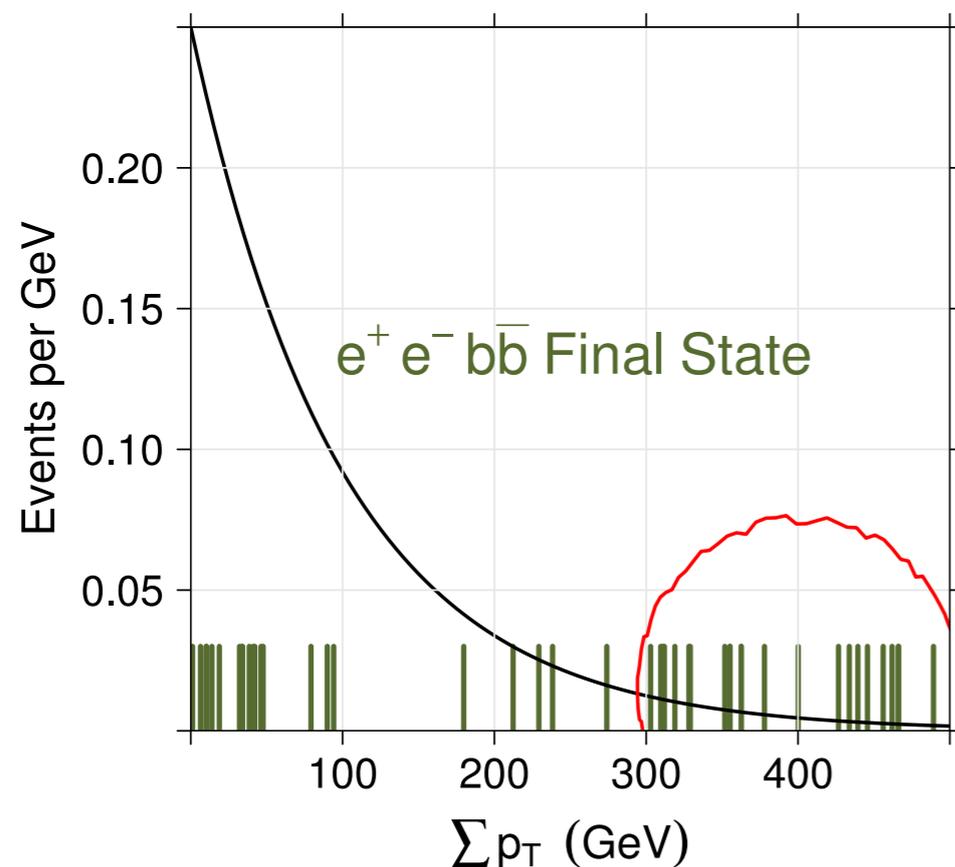
BARD: Interpreting New Frontier Energy Collider Physics



Bruce Knuteson*
MIT

Stephen Mrenna†
FNAL

In contemporary high energy physics experiments, it is not uncommon to observe discrepancies between data and Standard Model predictions. Most of these discrepancies have been explained away over time. To convincingly demonstrate that an observed effect is evidence of physics beyond the Standard Model, it is necessary to prove it is (1) not a likely statistical fluctuation, (2) not introduced by an imperfect understanding of the experimental apparatus, (3) not due to an inadequacy of the implementation of the Standard Model prediction, and (4) interpretable in terms of a sensible underlying theory. Those who object to (4) as being necessary fail to appreciate that most hypothesis development in science occurs before, rather than after, publication. This last criterion is essential, and will likely point the way to other discrepancies that must exist if the interpretation is correct.



Main idea behind this paper was to search systematically for a correlated set of deviations from the SM predictions and a possibility to explain them with a single NP hypothesis. With null search results from the LHC, this idea becomes extremely timely...

The original wishlist

Knuteson then came up with the “next-to-leading order (NLO) wishlist” (circa 2004), i.e. the list of processes whose reliable description he thought was instrumental for making his idea a reality. The appearance of the wishlist started a concerted effort by theorists to improve ways and means to perform NLO computations -- the beginning of the NLO revolution.

Note that we have ticked off one cross section from the first list

An experimenter’s wishlist

Run II Monte Carlo Workshop

Single Boson	Diboson	Triboson	Heavy Flavour
$W+ \leq 5j$	$WW+ \leq 5j$	$WWW+ \leq 3j$	$t\bar{t}+ \leq 3j$
$W + b\bar{b} \leq 3j$	$W + b\bar{b}+ \leq 3j$	$WWW + b\bar{b}+ \leq 3j$	$t\bar{t} + \gamma+ \leq 2j$
$W + c\bar{c} \leq 3j$	$W + c\bar{c}+ \leq 3j$	$WWW + \gamma\gamma+ \leq 3j$	$t\bar{t} + W+ \leq 2j$
$Z+ \leq 5j$	$ZZ+ \leq 5j$	$Z\gamma\gamma+ \leq 3j$	$t\bar{t} + Z+ \leq 2j$
$Z + b\bar{b}+ \leq 3j$	$Z + b\bar{b}+ \leq 3j$	$ZZZ+ \leq 3j$	$t\bar{t} + H+ \leq 2j$
$Z + c\bar{c}+ \leq 3j$	$ZZ + c\bar{c}+ \leq 3j$	$WZZ+ \leq 3j$	$t\bar{b} \leq 2j$
$\gamma+ \leq 5j$	$\gamma\gamma+ \leq 5j$	$ZZZ+ \leq 3j$	$b\bar{b}+ \leq 3j$
$\gamma + b\bar{b} \leq 3j$	$\gamma\gamma + b\bar{b} \leq 3j$		
$\gamma + c\bar{c} \leq 3j$	$\gamma\gamma + c\bar{c} \leq 3j$		
	$WZ+ \leq 5j$		
	$WZ + b\bar{b} \leq 3j$		
	$WZ + c\bar{c} \leq 3j$		
	$W\gamma+ \leq 3j$		
	$Z\gamma+ \leq 3j$		

Precision physics at the LHC?

Systematic precision studies at hadron colliders, aimed at discovering New Physics through indirect effects, have never been attempted before.

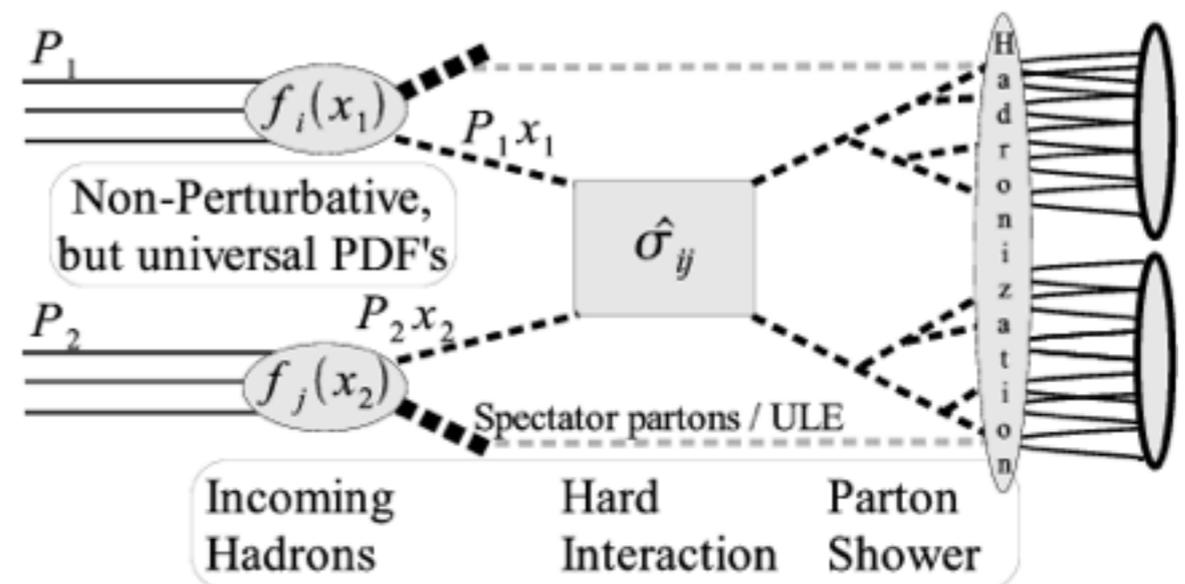
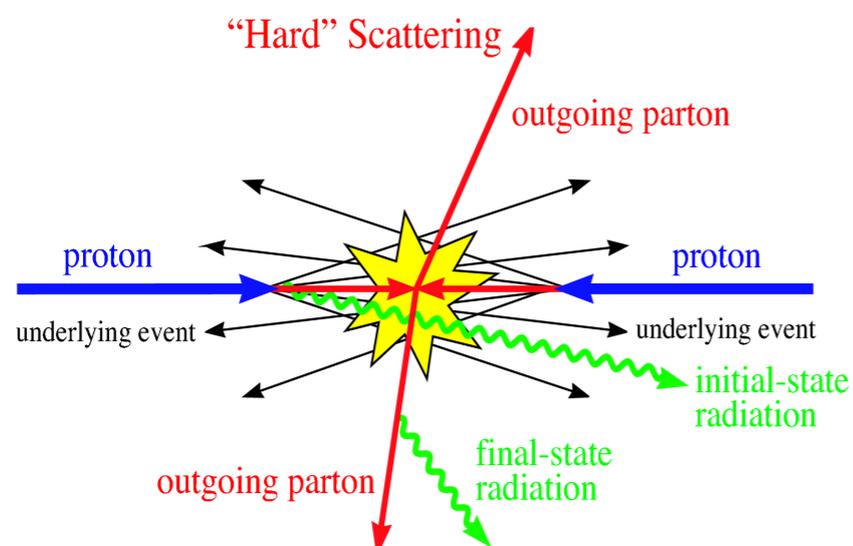
Indeed, hadrons are composite particles kept together by a poorly understood strong force. If we can't understand (compute) properties of one proton, can we confidently describe what happens when two protons collide?

