



## SUSY Search at LHC and beyond

Xuai Zhuang (庄胥爱)

xuai.zhuang@cern.ch

IHEP B305, Beijing, China

22 Nov. 2017

# Outline

- **SUSY Introduction**
- The LHC and ATLAS
- SUSY search strategy
- □ SUSY search @ LHC
- Longer term prospects @
  LHC and Future Colliders (14, 33, 100 TeV)
- Outlook and Summary







# Outline





## **SM and Beyond**





Photo: A. Mahmoud François Englert



Photo: A. Mahmoud Peter W. Higgs



The Nobel Prize in Physics 2013 François Englert, Peter Higgs

- Higgs boson observed, SM fits the experimental data very well → big success in EW scale
- While has problem in Planck scale:
  - Naturalness and "hierarchy" problem
  - Unification of gauge coupling
  - Dark Matter

Need a more fundamental theory in which SM is only a low-energy approximation → New Physics

### 在超出标准模型的新物理模型中, SUSY理论是唯一能够回答绝大多数Big Question的理论



#### SUSY search is one of the most hot topic at LHC and beyond



#### Establishes a symmetry between fermions (mater) and bosons (forces)

- Each particle has a super-partner
- Number of elementary particles doubled
- Spin differs by <sup>1</sup>/<sub>2</sub> between SUSY and SM particles
- Identical gauge numbers and couplings
- A more fundamental theory: compatible with SM in EW scale, solve most problems in Planck scale

# **SUSY Introduction (2)**



- Fermion and boson loops contribute with different signs to the Higgs radiative corrections
- Supersymmetric partner contributions to Higgs mass cancel SM contributions





## **SUSY Introduction (3)**



### Unification of gauge couplings

- New particle content changes running of couplings
- Requires SUSY masses at few TeV



Astronomers found that most of the matter in the Universe must be invisible Dark Matter



#### **'Supersymmetric' particles ?**



**J SUSY can provide perfect dark mater candidate – WIMP** (lightest neutralino in RPC models)

- Stable
- Electrically neutron
- Same density as DM

 $0.094 < \Omega_{CDM}h^2 < 0.136$  (95% CL)

# Outline







### Large Hadron Collider—LHC @CERN, Geneva



The Large Hadron Collider is a 27 km long collider ring housed in a tunnel about 100 m underground near Geneva <sup>11</sup>

### **CERN's particle accelerator chain**



### **Collisions at LHC**



### ATLAS and CMS detector @ LHC ATLAS and CMS: two multi-purpose detectors @LHC

### A Toroidal LHC ApparatuS

- 42m×22m, 7000 ton
- Solenoid + Toroidal magnet (2T)
- Fine granularity liquid Ar/Tile calorimeters

#### Compact Muon Spectrometer

ATLAS

- 21m×15m, 125000 ton
- All silicon trackers, 4T

solenoid magnet

- PbWO4+Tile calorimeters





- I Tracking (|η|<2.5, B=2T) :
  - Si pixels and strips
  - Transition Radiation Detector (e/π separation)
- Calorimetry (|η|<5) :
  - EM : Pb-LAr
  - HAD: Fe/scintillator (central), Cu/W-LAr (fwd)
- Muon Spectrometer (|η|<2.7) :
  - air-core toroids with muon chambers



## **MET: Missing Transverse Energy**

- At the LHC an unknown proportion of the energy of the colliding protons escapes down the beam-pipe
- Invisible particles (neutrinos, neutralinos?) are created their momentum can be constrained in the plane transverse to the beam direction

 $E_T^{\text{miss}} = -\sum p_T(i)$ 







# Outline







### Where do we start?

Huge parameter space, but guiding principles a



Early analyses
 dominated by broad and
 inclusive searches for
 gluino and squark
 production

Increasingluminositygaveaccesstoproductionchannels.AdditionalmotivationfromNaturalSUSYparadigm

 If 1<sup>st</sup> and 2<sup>nd</sup> squark and gluino is too heavy, EWK SUSY production may dominant in LHC SUSY searches strategy driven by cross section and luminosity -



## How do we start? - SUSY Signature

#### **Conserved R parity (RPC):** (originally introduced for stability of proton) R=+1 (SM)

$$R = (-1)^{3(B-L)+2S}$$

- R=-1 (SUSY) SUSY particles produced/annihilated in pairs
- Lightest SUSY particle (LSP) stable (DM candidate)
- Typical signature: jets/leptons/photons + MET (key signature: large MET)

### Violated R parity (RPV): no Dark Matter candidate



р

### **SUSY Signature & Search Strategy**

SUSY search strategy: search for <u>deviation from SM</u>

SUSY sensitive variables: Try to establish excess of events in some sensitive kinematic distribution (E<sub>T</sub><sup>miss</sup>, Meff, mT ...)

SM background: SUSY searches rely on accurate modeling of the Standard Model backgrounds





## **SUSY Sensitive Variables**



- **E**<sub>T</sub><sup>miss</sup> from escaping LSP, to suppress bg from mismeasured jets and oth. SM BG
- Related to the sparticle mass scale, like effective mass (**M**<sub>eff</sub>)

$$M_{\text{eff}} \equiv \sum_{i=1}^{N_{\text{jets}}} p_{\text{T}}^{\text{jet},i} + \sum_{j=1}^{N_{\text{lep}}} p_{\text{T}}^{\text{lep},j} + E_{\text{T}}^{\text{miss}}$$

**mT, mT2** (stransverse mass): suppress BG with Ws

$$m_{\mathrm{T2}} = \min_{\mathbf{q}_{\mathrm{T}}} \left[ \max \left( m_{\mathrm{T}}(\mathbf{p}_{\mathrm{T}}^{\ell 1}, \mathbf{q}_{\mathrm{T}}), m_{\mathrm{T}}(\mathbf{p}_{\mathrm{T}}^{\ell 2}, \mathbf{p}_{\mathrm{T}}^{\mathrm{miss}} - \mathbf{q}_{\mathrm{T}}) \right) \right]$$

Many others ...

