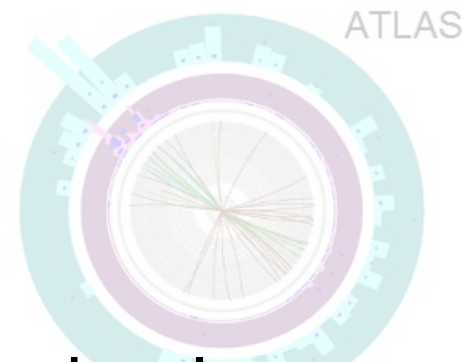
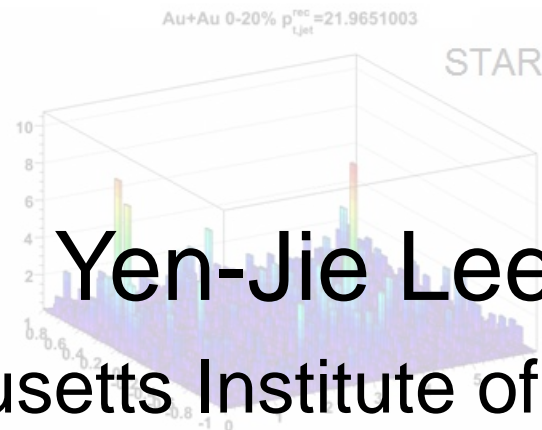


Jetting through the Quark Soup



Yen-Jie Lee

Massachusetts Institute of Technology

CMS
CMS Experiment at LHC, CERN
Data recorded: Sun Nov 14 19:31:39 2010 CEST
Run/Evnt: 131076 / 1328520
Lumi section: 249

CMS

The 7th Huada School on QCD
CCNU, Wuhan, China



Outline

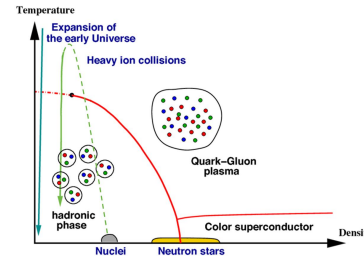
- Lecture 1
Why do we study relativistic heavy ion collisions?
- Lecture 2
How do we measure jets in heavy ion collisions?
- Lecture 3
Parton energy loss and its parton flavor dependence
- Lecture 4
Modification of jet substructure and medium response
- Lecture 5
Open questions and future direction

Lecture 1

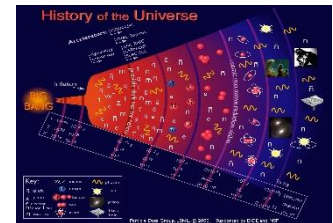
Why do we study
relativistic heavy ion collisions?

WHY?

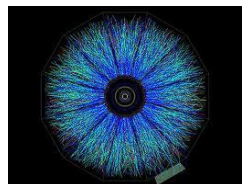
- Phase transition in QCD
 - “Melt the protons”



- Matter in the early universe: QCD in Cosmology
 - The super high temperature period which is not easily accessible by telescopes

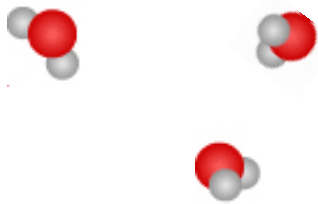


- To study the properties of the quantum matter
 - The earliest complex condense matter to form which is hard to compute without experimental guidance

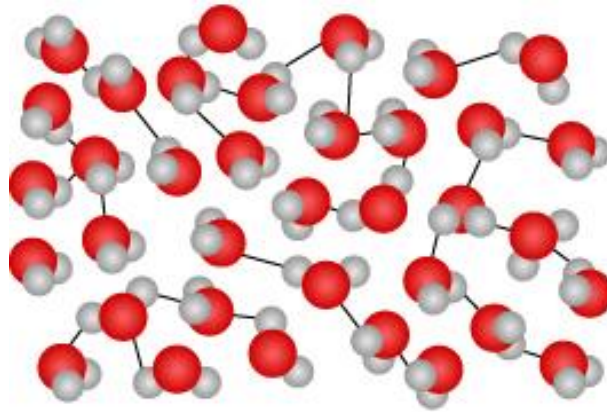


Phases of QED matter

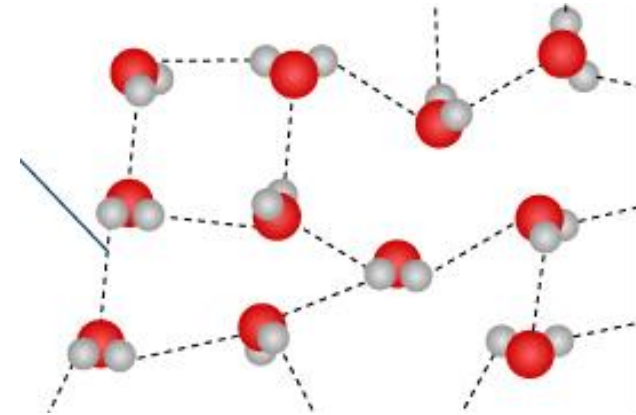
Vapor



Water

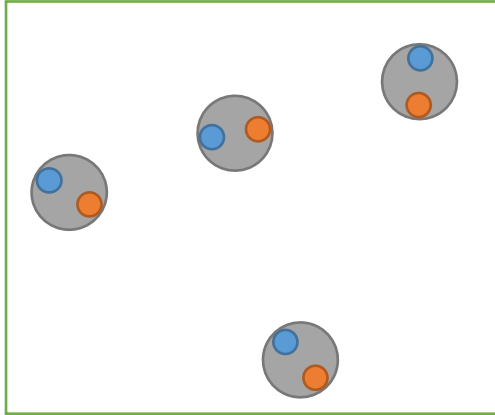


Ice

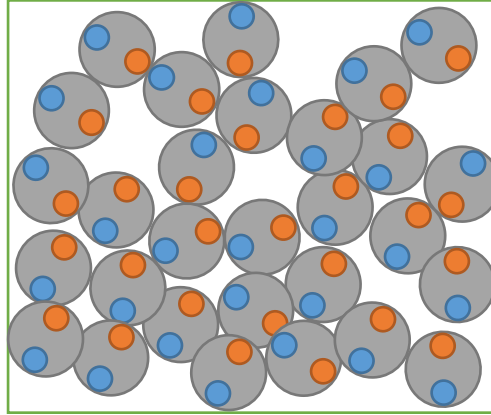


Ultra-dense QCD matter

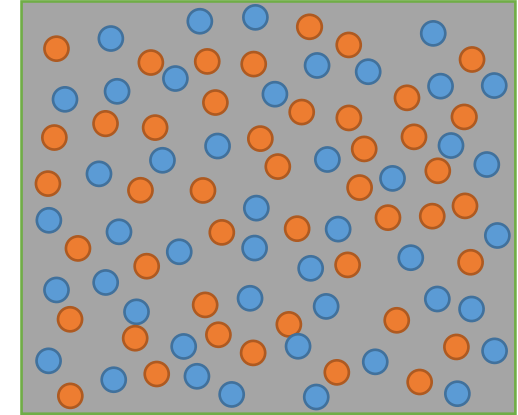
Increase the Temperature (T)



$$T < T_c$$



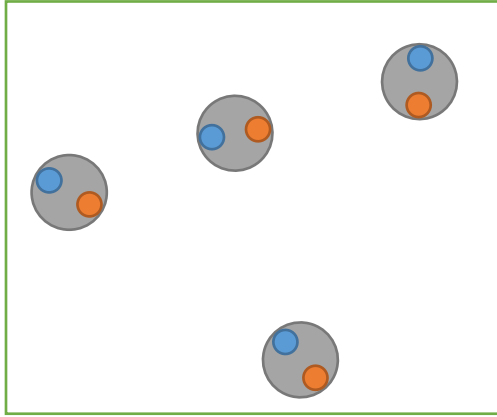
$$T \sim T_c$$



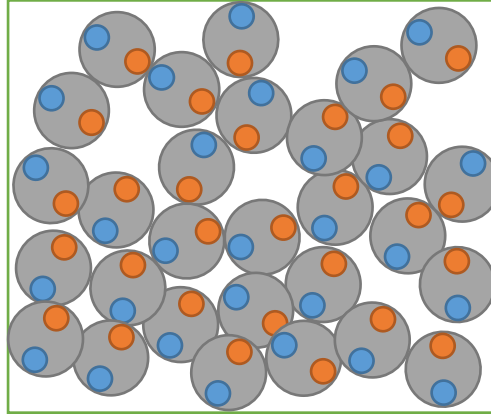
$$T > T_c$$

Ultra-dense QCD matter

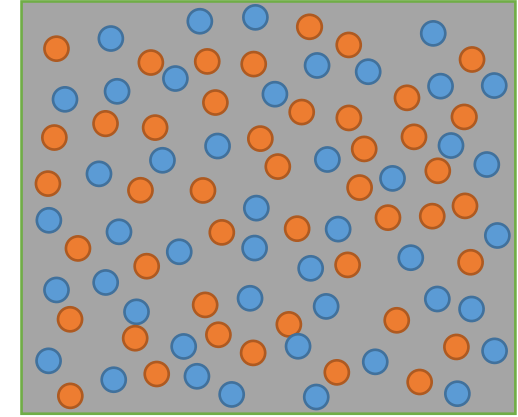
Increase the Temperature (T)



$$T < T_c$$

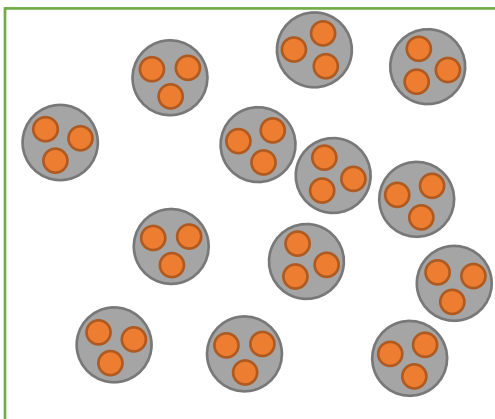


$$T \sim T_c$$

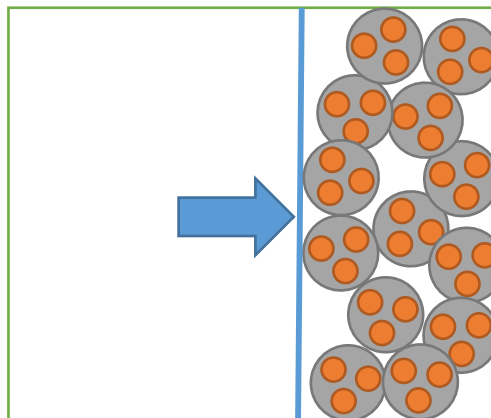


$$T > T_c$$

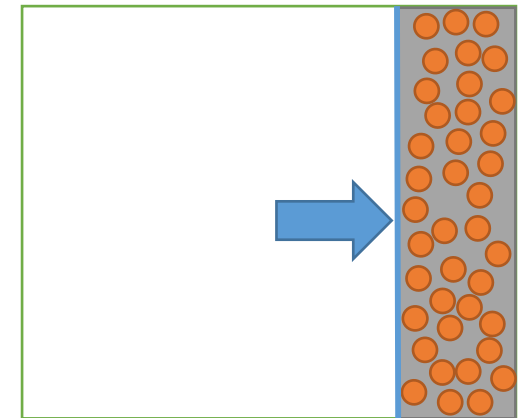
Increase the Density (ρ)