We believe that, to some extent, this can be done and that outcomes of certain (hard) hadron collisions can be understood starting directly from the SM Lagrangian.

Sufficiently inclusive hard hadron processes can be described by the collinear factorization formula.

Collins, Soper, Sterman

$$d\sigma = \int dx_1 dx_2 f_i(x_1) f_j(x_2) d\sigma_{ij}(x_1, x_2) F_J(1 + \mathcal{O}(\Lambda_{\text{QCD}}^n / Q^n)) \quad n \geq 1$$



LHC: the world of quarks and gluons

Hard scattering processes at the LHC can be understood in terms of quarks and gluons; very limited knowledge about protons is needed. Physics of quarks and gluons is governed by a non-abelian SU(3) gauge-field theory -- the QCD.

1) The Lagrangian

$$\mathcal{L}_{\text{QCD}} = \sum \bar{q}_j \left(i\hat{D} - m_j \right) q_j - \frac{1}{4} G_{\mu\nu}^a G^{a,\mu\nu}$$

$$D_\mu = \partial_\mu - ig_s T^a A_\mu^a, \quad G_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + ig_s f^{abc} A_\mu^b A_\nu^c. \quad [T^a, T^b] = if_{abc} T^c$$

2) **Degrees of freedom**: quarks (up, down, strange, charm, bottom, top) and gluons (also ghosts, see later).

3) **SU(3) group** --> **interaction charges (color)** --> each quark can appear in one of the three color states and a gluon in one of eight.

QCD Feynman rules

$$\begin{array}{c} a, \mu \quad p \quad b, \nu \\ \text{-----} \\ \text{-----} \end{array} = \frac{-i\delta^{ab}}{p^2} \left(-g_{\mu\nu} + \xi \frac{p_\mu p_\nu}{p^2} \right)$$

$$\begin{array}{c} j, \beta \quad i, \alpha \\ \text{-----} \\ p \end{array} = \left(\frac{i}{\hat{p}} \right)_{\alpha\beta} \delta^{ij}$$

$$\begin{array}{c} a, \mu \\ | \\ k \\ / \quad \backslash \\ b, \nu \quad p \quad c, \rho \quad q \end{array} = g_s f^{abc} \left(\begin{array}{l} g^{\mu\nu} (k-p)^\rho \\ + g^{\nu\rho} (p-q)^\mu \\ + g^{\rho\mu} (q-k)^\nu \end{array} \right)$$

$$\begin{array}{c} a, \mu \\ | \\ i \quad j \end{array} = ig_s \gamma^\mu T_{ij}^a$$

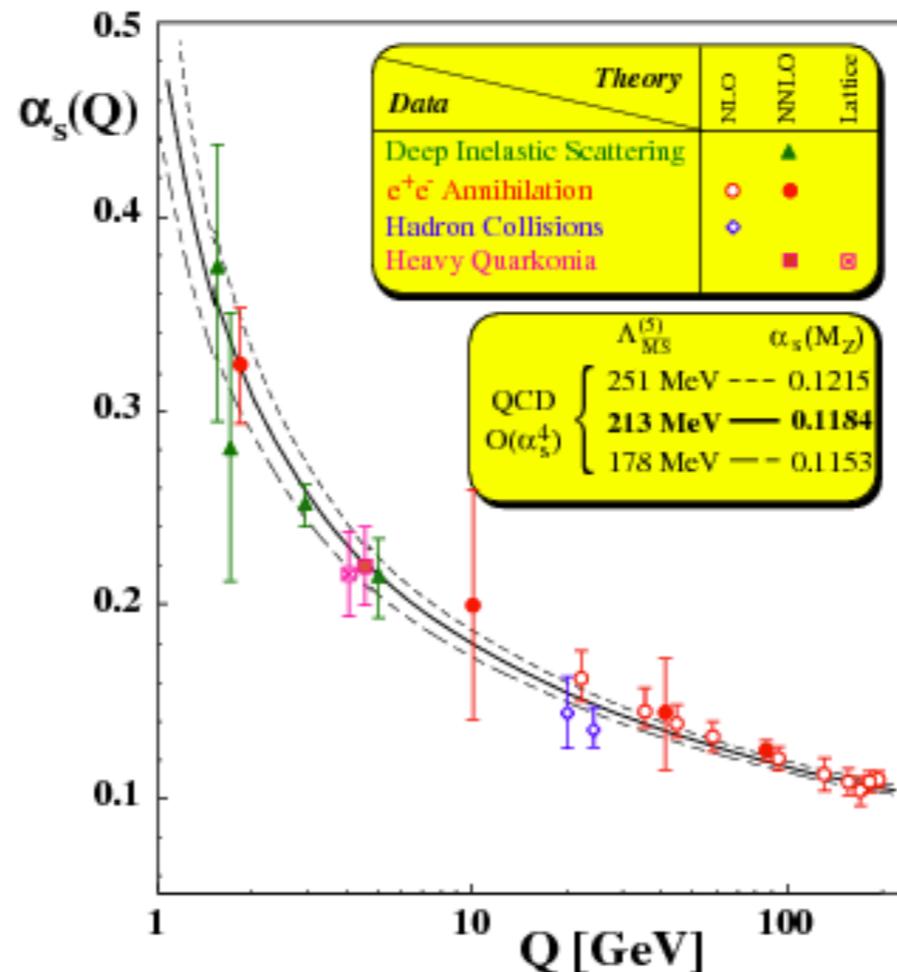
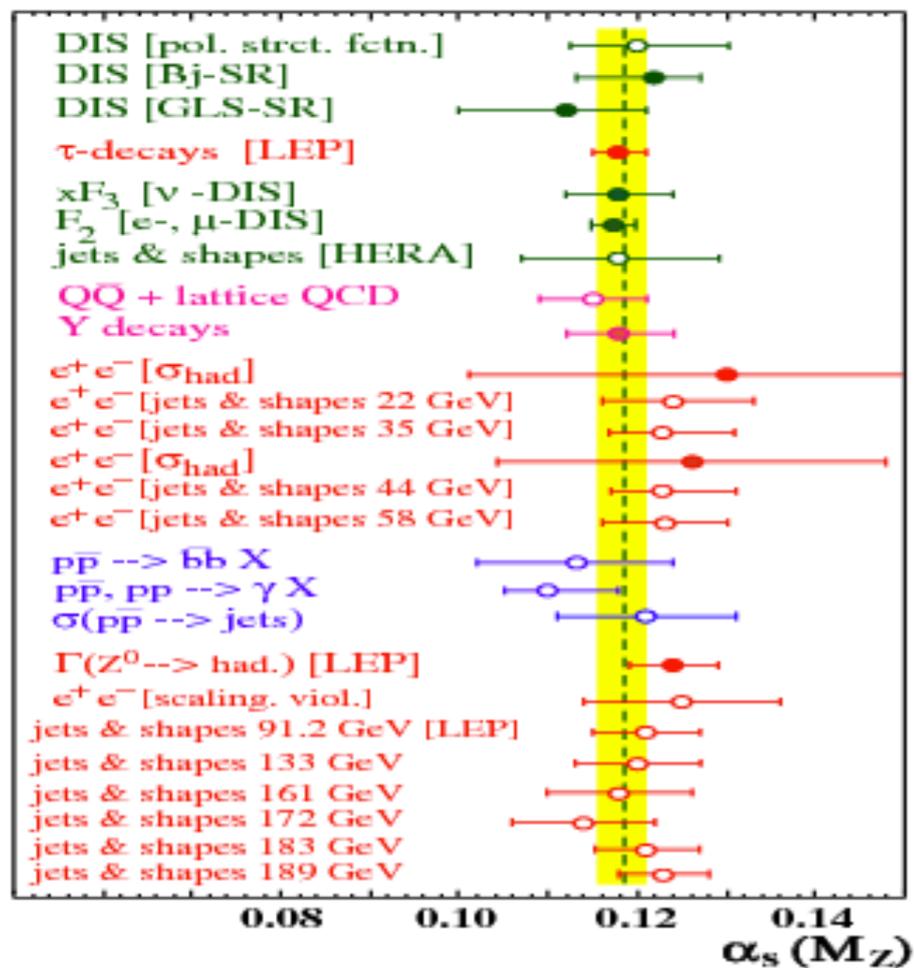
$$\begin{array}{c} a, \mu \quad b, \nu \\ \diagdown \quad \diagup \\ \diagup \quad \diagdown \\ c, \rho \quad d, \sigma \end{array} = -ig_s^2 \left[\begin{array}{l} f^{abe} f^{cde} (g^{\mu\rho} g^{\nu\sigma} - g^{\mu\sigma} g^{\nu\rho}) \\ + f^{ace} f^{bed} (g^{\mu\nu} g^{\rho\sigma} - g^{\mu\sigma} g^{\nu\rho}) \\ + f^{ade} f^{bce} (g^{\mu\nu} g^{\rho\sigma} - g^{\mu\rho} g^{\nu\sigma}) \end{array} \right]$$