#### Standard Model

Top, multijets V, VV, VVV, Higgs & combinations of these

Combined fit of all regions and bgs, and including systematic exp. and theory uncertainty as nuisance parameters

#### Reducible backgrounds

Determined from data Backgrounds and methods depend on analyses

#### Irreducible backgrounds

Dominant sources:

- Mulit-jet: data-driven
- Non-Multi-jet: normalise

MC in data control regions

□ Subdominant sources: MC

#### Validation

Validation regions used to cross check SM predictions with data

#### Signal regions

blinded

blinded









# Outline





# **SUSY Search @ LHC**

ppLPCC SUSY of WG 10 data SUSY) [pb] 8TeV LHC  $10^{-1}$ NLO(-NLL) o(pp-20 fb ĝĝ 10-2  $0^2$   $10^2$  #events in  $10^2$  $10^{2}$ 10<sup>-3</sup>⊧  $\widetilde{\chi}^{\star}\widetilde{\chi}$  $10^{-10}$ 200 400 600 800 1000 1200 1400 1600 SUSY sparticle mass [GeV]  $\tilde{\chi}_1^{\pm}$ parXiv:1206.2892  $ilde{\chi}^{0}_{2}$ 

#### Strong production:

targeting gluinos and 1<sup>st</sup> and 2<sup>nd</sup> generation squarks

ATLAS public link

CMS public link

by far largest cross-sections

### 3<sup>rd</sup> generation:

- □ targeting stop and sbottoms
- Should be lowest mass squarks for naturalness reasons

### **Electroweak production**:

- Largeting Electroweakinos, sleptons
- Lowest mass sparticles, clean signature

### RPV/LL:

- targeting R-parity violating models and long lived sparticles
- More exotic models

## **SUSY Search @ LHC**



#### Strong production:

targeting gluinos and 1<sup>st</sup> and 2<sup>nd</sup> generation squarks

#### by far largest cross-sections

3<sup>rd</sup> generation:

- targeting stop and sbottoms
- Should be lowest mass squarks for naturalness reasons

#### **Electroweak production:**

- Largeting Electroweakinos, sleptons
- Lowest mass sparticles, clean signature

#### RPV/LL:

- targeting R-parity violating models and long lived sparticles
- More exotic models

### **Electroweakino Production**



### **Strong Production: All hadronic**

- □ ATLAS has two searches based on different scenarios: <u>ATLAS-CONF-</u> <u>2017-022</u>(SS,GG), <u>arXiv:1711.01901</u> (Gtt, Gbb), <u>arXiv:1708.02794</u> (GG, 7-11jets)
- □ Signal/BG discrimination based on:
  - Large/medium mass split: Meff, MET,
  - Compressed region: R Jigsaw
- Huge #signal regions (SRs) defined targeting above different search scenarios and phase space

- CMS has several searches based on different kinematics:
- $H_T^{miss}$  (PhysRevD.96.032003):
  - ✓ binned in jet and b-jet multiplicity.
  - ✓ In each bin, bin further in  $H_T$  and  $H_T^{miss}$
- mT<sub>2</sub> search (Eur. Phys. J C 77 (2017) 710):
  - ✓ binned in  $H_T$ , jet and b-jet multiplicity.
  - ✓ In each bin, look at tails of mT2





### All hadronic: ATLAS | CMS results





### All hadronic: ATLAS | CMS results





### All hadronic: ATLAS | CMS results








### **Strong Production: leptonic signatures**

- We can use lepton(s) to probe for strong production signals
  - $\rightarrow$  suppressed hadronic backgrounds
  - $\rightarrow$  single lepton, OS, SS di-leptons, multi-leptons (+ jets+ MET)



### **Strong Production: leptonic signatures**

- We can use lepton(s) to probe for strong production signals
  - $\rightarrow$  suppressed hadronic backgrounds
  - $\rightarrow$  single lepton, OS, SS di-leptons, multi-leptons (+ jets+ MET)



### **Strong Production: photonic signatures**

- We can use photon(s) to probe for strong production signals
  - $\rightarrow$  suppressed hadronic backgrounds



39

# **Strong Production (summary)**





# **Prospects at HL-LHC**



### **For exclusions:**

- Gluinos: ~ 2 TeV now
- 2.5 TeV @ 300 fb-1
- 3 TeV @ 3000 fb-1

# **Places to hunt for SUSY?**





For discovery potential:

- O We explored 85%
   of our mass reach
   (2 → 2.5 TeV)
- with a small
   window remaining
   between 1-2 TeV

300/fb 5σ
 discovery case is
 practically
 excluded

# **SUSY Search @ LHC**



models and long lived sparticles

A More exotic models

### **3rd Generation: stop**

□ Search for stop directly from ~t~t production

□ Large spectrum of possible stop decays, covering range from low to heavy stop mass, various decay modes.