$$\rho < \rho_c$$

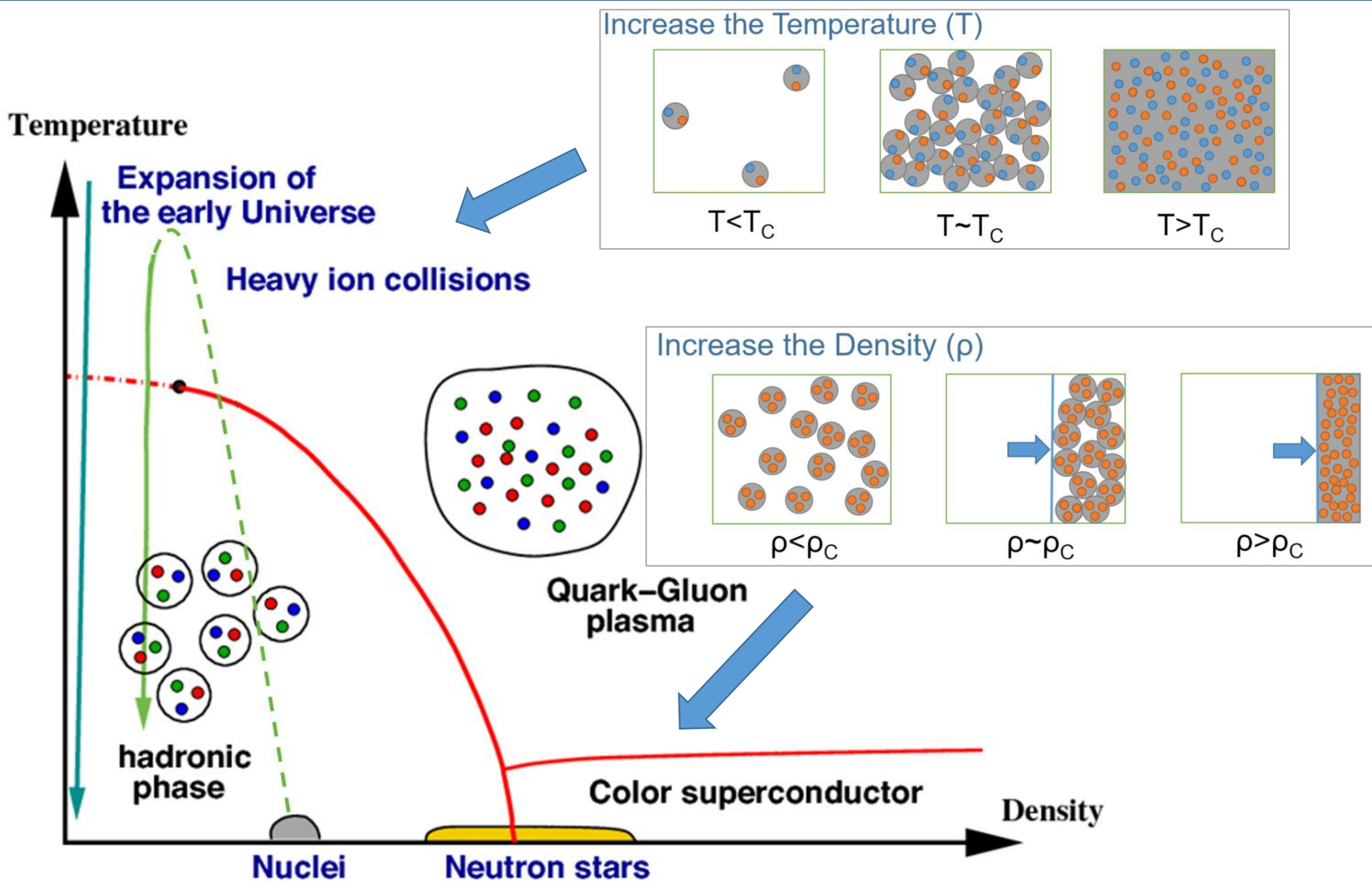


$$\rho \sim \rho_c$$

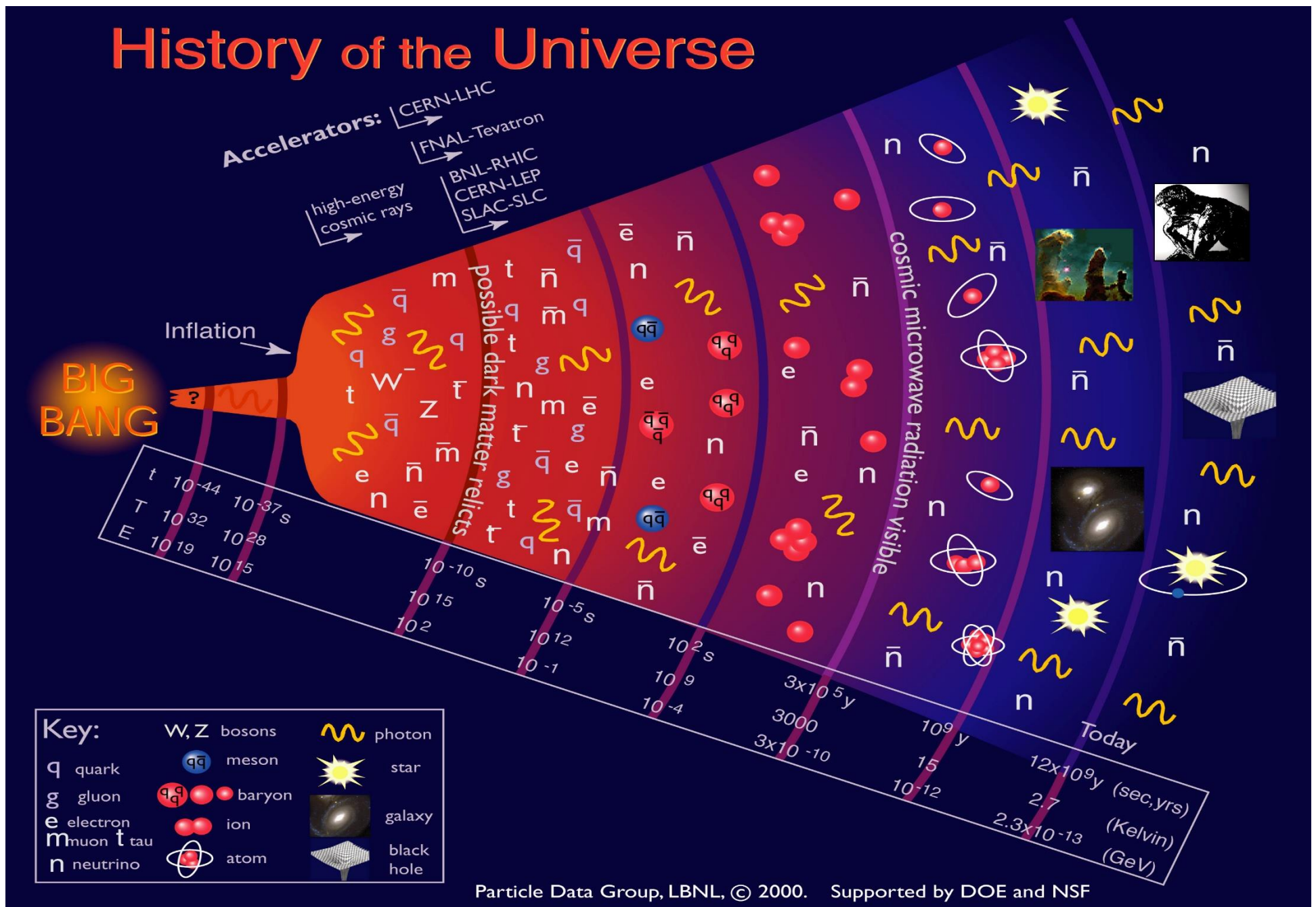


$$\rho > \rho_c$$

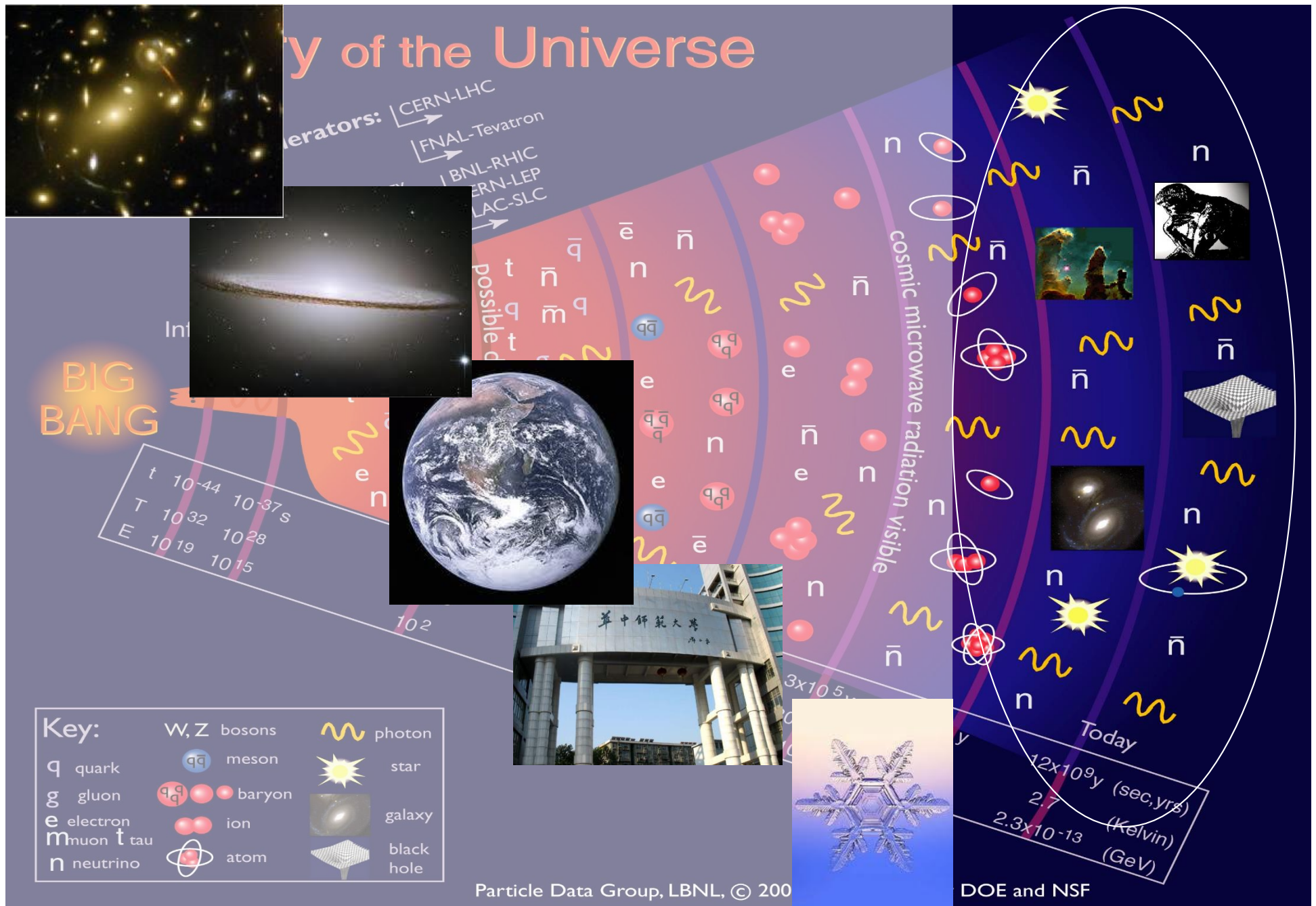
QCD Phase Diagram



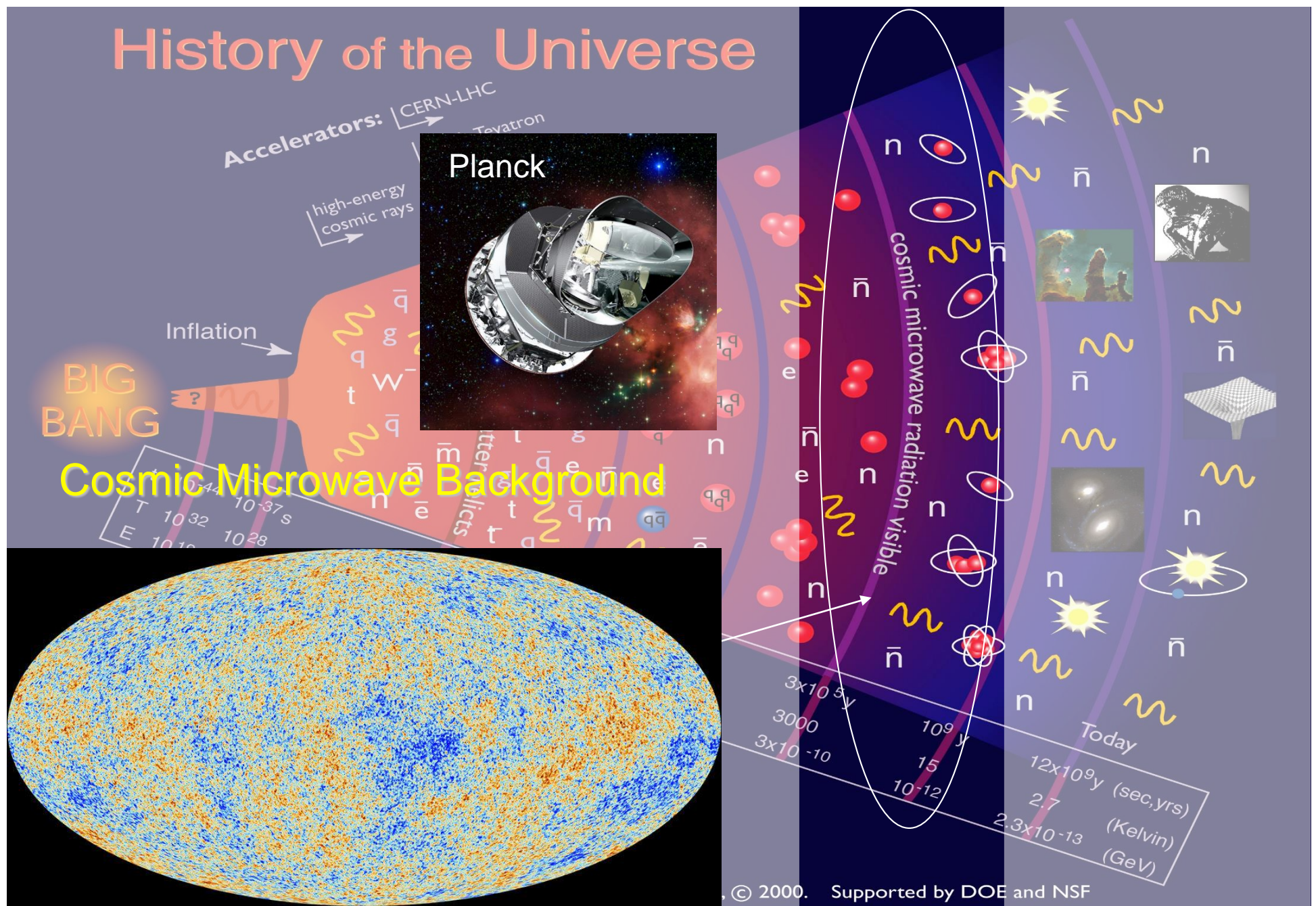
History of the Universe



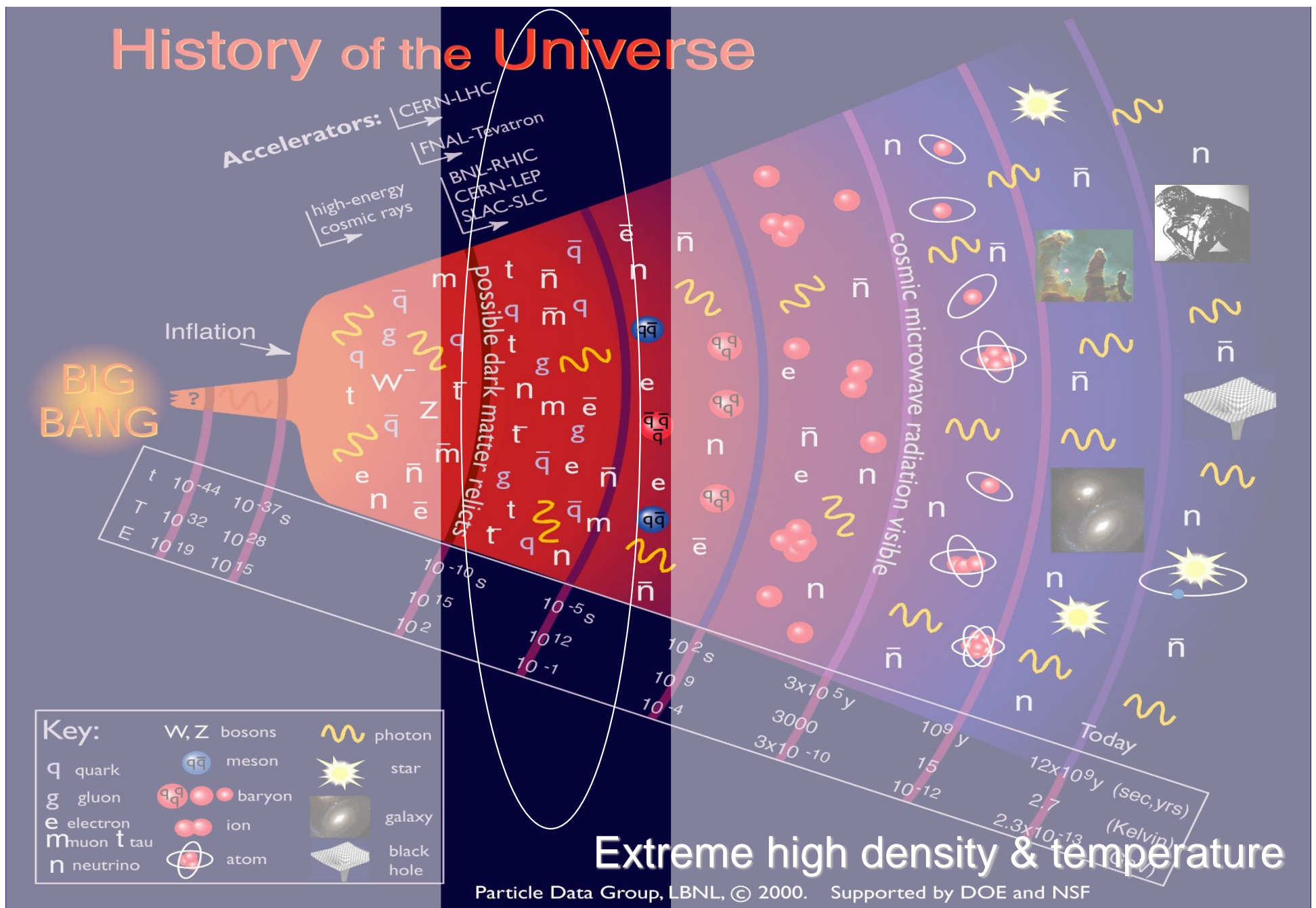
History of the Universe



History of the Universe



History of the Universe

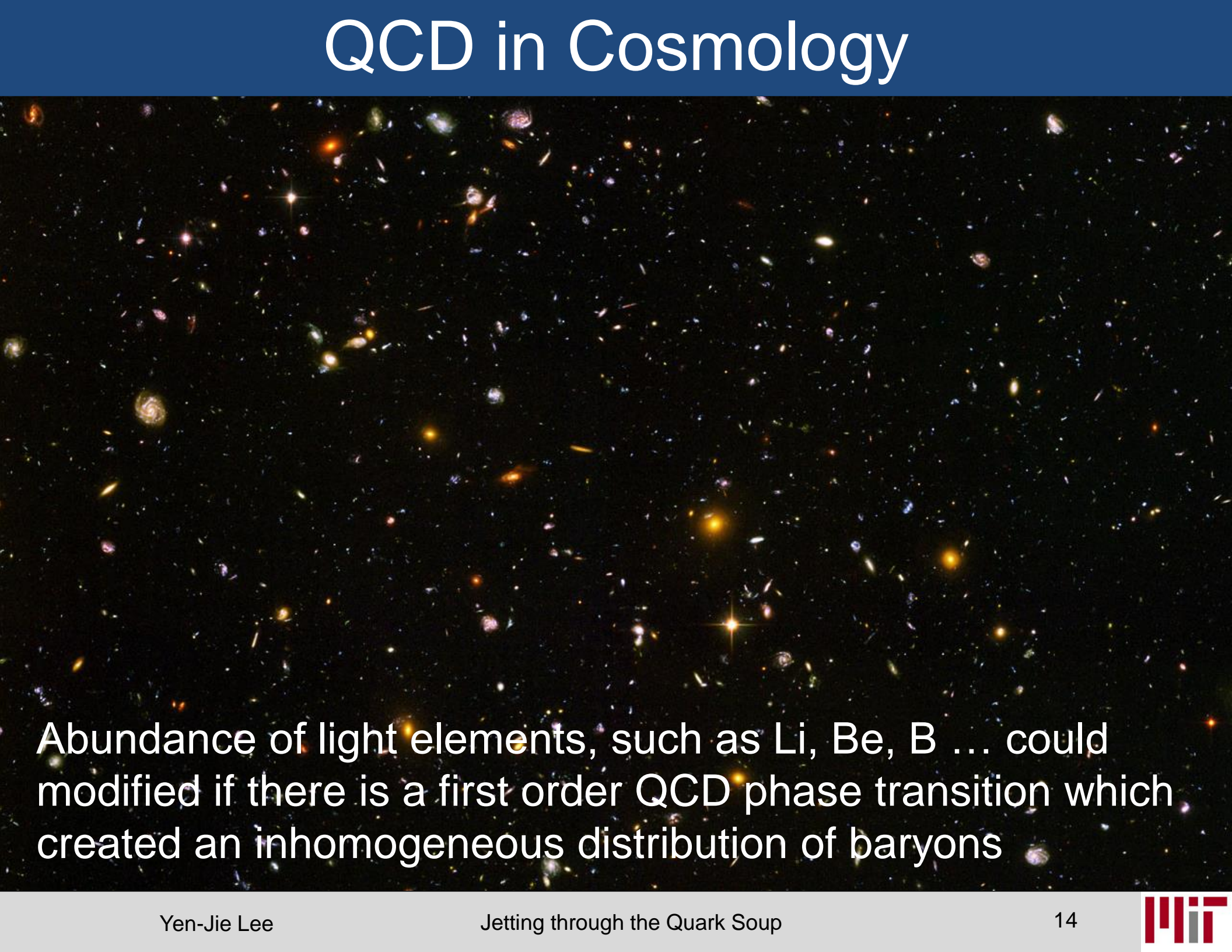


First order transition between water and ice



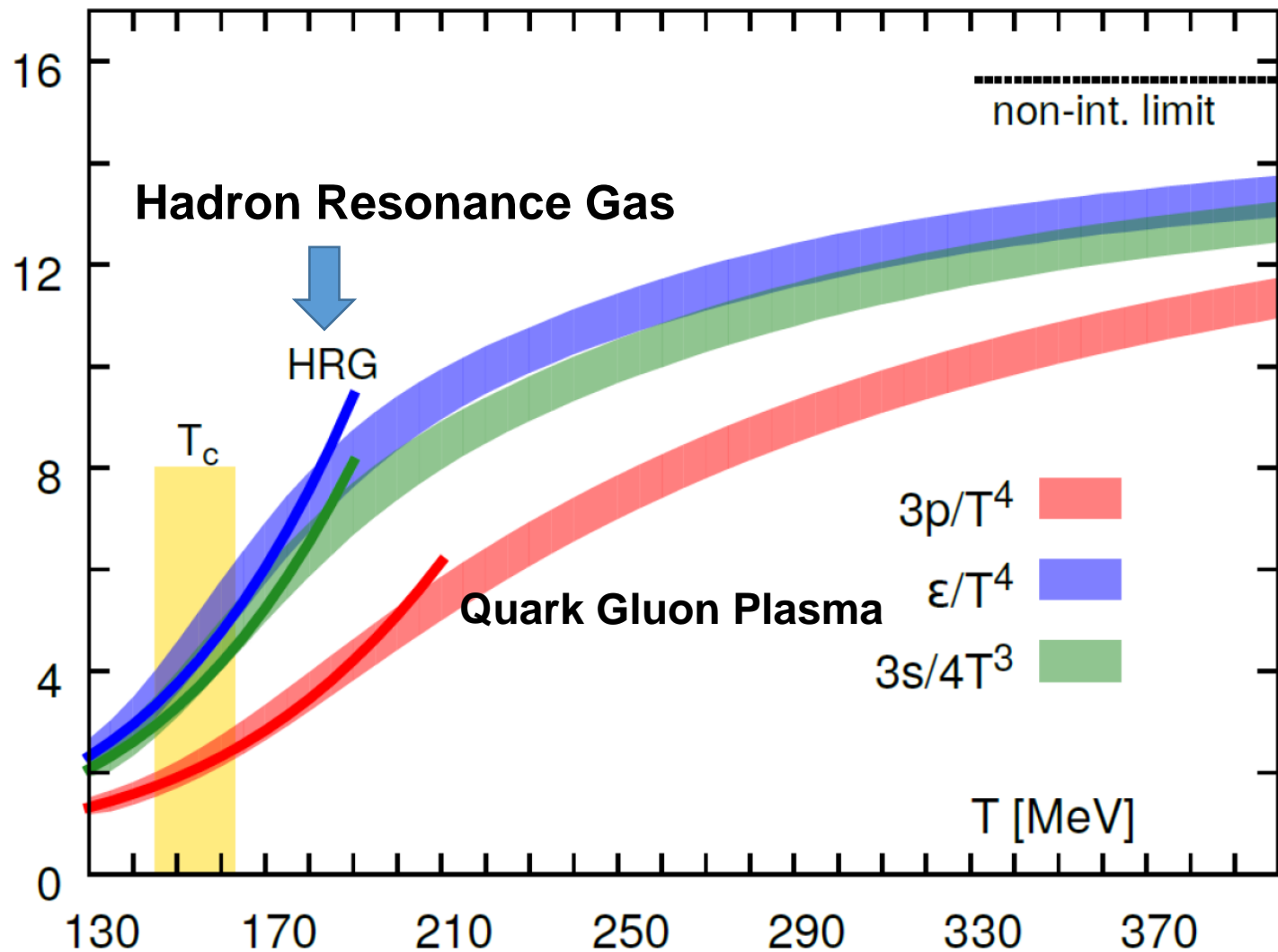
<https://i7wen.com/zh-tw/articles/5041>

QCD in Cosmology



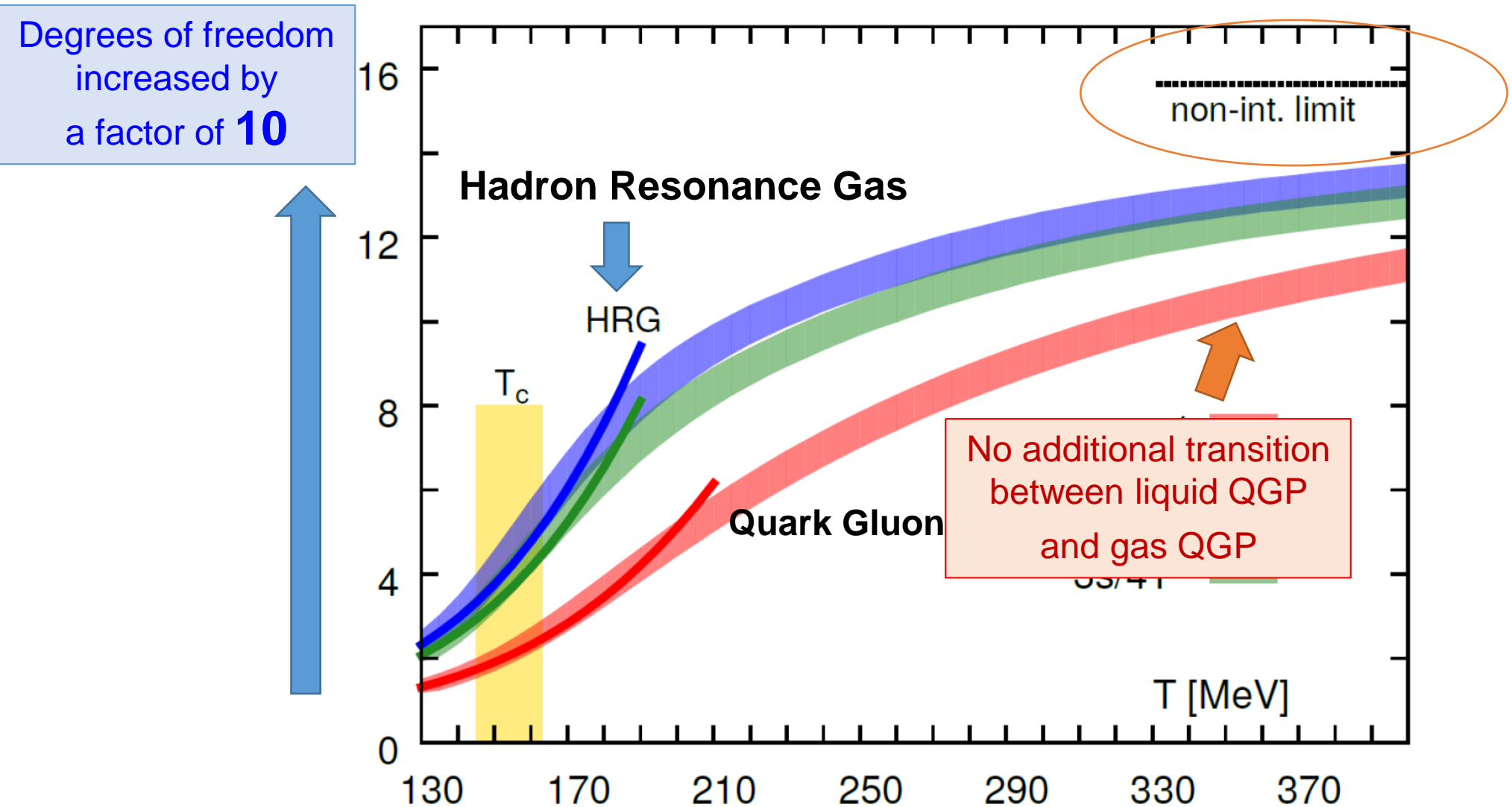
Abundance of light elements, such as Li, Be, B ... could be modified if there is a first order QCD phase transition which created an inhomogeneous distribution of baryons

QCD Equation of State at $\mu_B=0$



Lattice QCD predicts a **continuous cross-over** between hadron gas and quark gluon plasma

QCD Equation of State at $\mu_B=0$



Lattice QCD predicts a **continuous cross-over** between hadron gas and quark gluon plasma

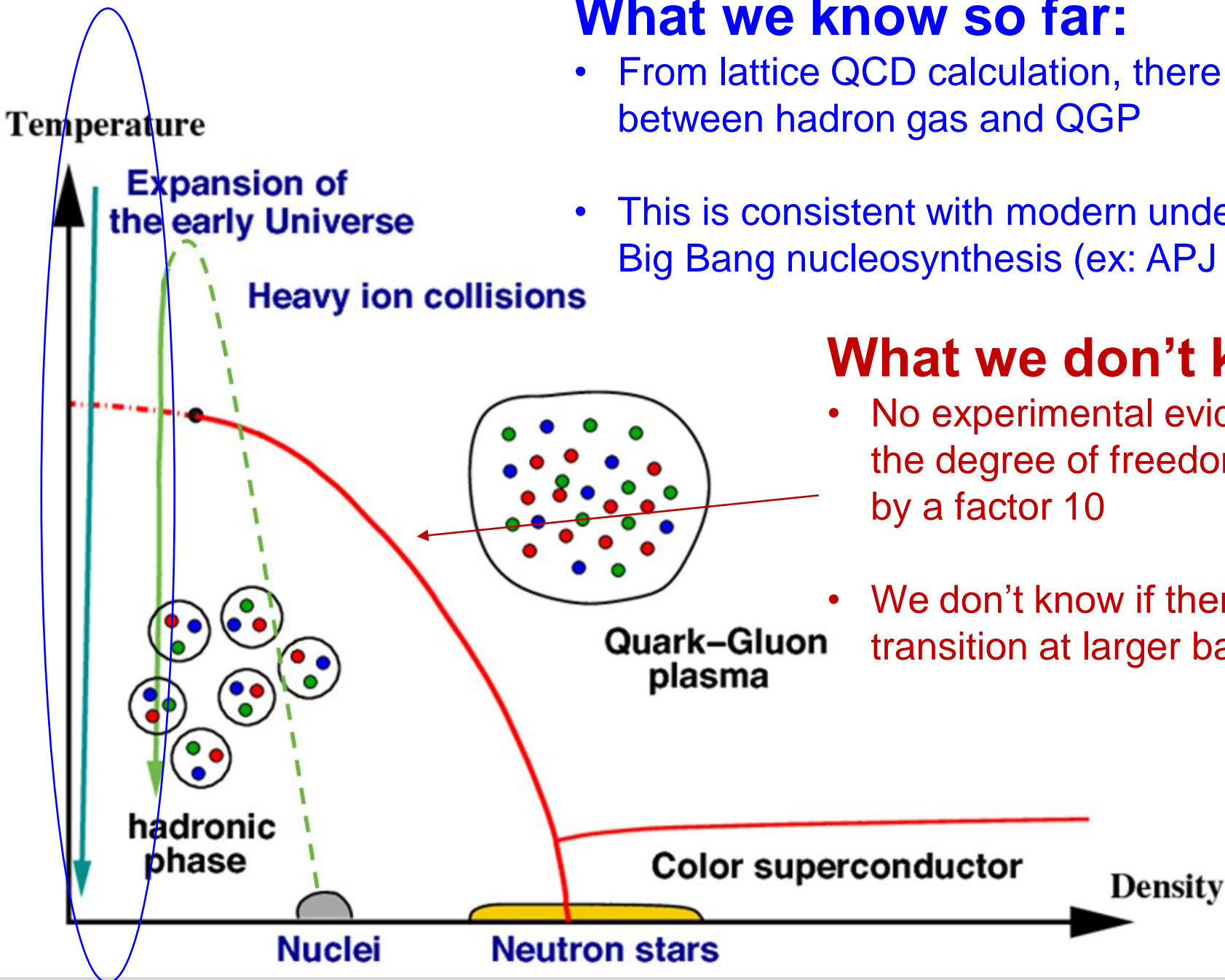
QCD Phase Diagram

What we know so far:

- From lattice QCD calculation, there is a cross-over between hadron gas and QGP
- This is consistent with modern understanding of Big Bang nucleosynthesis (ex: APJ 430, 291 (1994))

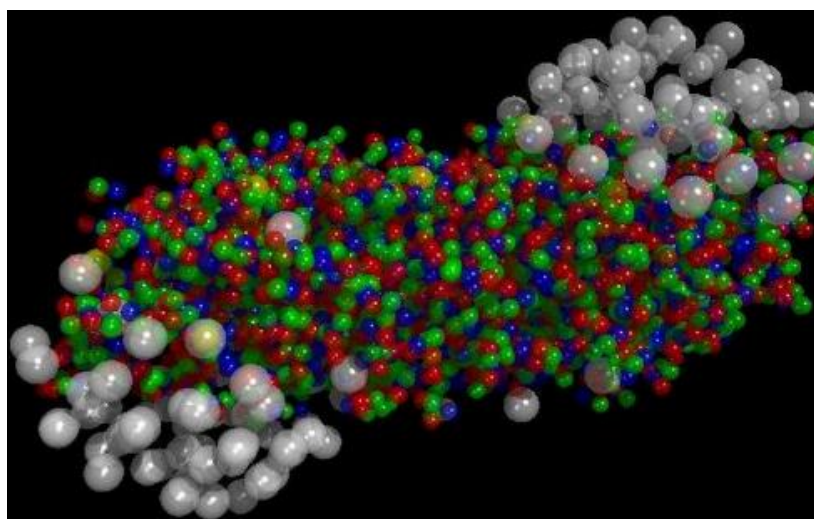
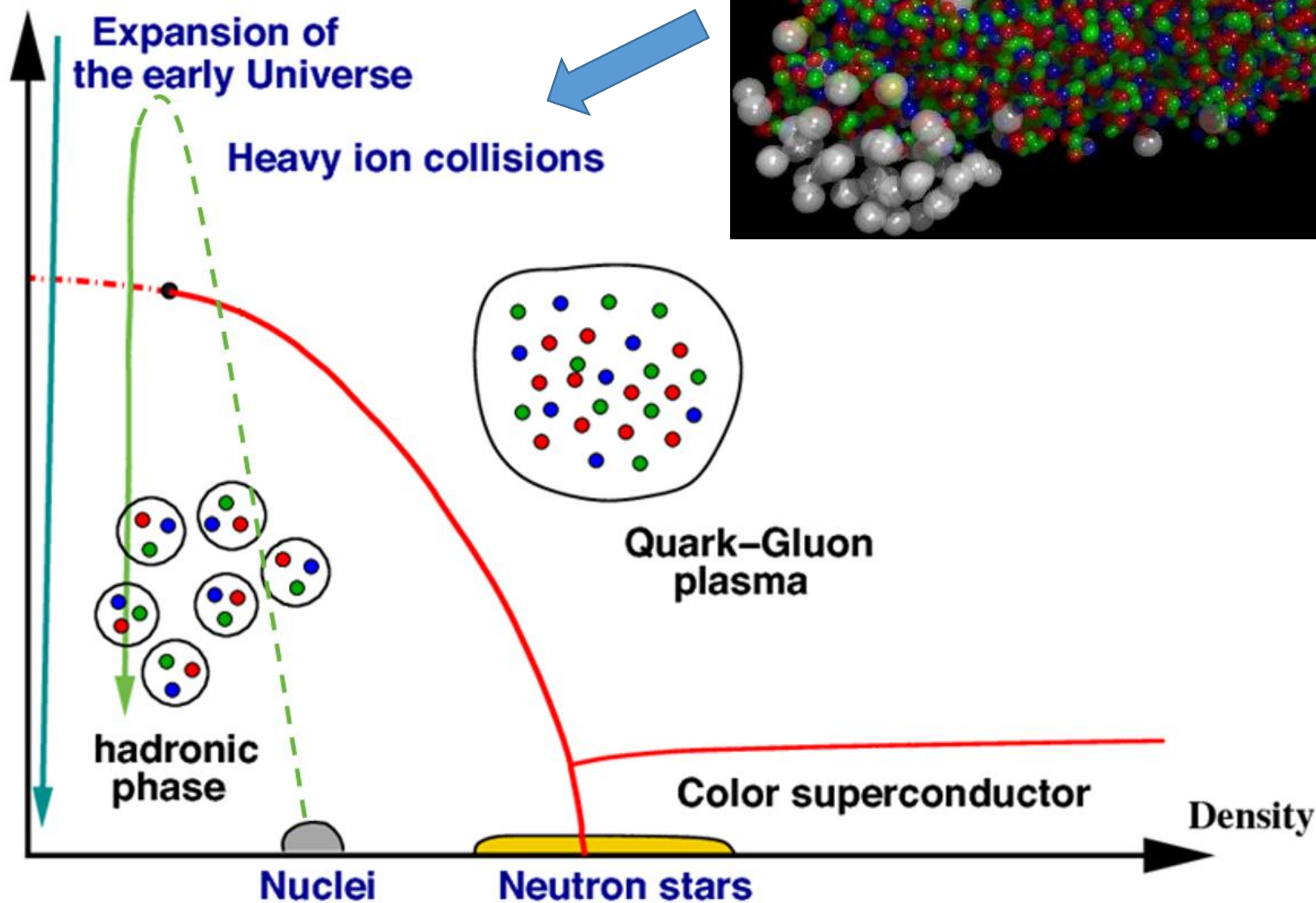
What we don't know yet:

- No experimental evidence yet that the degree of freedom has increased by a factor 10
- We don't know if there is a first order transition at larger baryon density



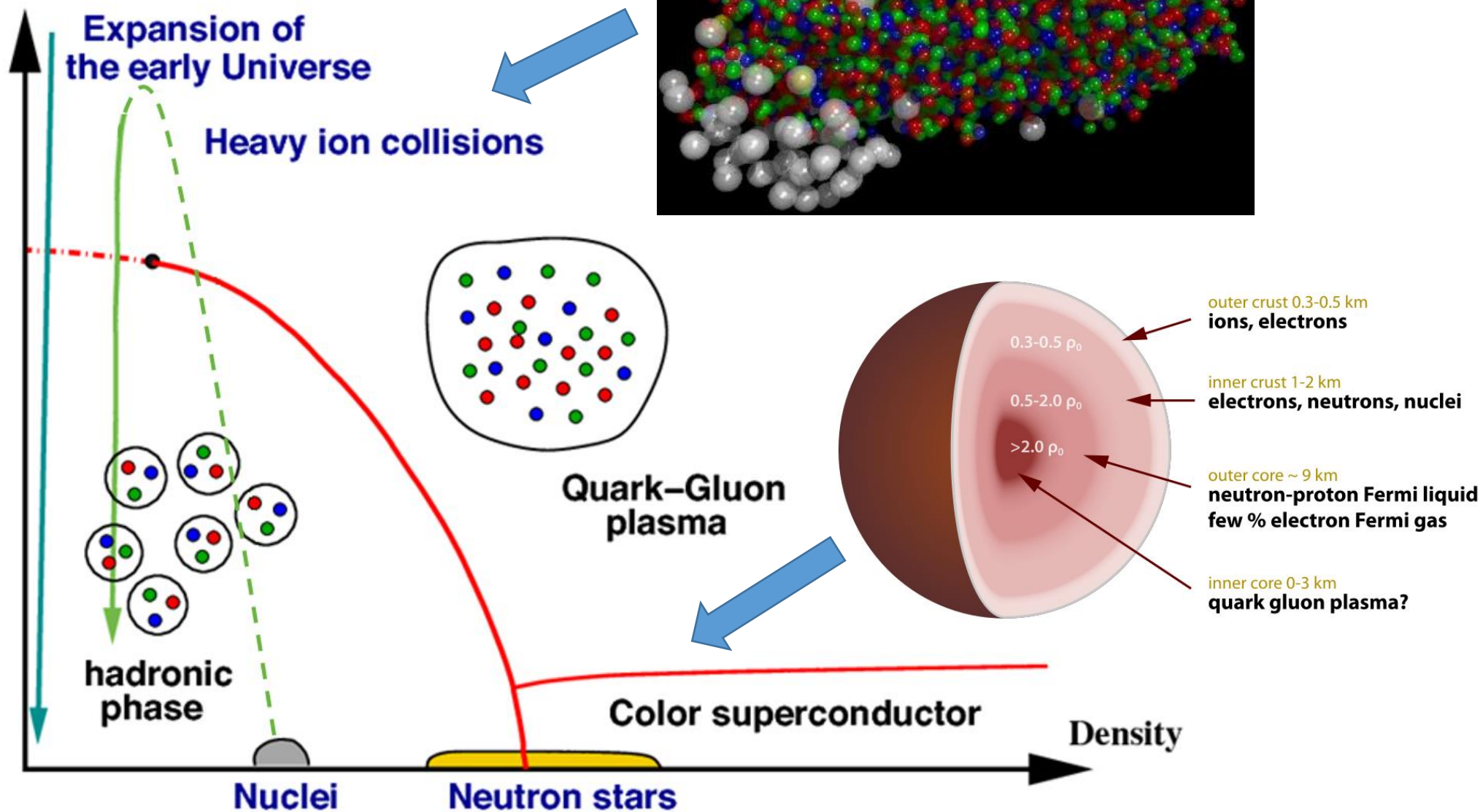
QCD Phase Diagram

Temperature



QCD Phase Diagram

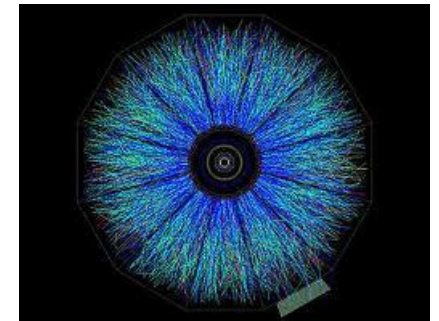
Temperature



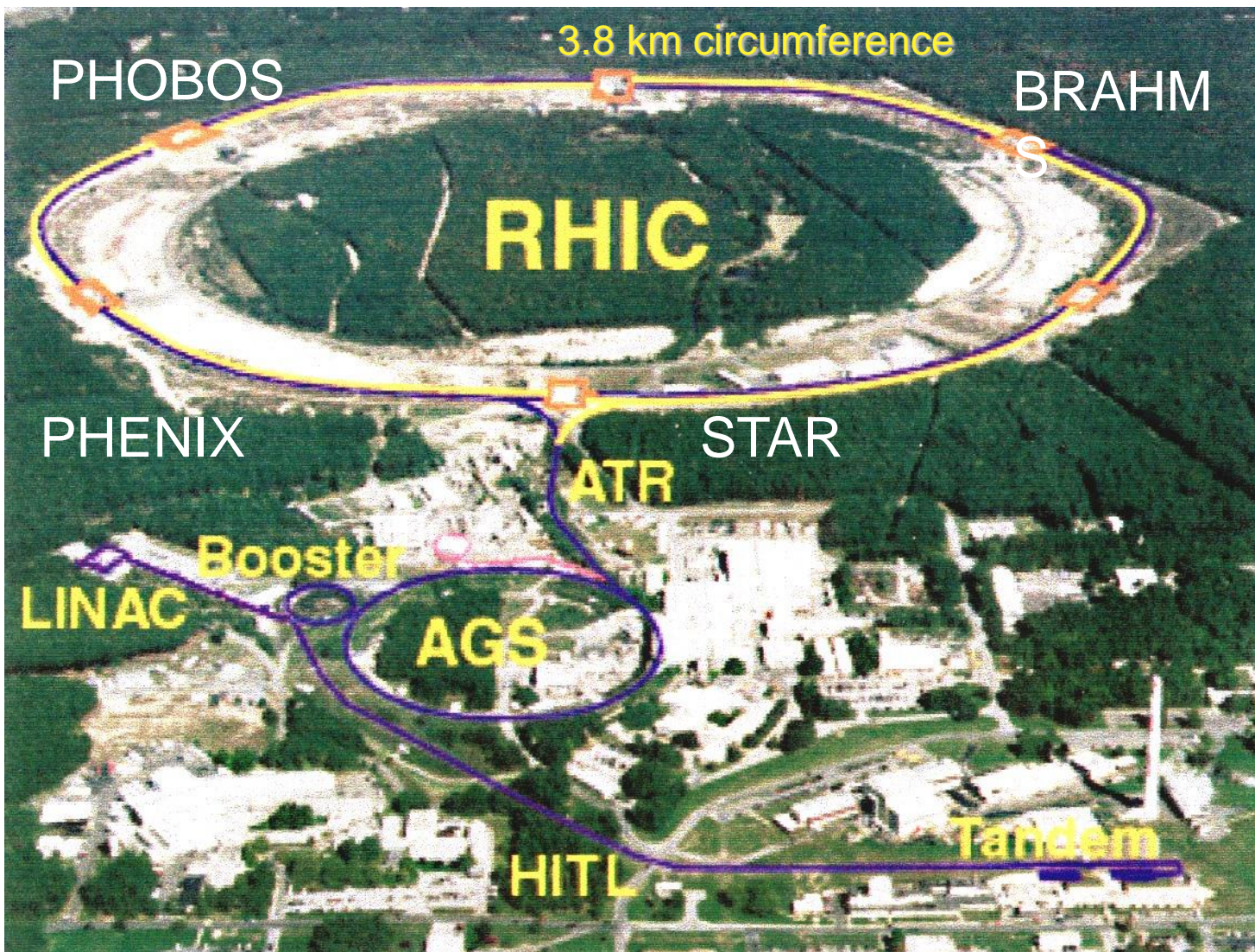
The First Dedicated Heavy Ion Collider

Relativistic Heavy Ion Collider

Gold+Gold
7.7 - 200 GeV



Since 2000~



The High Energy Frontier

Large Hadron Collider

Lead+Lead collisions

2010-11: 2.76 TeV

2015: 5.02 TeV

25x jump with respect to RHIC!

27 km circumference

Lake Geneva

CMS

France

LHCb

ALICE

ATLAS

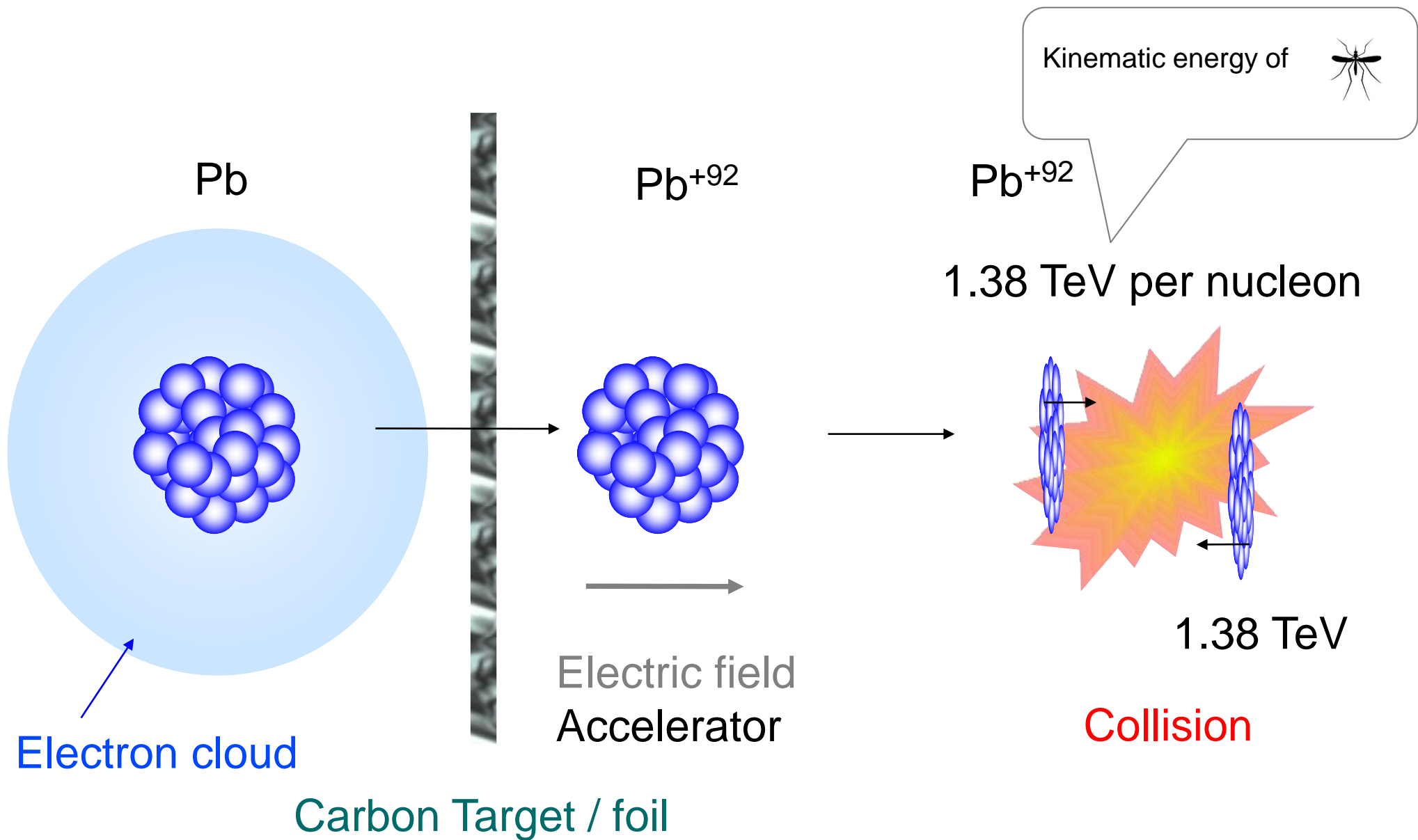
Switzerland



RHIC



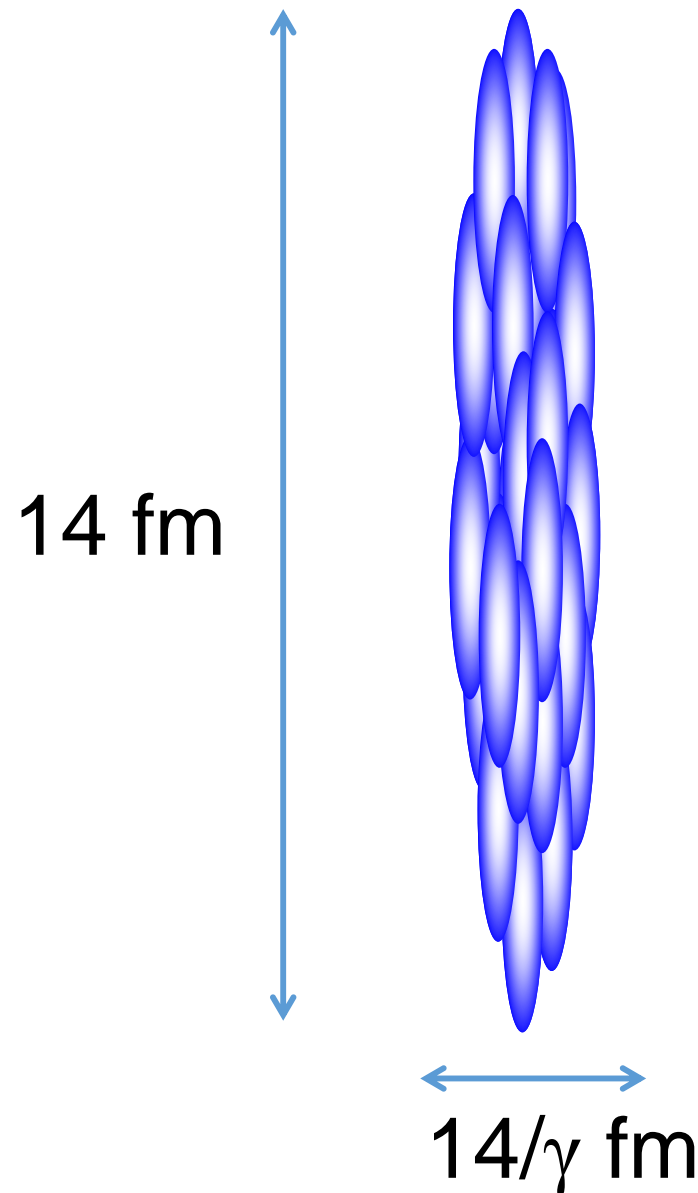
How do we collide ions?



Each ion carry $1.38 \text{ TeV} \times 208 \sim 50 \mu\text{J}$

Lorentz Contracted Discs

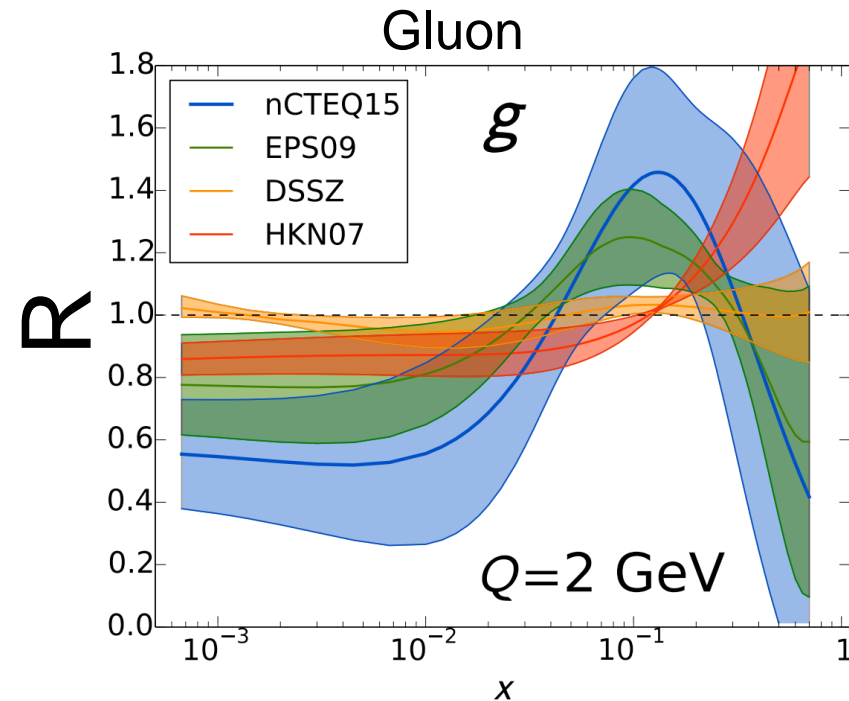
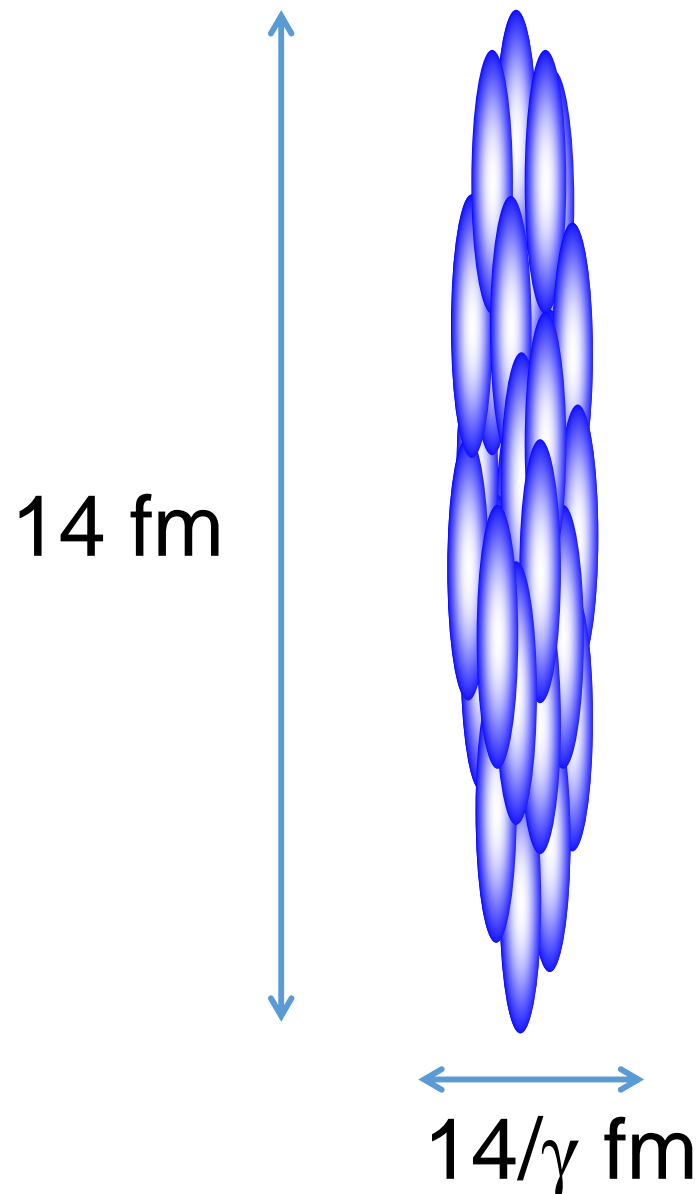
For large nuclei such as Gold (Au) and Lead (Pb):



$\gamma = 100$ at RHIC
2500 at LHC

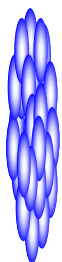
Parton Distribution Function (PDF)

For large nuclei such as Gold (Au) and Lead (Pb):



$$R = \frac{nPDF}{PDF}$$

Relativistic Heavy Ion Collisions



Heavy Ion



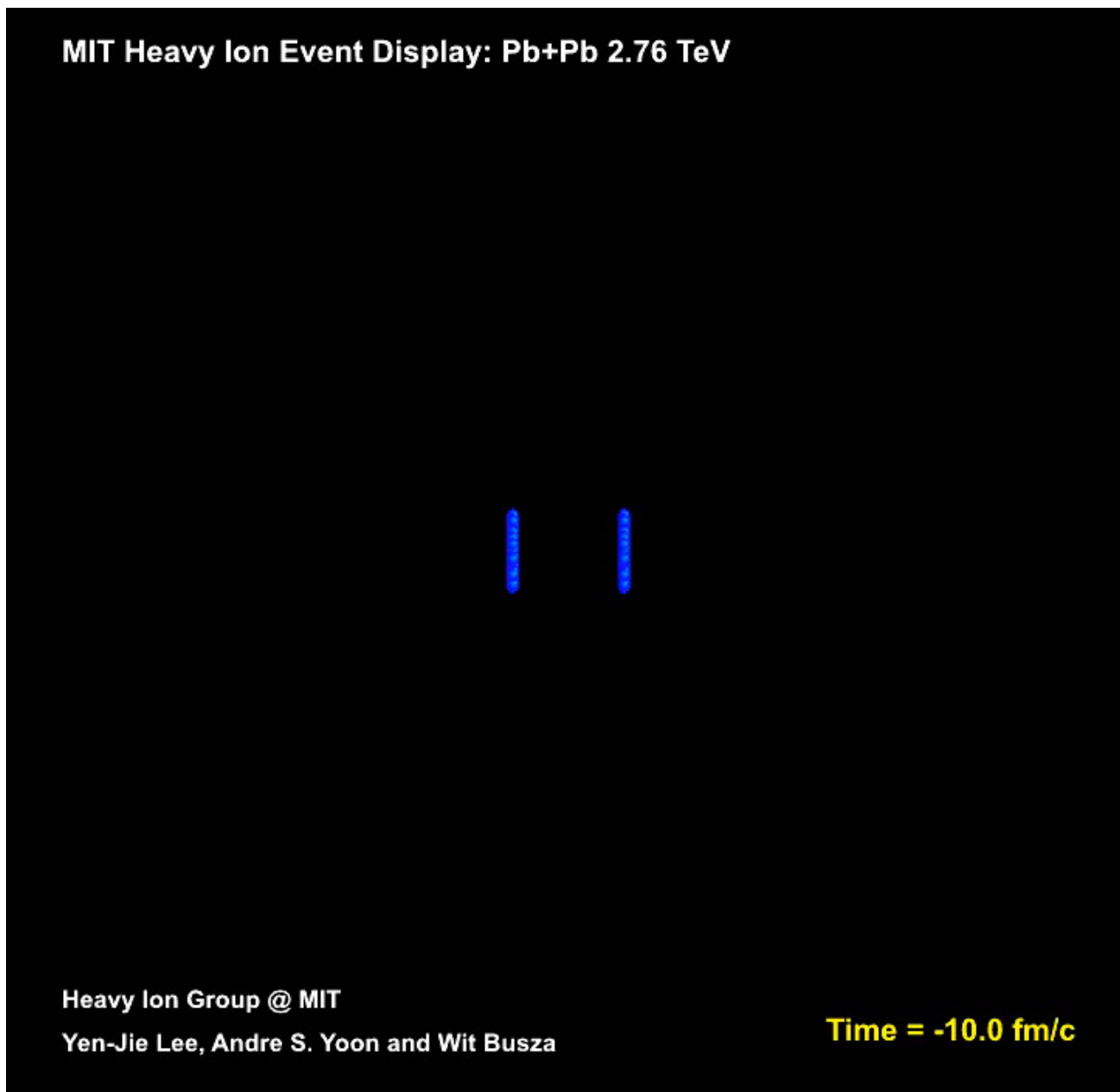
Particles before hadronization
(Quark Gluon Plasma)