$$\begin{array}{c} a \quad b \\ \text{-----} \\ p \end{array} = \frac{i\delta^{ab}}{p^2}$$

$$\begin{array}{c} b, \mu \\ | \\ a, p \quad c \end{array} = -g_s f^{abc} p^\mu$$

The running coupling constant

The effective coupling constant in QCD changes in such a way that it decreases at large momenta transfers (short distances). This phenomenon, known as the asymptotic freedom, enables us to describe perturbatively hard scattering processes at the LHC.



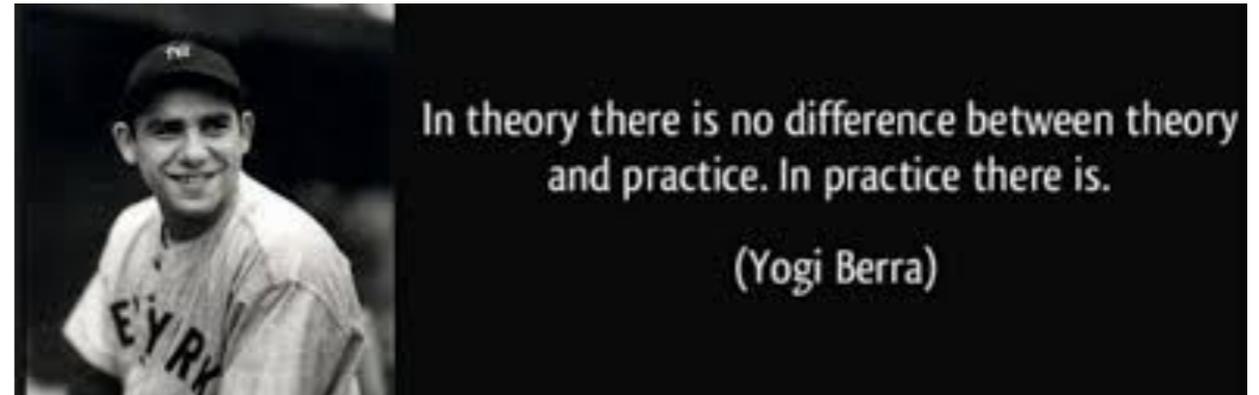
$$\alpha_s(\mu) = \frac{1}{\beta_0 \log \frac{\mu^2}{\Lambda^2}},$$

$$\beta_0 = \frac{33 - 2n_f}{12\pi} \sim 0.5|_{n_f=5}$$

$$\Lambda \approx 300 \text{ MeV}$$

Perturbation theory for cross sections

Starting from the QCD Feynman rules -- and fixing initial and final states for which we would like to know the scattering amplitudes and the order in perturbation theory, we can put together Feynman diagrams and calculate them using standard rules of perturbative QFTs.



Sounds simple but not quite:

- need to deal with (very) large number of diagrams and integrals (algebraic complexity);

Badger

- need to be able to compute complicated loop integrals (analytic complexity);

Tancredi

- need to understand complex interplay between final states with different multiplicities, to arrive at predictions that are insensitive to long-distance (non-perturbative) effects.

Röntsch

Cross sections and observables

In addition, quite often we have to go **beyond fixed order perturbation theory**:

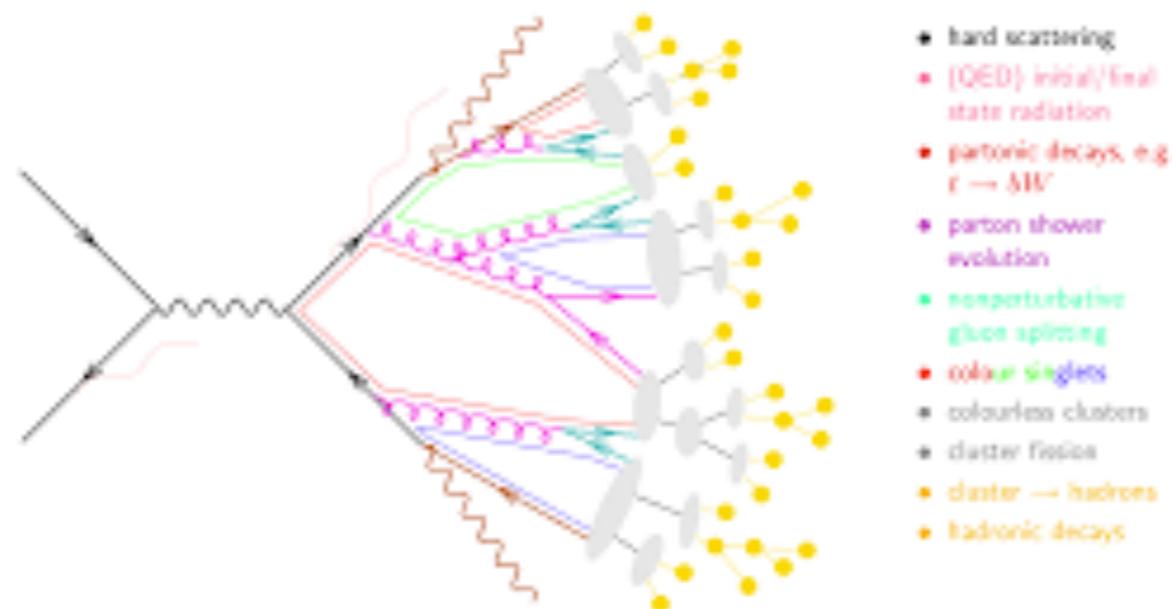
- phase-space regions where fixed order perturbation theory fails but long-distance effects are still suppressed;

$$\frac{d\sigma}{dp_{\perp}} = \sum_k^{\infty} \alpha_s^k \ln^{2k} \frac{p_{\perp}}{m} c_k$$

Monni

- hadrons, not quarks and gluons hit LHC detectors; need to model the transition to connect theoretical computations with experimental measurements.