### 3<sup>rd</sup> Generation: stop (fully hadronic)



45

### 3<sup>rd</sup> Generation: stop (leptonic)



# 3<sup>rd</sup> Generation (summary)



□ For bottom squarks: exclusion limits beyond 1 TeV (CMS-PAS-SUS-16-032)
 □ Still <600 GeV for compressed region, also for stop→charm+MET (ATLAS-CONF-2017-038)</li>

48

## **Prospects at HL-LHC**



### For exclusions (all hadronic):

- stop: ~ 1 TeV now
- 1.2 TeV @ 300 fb-1
- o 1.4 TeV @ 3000 fb-1

#### For discovery:

- 1.0 TeV @ 300 fb-1
- o 1.2 TeV @ 3000 fb-1

#### ATL-PHYS-PUB-2013-011



# **SUSY Search @ LHC**



### **Electroweakino Production**



### **Chargino-Neutralino Pairs: via sleptons**



### **Chargino-Neutralino Pairs: via WZ or WH**



52

### **Chargino-Chargino Pairs: via sleptons**



### **Direct slepton pair**



# **EWK Production (summary)**



Powerful exclusions in decays via sleptons (C1/N2 up to 0.6-1.1 TeV)

- **Exclusions is not so large in decays via <b>bosons** (up to 150-400 GeV)
- □ Mass limit on selectron/smuon up to 500 GeV, not yet on staus

#### ATL-PHYS-PUB-2014-010



# **Prospects at HL-LHC**



#### For exclusions:

- chargino: ~ 0.45 TeV now
- 0.8 TeV @ 300 fb-1
- o 1.1 TeV @ 3000 fb-1

#### For discovery:

- 0.55 TeV @ 300 fb-1
- 0.8 TeV @ 3000 fb-1

Some room for discovery



# **Direct** stau



#### For exclusions:

- stau: ~ 0.1 TeV now
- 0.7 TeV @ 3000 fb-1

#### For discovery:

0.5 TeV @ 3000 fb-1

### **Very promising at LHC**

# 2015+2016 – A milestone for SUSY



#### **This means:**

- We explored 85% of our mass reach for gluino pair production, about 75% for stop
- ~60% for gauginos, and just above 50% for higgsinos

# Outline



Outlook and Summary

# **Prospects at Future Proton Colliders**

Machine	$\sqrt{s}$	Final Integrated Luminosity 300 fb <sup>-1</sup> 3000 fb <sup>-1</sup>		
LHC Phase I	14 TeV			
HL-LHC or LHC Phase II	14 TeV			
HE-LHC	33 TeV	$3000 \text{ fb}^{-1}$		
VLHC	100 TeV	$3000 \text{ fb}^{-1}$		

arXiv:1311.6480, 1406.4512, 1410.6287

- Long term prospects for 4 collider scenarios have been studied (14, 33, 100 TeV @3000 fb<sup>-1</sup>)
- Use same search strategy as 8-13TeV @LHC
- Use simple analysis strategies, avoid assumption on detector design, pileup sensitivity, etc

## Gluino pair (Gqq) – light flavor



![](_page_60_Figure_2.jpeg)

61

### Gluino pair (Gtt) – heavy flavor

![](_page_61_Figure_1.jpeg)

![](_page_61_Figure_2.jpeg)

62

# Squark pair

![](_page_62_Figure_1.jpeg)

![](_page_62_Figure_2.jpeg)

# **Stop pair**

Search for stop with on-shell top decay @ 100 TeV
 Tops are very boosted at 100 TeV, the search strategy with isolated leptons, b-tags or top-tagging are sensitive to detector granularity/performance, pileup conditions etc. → So use simple handles with hard jets, MET and muon-in-jet
 Two scenarios with high mass stop and

![](_page_63_Figure_2.jpeg)

Discover (Exclude) 6.5 (8) TeV stop @100TeV

compressed region designed

![](_page_63_Figure_4.jpeg)

# **EWK Signature**

![](_page_64_Figure_1.jpeg)

- Potential search for electroweakinos with multi-lepton signatures: 3L || OS2L || SS2L @100 TeV
- Discover (Exclude) 2.1 (3.2) TeV electroweakinos for wino NLSP

![](_page_64_Figure_4.jpeg)

![](_page_64_Figure_5.jpeg)

#### ATL-PHYS-PUB-2013-011

# Outline

SUSY Introduct **LHC and Futur** (14, 33, 100 Te

![](_page_65_Picture_2.jpeg)

Outlook and Summary

# **Summary and Outlook**

- □ [LHC 13 TeV ~36 fb<sup>-1</sup>] In canonical scenarios, sensitivity is achieved to ~2 TeV gluinos, ~1 TeV stops and ~400-1000 GeV EWKinos
- □ [LHC 14 TeV ~3000 fb<sup>-1</sup>] Discovery potential up to 2.2 TeV gluinos, 1.2 TeV stop and 800 GeV EWKinos
- □ [100 TeV ~3000 fb<sup>-1</sup>] Discover (Exclude) 11 (13) TeV gluino, 6.5 (8) TeV stop, 2.1 (3.2) TeV EWKinos
- □ LHC is a discovery machine, new physics may come at any time , stay tuned!