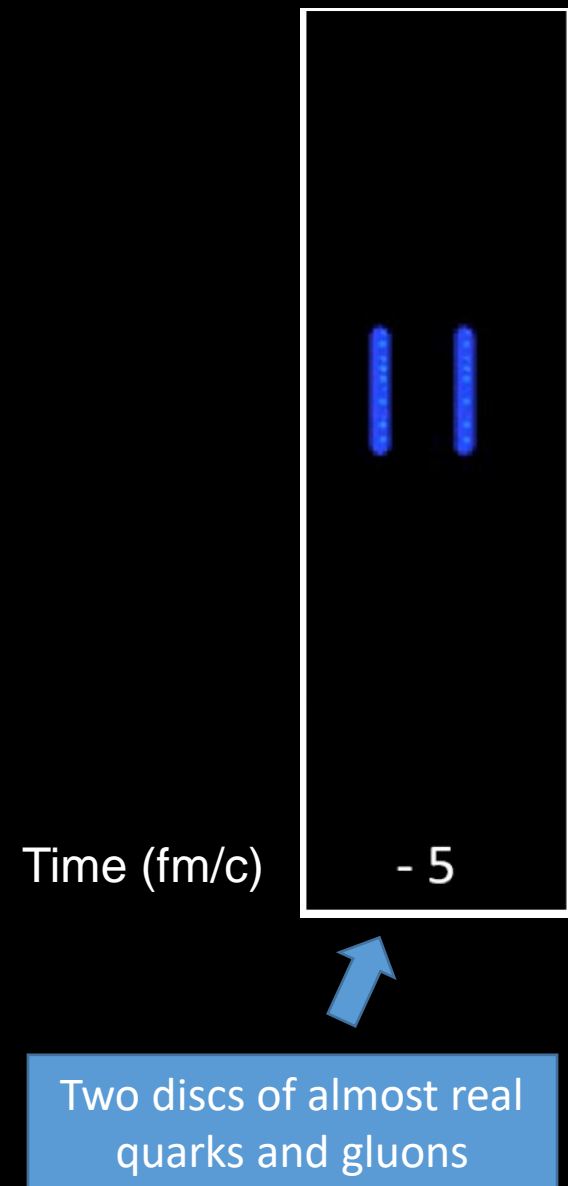
Meson



Baryon

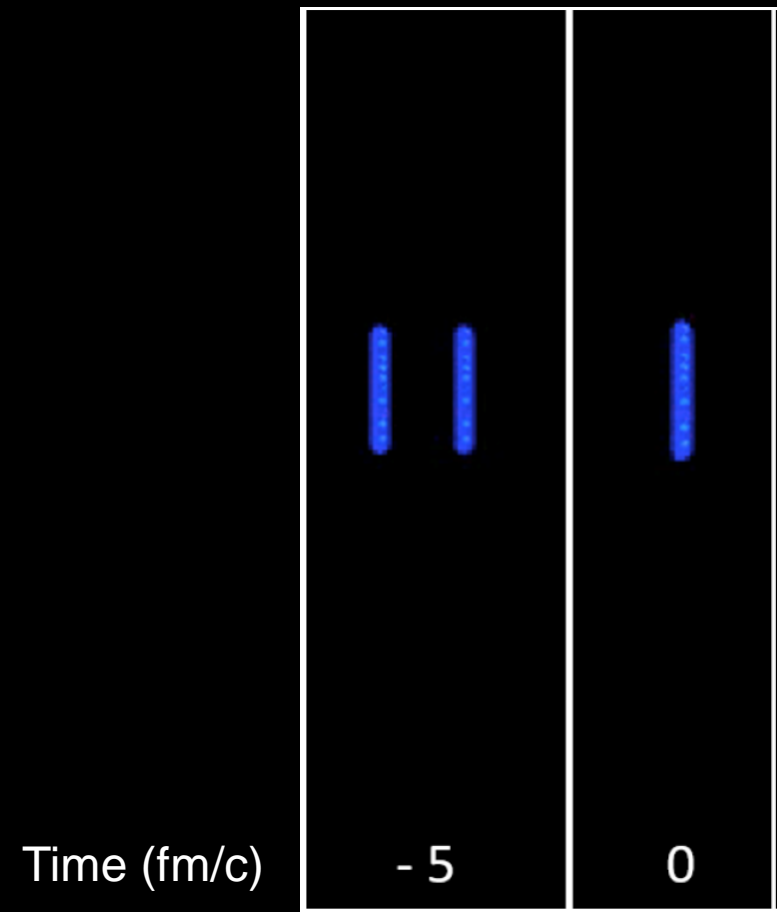


Relativistic Heavy Ion Collisions



Relativistic Heavy Ion Collisions

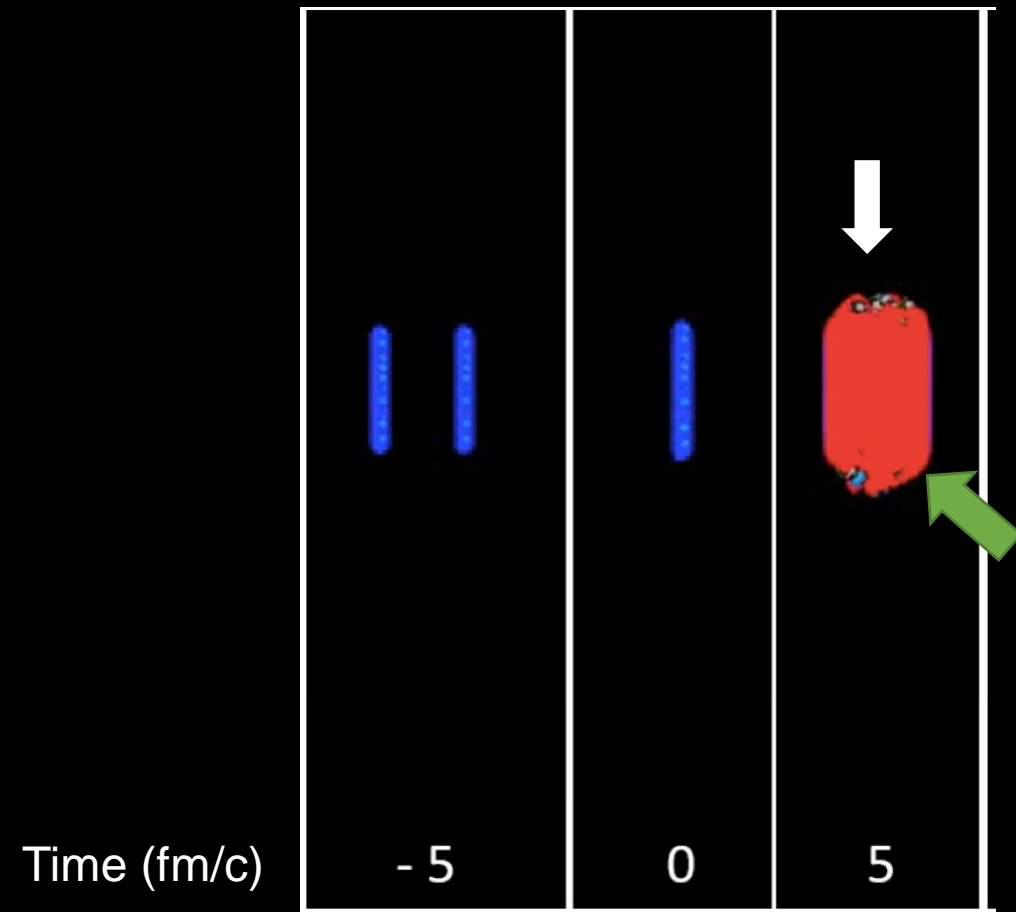
Collision! Highest energy density state. Huge amount of soft (low momentum transfer) scatterings.



Two discs of almost real quarks and gluons

Relativistic Heavy Ion Collisions

Collision! Highest energy density state. Huge amount of soft (low momentum transfer) scatterings.



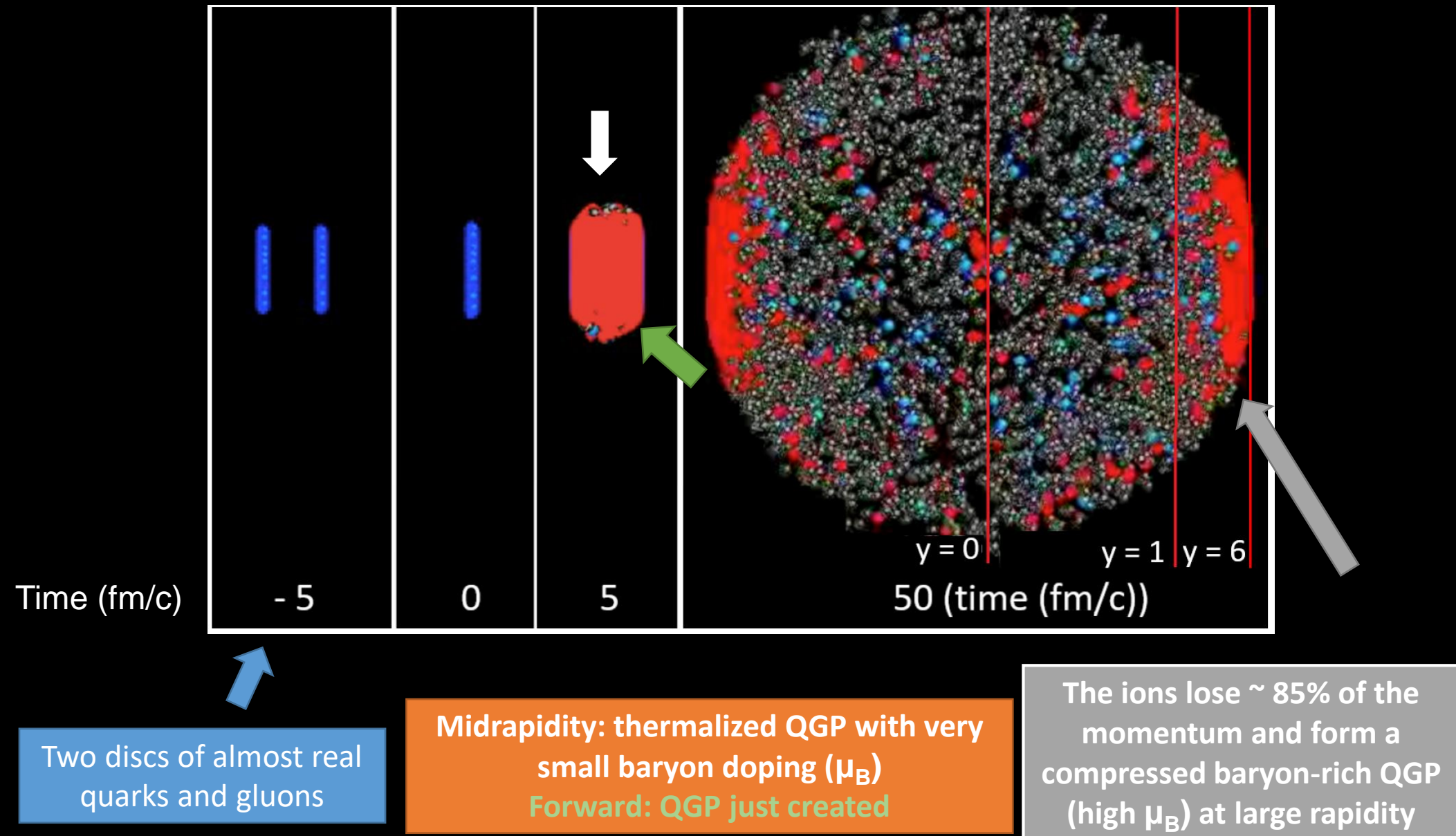
Two discs of almost real quarks and gluons

Midrapidity: thermalized QGP with very small baryon doping (μ_B)
Forward: QGP just created

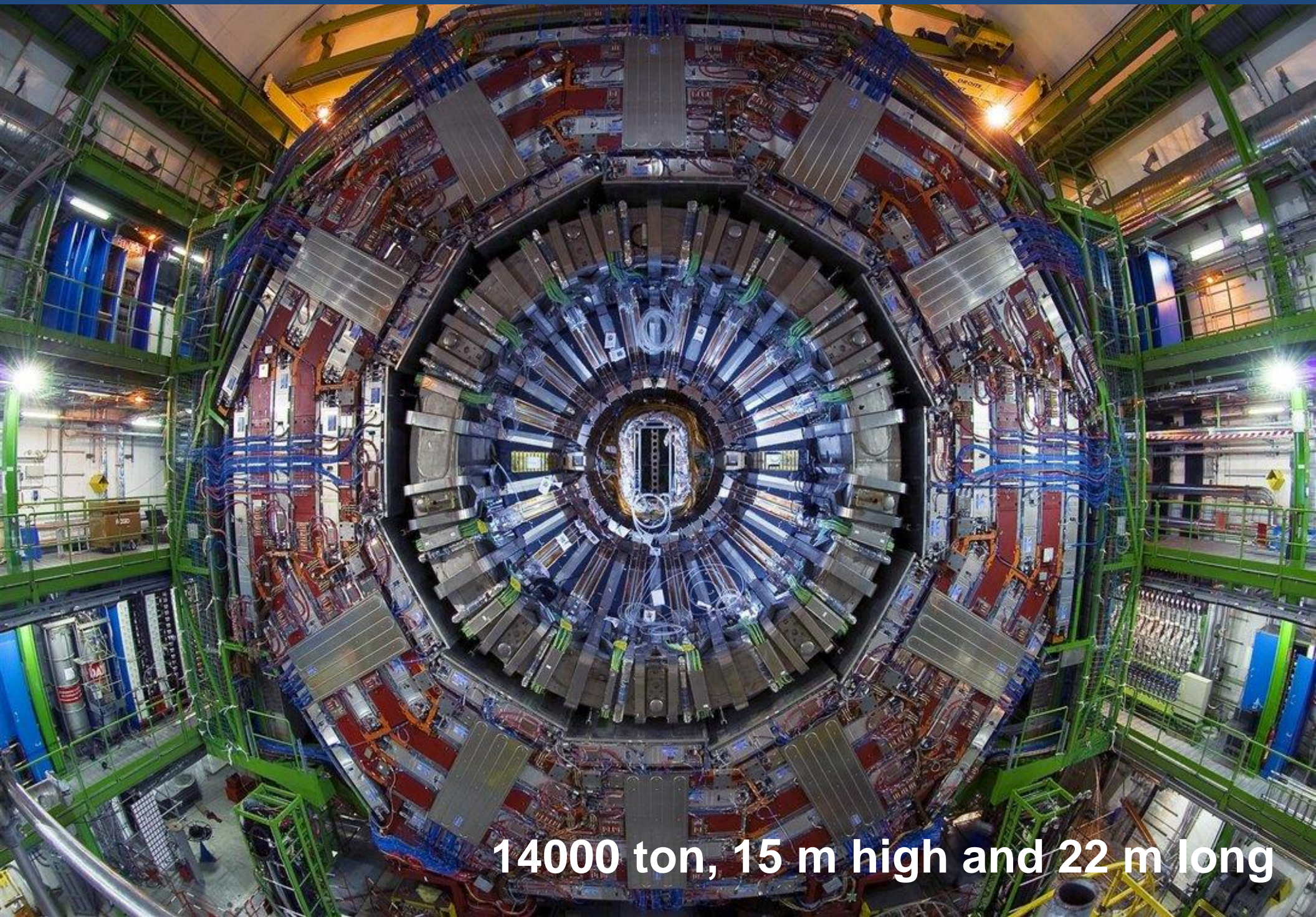
Relativistic Heavy Ion Collisions

Collision! Highest energy density state. Huge amount of soft (low momentum transfer) scatterings.

Hadronization of QGP, different from elementary collisions like e^+e^- or pp collisions



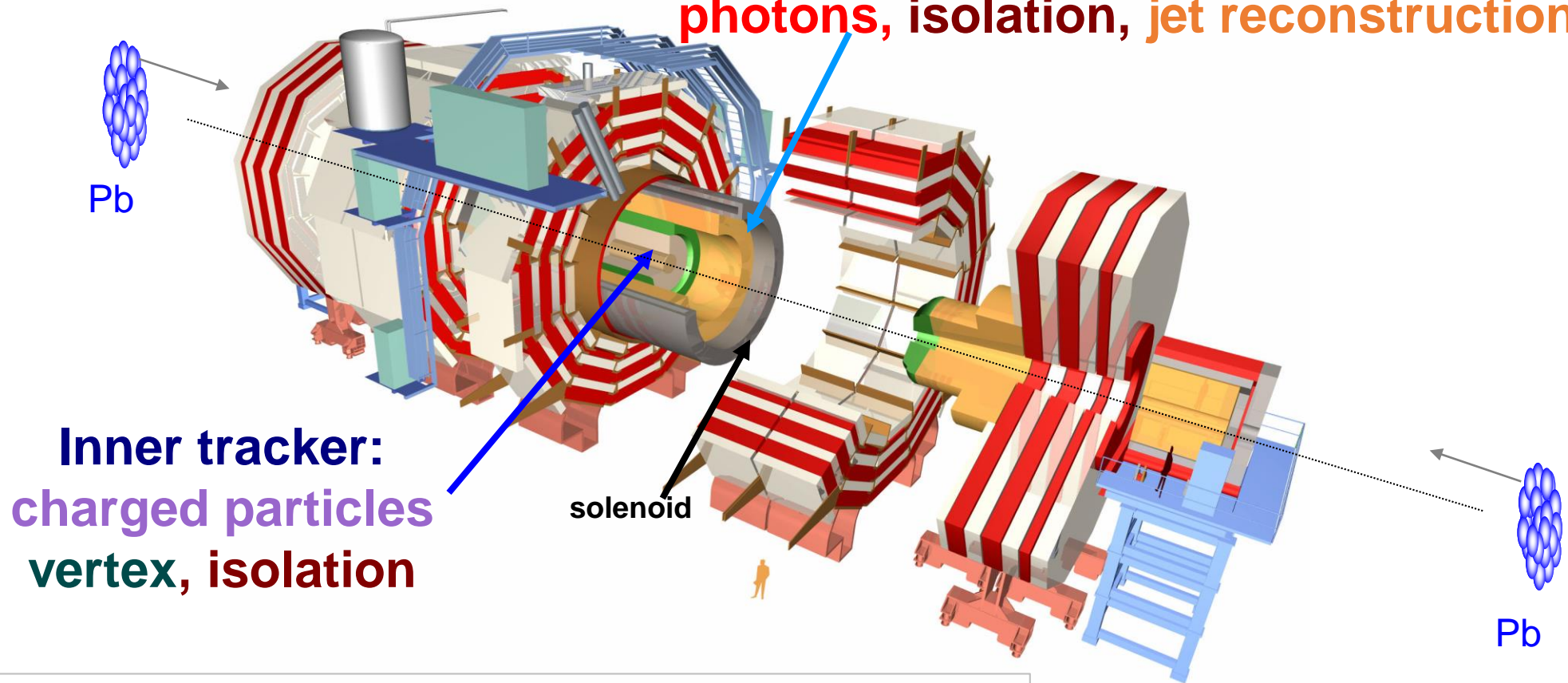
A QGP detector: CMS (2008)



14000 ton, 15 m high and 22 m long

A QGP detector: CMS

EM and **Hadron** calorimeters
photons, isolation, jet reconstruction



Muon

$|\eta| < 2.4$

HCAL

$|\eta| < 5.2$

ECAL

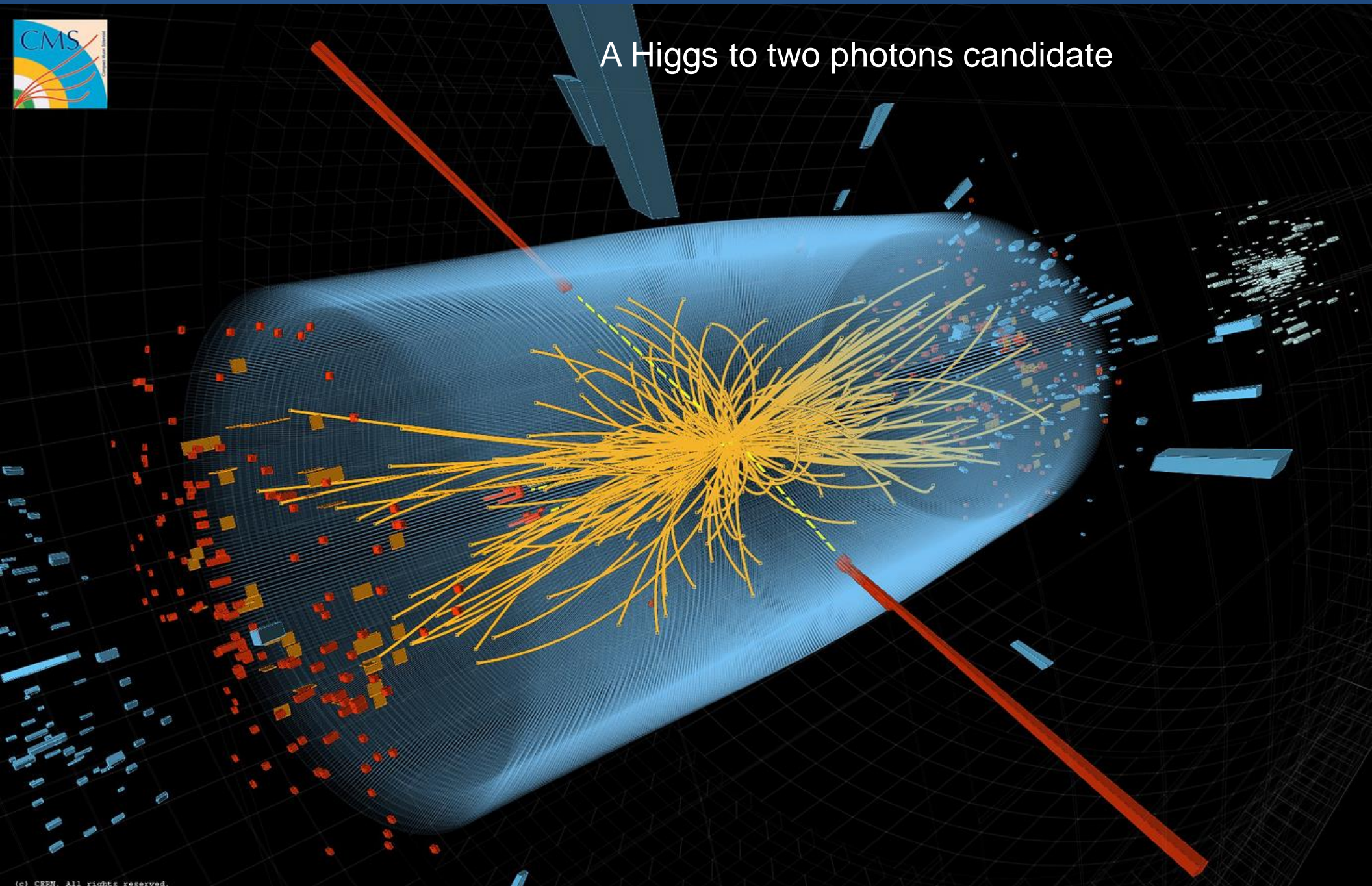
$|\eta| < 3.0$

Tracker

$|\eta| < 2.5$

Optimized for high p_T physics:
a perfect jet detector for the
studies of QGP

Proton-Proton Collision Recorded by the CMS detector



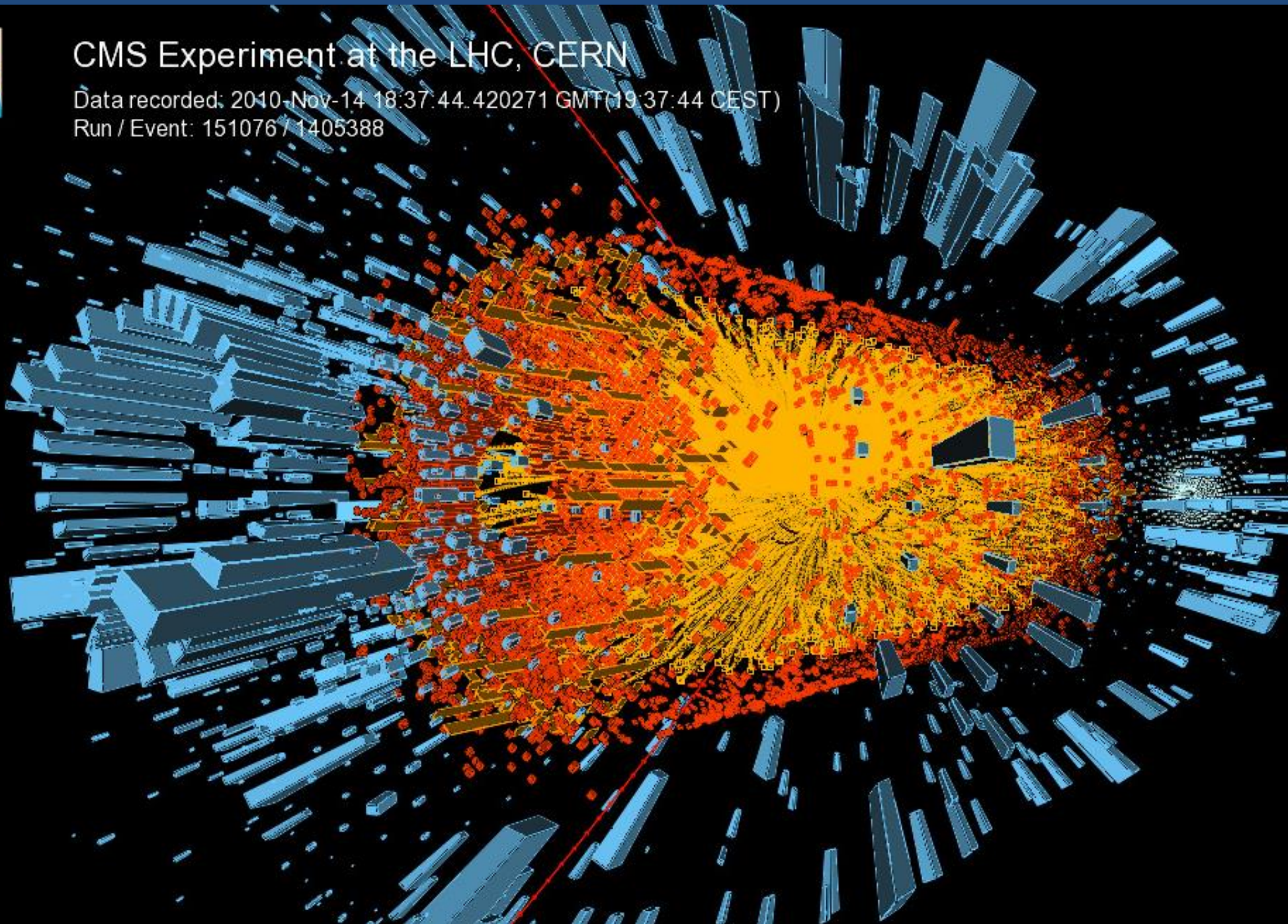
Lead-Lead Collision Recorded by the CMS Detector (2010)