Hoeche



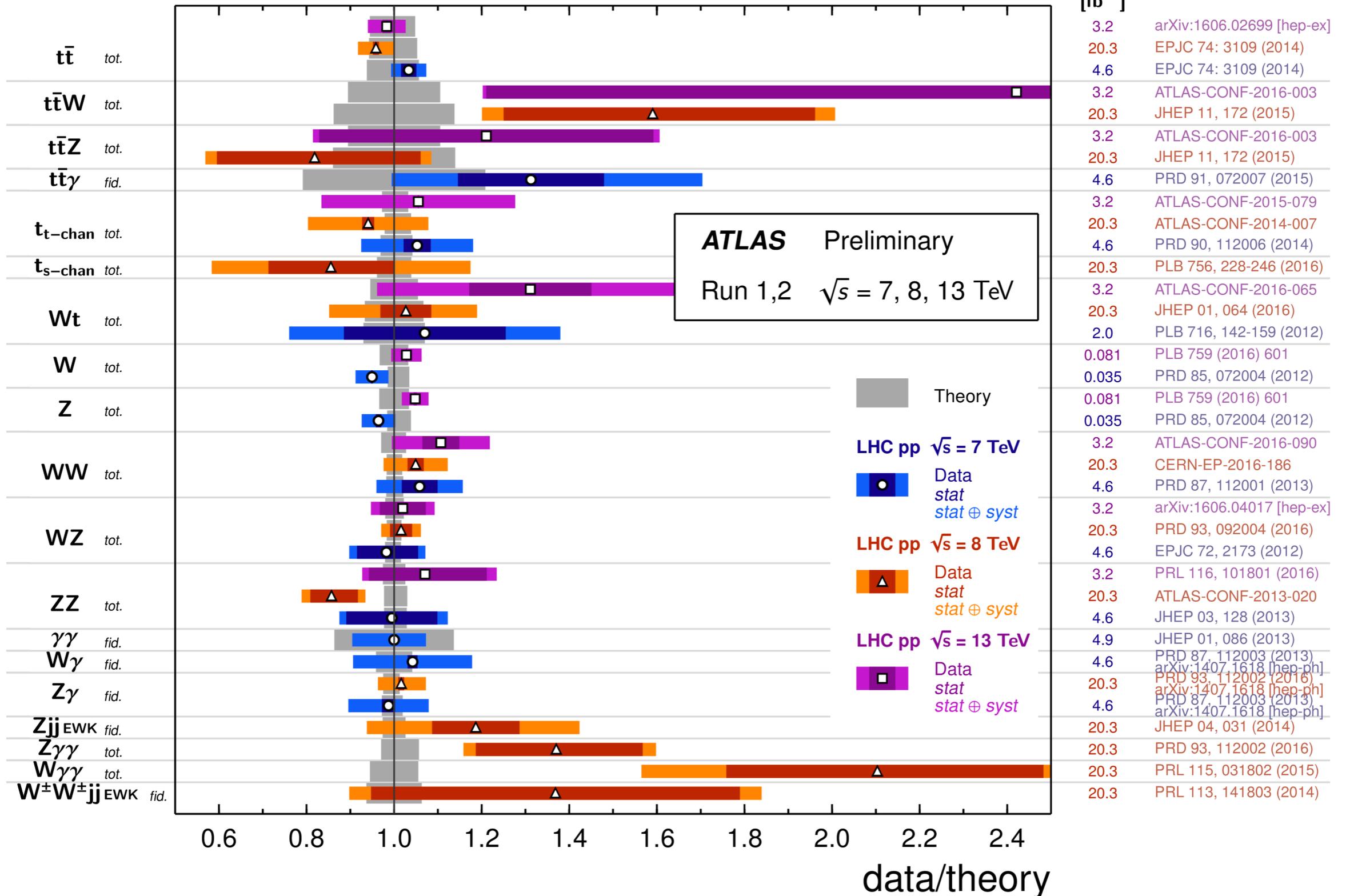
What has been studied and how well?

What has been studied and how well?

Standard Model Production Cross Section Measurements

Status: August 2016 $\int \mathcal{L} dt$
[fb⁻¹]

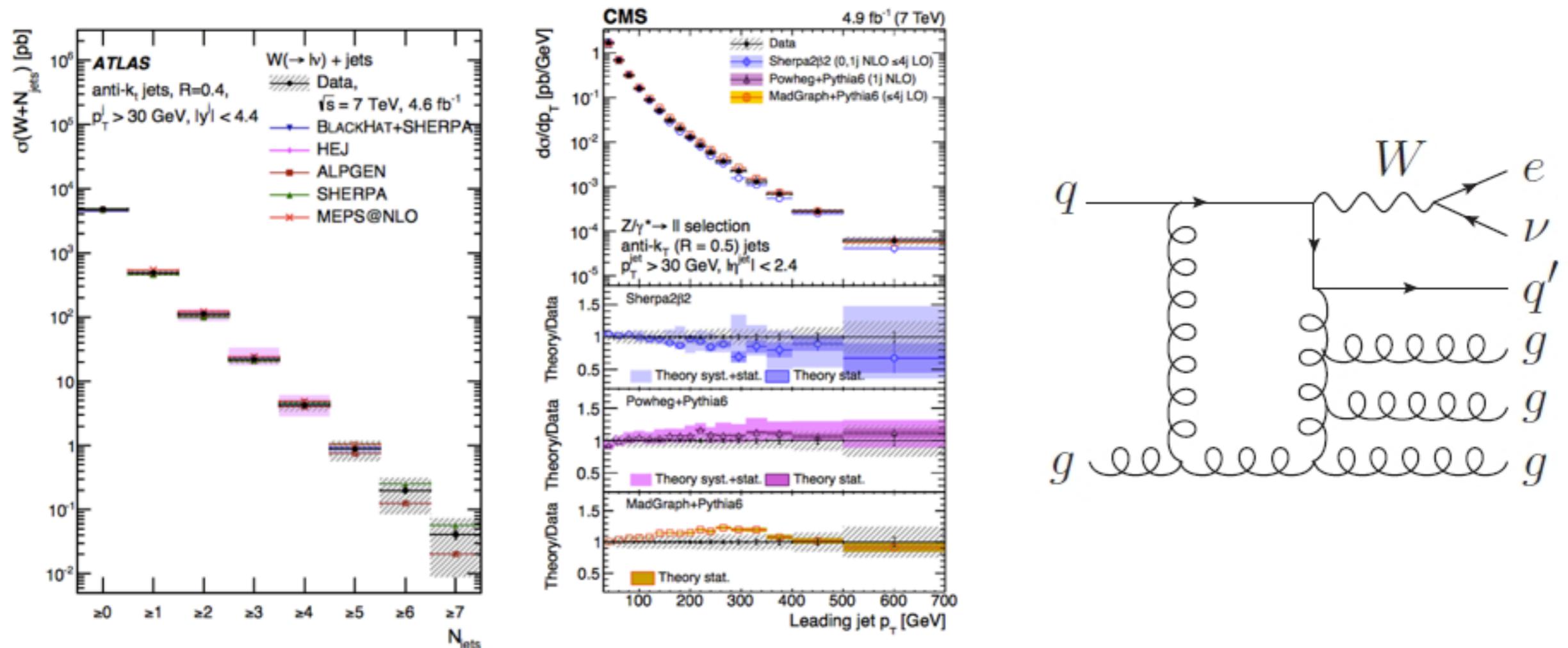
Reference



What has been studied and how well?

Vector boson production in association with jets (up to seven!). Need to describe kinematic properties of jets; sufficient statistics to be sensitive to O(1-10) percent accuracy.

Physics: PDFs; backgrounds to BSM searches with complex signatures.