# **Thanks for your attention!**

<b>A</b> N	TLAS SUSY Sea	rches*	- 95%	6 C	L Lov	ver Limits			<b>ATLAS</b> Preliminary $\sqrt{s} = 7, 8, 13$ TeV
	Model	$e, \mu, \tau, \gamma$	⁄ Jets	E <sup>miss</sup> T	∫ <i>L dt</i> [fb	<sup>-1</sup> ] Mass limit	$\sqrt{s} = 7$	<b>TeV</b> $\sqrt{s} = 13$ TeV	Reference
nclusive Searches	$ \begin{array}{l} MSUGRA/CMSSM \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{q}q, \tilde{q} \rightarrow \tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{1} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{g} \rightarrow qq\ell(\ell/\nu\nu)\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{s}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \\ GMSB(\ell  NLSP) \\ GGM(bino  NLSP) \\ GGM(higgsino-bino  NLSP) \end{array} $	$\begin{array}{c} 0-3 \ e, \mu/1-2 \ \tau \\ 0 \\ mono-jet \\ 0 \\ 3 \ e, \mu \\ 0 \\ 1-2 \ \tau + 0-1 \\ 2 \ \gamma \\ \gamma \end{array}$	2-10 jets/3 2-6 jets 1-3 jets 2-6 jets 2-6 jets 4 jets 7-11 jets ℓ 0-2 jets - 1 b	b Yes Yes Yes Yes Yes Yes Yes Yes	20.3 36.1 3.2 36.1 36.1 36.1 36.1 3.2 3.2 20.3		1.85 TeV 1.57 TeV 2.02 TeV 2.01 TeV 1.825 TeV 1.8 TeV 2.0 TeV 1.65 TeV 3.7 TeV	$\begin{split} &m(\tilde{q})\!\!=\!m(\tilde{g}) \\ &m(\tilde{x}_{1}^{0})\!<\!200~\text{GeV},~m(1^{\text{tr}}~\text{gen},\tilde{q})\!=\!m(2^{\text{nd}}~\text{gen},\tilde{q}) \\ &m(\tilde{q})\!-\!m(\tilde{\chi}_{1}^{0})\!<\!5~\text{GeV} \\ &m(\tilde{\chi}_{1}^{0})\!<\!200~\text{GeV},~m(\tilde{\chi}^{+})\!=\!0.5(m(\tilde{\chi}_{1}^{0})\!+\!m(\tilde{g})) \\ &m(\tilde{\chi}_{1}^{0})\!<\!400~\text{GeV} \\ &m(\tilde{\chi}_{1}^{0})<\!400~\text{GeV} \\ &cr(NLSP)\!<\!0.1~\text{mm} \\ &m(\tilde{\chi}_{1}^{0})\!<\!\!950~\text{GeV},~cr(NLSP)\!<\!0.1~\text{mm},~\mu\!<\!0 \end{split}$	1507.05525 ATLAS-CONF-2017-022 1604.07773 ATLAS-CONF-2017-022 ATLAS-CONF-2017-022 ATLAS-CONF-2017-030 ATLAS-CONF-2017-033 1607.05979 1606.09150 1507.05493
4	GGM (higgsino-bino NLSP) GGM (higgsino NLSP) Gravitino LSP	γ 2 <i>e</i> ,μ (Z) 0	2 jets 2 jets mono-jet	Yes Yes Yes	13.3 20.3 20.3	š         900 GeV           F <sup>1/2</sup> scale         865 GeV	1.8 TeV	$\begin{split} & m(\tilde{\chi}_{j}^{U}){>}680~GeV,~c\tau(NLSP){<}0.1~mm,~\mu{>}0\\ & m(NLSP){>}430~GeV\\ & m(\tilde{G}){>}1.8\times10^{-4}~eV,~m(\tilde{g}){=}m(\tilde{q}){=}1.5~TeV \end{split}$	ATLAS-CONF-2016-066 1503.03290 1502.01518
3 <sup>rd</sup> gen ẽ med.	$\begin{array}{l} \tilde{g}\tilde{g}, \ \tilde{g} \rightarrow b\bar{b}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \ \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \ \tilde{g} \rightarrow b\bar{t}\tilde{\chi}_{1}^{+} \end{array}$	0 0-1 <i>e</i> ,μ 0-1 <i>e</i> ,μ	3 b 3 b 3 b	Yes Yes Yes	36.1 36.1 20.1	iš	1.92 TeV 1.97 TeV .37 TeV	$m(\tilde{\chi}_1^0) < 600 \text{ GeV}$ $m(\tilde{\chi}_1^0) < 200 \text{ GeV}$ $m(\tilde{\chi}_1^0) < 300 \text{ GeV}$	ATLAS-CONF-2017-021 ATLAS-CONF-2017-021 1407.0600
3 <sup>rd</sup> gen. squarks direct production	$ \begin{split} \tilde{b}_{1} \tilde{b}_{1} \cdot \tilde{b}_{1} \to b \tilde{x}_{1}^{0} \\ \tilde{b}_{1} \tilde{b}_{1} \cdot \tilde{b}_{1} \to t \tilde{x}_{1}^{1} \\ \tilde{r}_{1} \tilde{t}_{1} \cdot \tilde{t}_{1} \to t \tilde{x}_{1}^{1} \\ \tilde{r}_{1} \tilde{t}_{1} \cdot \tilde{t}_{1} \to b \tilde{x}_{1}^{0} \\ \tilde{r}_{1} \tilde{t}_{1} \cdot \tilde{t}_{1} \to \tilde{x}_{1}^{0} \\ \tilde{r}_{1} \tilde{t}_{1} \cdot \tilde{t}_{1} \to \tilde{x}_{1}^{0} \\ \tilde{r}_{1} \tilde{t}_{1} (natural GMSB) \\ \tilde{t}_{2} \tilde{t}_{2} \cdot \tilde{t}_{2} \to \tilde{t}_{1} + Z \\ \tilde{t}_{2} \tilde{t}_{2} \cdot \tilde{t}_{2} \to \tilde{t}_{1} + h \end{split} $	0 2 $e, \mu$ (SS) 0-2 $e, \mu$ 0-2 $e, \mu$ 0 2 $e, \mu$ (Z) 1-2 $e, \mu$	2 b 1 b 1-2 b 0-2 jets/1-2 mono-jet 1 b 1 b 4 b	Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 