

Integral luminosity measurement at CEPC

- update on physics background -



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- Reminder on \mathcal{L} measurement as a counting experiment
- What do we know so far on systematics in \mathcal{L} measurement at CEPC
 - Effects from mechanical uncertainties and MDI
- Physics processes as background to the Bhabha count
 - t-channel 2-photon exchange
 - First estimates at 240 GeV
- Conclusion

- Integral luminosity measurement based on Bhabha scattering is a counting experiment

$$\mathcal{L} = N_{\text{Bh}} / \sigma$$

- N_{Bh} is Bhabha count in the certain phase space and within the detector acceptance (fiducial) region
- σ is the theoretical cross-section in the same geometrical and phase space
- Both N_{Bh} and σ have to be known at the 10^{-3} (or -4) level

But, $N_{\text{Bh}} \rightarrow N_x$

- In N_{Bh} , miscounts due to various effects are contained:
 - detector resolution
 - mechanics (positioning and alignment) ✓
 - center-of-mass energy, beam synchronization, IP displacements ✓
 - physics background ✓
 - beam-induced processes (off-momentum electrons)



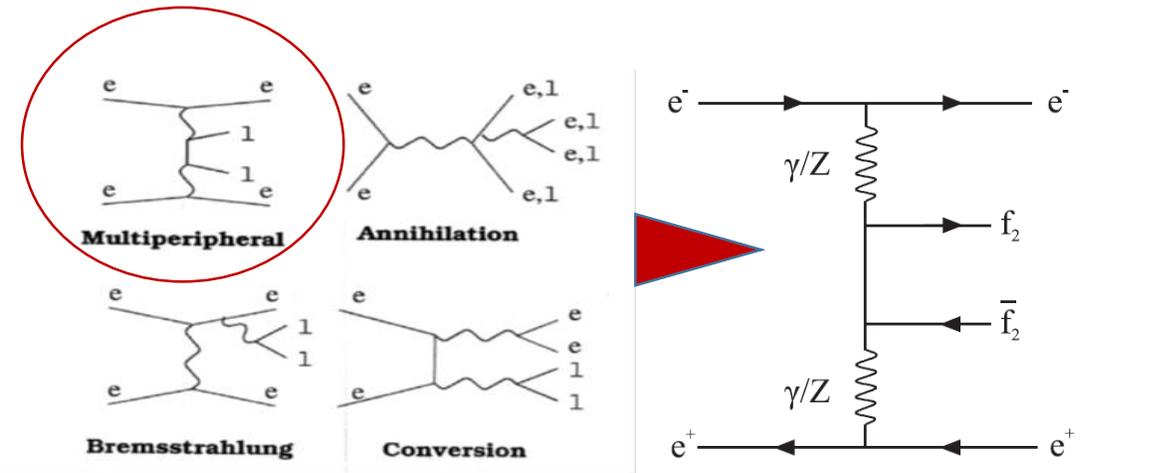
Addressed in CDR

To correct for it (recover N_{Bh}) implies that effects have to be known at 10^{-3} (or -4) level