Gigantic number of tree-level and one-loop diagrams needed to compute cross-sections and kinematic distributions. Very hard if traditional methods are used -- **modern unitary-based methods for one-loop computations.**

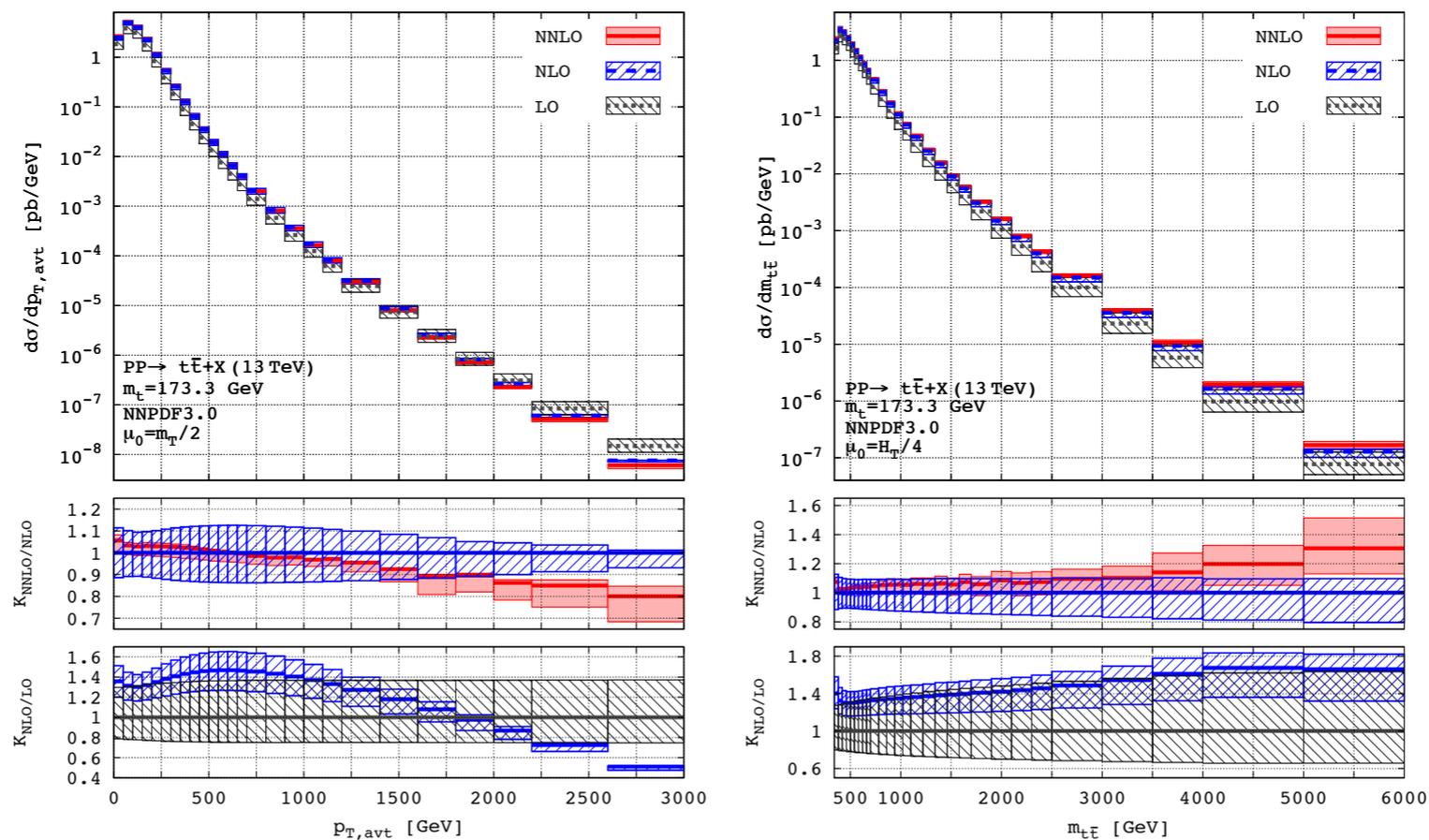
What has been studied and how well?

Top quark pair production cross section is currently measured to about 3 percent -- need NNLO QCD accuracy.

Physics: backgrounds to top-like BSM physics, gluon PDF, top quark mass.

- ▶ fully-differential NNLO-QCD predictions for $t\bar{t}$ production

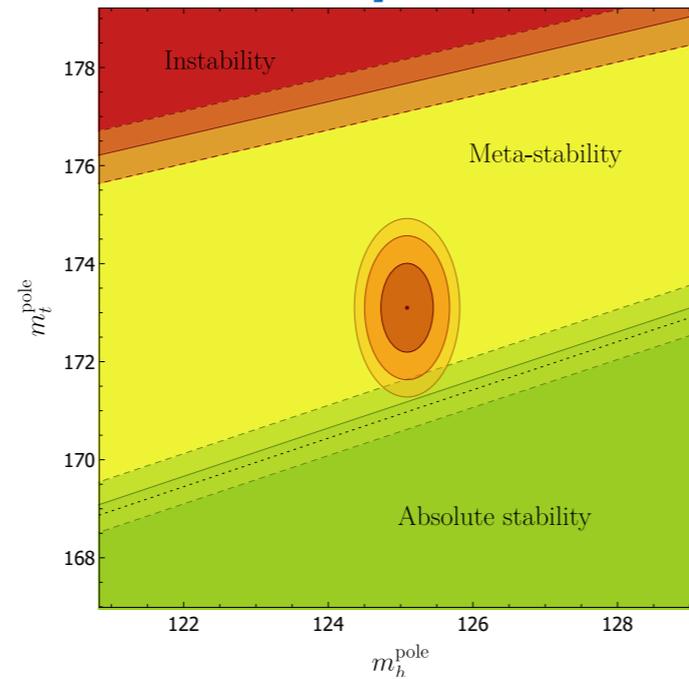
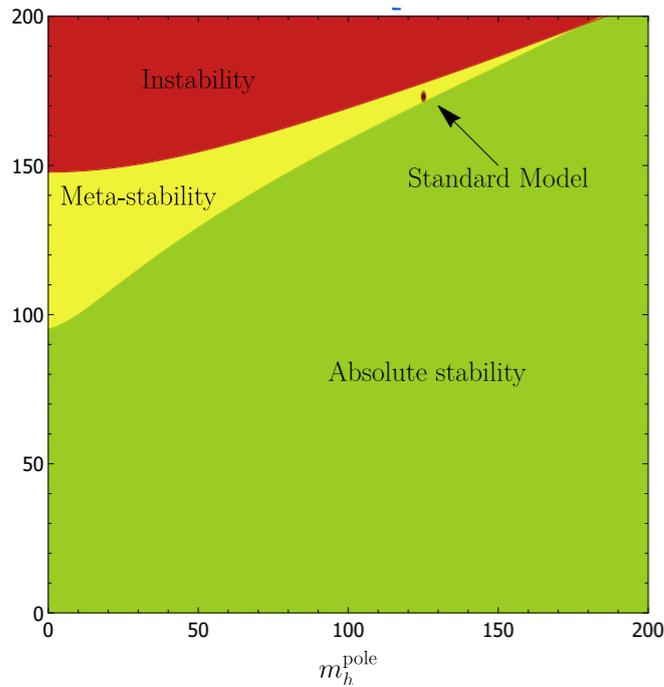
[Czakon, Heymes, Mitov '16]



Very complex loop diagrams with massive particles inside -- we still do not know how to compute them analytically. Complicated subtraction scheme to put together final states with different multiplicities.