4.7/13.3 20.3/36.1 3.2 20.3 36.1 36.1	b1         950 GeV           b1         275-700 GeV           i1         117-170 GeV         200-720 GeV           i1         90-198 GeV         205-950 GeV           i1         90-323 GeV         205-950 GeV           i1         90-323 GeV         200-720 GeV           i2         200-790 GeV         200-790 GeV           i2         320-880 GeV         320-880 GeV		$\begin{split} &m(\tilde{x}_{1}^{0}){<}420GeV \\ &m(\tilde{x}_{1}^{0}){<}200GeV, m(\tilde{x}_{1}^{+}){=}m(\tilde{x}_{1}^{0}){+}100GeV \\ &m(\tilde{x}_{1}^{-}){=}2m(\tilde{x}_{1}^{0}),m(\tilde{x}_{1}^{0}){=}55GeV \\ &m(\tilde{x}_{1}^{-}){=}1GeV \\ &m(\tilde{x}_{1}^{-}){=}15GeV \\ &m(\tilde{x}_{1}^{0}){=}15GeV \\ &m(\tilde{x}_{1}^{0}){=}0GeV \\ &m(\tilde{x}_{1}^{0}){=}0GeV \end{split}$	ATLAS-CONF-2017-038 ATLAS-CONF-2017-030 1209.2102, ATLAS-CONF-2016-077 1506.08616, ATLAS-CONF-2017-020 1604.07773 1403.5222 ATLAS-CONF-2017-019 ATLAS-CONF-2017-019
EW direct	$ \begin{array}{l} \tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0 \\ \tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\ell} \nu(\ell \tilde{\nu}) \\ \tilde{\chi}_1^+ \tilde{\chi}_1^- / \tilde{\chi}_2^0, \tilde{\chi}_1^+ \rightarrow \tilde{\nu}(\tau \tilde{\nu}), \tilde{\chi}_2^0 \rightarrow \tilde{\tau} \tau(\nu \tilde{\nu}) \\ \tilde{\chi}_1^+ \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_1 \nu \tilde{\ell}_1 \ell(\tilde{\nu}), \ell \tilde{\nu} \tilde{\ell}_1 \ell(\tilde{\nu} \nu) \\ \tilde{\chi}_1^+ \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \ell \tilde{\chi}_1^0 \\ \tilde{\chi}_1^+ \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \ell \tilde{\chi}_1^0, h \rightarrow b \tilde{b} / W W / \tau \tau / \gamma \gamma \\ \tilde{\chi}_2^+ \tilde{\chi}_2^0 \rightarrow \tilde{\chi}_2^0 \delta \rightarrow \tilde{\ell}_R \ell \\ \begin{array}{c} \text{GGM (wino NLSP) weak prod., } \tilde{\chi}_1^0 \rightarrow \\ \text{GGM (bino NLSP) weak prod., } \tilde{\chi}_1^0 \rightarrow \end{array} \end{array} $	$\begin{array}{c} 2 \ e, \mu \\ 2 \ e, \mu \\ 2 \ \tau \\ 3 \ e, \mu \\ 2 \ 3 \ e, \mu \\ e, \mu, \gamma \\ 4 \ e, \mu \\ \gamma \tilde{G} \ 1 \ e, \mu + \gamma \\ \gamma \tilde{G} \ 2 \ \gamma \end{array}$	0 0  0-2 jets 0-2 <i>b</i> 0 - -	Yes Yes Yes Yes Yes Yes Yes Yes	36.1 36.1 36.1 36.1 20.3 20.3 20.3 20.3	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	e <b>V</b> m(k¯1)= m(k <sub>2</sub> )=	$\begin{split} &m(\tilde{x}_{1}^{0}){=}0 \\ &m(\tilde{x}_{1}^{0}){=}0, \ m(\tilde{\ell},\tilde{\nu}){=}0.5(m(\tilde{k}_{1}^{+}){+}m(\tilde{k}_{1}^{0})) \\ &m(\tilde{k}_{1}^{0}){=}0, \ m(\tilde{\tau},\tilde{\nu}){=}0.5(m(\tilde{k}_{1}^{+}){+}m(\tilde{k}_{1}^{0})) \\ &m(\tilde{k}_{1}^{0}){-}m(\tilde{k}_{2}^{0}){-}m(\tilde{\kappa}_{1}^{0}){=}0, \ \tilde{\ell} \text{ decoupled} \\ &m(\tilde{k}_{1}^{0}){-}m(\tilde{k}_{2}^{0}), \ m(\tilde{k}_{1}^{0}){=}0, \ \tilde{\ell} \text{ decoupled} \\ &m(\tilde{k}_{1}^{0}){-}m(\tilde{k}_{2}^{0}){-}m(\tilde{\ell},\tilde{\nu}){=}0.5(m(\tilde{k}_{2}^{0}){+}m(\tilde{k}_{1}^{0})) \\ &c\tau{<}1 \ mm \\ \\ &cr{<}1 \ mm \end{split}$	ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 ATLAS-CONF-2017-035 ATLAS-CONF-2017-039 ATLAS-CONF-2017-039 1501.07110 1405.5086 1507.05493 1507.05493