CMS Experiment at the LHC, CERN

Data recorded: 2010-Nov-14 18:37:44.420271 GMT(19:37:44 CEST)

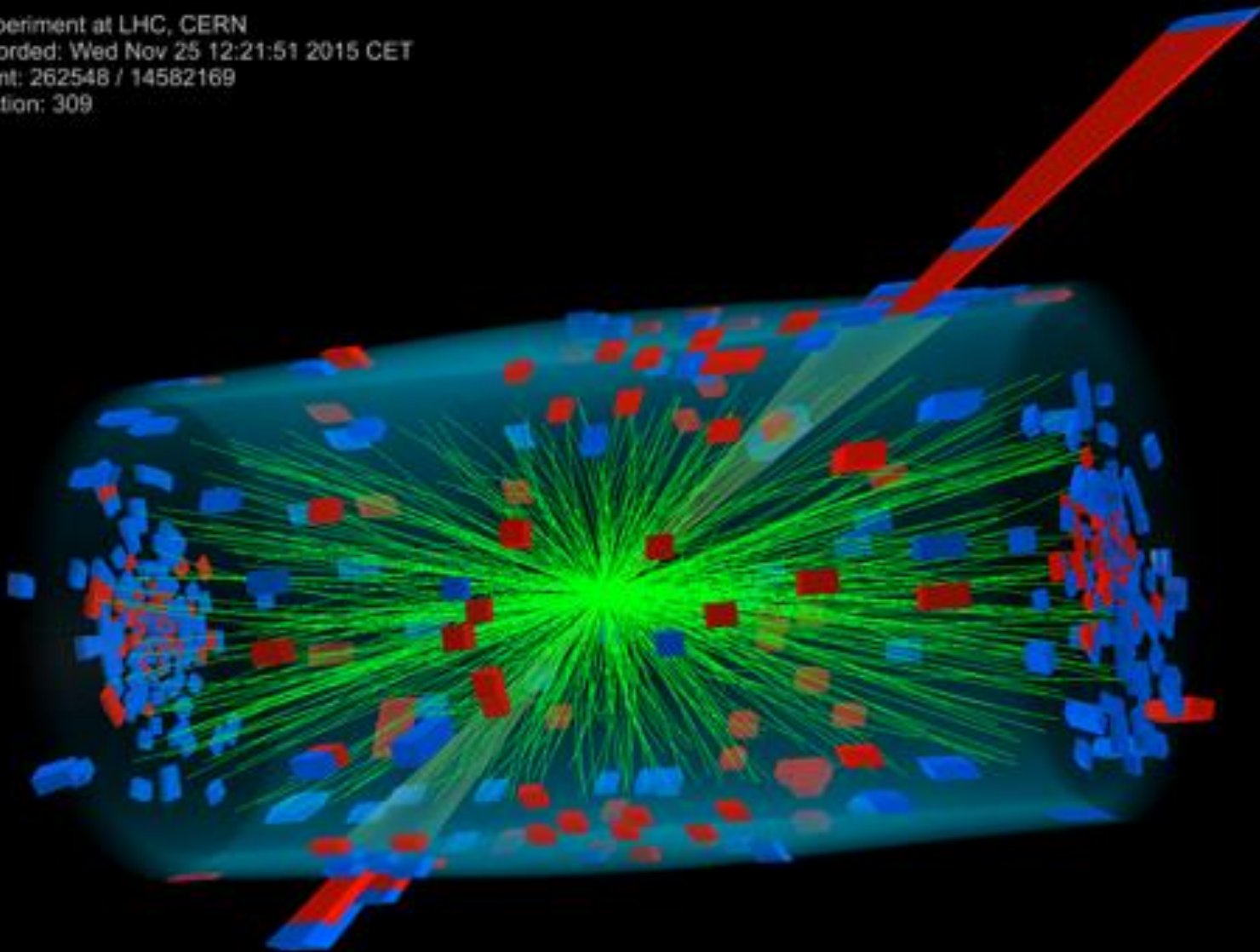
Run / Event: 151076 / 1405388



A beautiful dijet event in PbPb collisions

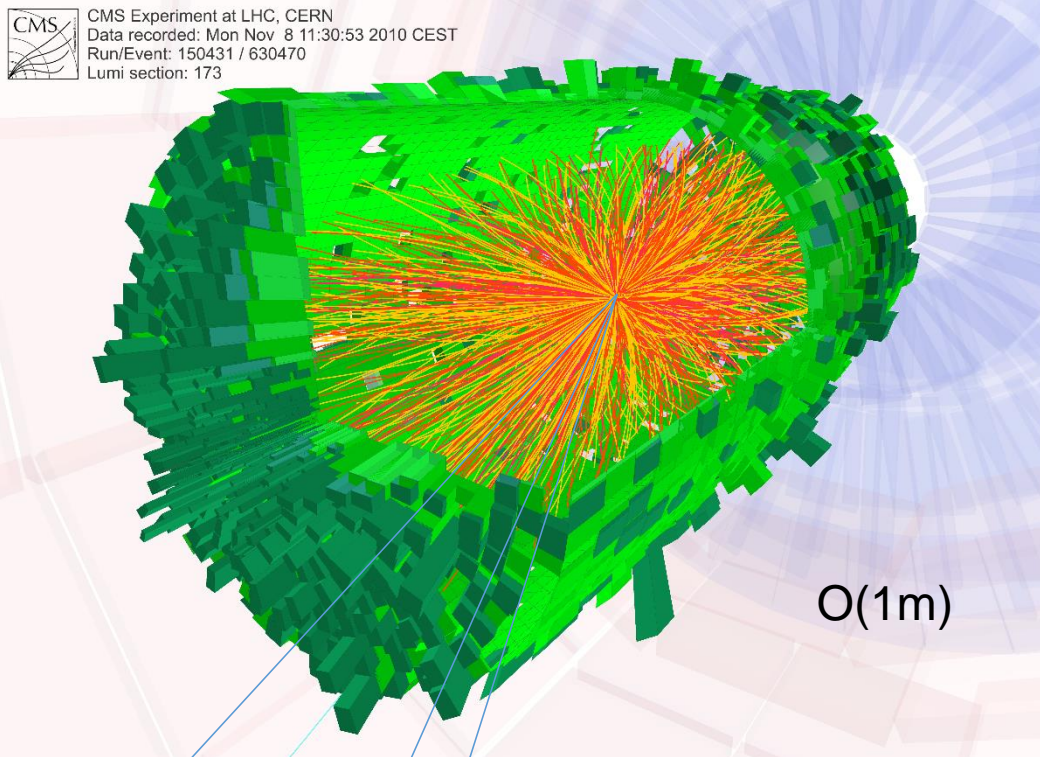


CMS Experiment at LHC, CERN
Data recorded: Wed Nov 25 12:21:51 2015 CET
Run/Event: 262548 / 14582169
Lumi section: 309

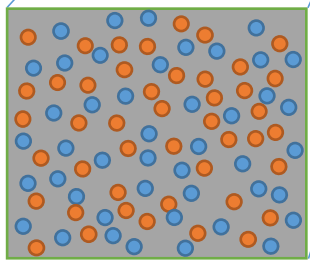


PbPb @ 5 TeV (2015)

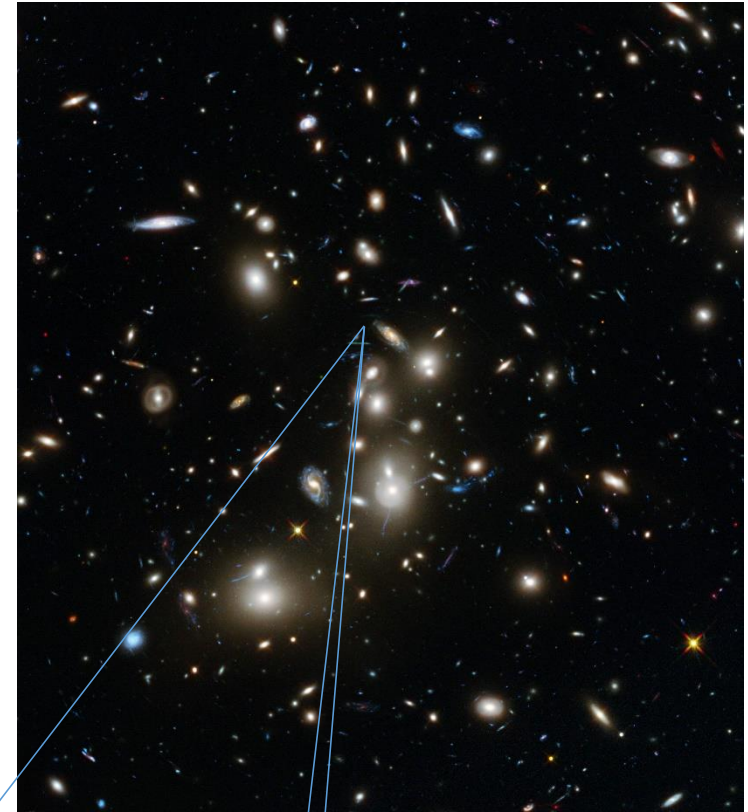
“Reconstruct the QGP” with CMS detector



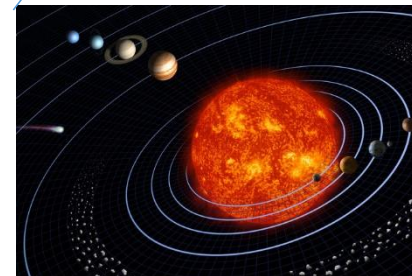
$O(1\text{m})$



$O(10^{-15}\text{ m})$



$O(10^{27}\text{ m})$

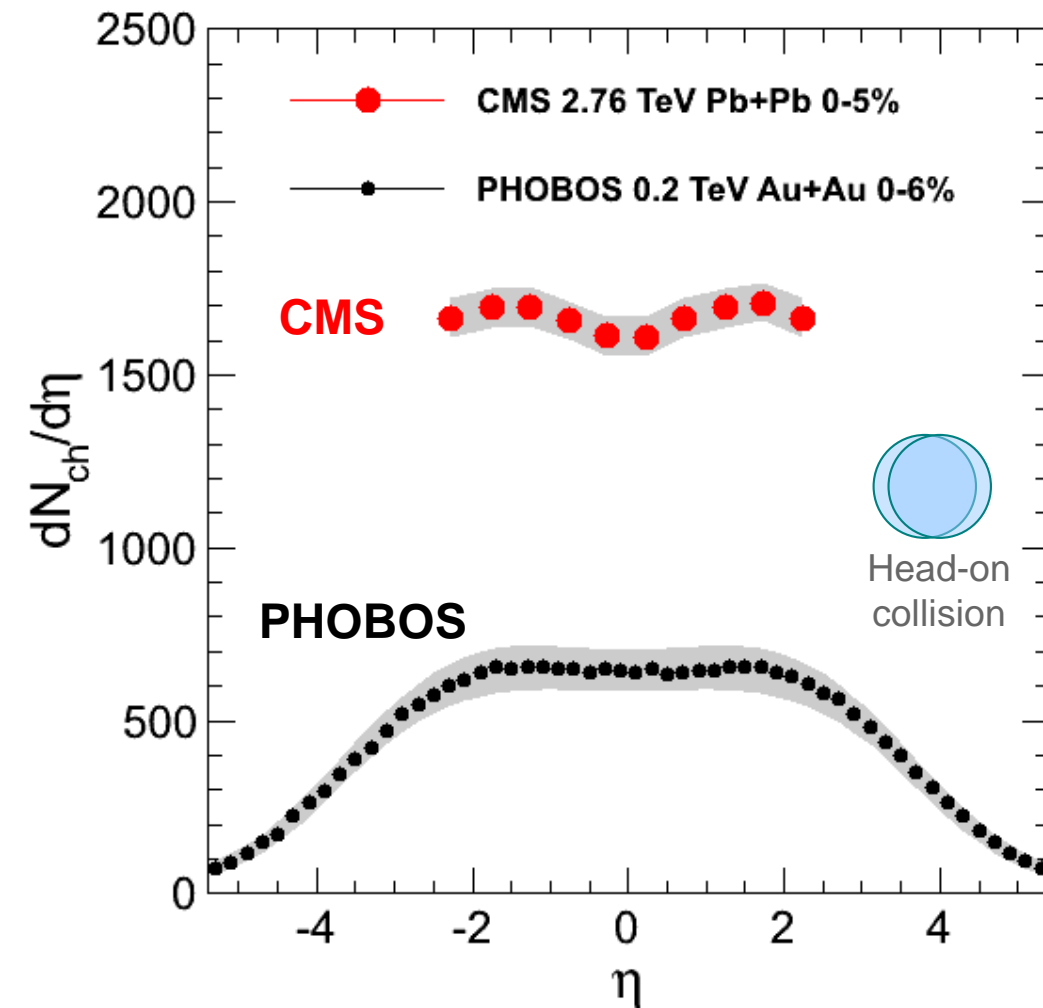


$O(10^{13}\text{ m})$

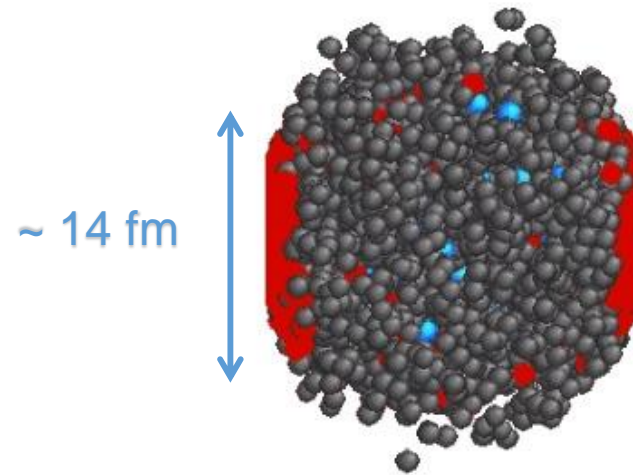
Charged Particle Counting

Number of particles produced in the collisions.

Particle density in Lead+Lead $\sim 400\times$ of that in proton+proton



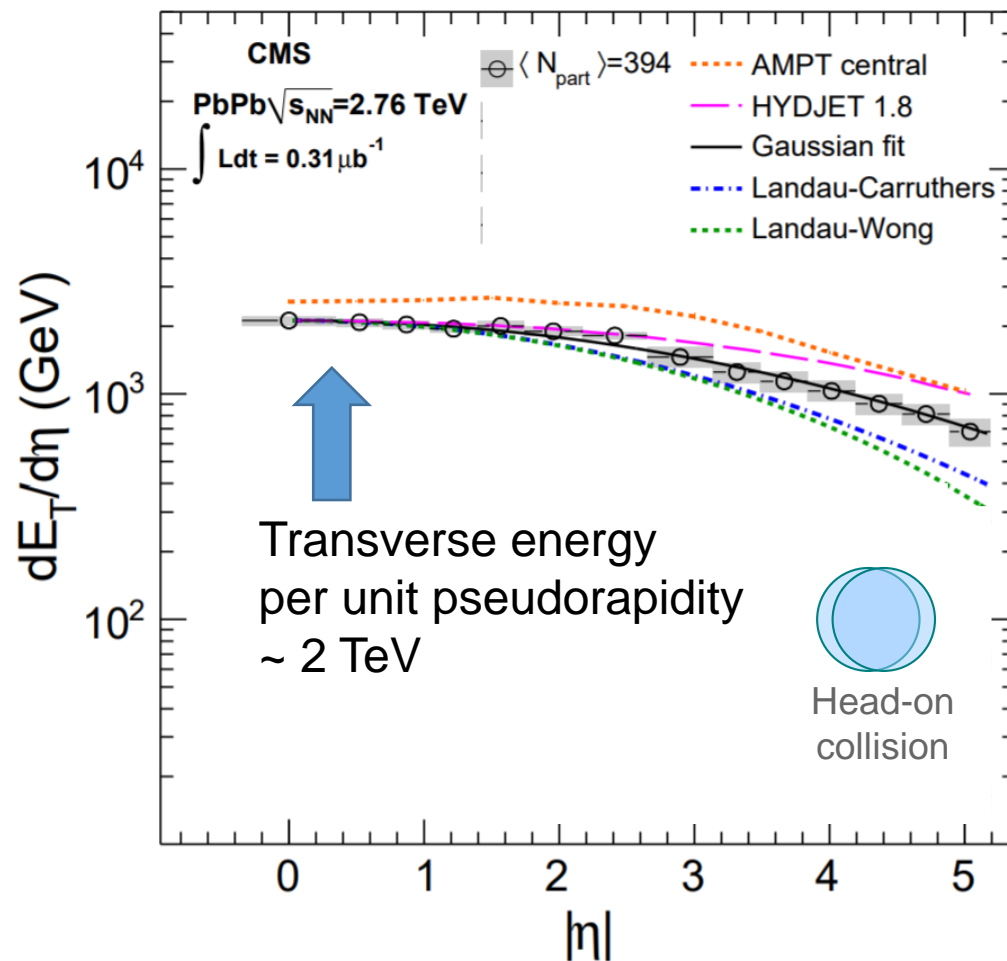
Consider the situation at $t=1$ fm/c:



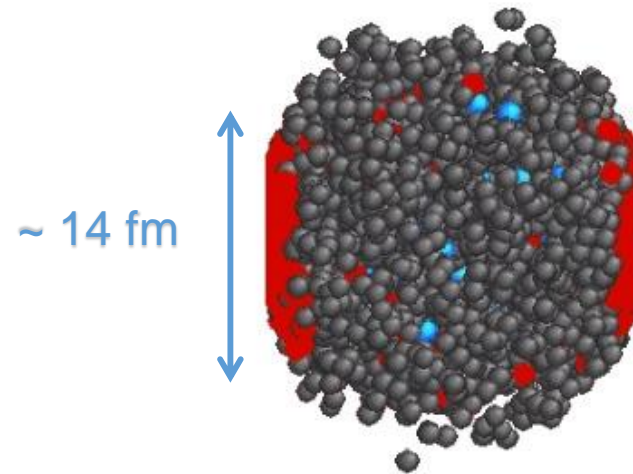
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Consider the situation at $t=1$ fm/c:



Energy density of this medium at this time:
 $2 \text{ TeV} / [\pi (7 \text{ fm})^2 (1 \text{ fm})] = 13 \text{ GeV} / \text{fm}^3$

Hadron: $\sim 0.5 \text{ GeV} / \text{fm}^3$

$>10\times$ denser than the proton!

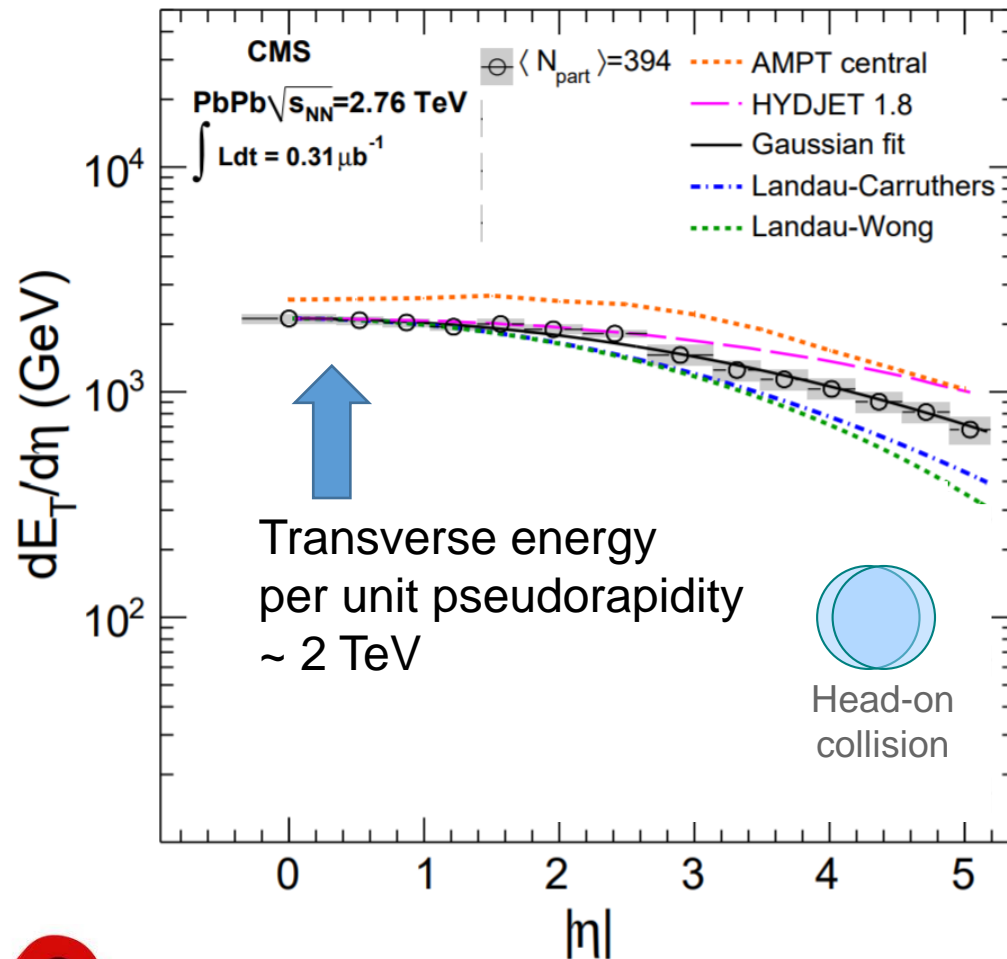
At **early time** of the collision, the system can not be described by hadrons

PRL 109 (2012) 152303

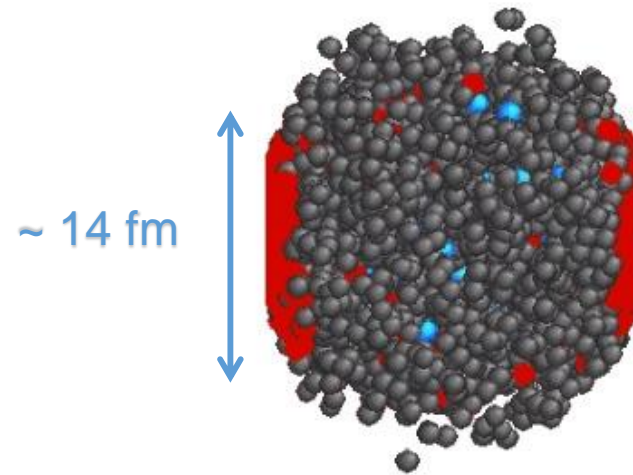
Charged Particle Counting

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$>10\times$ denser than the proton!

Wait! How do we know this system is not just a lot of incoherent NN collisions overlapping each other?

PRL 109 (2012) 152303

Hydrodynamization and Isotropization

Can the debris be described by hydrodynamics?