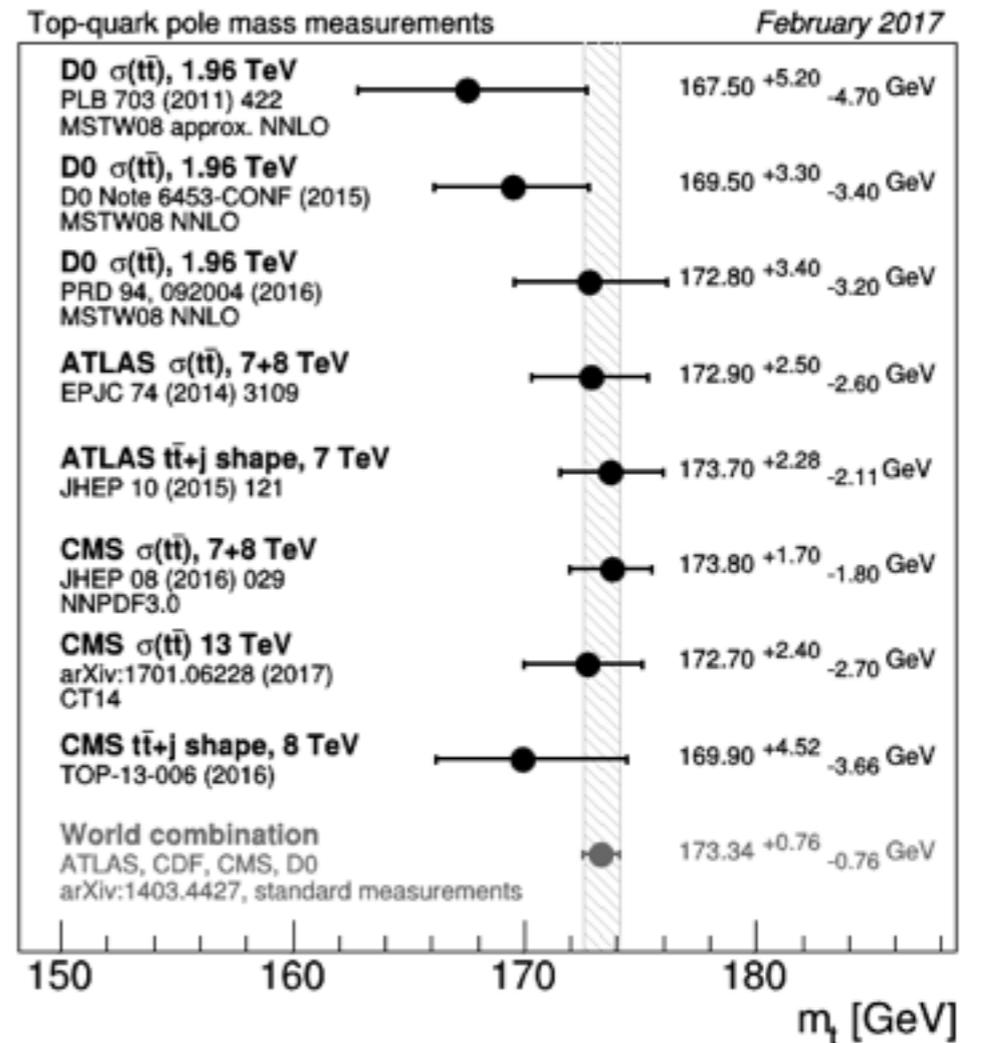
What has been studied and how well?

Top quark mass may shine brightly, as it holds keys to the fate of our Universe, but it has a past.. a dark connection to the underworld of non-perturbative QCD.



$$\tau_{\text{SM}} = \left(\frac{\Gamma}{V} \right)^{-1/4} = 10^{139+102}_{-51} \text{ years}$$

Uncertainty equal parts m_t ,
 α_s , threshold corrections

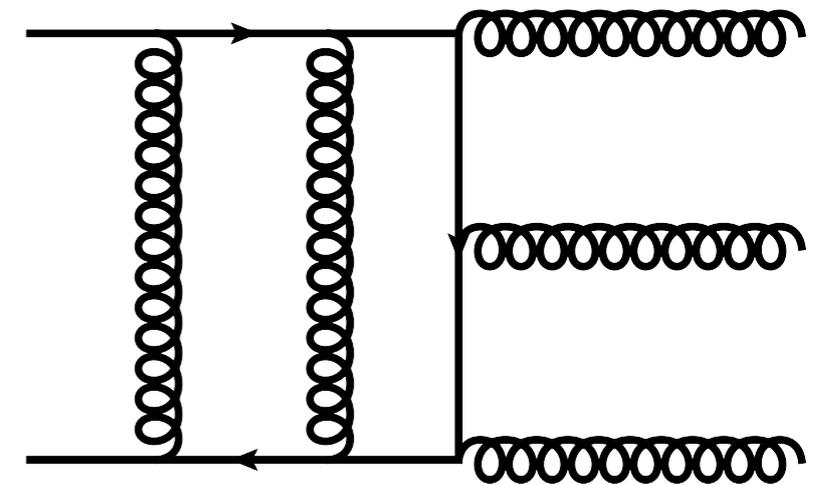
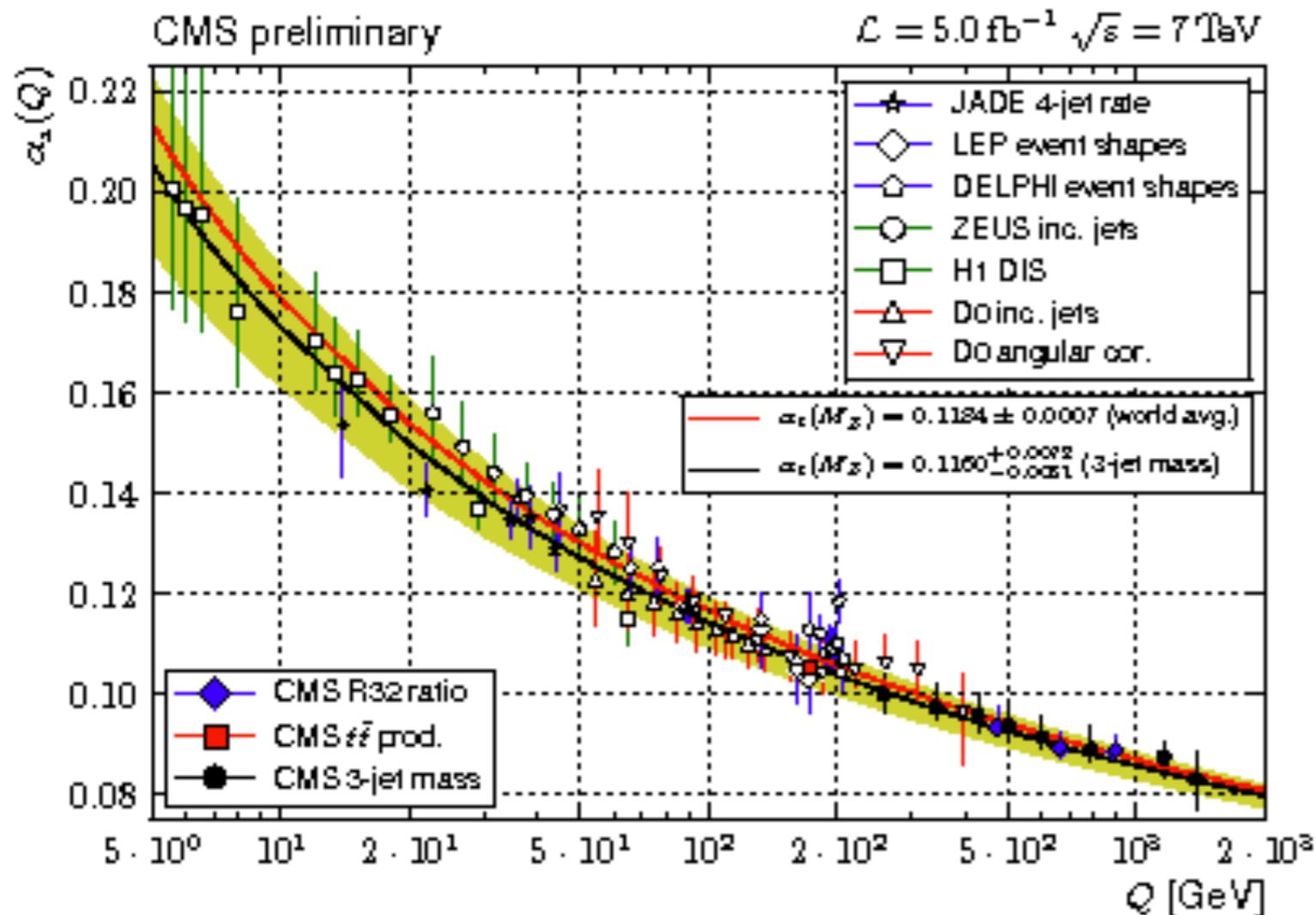


The most precise measurements of the top quark pole mass rely on parton shower event generators, to correlate kinematics of complex final states with the value of the top quark mass and to estimate non-perturbative effects.