What do we know so far on systematics in \mathcal{L} measurement at CEPC

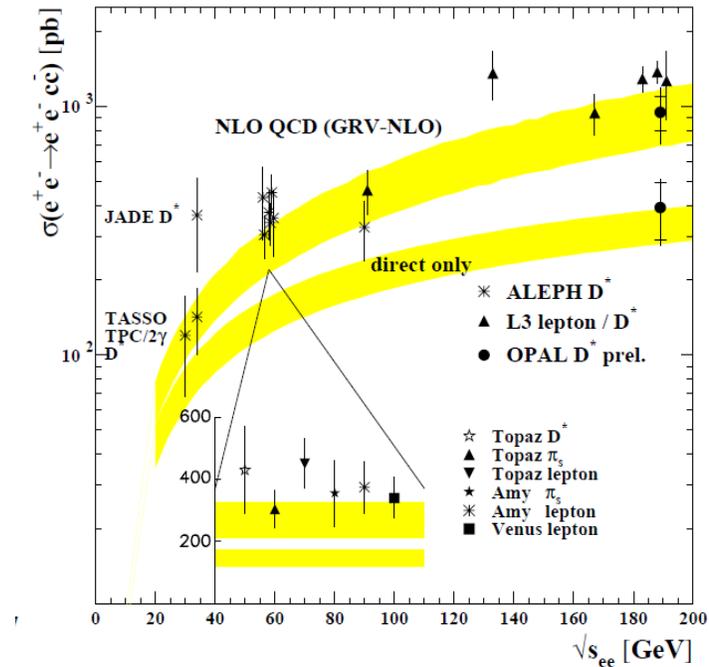
Parameter	unit		limit (LEP style)
ΔE_{CM}	MeV		120
$E_{e^+} - E_{e^-}$	MeV		240
$\frac{\delta\sigma_{E_{\text{beam}}}}{\sigma_{E_{\text{beam}}}}$			Effect cancelled
ΔX_{IP}	mm	Higgs treshold	1
ΔZ_{IP}	mm		10
Beam synchronisation	ps	$\Delta\mathcal{L} / \mathcal{L} = 10^{-3}$	15
$\sigma_{X_{\text{IP}}}$	mm		1
$\sigma_{Z_{\text{IP}}}$	mm		10
r_{in}	μm		10
$\sigma_{r_{\text{shower}}}$	mm		1
Δd_{IP}	mm		0.5
$\Delta\varphi_{\text{tilt}}$	mrad		6

Dominant effects from mechanics and MDI comes from the uncertainty of the available center of mass energy and the inner radius of the luminometer

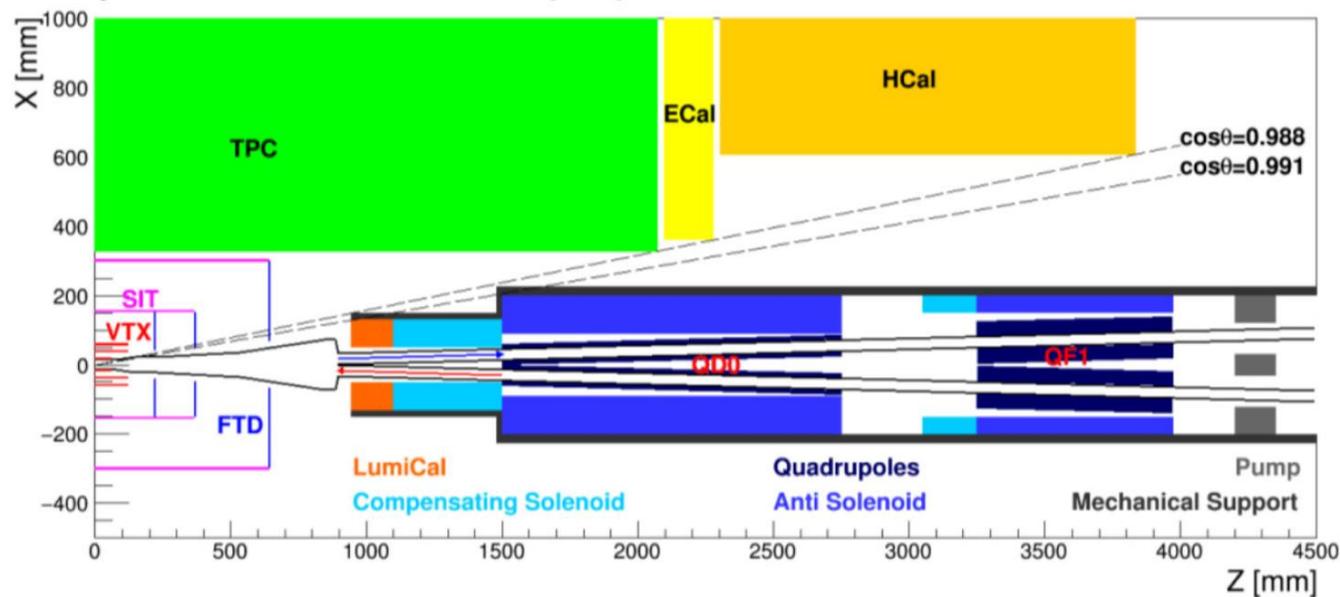


- Multiperipheral process dominates the x-section
- Cross-sections are large (\sim nb) saturating at higher energies
- **High energy e^\pm spectators can fake the signal**
- ... although most of spectators go below luminometer acceptance