Hydrodynamization

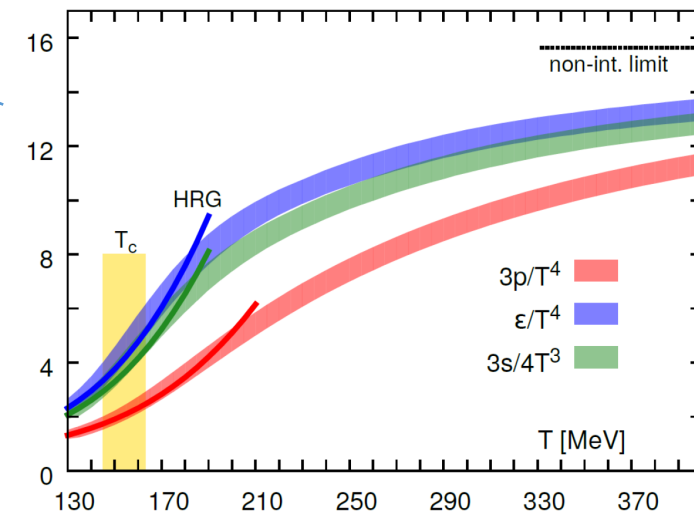
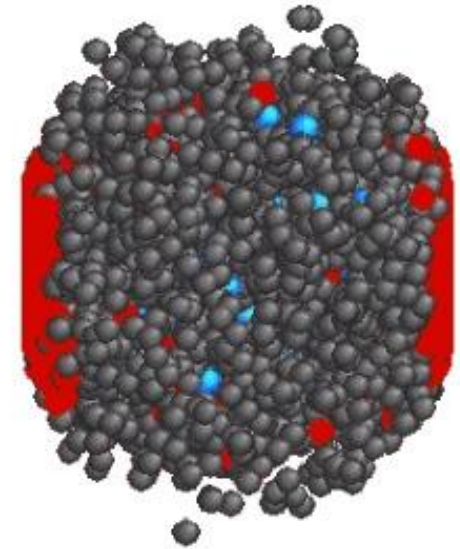
$$T_{\mu\nu} = \varepsilon u_\mu u_\nu + p[\varepsilon] \Delta_{\mu\nu} - \eta[\varepsilon] \sigma_{\mu\nu} - \zeta[\varepsilon] \Delta_{\mu\nu} \nabla_\mu u^\mu + \mathcal{O}(\partial^2), \text{ where}$$
$$\sigma_{\mu\nu} = \Delta_{\mu\alpha} \Delta_{\nu\beta} (\nabla^\alpha u^\beta + \nabla^\beta u^\alpha) - \frac{2}{3} \Delta_{\mu\nu} \Delta_{\alpha\beta} \nabla^\alpha u^\beta,$$
$$\Delta_{\mu\nu} = g_{\mu\nu} + u_\mu u_\nu,$$

QCD come in here

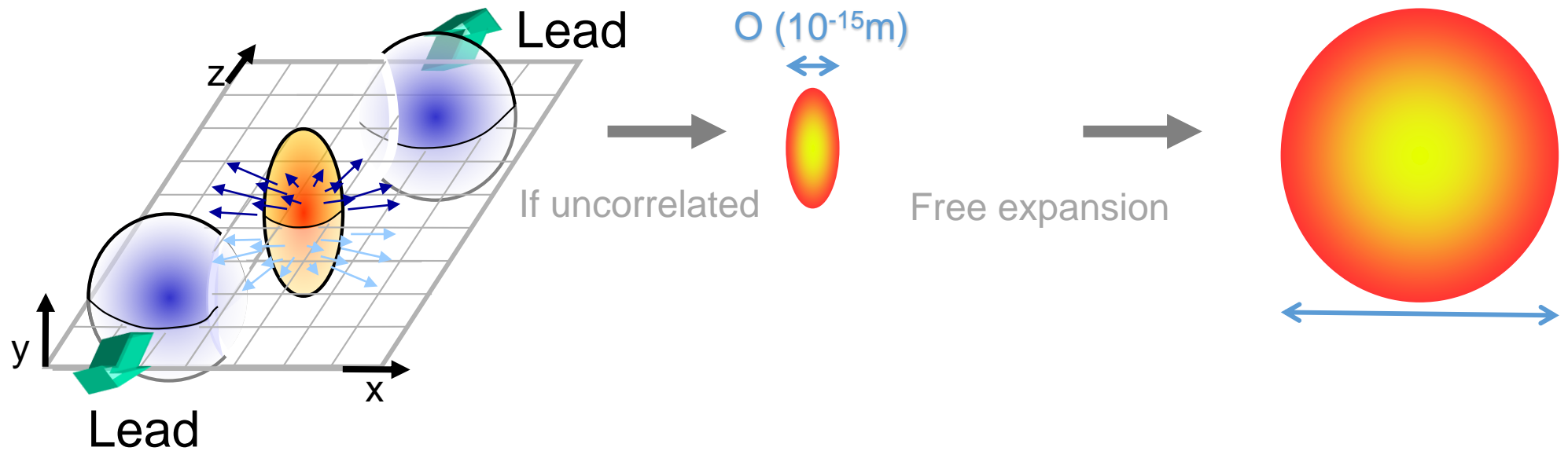
hydrodynamic evolution equations, $\nabla_\mu T^{\mu\nu} = 0$

Can the debris isotropize (i.e., isotropic in the local QGP rest frame) and thermalize?

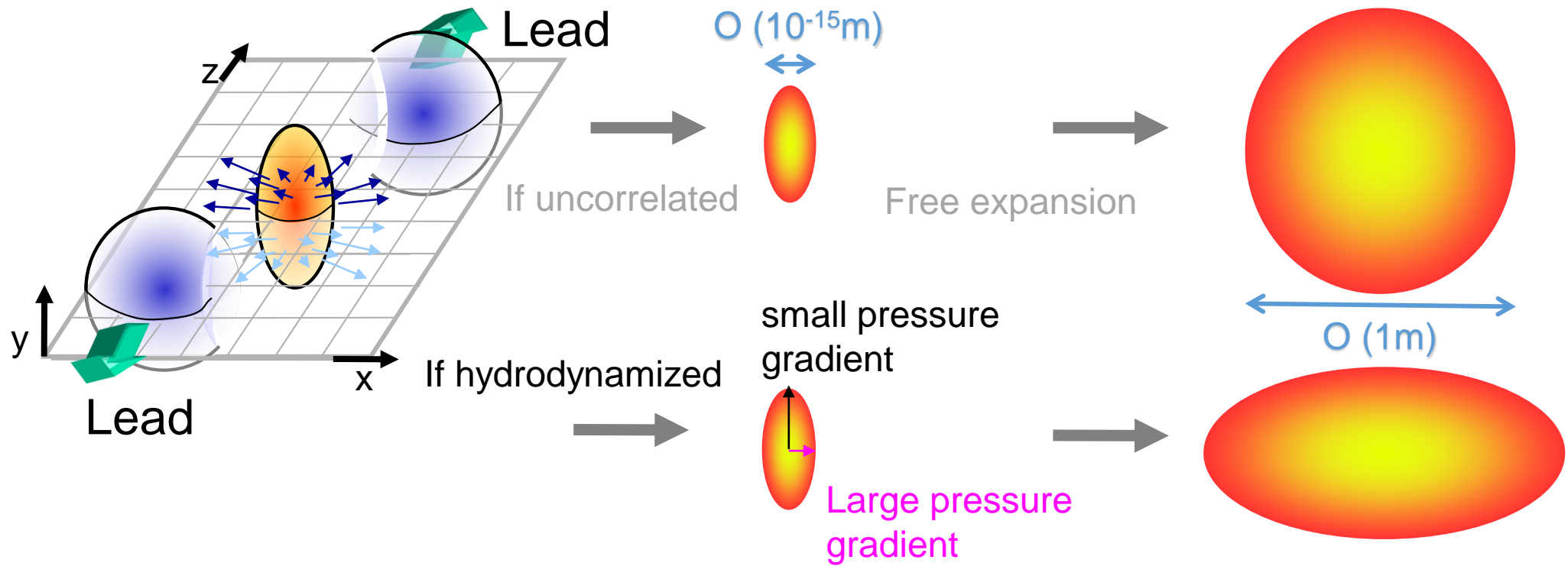
Isotropization



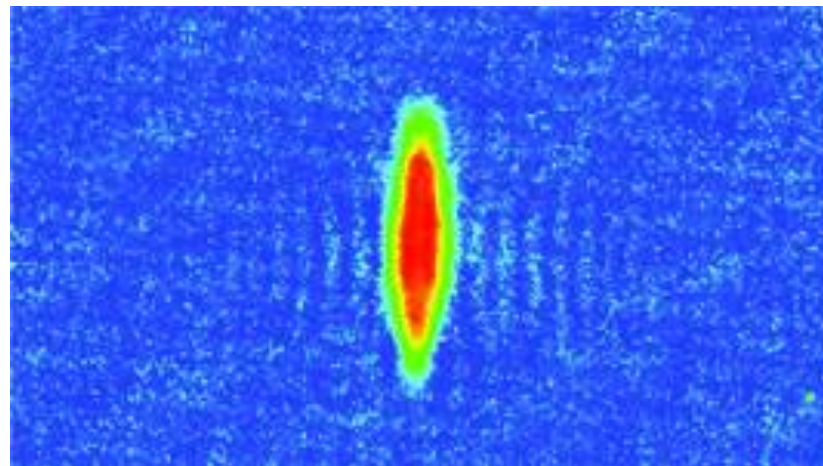
Free streaming vs. strong interacting



Free Streaming vs. Strongly Interacting

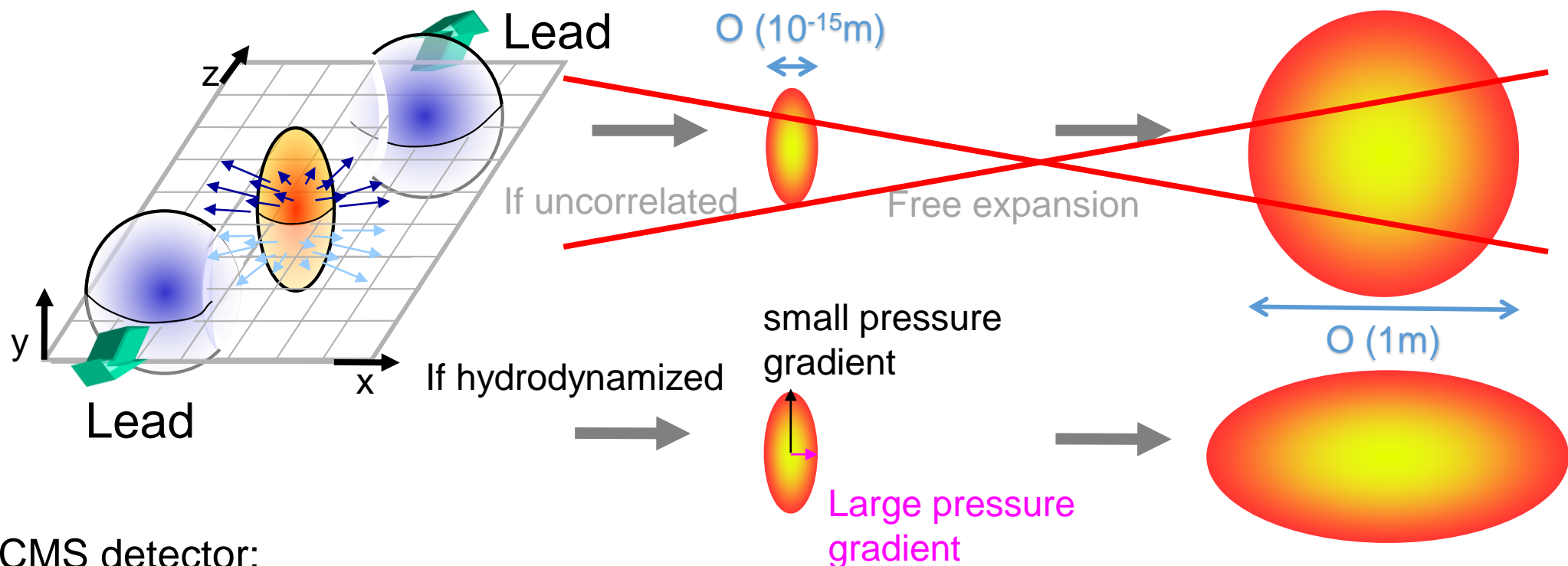


Expansion of Ultra-cold atoms released from trap



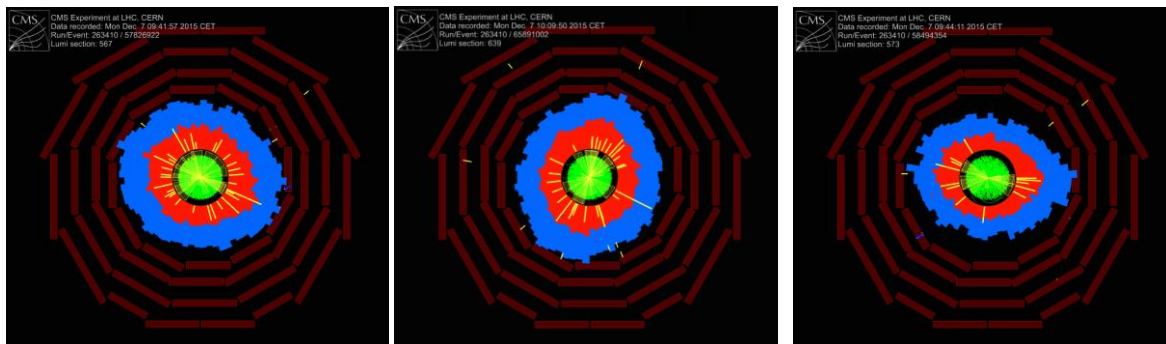
$100 \mu\text{s}$

Free Streaming vs. Strongly Interacting

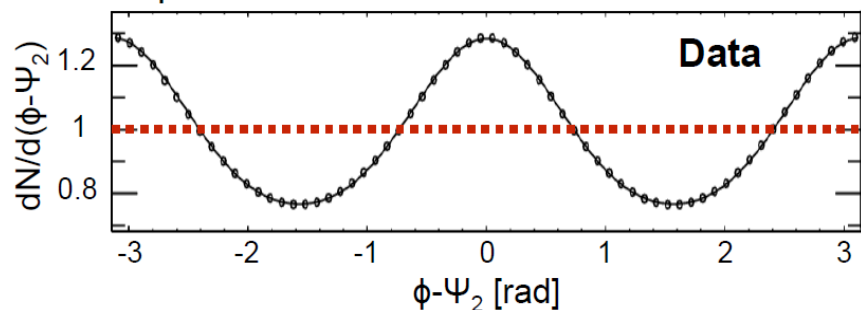


CMS detector:

Electromagnetic, **Hadronic** energy and **charged particles**



Anisotropic azimuthal distribution:

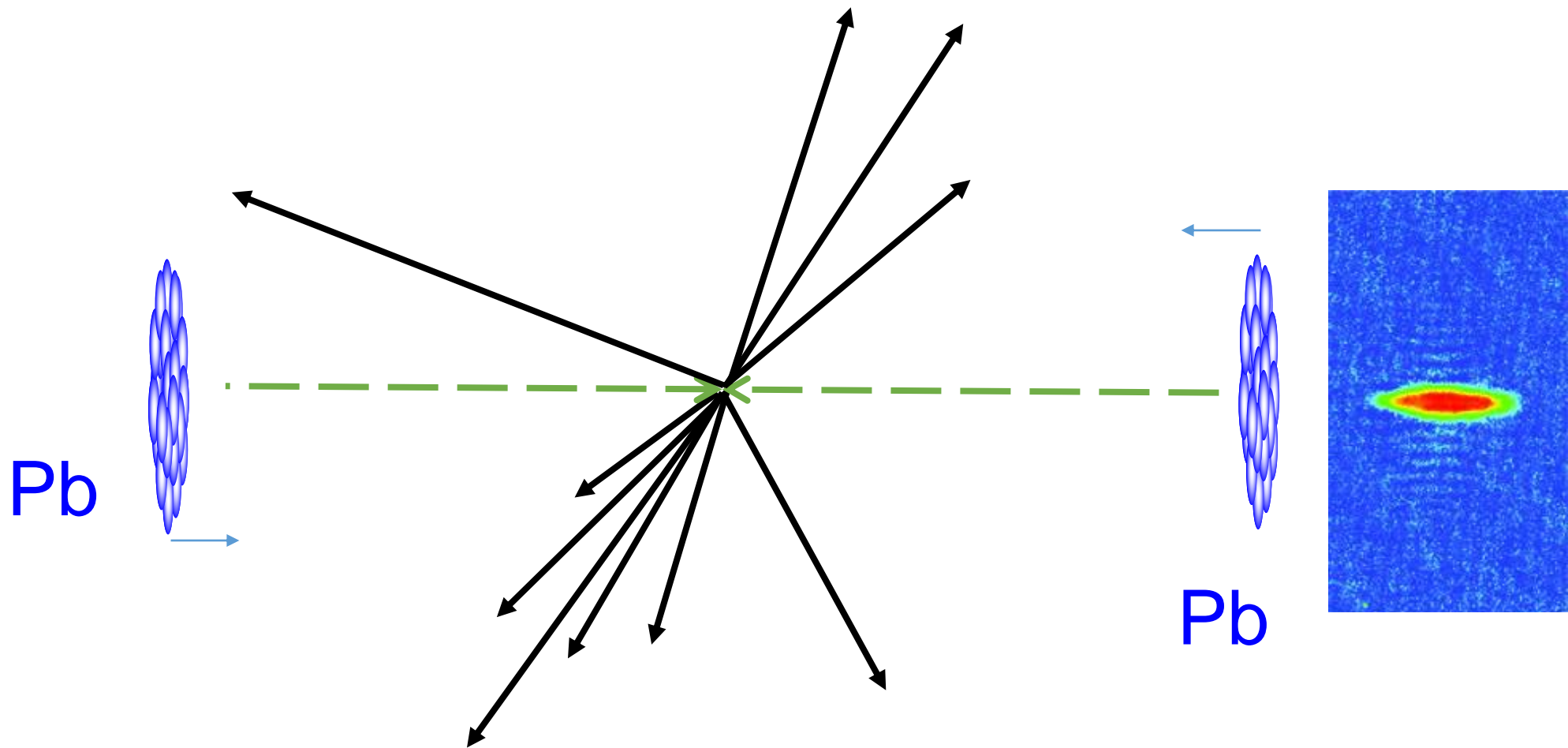


$$dN/d(\phi-\Psi_2) \sim 1 + 2v_2 \cos[2(\phi-\Psi_2)] \quad v_2: \text{elliptic flow}$$

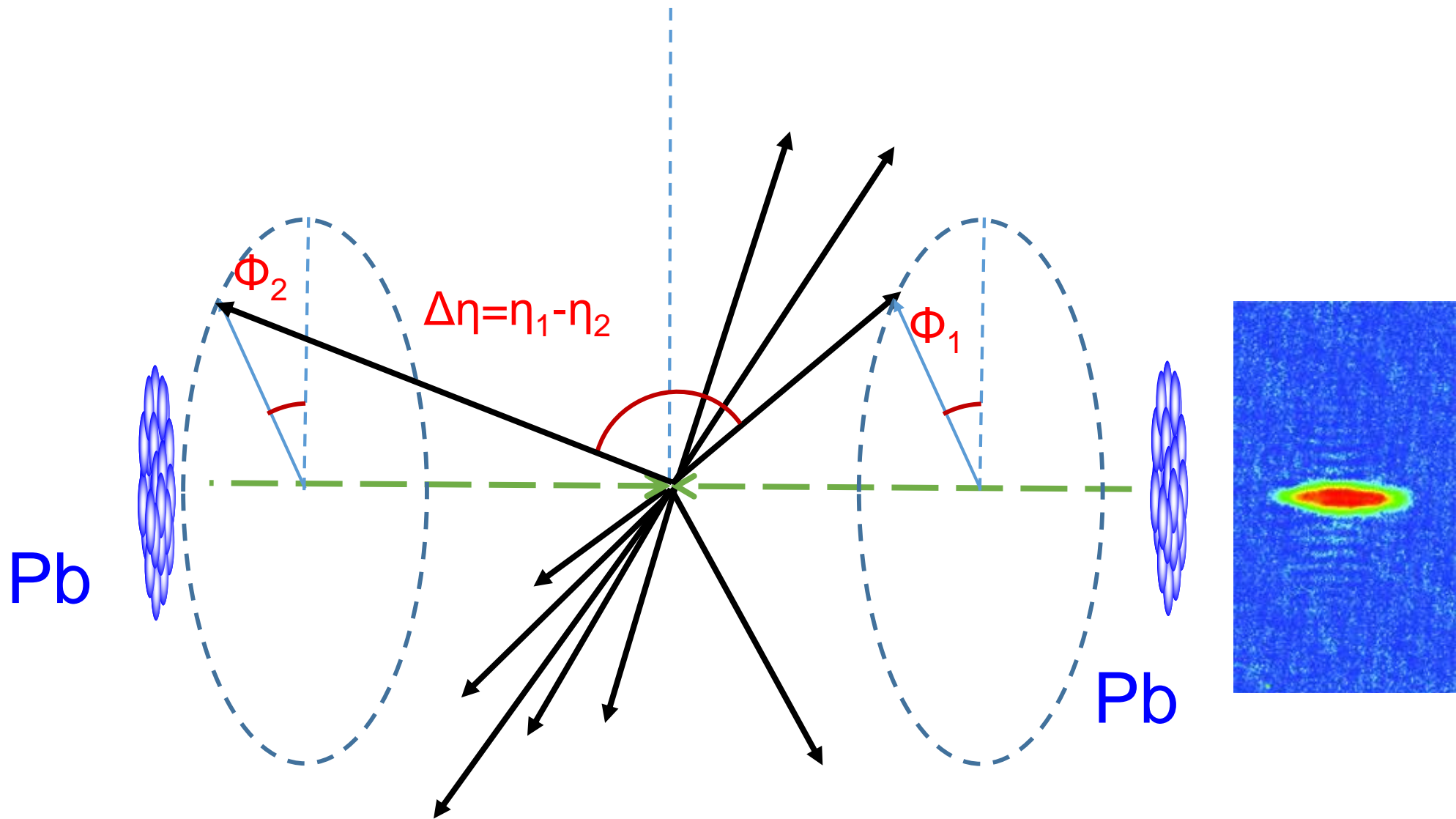
Collective motion is observed in the particle azimuthal distributions!! → **Early hydrodynamization**

Particle azimuthal distributions

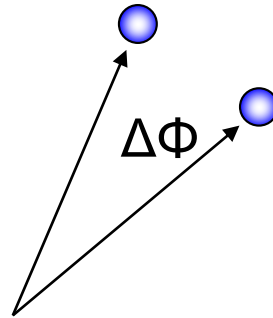
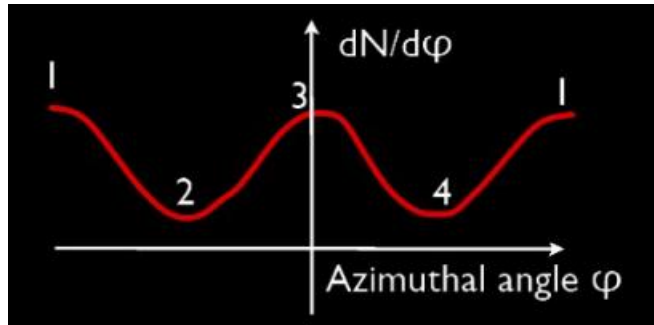
Particle Angular Distribution



Particle Angular Distribution



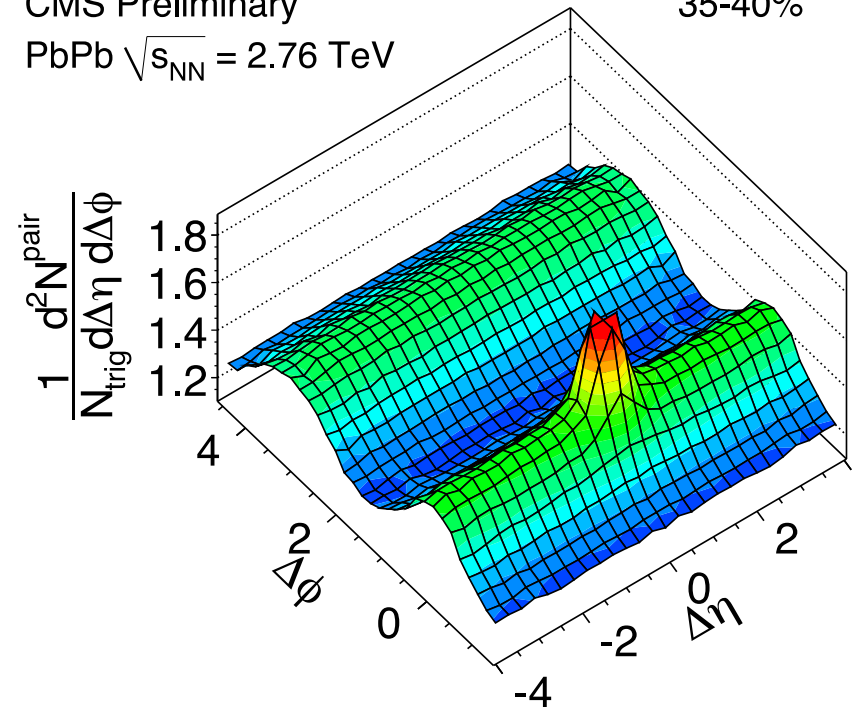
Have we created the Quark-Gluon Plasma?