Where do we go from here ?

The strong coupling constant

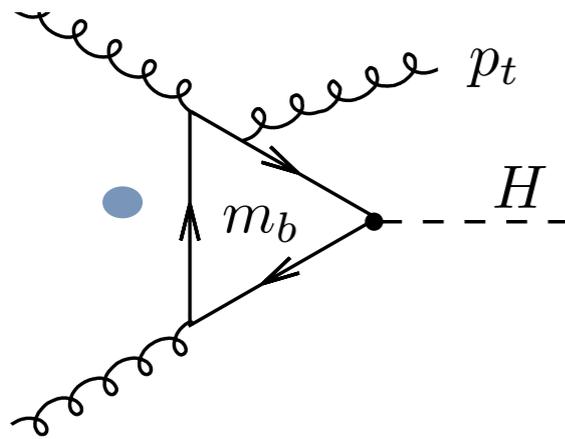
Measurement of the strong coupling constant at the shortest distances. Statistically unlimited. Need NNLO QCD corrections to three jet production at the LHC.



Very complicated loop integrals; enormous amount of algebra. Will the current attempts to compute the 3-jet rate lead to new groundbreaking methods for multi-loop computations?

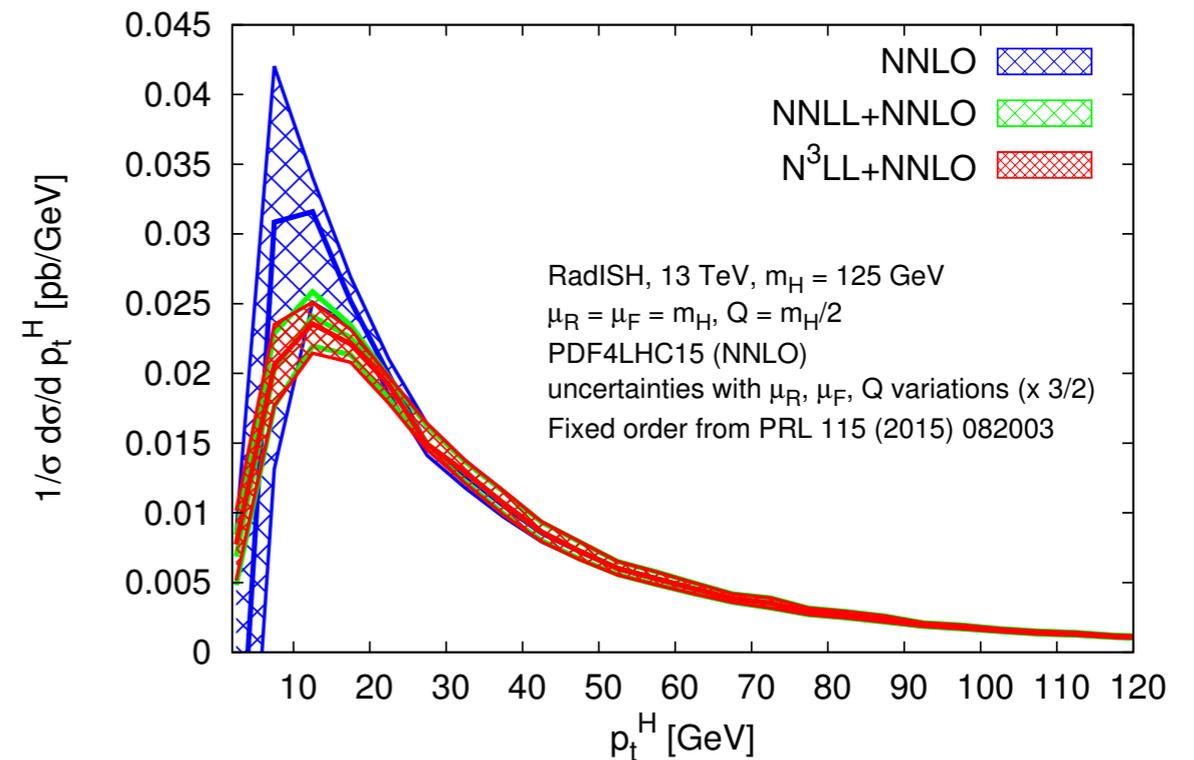
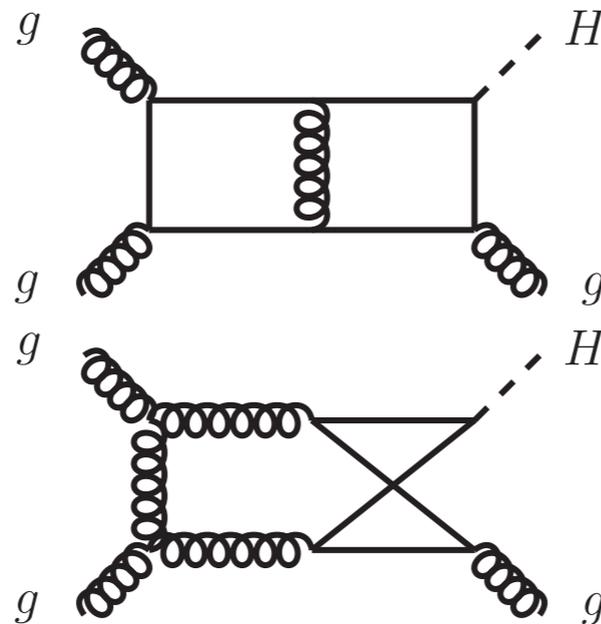
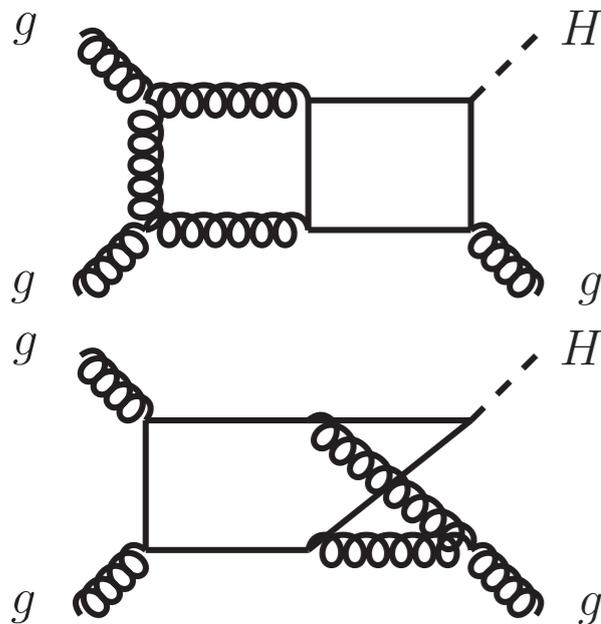
Light quark Yukawa couplings

Higgs transverse momentum distribution can be used to study Yukawa couplings of light quarks (b,c). Unconventional logarithms that depend on the light quark masses. **Re-summation** of these logarithms is not understood. Impacts precision of the transverse momentum distribution.