In principle, uncertainty of the theoretical x-section for $ee \rightarrow eeff$ has to be known at CEPC energies in order to correct for the miscount



First estimates at 240 GeV



Simulation:

Particle tracks are projected to the front LumiCal plane

- $ee \rightarrow ee\mu\mu$ WHIZARD V2.6
- 10^5 events, $|\cos(\theta)| < 0.999$
- $\sigma_{\text{eff}} \sim 0.3$ pb (in the LumiCal fiducial volume)
- $\delta(\sigma) \sim 1\%$

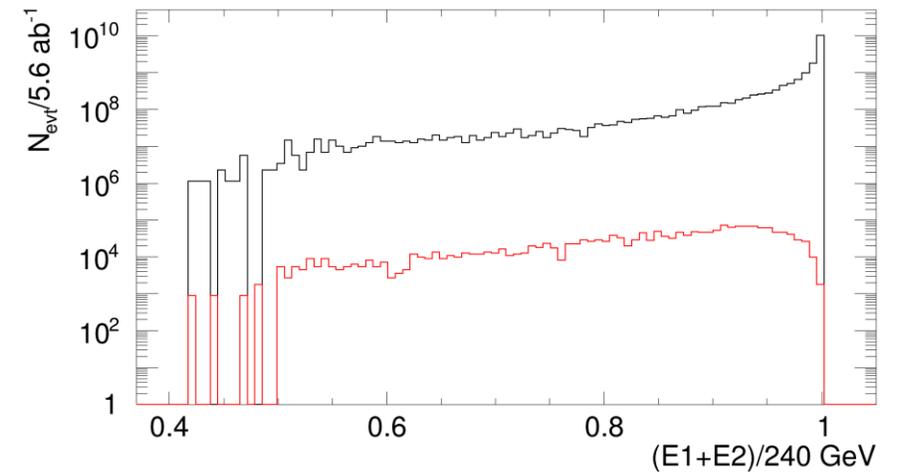
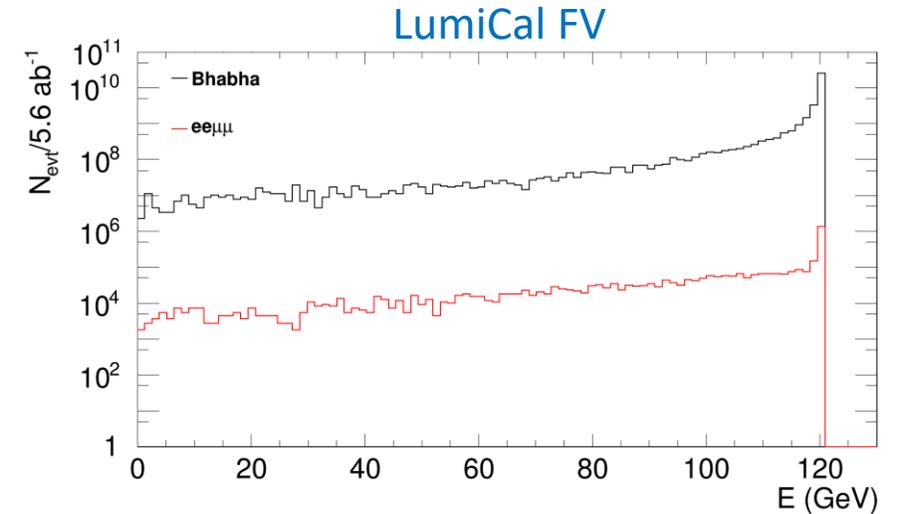
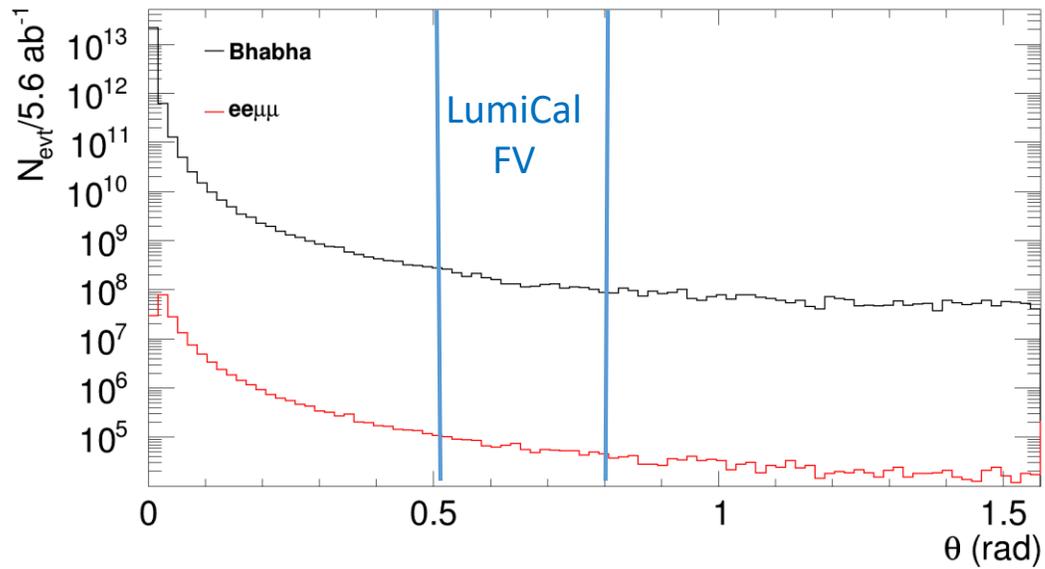
Normalization: $5.6 \text{ ab}^{-1}/7\text{y}$.