Azimuthal angle difference
between two charged particles

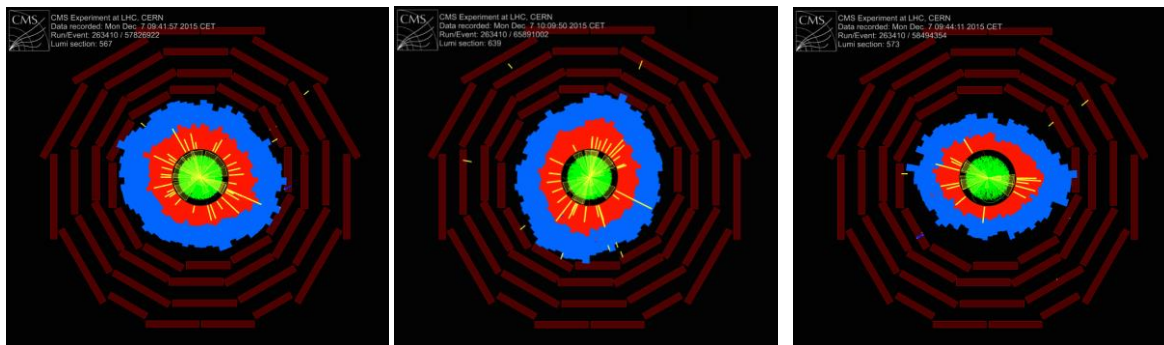
CMS Preliminary
PbPb $\sqrt{s_{NN}} = 2.76$ TeV

35-40%



Particle azimuthal distributions:

$$dN/d\phi \propto 1 + 2 \mathbf{V}_2 \cos(2\phi - \phi_0)$$



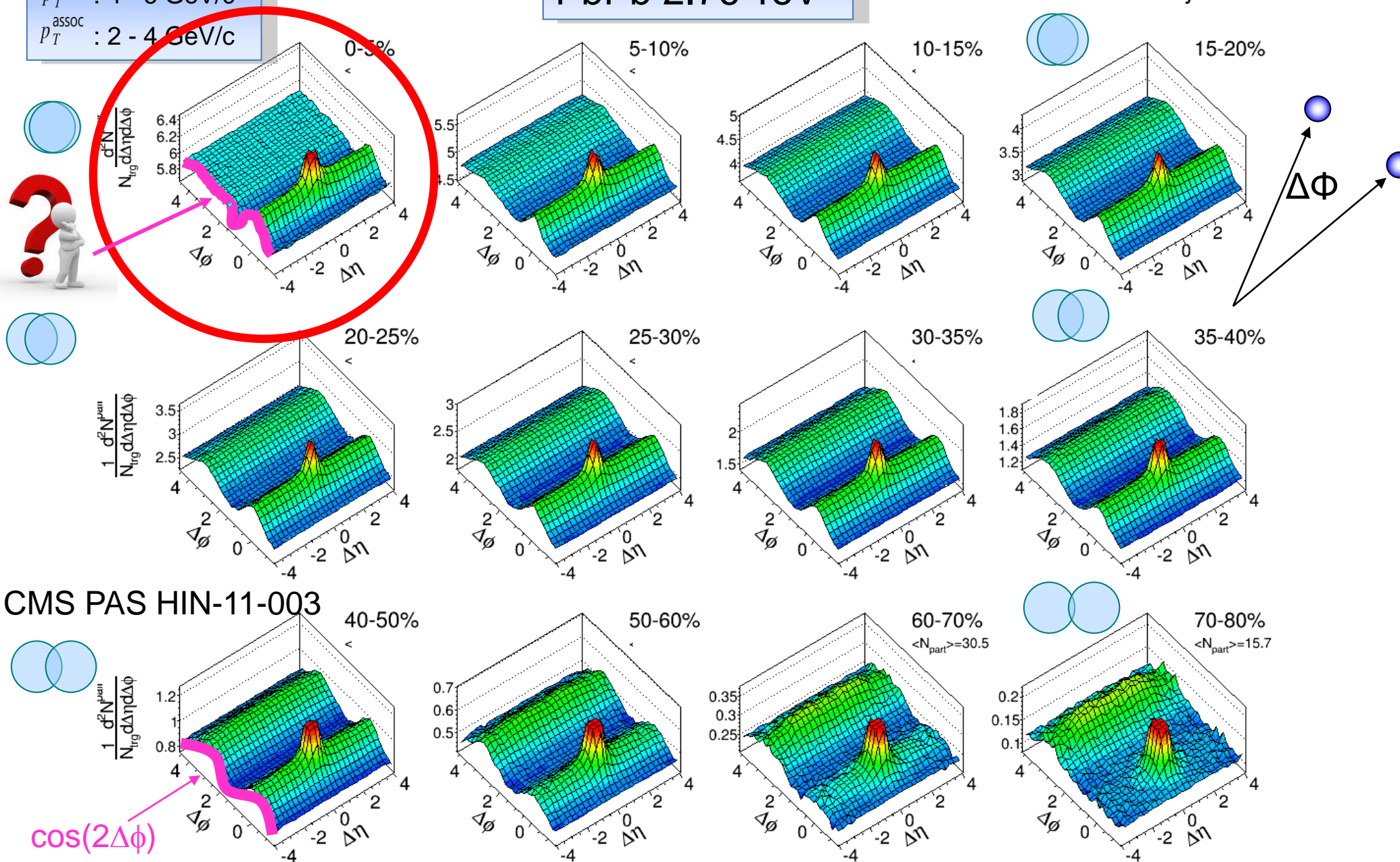
Collective motion is seen in the particle
azimuthal distributions → Early hadrodynamization

Two-particle correlation: Centrality dependence in PbPb

PbPb 2.76 TeV

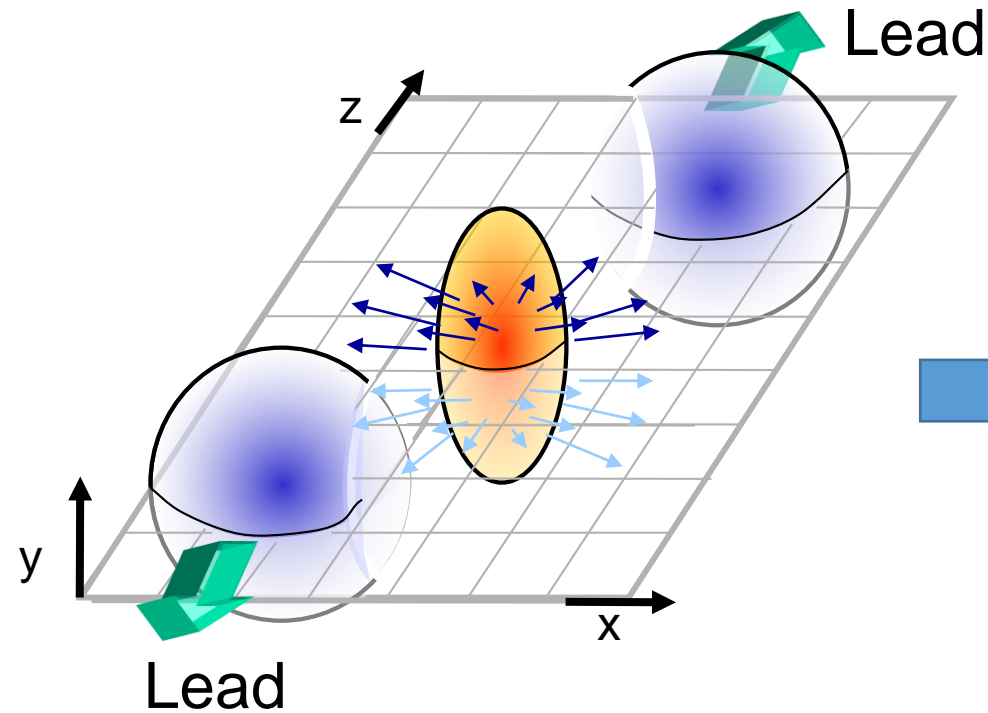
$p_T^{\text{trig}} : 4 - 6 \text{ GeV/c}$
 $p_T^{\text{assoc}} : 2 - 4 \text{ GeV/c}$

CMS Preliminary

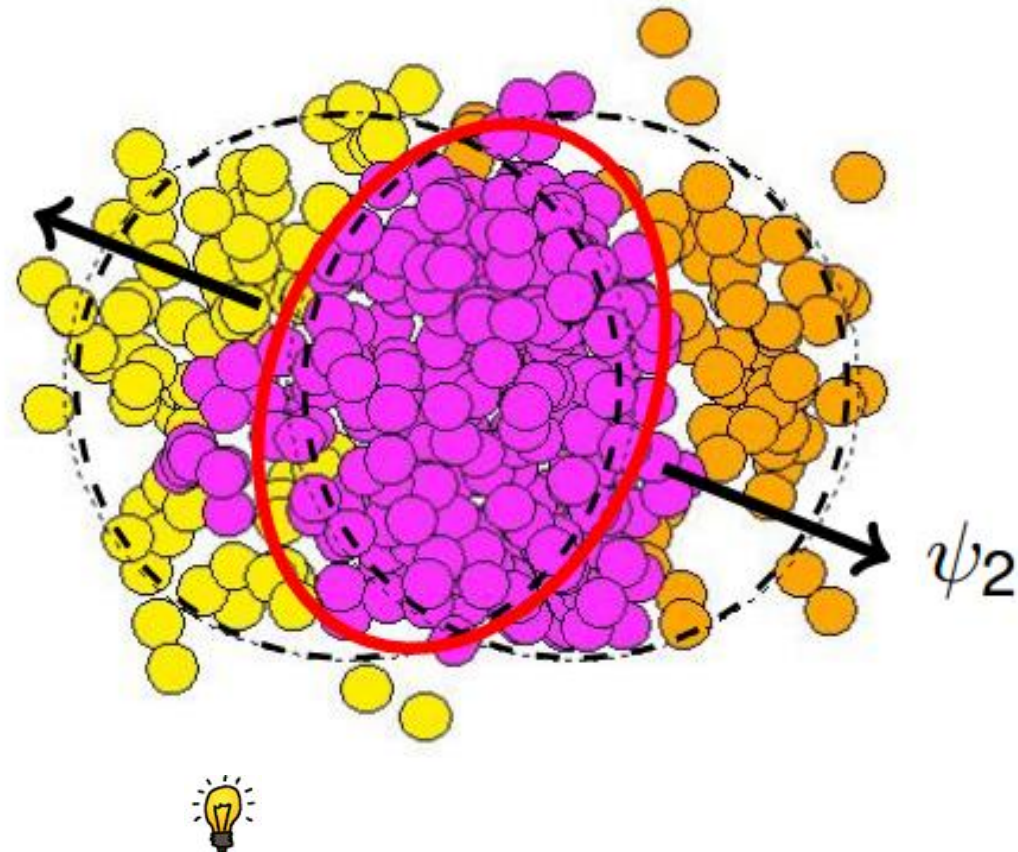


Initial Condition

“Classic view”



“Modern view”

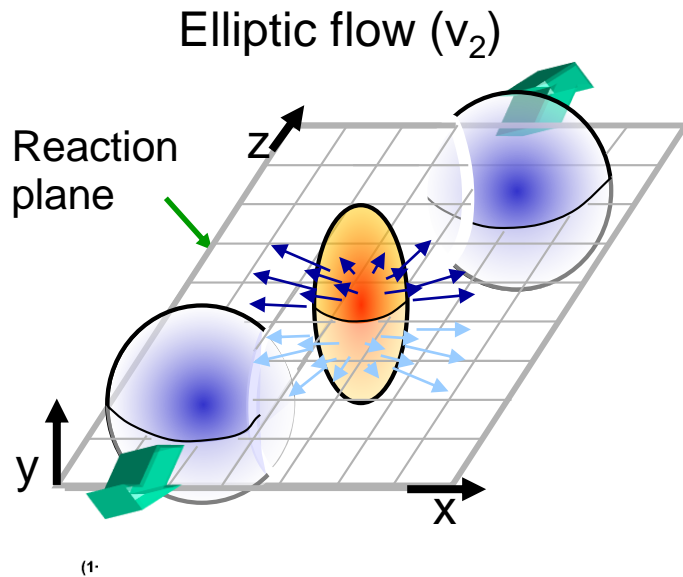


B. Alver & G. Roland

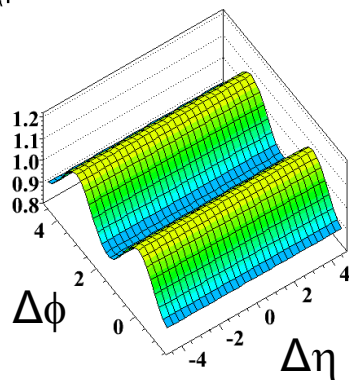
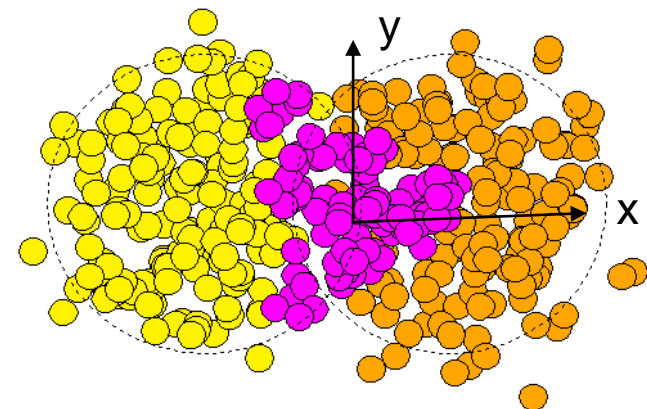
*Phys. Rev. C*81:054905, 2010

Fourier Decomposition

It was recently realized that the ridge may be induced just by higher order flow terms ($v_2, v_3, v_4, v_5, \dots$)



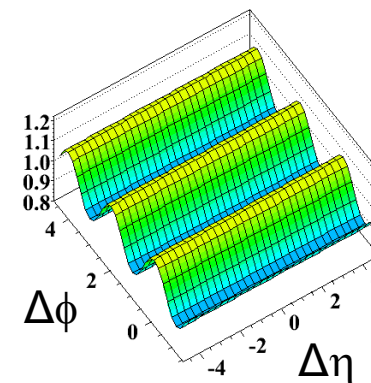
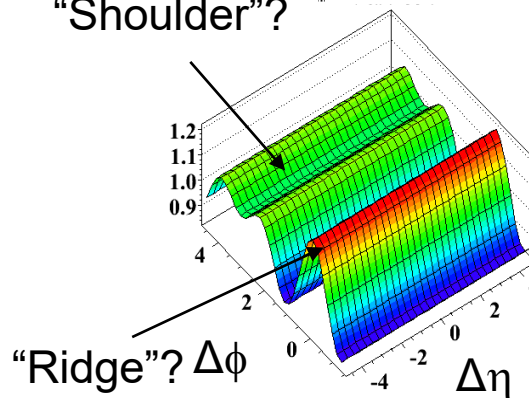
Triangular flow (v_3) from event-by-event fluctuation



$$\sim V_2 \cos(2\Delta\phi)$$

Add V_2 and V_3 together

“Shoulder”?



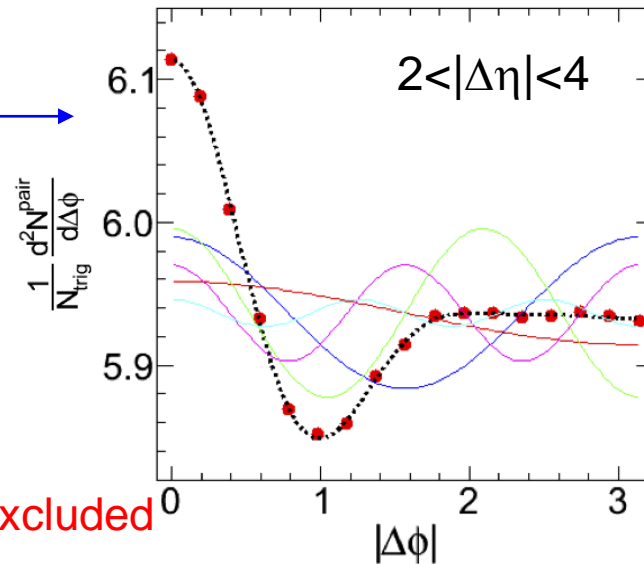
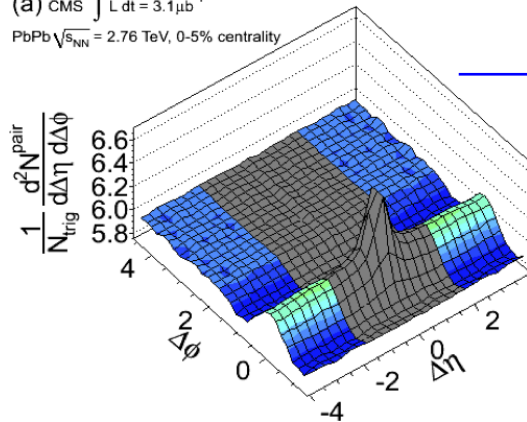
$$\sim V_3 \cos(3\Delta\phi)$$

Fourier Decomposition of $\Delta\phi$ correlations

Fourier decomposition

$$\frac{1}{N_{\text{trig}}} \frac{dN^{\text{pair}}}{d\Delta\phi} = \frac{N_{\text{assoc}}}{2\pi} \left\{ 1 + \sum_{n=1}^{\infty} 2V_{n\Delta} \cos(n\Delta\phi) \right\}$$

(a) CMS $\int L dt = 3.1 \mu\text{b}^{-1}$
PbPb $\sqrt{s_{\text{NN}}} = 2.76 \text{ TeV}$, 0-5% centrality



Flow driven correlations:

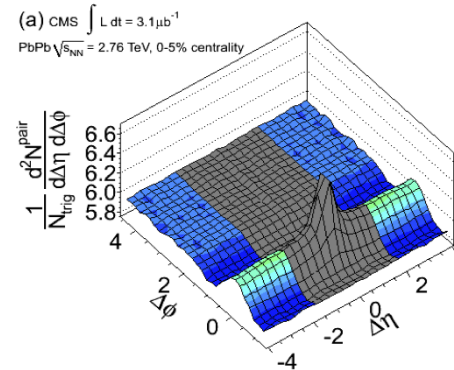
$$V_{n\Delta} \sim v_n^{\text{trig}} \times v_n^{\text{assoc}}$$

(f: Fourier analysis of long-range dihadron correlations)

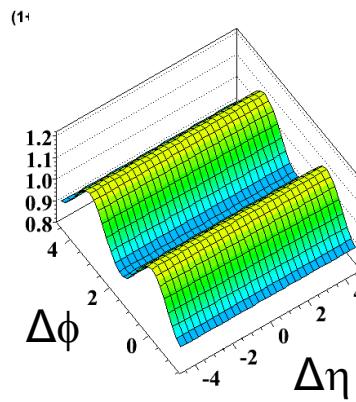
ArXiv:1105.2438
JHEP 1107:076,2011

Short-range non-flow effects excluded

(a) CMS $\int L dt = 3.1 \mu\text{b}^{-1}$
PbPb $\sqrt{s_{\text{NN}}} = 2.76 \text{ TeV}$, 0-5% centrality

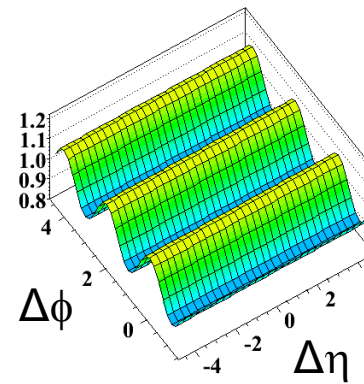


=



$\sim V_2 \cos(2\Delta\phi)$

+



$\sim V_3 \cos(3\Delta\phi)$

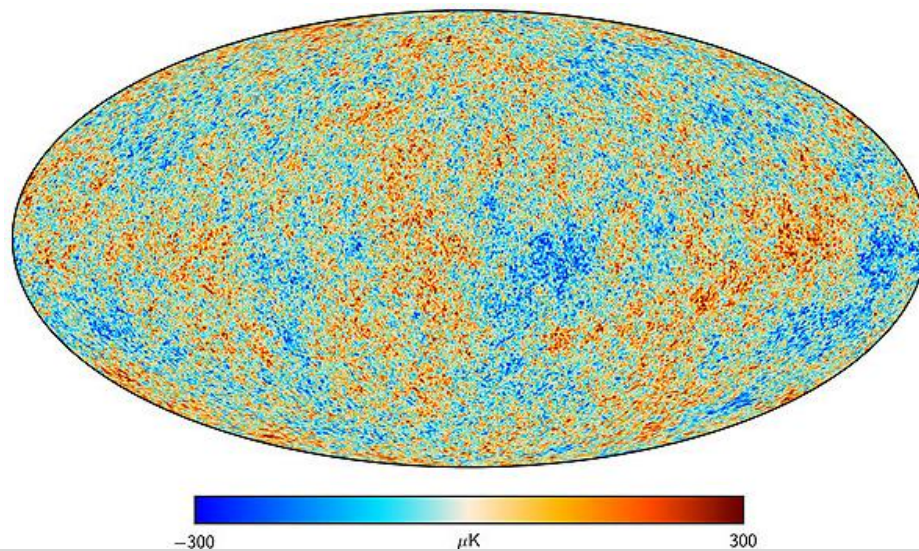
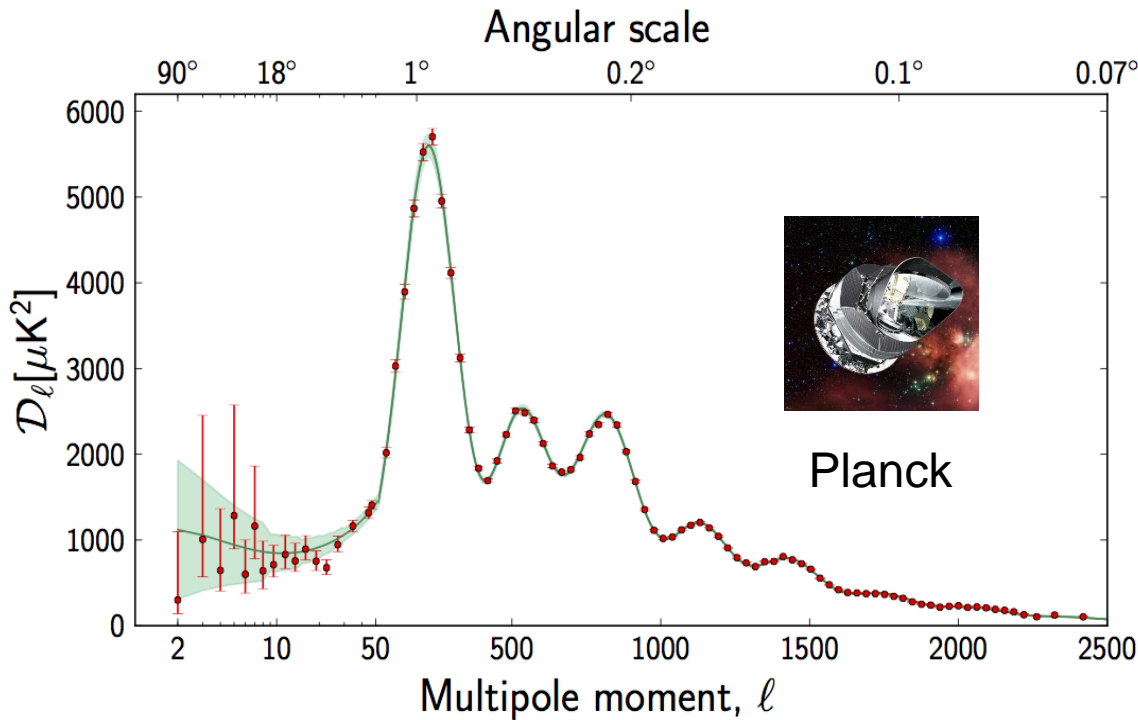
+

...