Long-lived particles	Direct $\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}$ prod., long-lived $\tilde{X}_{1}^{\pm}$ Direct $\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}$ prod., long-lived $\tilde{X}_{1}^{\pm}$ Stable, stopped $\tilde{g}$ R-hadron Stable $\tilde{g}$ R-hadron Metastable $\tilde{g}$ R-hadron GMSB, stable $\tilde{\tau}, \tilde{X}_{1}^{0} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$ GMSB, $\tilde{X}_{1}^{0} \rightarrow \gamma \tilde{G}$ , long-lived $\tilde{X}_{1}^{0}$ $\tilde{g}\tilde{g}, \tilde{X}_{1}^{0} \rightarrow eev/e\muv/\mu\nuv$ GGM $\tilde{g}\tilde{g}, \tilde{X}_{1}^{0} \rightarrow Z\tilde{G}$	Disapp. trk dE/dx trk 0 trk dE/dx trk $1-2 \mu$ $2 \gamma$ displ. $ee/e\mu/\mu$ displ. vtx + je	x 1 jet - 1-5 jets - - - μμ - ets -	Yes Yes - - Yes - -	36.1 18.4 27.9 3.2 19.1 20.3 20.3 20.3	$\begin{array}{c c} \ddot{x}_{1}^{+} & 430  \text{GeV} \\ \hline \ddot{x}_{1}^{+} & 495  \text{GeV} \\ \hline \ddot{s} & 850  \text{GeV} \\ \hline \ddot{s} $	1.58 TeV 1.57 TeV	$\begin{split} &m(\tilde{k}_1^+) \cdot m(\tilde{k}_1^0) \sim 160 \; MeV, r(\tilde{k}_1^+) = 0.2 \; ns \\ &m(\tilde{k}_1^+) \cdot m(\tilde{k}_1^0) \sim 160 \; MeV, r(\tilde{k}_1^+) < 15 \; ns \\ &m(\tilde{k}_1^0) = 100 \; GeV, 10 \; \mu s < r(\tilde{g}) < 1000 \; s \\ &m(\tilde{k}_1^0) = 100 \; GeV, r > 10 \; ns \\ &10 < tan\beta < 50 \\ &1 < r(\tilde{k}_1^0) < 3 \; ns, SPS8 \; model \\ &7 < cr(\tilde{k}_1^0) < 740 \; nm, m(\tilde{g}) = 1.3 \; TeV \\ &6 < cr(\tilde{k}_1^0) < 480 \; nm, m(\tilde{g}) = 1.1 \; TeV \end{split}$	ATLAS-CONF-2017-017 1506.05332 1310.6584 1606.05129 1604.04520 1411.6795 1409.5542 1504.05162 1504.05162
RPV	$ \begin{array}{l} LFV pp \rightarrow \widetilde{v}_\tau + X, \widetilde{v}_\tau \rightarrow e\mu/e\tau/\mu\tau \\ Bilinear \; RPV \; CMSSM \\ \widetilde{\chi}_1^+ \widetilde{\chi}_1^-, \widetilde{\chi}_1^+ \rightarrow W\widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow eev, e\mu\nu, \mu\mu\nu \\ \widetilde{\chi}_1^+ \widetilde{\chi}_1^-, \widetilde{\chi}_1^+ \rightarrow W\widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow trv_e, e\tau\nu_\tau \\ \widetilde{g}\widetilde{s}, \widetilde{s} \rightarrow qqq \\ \widetilde{g}\widetilde{s}, \widetilde{g} \rightarrow q\widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow qqq \\ \widetilde{g}\widetilde{s}, \widetilde{g} \rightarrow d\widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow qqq \\ \widetilde{g}\widetilde{s}, \widetilde{g} \rightarrow d\widetilde{\chi}_1^0, \widetilde{\chi}_1 \rightarrow bs \\ \widetilde{t}_1 \widetilde{t}_1, \widetilde{t}_1 \rightarrow bs \\ \widetilde{t}_1 \widetilde{t}_1, \widetilde{t}_1 \rightarrow b\ell \end{array} $	$e\mu,e\tau,\mu\tau$ 2 e, $\mu$ (SS) 4 e, $\mu$ 3 e, $\mu$ + $\tau$ 0 4 1 e, $\mu$ 1 e, $\mu$ 0 4 2 e, $\mu$	- 0-3 b - - 4-5 large- <i>R</i> je 8-10 jets/0-4 8-10 jets/0-4 2 jets + 2 b 2 b	- Yes Yes ets - ets - + b - + b - - -	3.2 20.3 13.3 20.3 14.8 14.8 36.1 36.1 15.4 36.1	$\tilde{r}_{\tau}$ $\tilde{q}, \tilde{g}$ $\tilde{\chi}_{1}^{+}$ 1.14 Te $\tilde{\chi}_{1}^{+}$ 450 GeV $\tilde{g}$ 1.08 Te $\tilde{g}$ 1.08 Te $\tilde{g}$ $\tilde{g}$ $\tilde{l}_{1}$ 410 GeV 450-510 GeV $\tilde{l}_{1}$ 0.4	1.9 TeV 1.45 TeV 1.55 TeV 2.1 Te 1.65 TeV	$\begin{split} \lambda'_{311} = 0.11,  \lambda_{132/133/233} = 0.07 \\ m(\bar{q}) = m(\bar{g}),  c\tau_{LSP} < 1  \text{mm} \\ m(\bar{\chi}_1^0) > 400 \text{GeV},  \lambda_{12k} \neq 0  (k = 1, 2) \\ m(\bar{\chi}_1^0) > 0.2 \times m(\bar{\chi}_1^T),  \lambda_{133} \neq 0 \\ \text{B}((r) = \text{B}(k)) = \text{B}(c) = 0\% \\ m(\bar{\chi}_1^0) = 100  \text{GeV} \\ m(\bar{\chi}_1^0) = 1  \text{TeV},  \lambda_{112} \neq 0 \\ m(\bar{\chi}_1) = 1  \text{TeV},  \lambda_{233} \neq 0 \\ \text{B}R(\bar{t}_1 \rightarrow be/\mu) > 20\% \end{split}$	1607.08079 1404.2500 ATLAS-CONF-2016-075 1405.5086 ATLAS-CONF-2016-057 ATLAS-CONF-2016-057 ATLAS-CONF-2017-013 ATLAS-CONF-2017-013 ATLAS-CONF-2017-013 ATLAS-CONF-2016-022, ATLAS-CONF-2016-084 ATLAS-CONF-2017-036
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 <i>c</i>	Yes	20.3	č 510 GeV		m(𝐺̃ <sup>0</sup> <sub>1</sub> )<200 GeV	1501.01325
*Only	a selection of the available ma	ass limits on	new state	es or	1	$0^{-1}$ 1	7	Mass scale [TeV]	68

phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

Mass scale [lev]

#### Selected CMS SUSY Results\* - SMS Interpretation

ICHEP '16 - Moriond '17

![](_page_68_Figure_2.jpeg)

Only a selection of available mass limits. Probe \*up to\* the quoted mass limit for mg ≈0 GeV unless stated otherwise

69

![](_page_69_Picture_0.jpeg)

# **ATLAS IHEP Team Members**

#### **Total of 43 members:**

13 physicist staff8-10 Post-doc16 PhD students4 engineer staff

**Staffs**: Joao Guimaraes da Costa, Xuai Zhuang, Hongbo Zhu, Yaquan Fang, Xinchou Lou, Qun Ouyang, Feng Lu, Lianyou Shan, Zhijun Liang, Yanping Huang, Da Xu, Xin Shi

#### **General Management and Boards**

Xinchou LouATLAS ITK Steering Committee memberXuai ZhuangATLAS CB Chair Advisory Group member (2016 - 2018)

#### **Detector Operations and R&D**

JuanAn Pascual JuanAn Pascual Jason Mansour

Pixel DAQ coordinator (2016.5 – 2018.1) Deputy Run Coordinator (2017.11 – 2018. 10) Inner detector online and DAQ coordinator

#### **Physics and Performance**

Xuai ZhuangEWK SUSY Group Convener (2015.4 - 2016.4)Joao CostaHWW Group Convener (2016.4 - 2017.3)Yanping HuangPhoton ID group convener (2016. 10 - now)

# **Recent group activities**

### Detector Upgrade

• Phase II ITK strip upgrade

### Detector operation, Core software and Performance

- Pixel Detector DAQ and Operations, offline studies
- Inner detector data quality and tracking CP
- Photon ID
- Tau reconstruction and data quality
- Second vertex finder
- MET performance

### Physics Analysis

- Higgs:  $H \rightarrow \gamma \gamma$ ,  $Z\gamma$ , WW, bb, ttH(bb), di-Higgs in WW $\gamma \gamma$ , WWWW, bbbb
- New resonance search:  $\gamma\gamma$ ,  $Z\gamma$
- SUSY searches : di-tau, 1L, SS/2L&3L, tau+jets
- SM measurement: Zy VBS, Z+jet
# **IHEP SUSY Group**







Xuai ZHUANG

Shan JIN

Da XU

#### Feng LU



Mohamad Huan REN Huajie Cheng Peng Zhang Yang Liu Kassem AYOUB

Chenzheng Zhu <sup>73</sup>

# **SUSY Search Topics @ IHEP**



 $\begin{array}{c} 0 \\ \text{#events in 20} \end{array}$ 

# **SUSY models: good sale in market**

#### □ Simplified Models:

- Not really a model (Br~100%, most masses fixed at high scales)
- Important tool for interpretation

#### □ Phenomenological models:

- pMSSM: captures "most" of phenomenologic features of Rparity conserving MSSM
  - 19 free parameters: M1,M2,M3 ; tan β, μ and m<sub>A;</sub> 10 sfermion mass parameters; A<sub>t</sub>, A<sub>b</sub> and A<sub>τ</sub>
  - Comprehensive and computationally realistic approximation of the MSSM with neutralino LSP
- GGM (gravitino)

Complete SUSY models: mSUGRA, GMSB ...

# A change of perspective

- Develop more complex analysis (compressed, small cross section ...)
- □ Are there topologies of strong production (including 3<sup>rd</sup> gen.) not covering?
  - PRV signatures? Long-lived signatures? ...
- EW signatures have still more room to be discovered
  - Significant gain in sensitivity still expected for conventional C1N2 via WZ, WH decay
  - No proven sensitivity yet in run2 for a number of well-motivated signatures: direct stau, C1C1 via WW ... (<u>PhysRevD.93.052002</u>)
  - Higgsino LSP signatures
  - RPV signatures of EW production

```
o ...
```

Excellent LHC performance is a (nice) challenge for the experiment:

- Trigger

- Pile-up

 Maintain accuracy of the the measurements in this environment





Inner Detector for a Z  $\rightarrow \mu\mu$  event with 25 primary vertices

### 希腊神话中的怪物 "Uroboros" 与格拉肖的 "宇宙圈"



# An attempt to map out the SUSY model space with all the ATLAS analyses, giving an impression of where SUSY could still hide ...