- $Bhabha$ BHLUMI V4.04
- 10^7 events, $\theta > 3$ mrad
- $\sigma_{\text{eff}} \sim 3.3$ nb (in the LumiCal fiducial volume)
- $\delta(\sigma) \sim 1.7 \cdot 10^{-4}$

$B/S \sim 10^{-4}$

Geometry:

- Geometrical coverage: $r_{\text{in}} = 25$ mm; $r_{\text{out}} = 100$ mm, (26 - 105) mrad
- Fiducial volume: $r_{\text{in},f} = 50$ mm; $r_{\text{out},f} = 75$ mm, that translates into θ_{FV} : (53-79) mrad
- $d_{\text{IP}} = 950$ mm



Main features:

- Most spectator electrons goes below the LumiCal
- Initial contamination (without any selection) of the detector volume is $\sim 10^{-4}$ w.r.t. the signal
- B/S ~ 10 times smaller than at 500 GeV ILC. This is mostly due to the Bhabha x-section dependence as $1/s$, while $2\text{-}\gamma$ x-section is scaling like $\ln^2(s)$

5.6 ab ⁻¹ /240 GeV	4-f	Bhabha	B/S
LumiCal FV	$1.8 \cdot 10^6$	$1.8 \cdot 10^{10}$	$1 \cdot 10^{-4}$
LumiCal + E _{rel}	$1.3 \cdot 10^6$	$1.7 \cdot 10^{10}$	$7.6 \cdot 10^{-5}$

- The total amount of background should be scaled by a factor ≤ 3 with flavor integration amounting to $B/S \leq 3 \cdot 10^{-4}$ without any selection
- Energy cut on relative energy $E_{rel} = (E_+ + E_-) > 0.8 \cdot s$ rejects $\sim 30\%$ of background, but is also important in a treatment of radiative Bhabha events and off-momentum background
- Refinements are possible with the coplanarity request between left and right detector arms ($|\phi_+ - \phi_-|$), also useful to suppress off-momentum particles
- Finally, physics background can be taken as a correction to the count (\mathcal{L} systematics comes from the x-section uncertainty)

- First estimates of the contribution of physics background to the luminosity systematics has been done
- Physics background is estimated to be present w.r.t the signal at the level of $\leq 3 \cdot 10^{-4}$ in the luminometer fiducial volume (what is more favorable than at higher cm energies and/or closer luminometer to the IP)
- Other refinements are possible in terms of:
 - Detector simulation,
 - Simulation of off-momentum background
 - Application of the asymmetric acceptance in θ (needed to suppress other sources of L-R symmetric systematics),
 - Introduction of the coplanarity requirement in the selection,
what should all improve B/S ratio further
- The ultimate uncertainty of \mathcal{L} from physics background will come from the uncertainty of the cross-section of 4-f processes. For that, some theoretical effort is needed.