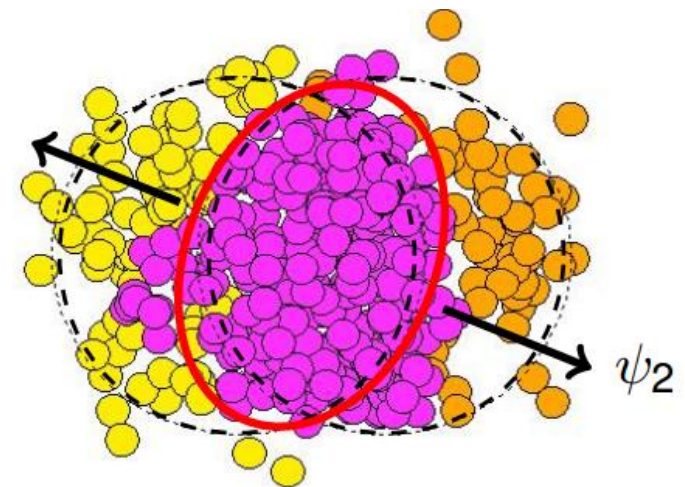
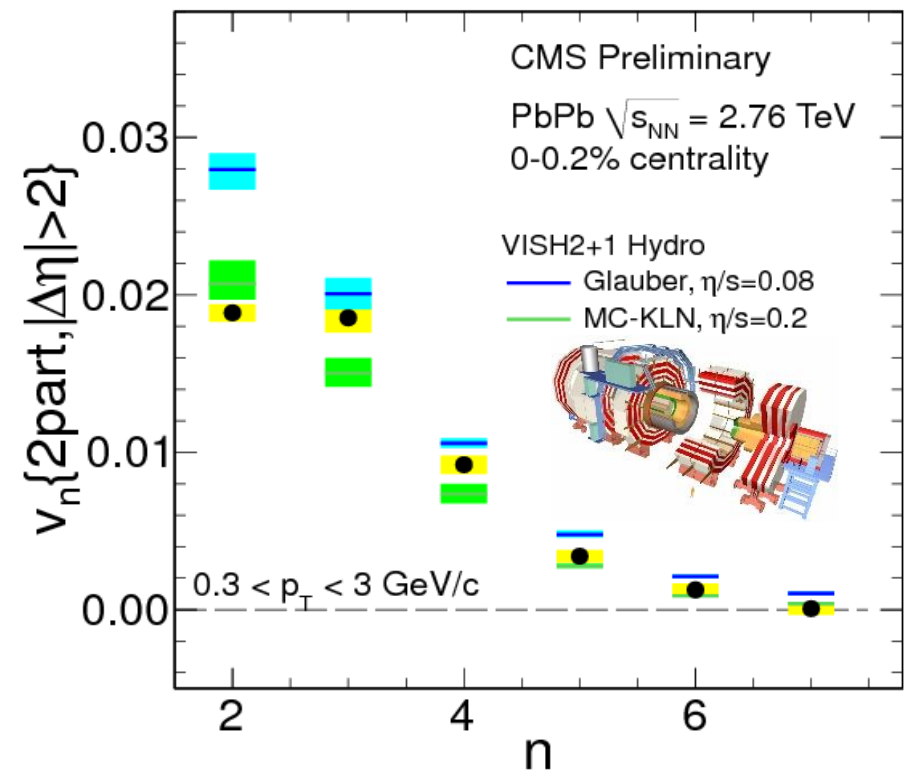
Higher order terms

Density Fluctuation

Power spectrum of the Big Bang

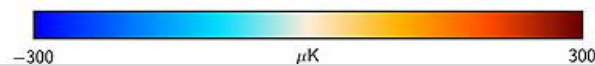
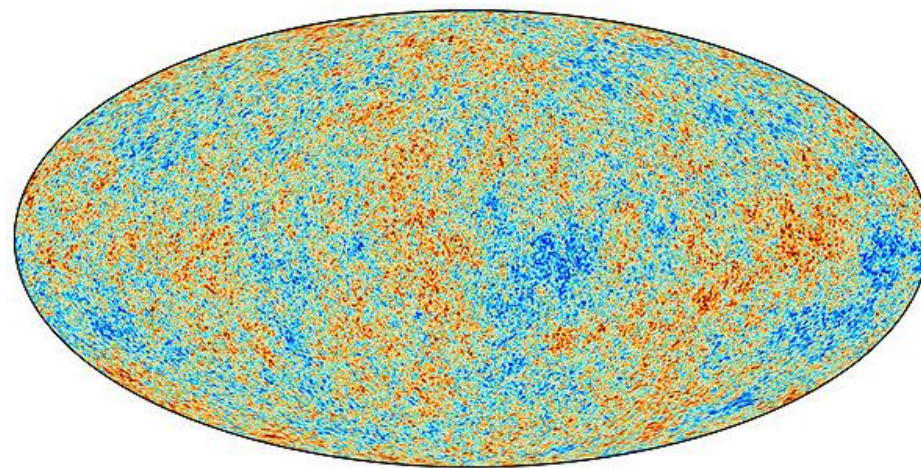
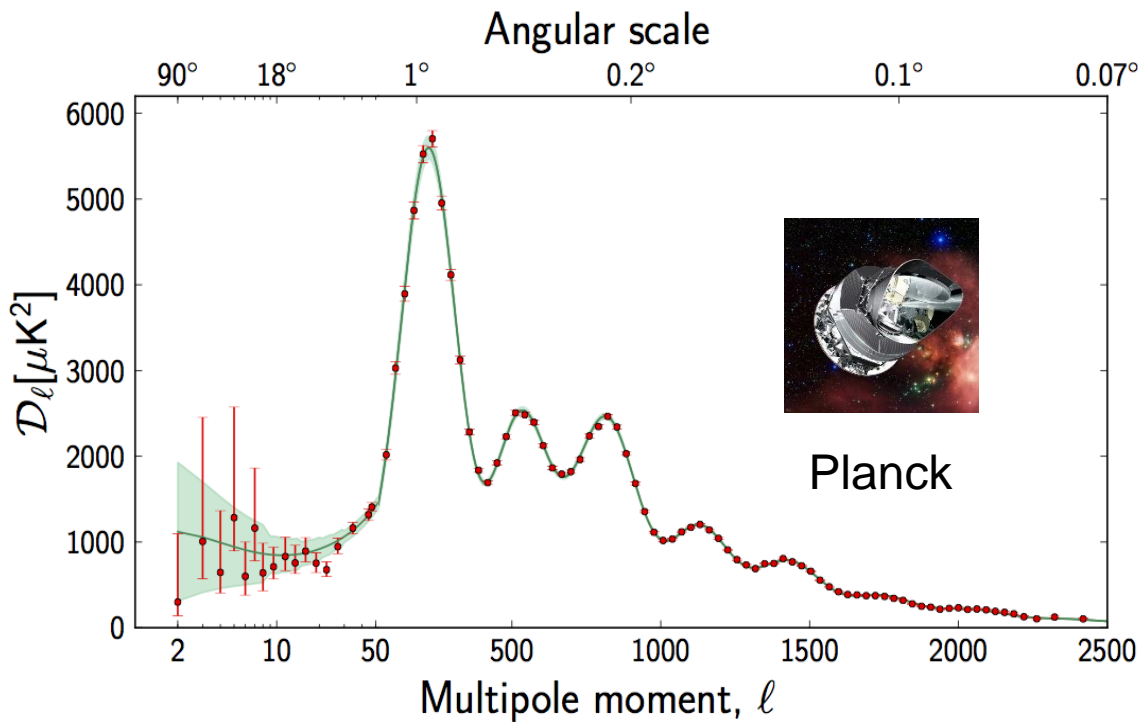


Fourier analysis of many little bangs

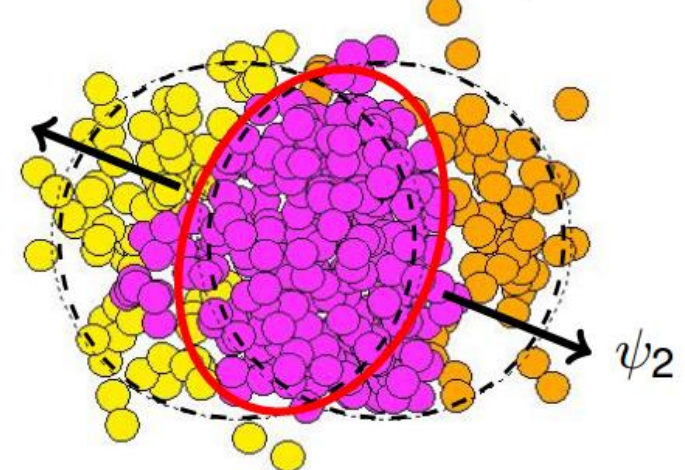
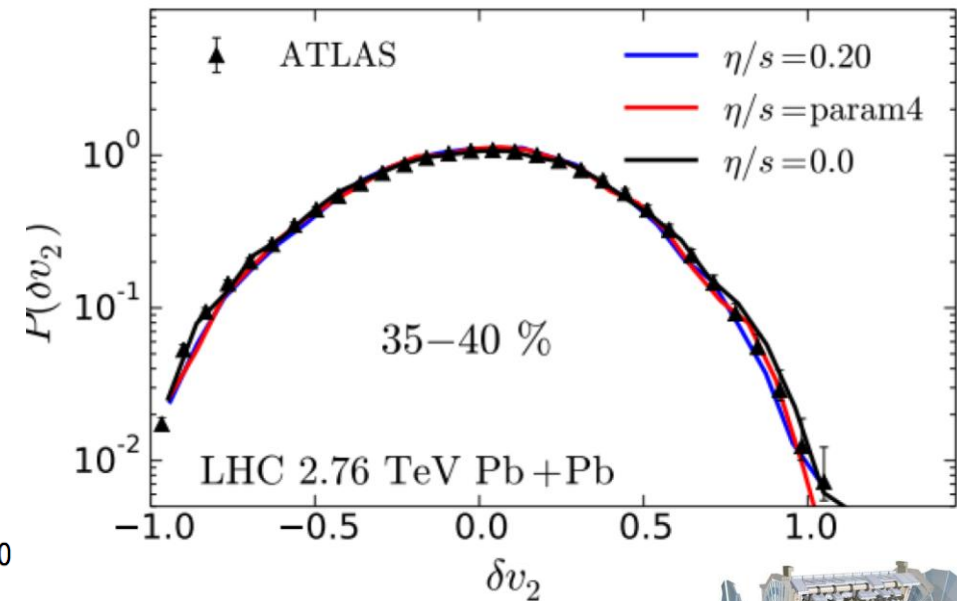


Density Fluctuation

Power spectrum of the Big Bang



Fourier analysis of many little bangs Event by event!!!

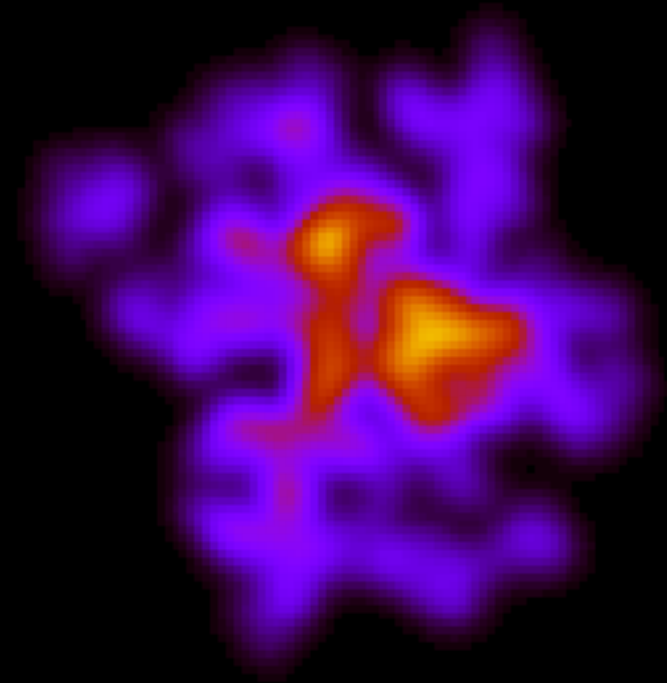
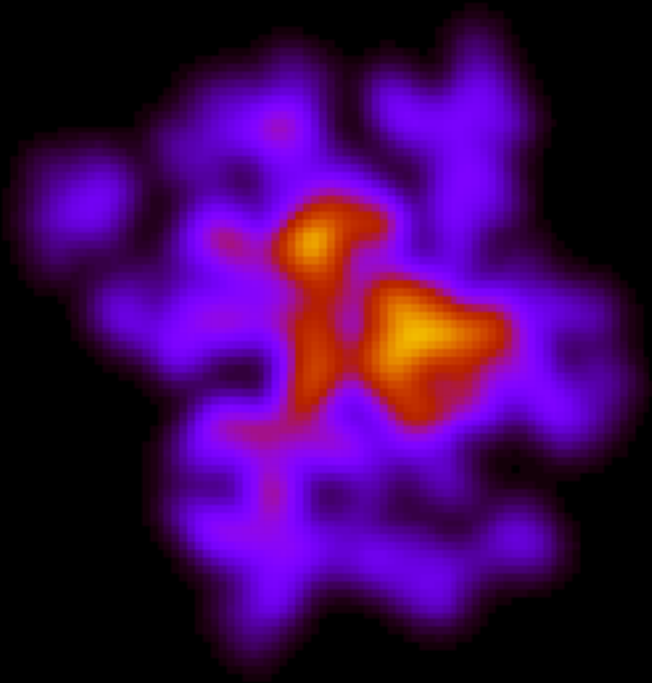


Effect of Shear Viscosity

Animation made by Bjorn Schenke
<https://quark.phy.bnl.gov/~bschenke/>

Ideal hydrodynamics

Viscous hydrodynamics



$t = 0.5 \text{ fm/c}$

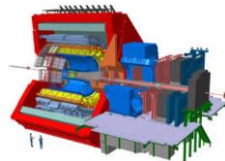
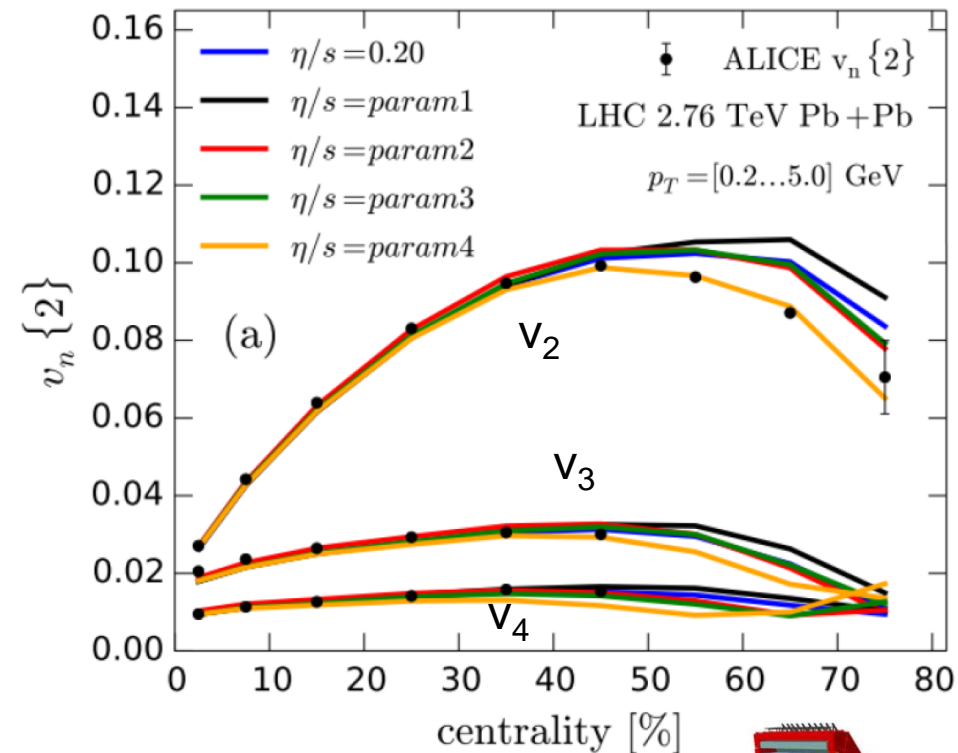
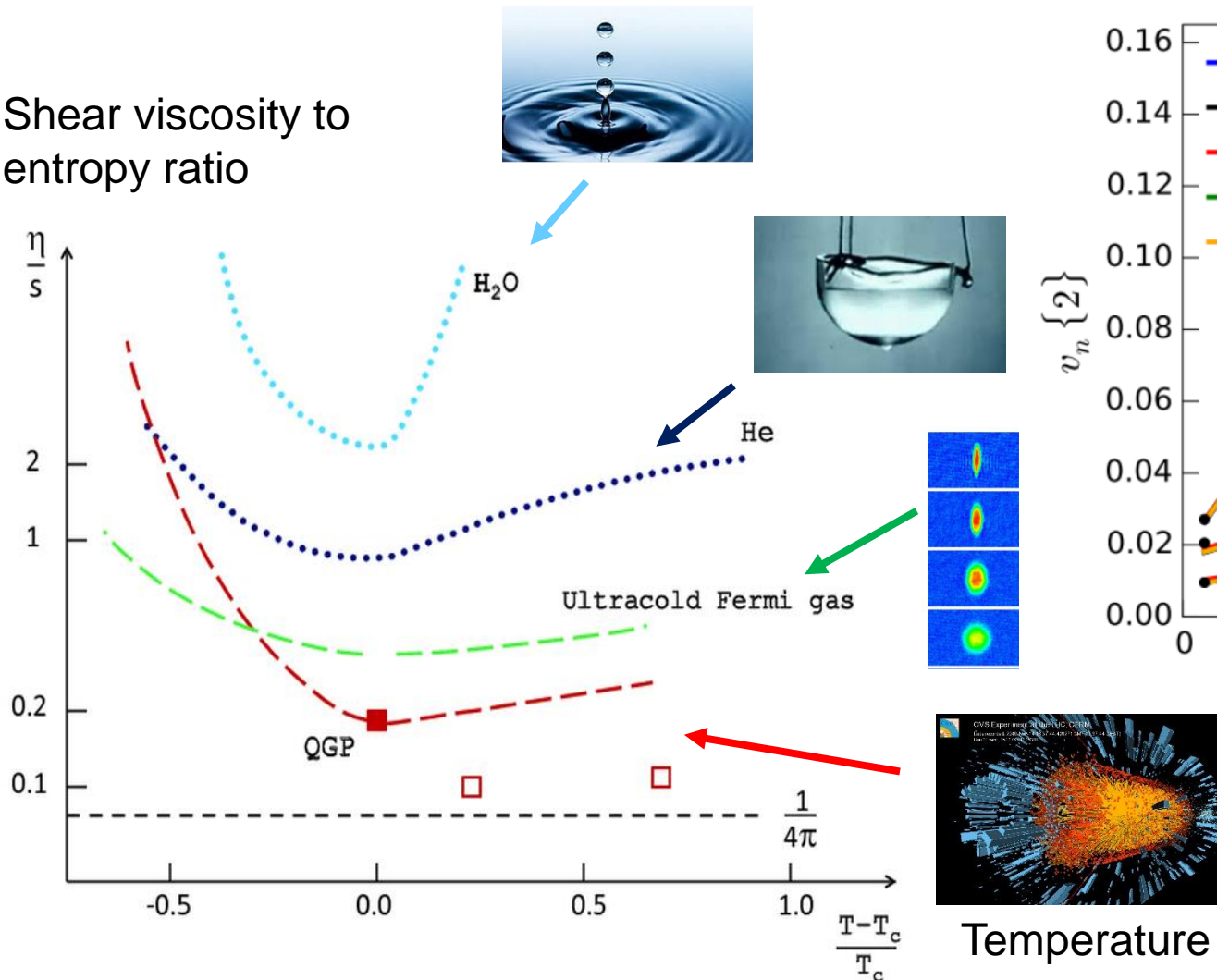
Alver and Roland (MITHIG)
“Collision geometry fluctuation”
PRC82 (2010) 039903

Flow coefficients (v_n^f) vs centrality and viscosity

$$\frac{1}{N_{\text{trig}}} \frac{dN^{\text{pair}}}{d\Delta\phi} = \frac{N_{\text{assoc}}}{2\pi} \left\{ 1 + \sum_{n=1}^{\infty} 2V_{n\Delta} \cos(n\Delta\phi) \right\}$$

Can be explained by
hydrodynamics calculation
+ Initial state fluctuation

Shear viscosity to
entropy ratio



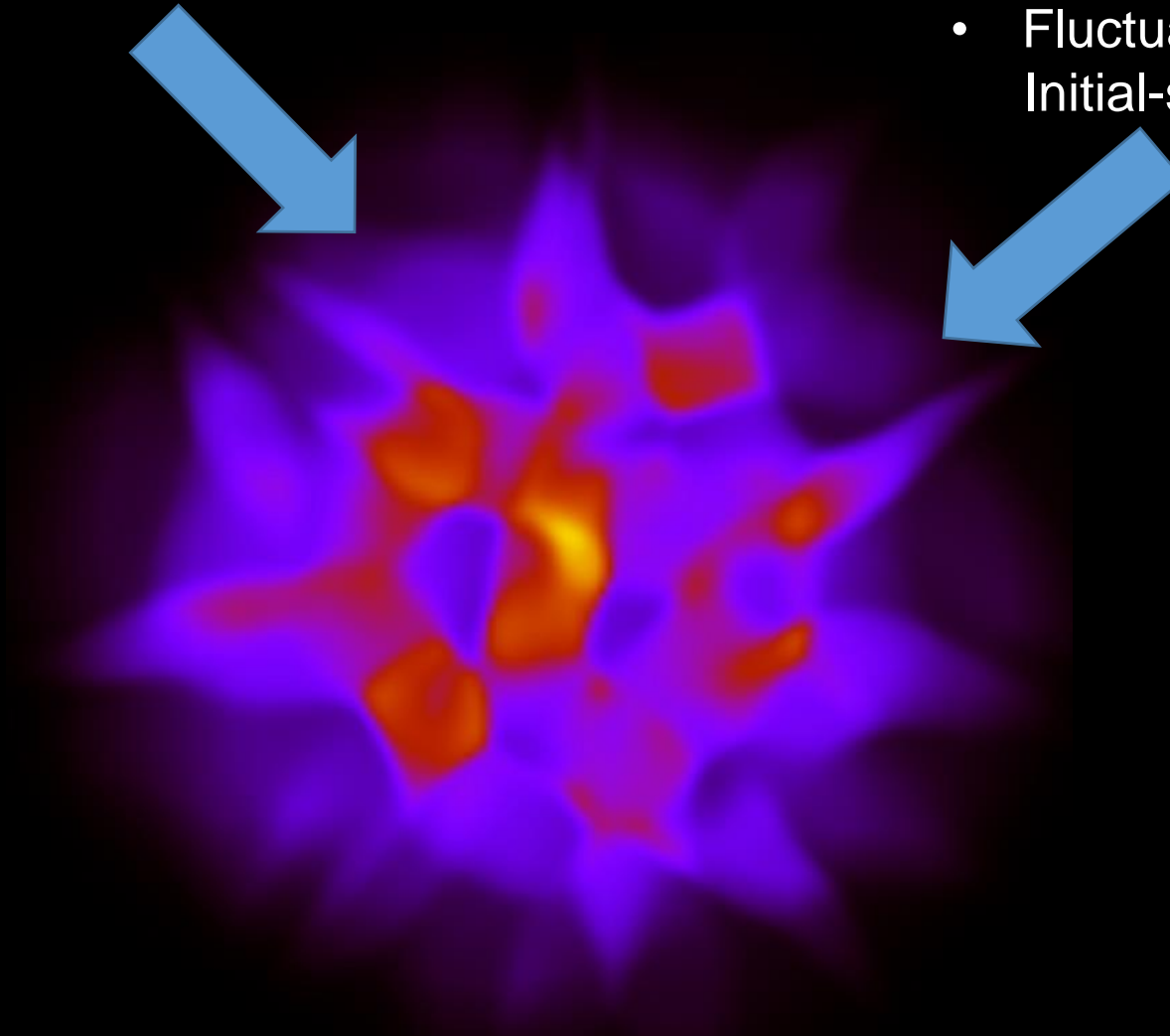
Prepare the Quark Soup

Particle Multiplicity

- Collision impact parameter
- Energy density

Azimuthal anisotropy

- Early thermalization < 1 fm/c
- Shear viscosity
- Fluctuation of v_N coefficients:
Initial-state geometry fluctuation



What have we learned so far?

- The produced medium is strongly interacting
- Based on hydrodynamics models with the equation of state from lattice QCD calculation:
 - Initial state fluctuations (leads to elliptic flow (v_2) triangular flow (v_3) ...)
 - The system has to hydrodynamize extremely fast start from around **0.2 – 1 fm/c** to produce the observed azimuthal anisotropy
 - Isotropization / complete thermalization could take as long as **a few fm/c**
 - Both shear viscosity (η) and the entropy (s) of the produced medium are **HUGE**
 - The extracted η/s ratio is **extremely small** (smaller than any known materials), close to ideal fluid

What we still don't know very well

- Why is the system hydrodynamize so fast? How does the strongly interacting medium emerge from an asymptotic free theory?
- What is the initial transverse fluid velocity and the role of the pre-hydrodynamization phase? When is QGP formed?
- How does the QGP hadronize?
- Can we see quasi particles (quarks and gluons) in the Quark Gluon Plasma ? Can we see medium response?
- Are we completely wrong?!



What we still don't know very well

- Why is the system hydrodynamize so fast? How does the strongly interacting medium emerge from an asymptotic free theory?

Start from “un-thermalized” objects and see how they are hydrodynamized / thermalized in the Quark Soup

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Final a way to “turn off” temporarily the interaction with QGP**

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Shoot colored objects through the QGP



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- Are we completely wrong? **Extract medium properties with probes**