JHEP 10 (2015) 134

#### LHC / HL-LHC Plan

2015: FTK deployment



New L0/L1 trigger scheme

Upgrade muon/calorimeter

Upgrade of DAQ detector readout

electronics



cope with higher rates

80

# Long term prospects

 ATLAS studied long term prospects for the (HL-)LHC with 300, 3000 fb<sup>-1</sup>@14TeV
 Discovery potential up to 2.5 TeV gluinos, 1.3 TeV squarks/sbottom and 800 GeV Electroweakinos, 500 GeV stau (IHEP)



### C1N2 via slepton: dependent with x



### The CLs significance as a function of x



- a) When only C1C1 production is considered, the benchmark scenario with <u>large mass-splitting (600,0)</u> can be excluded for **x up** to 0.75. For larger values of x, the pT spectra of the tau from the chargino decay become very soft.
- b) The <u>compressed benchmark scenario</u> (250,100) can only be excluded for the extreme cases with x = 0.05 or x = 0.95 since the mT2 requirement is more effective for models with large masssplittings between C1 or the staus and N1.

### The CLs significance as a function of x

#### arXiv:1708.07875

 $m(\tilde{\tau}_{\mathrm{L}}) = m(\tilde{\chi}_{1}^{0}) + x \cdot \left(m(\tilde{\chi}_{1}^{\pm}) - m(\tilde{\chi}_{1}^{0})\right)$ 



For <u>combined production of C1C1 and C1N2</u>, the same general features are observed, but due to the higher signal yields with respect to C1C1 production alone, both benchmark scenarios can be excluded for all considered values of x.





the complementary search using the **Recursive Jigsaw Reconstruction (RJR) techniques** in the construction of a discriminating variable set ('RJR-based search'). By using a dedicated set of selection criteria, the RJR-search improve the sensitivity to supersymmetric models with small mass splittings between the sparticles (models with compressed spectra).

#### **Recursive jigsaw reconstruction**

- based on assumption of decay tree
- fix set of rules to resolve combinatorics and unknowns in invisible system
- can form set of variables in the rest frame of each level in the decay tree



Inner Detector: Highly segmented silicon strips, determine very accurately charged particles trajectories

Solenoid Magnet: Solenoid coil that generates a 2T magnetic field in the region of the Inner Detector

Electromagnetic Calorimeter: Electron and photon energies are measured through electromagnetic showers

- Hadronic Calorimeter: Hadrons interact with dense material and produce a shower of charged particles
- Toroid Magnets: 8 toroidal coils that create a 0,4T magnetic field in the area of the Muon Spectrometer
- Muon Spectrometer: Muons traverse the rest of the detector and are measured in its outer layers

### Data-driven background estimation

**One approach** to data-driven bg **estimation** is to use uncorrelated model-independent variables to *extrapolate* the background from a background-dominated control region to the signal region.



```
Nbg in signal region D = (A/B)*C
```

```
Normalize Factor
Control Sample
```

"ABCD"

Method

#### Key points:

- <u>The two variables</u> should have good discrepancy and uncorrelated
- <u>Control Sample selection:</u> enough statistics;lower susy contamination; unbiased estimation of SM background
- <u>Normalization region selection</u>: enough statistics; lower susy contamination; flat ratio(A/B) distribution with ETmiss

# **Background Estimation Strategy**

#### ATLAS-CONF-2013-062



Other small BGs (diboson, single top etc) are directly estimated from MC.



Figure 12: Expected  $5\sigma$  discovery contours for the  $\sqrt{s} = 14$  TeV LHC [left] and a 100 TeV proton collider [right] with 3000 fb<sup>-1</sup>. The different curves correspond to various assumptions for the systematic uncertainty on the background: 5% [green], 10% [red], 20% [blue], and 30% [black].

- It is likely that the experiments will significantly reduce these uncertainties with larger datasets and an improved understanding of their detectors
- Varying the systematic background uncertainty from 30% to 5%, the discovery reach increases by roughly 600 GeV (3.4 TeV) in m(~g) at 14 TeV (100 TeV) and the coverage in LSP direction is roughly doubled

### **Impact of Pileup**



Figure 14: Discovery contours [right] and expected limits [left] for the analyses performed with [red, dotted] and without [black, solid] pileup at the 14 TeV LHC with 3000 fb<sup>-1</sup> integrated luminosity.

- Compared the results with 140 additional minimum-bias interactions
- The Delphes based Snowmass simulation includes a pileup suppression algorithm that primarily impacts the Emiss resolution (Snowmass detector:ArXiv:1309.1057)
- Given that the HT and ETmiss distributions are effectively unchanged, it is not surprising that the results are very similar with and without pileup

# Interpretation strategy

Based on the number of observed, expected events in all regions with all uncertainties: Probability density function (PDF)

Likelihood function: L(μ,θ) μ: signal strength (POI); θ: nuisance parameters(NP) Profile Likelihood: constrain uncertainty (NP) as part of a likelihood fit

Construct test statistics  $t_{\mu}$  based on likelihood ratio  $\lambda$ :







If CLs<0.05: the value

of signal is excluded at

95% CL.....

 $\mathrm{CL}_{s} = \frac{\mathrm{CL}_{s+b}}{\mathrm{CL}_{b}}$ 