$$m_b/p_{\perp} \ll 1$$

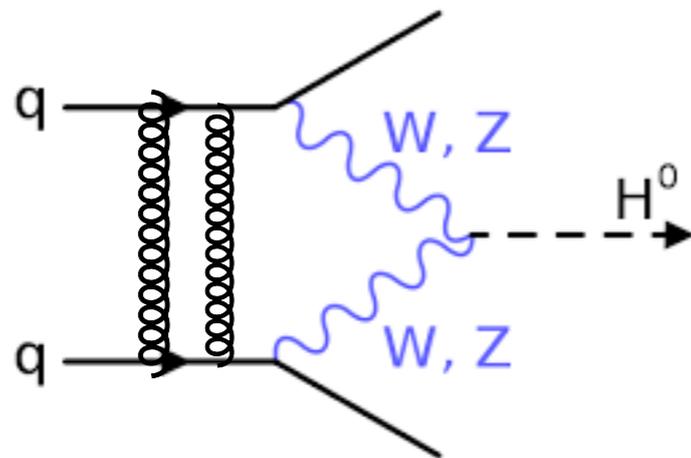
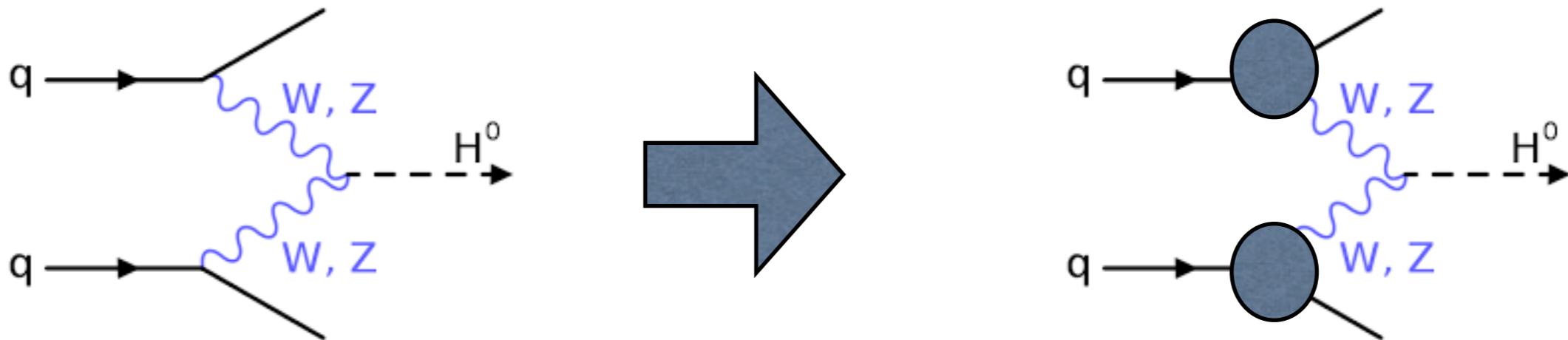
$$\alpha_s \ln^2 \frac{p_{\perp}}{m_b} \gg 1$$



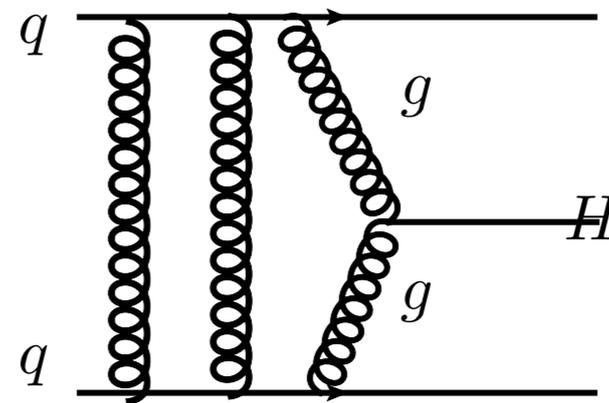
Bizon, Monni, Re, Rottoli, Torrielli

Higgs gauge coupling

Weak boson fusion gives access to Higgs coupling to vector boson -- same coupling that breaks the electroweak symmetry? Very precise predictions are available -- [in the factorization approximation](#). Time to go beyond it. Computing H+2jets at NNLO for weak boson fusion kinematics will help further to improve the precision of the coupling extraction.



Beyond the factorization approximation



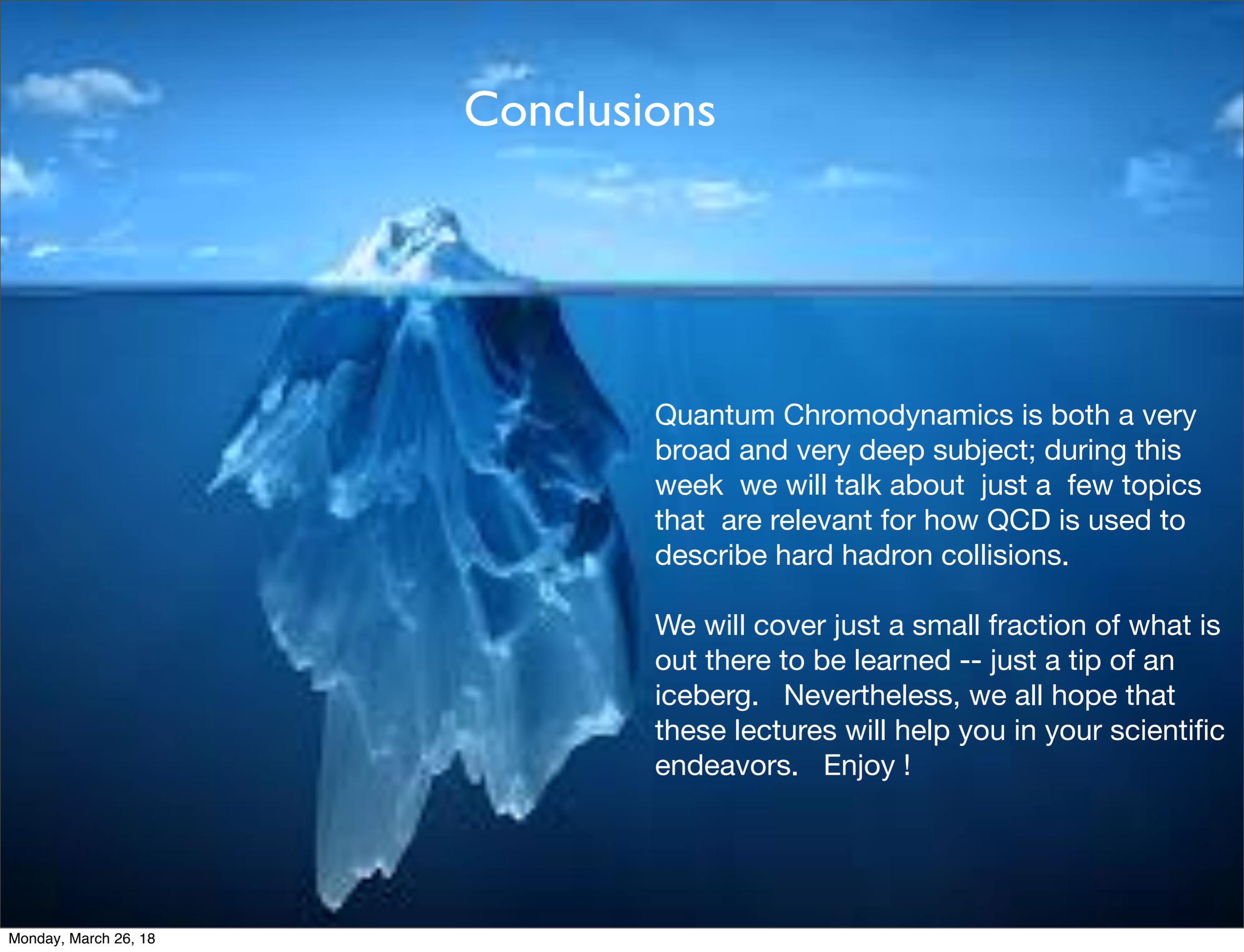
Background at NNLO

What will we need from pQCD

For this research program to succeed, it will be important:

- to understand better the role of non-perturbative effects in collider QCD and their impact on observables;
- to make parton showers a systematically-improvable approximation;
- to understand re-summations for more exclusive quantities;
- to learn how to get around the exponential growth in algebraic complexity in multi-loop multi-parton computations ;
- to find robust ways to compute Feynman integrals that do not rely on simplifications allowed by special cases (multi-mass cases etc.);
- to design physically-transparent and efficient subtraction schemes at NNLO. Connect them to resummations and parton showers.

Conclusions

The background of the slide is a blue-tinted photograph of an iceberg. The iceberg is mostly submerged, with only a small, jagged peak visible above the dark blue water. The sky is a lighter blue with some wispy clouds. The overall mood is serene and contemplative.

Quantum Chromodynamics is both a very broad and very deep subject; during this week we will talk about just a few topics that are relevant for how QCD is used to describe hard hadron collisions.

We will cover just a small fraction of what is out there to be learned -- just a tip of an iceberg. Nevertheless, we all hope that these lectures will help you in your scientific endeavors. Enjoy !