Even taken as a full size effect, conclusion for CEPC is optimistic

BACKUP

A long list of sources of integral luminosity systematic uncertainties:

1. Beam related:

- Uncertainty of the average net CM energy
- Uncertainty of the asymmetry in energy of the e^+ and e^- beam
- Uncertainty of the beam energy spread
- IP position displacement and fluctuations w.r.t. the LumiCal, finite beam sizes at the IP
- Uncertainty of the (eventual) beam polarization

2. Detector related:

- Uncertainty of the LumiCal inner radius
- Positioning of the LumiCal (longitudinal L-R distance)
- Mechanical fluctuations of the LumiCal position w.r.t the IP (vibrations, thermal stress)
- Tilt and twist of the calorimeters
- Uncertainty of the sampling term
- Detector performance: energy and polar angle resolution

3. Physics interactions:

- Bhabha and physics background cross-section (uncertainty of the count)
- Bhabha acolinearity – other sources of the acceptance losses (ISR and FSR, Beamstrahlung)
- Machine-related backgrounds (off-momentum electrons from the beam-gas scattering)

Uncertainty of count is based on:

- Modification of the acceptance region (either directly or through the loss of colinearity of Bhabha events via longitudinal boost)
- Effect on the Bhabha cross-section calculation (modification of the phase space and E_{CM})
- Sensitivity of selection based observables (reconstructed energy, polar and azimuthal angles)

- Instrumentation of the very forward region is very important for the realization of the CepC physics program. Luminosity measurement uncertainty can affect:
 - Precision of the cross-section measurements
 - Anomalous TGCs measurement
 - Single-photon production with E_{mis} (BSM, dark matter)
 - Di-photon production (various BSM models)
 - Extended theories (Z') at high energies
 - Precision EW observables at Z^0 pole
- In most cases 10^{-3} precision of luminosity should be sufficient
- In particular, 10^{-4} uncertainty of integral luminosity comes from:
 - Fermion-pair production cross-section - access to the higher order corrections
 - W-pair production cross-section
 - Z^0 total hadronic cross-section at Z^0 pole
- This a 'common knowledge', 10^{-4} sensitivity should be proven through the dedicated physics analyses

CEPC Parameters

	<i>Higgs</i>	<i>W</i>	<i>Z</i>
Number of IPs	2		
Energy (GeV)	120	80	45.5
Circumference (km)	100		
SR loss/turn (GeV)	1.68	0.33	0.035
Half crossing angle (mrad)	16.5		
Piwinski angle	2.96	4.74	11.7
N_e /bunch (10^{10})	12.9	3.6	1.6
Bunch number	304	5230	11720
Beam current (mA)	18.8	90.5	90.1
SR power /beam (MW)	31.7	30	3.1
Bending radius (km)	10.9		
Momentum compaction (10^{-5})	1.14		
β_{IP} x/y (m)	0.36/0.002		
Emittance x/y (nm)	1.21/0.0036	0.54/0.0018	0.17/0.0029
Transverse σ_{IP} (um)	20.9/0.086	13.9/0.060	7.91/0.076
ξ_x/ξ_y /IP	0.021/0.088	0.008/0.051	0.0034/0.023
RF Phase (degree)	128	134.4	138.6
V_{RF} (GV)	2.14	0.465	0.053
f_{RF} (MHZ) (harmonic)	650		
Nature bunch length σ_z (mm)	2.72	2.98	3.67
Bunch length σ_z (mm)	3.75	4.0	5.6
HOM power/cavity (kw)	0.47 (2cell)	0.31 (2cell)	0.08 (2cell)
Energy spread (%)	0.098	0.066	0.037
Energy acceptance requirement (%)	1.12		
Energy acceptance by RF (%)	2.06	1.48	0.75
Photon number due to beamstrahlung	0.25	0.11	0.08
Lifetime due to beamstrahlung (hour)	1.0		
F (hour glass)	0.93	0.96	0.986
L_{max} /IP ($10^{34}\text{cm}^{-2}\text{s}^{-1}$)	2.0	3.9	1.0