The 7th Asian Triangle Heavy-Ion Conference(ATHIC-2018)

Rotation-induced chiral condensate

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Outline

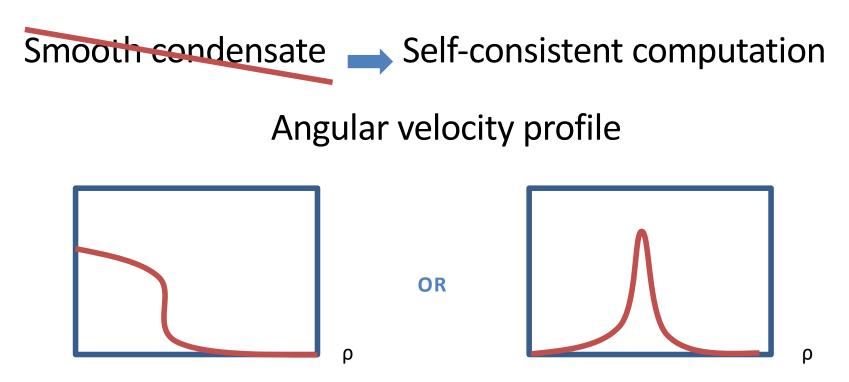
- Motivation
- Introduction of the BdG framework
- Inhomogeneous condensate
- Phase diagram
- Conclusions and outlook

Motivation

- Rotation or vortex is a common phenomenon in cold atom, heavy ion collision and neutron star.
- Non-triviality by considering the limit of speed and centrifugal effects.
- Requiring a unified framework to deal with finite size and inhomogeneous effects.

Assumptions and scheme

- Speed limit: $v = 2 r^{-1} \int_0^r d\rho \rho \omega(\rho) < 1.$
- Space dependence or finite size.

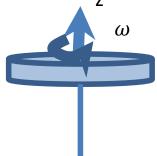


NJL model with rotation

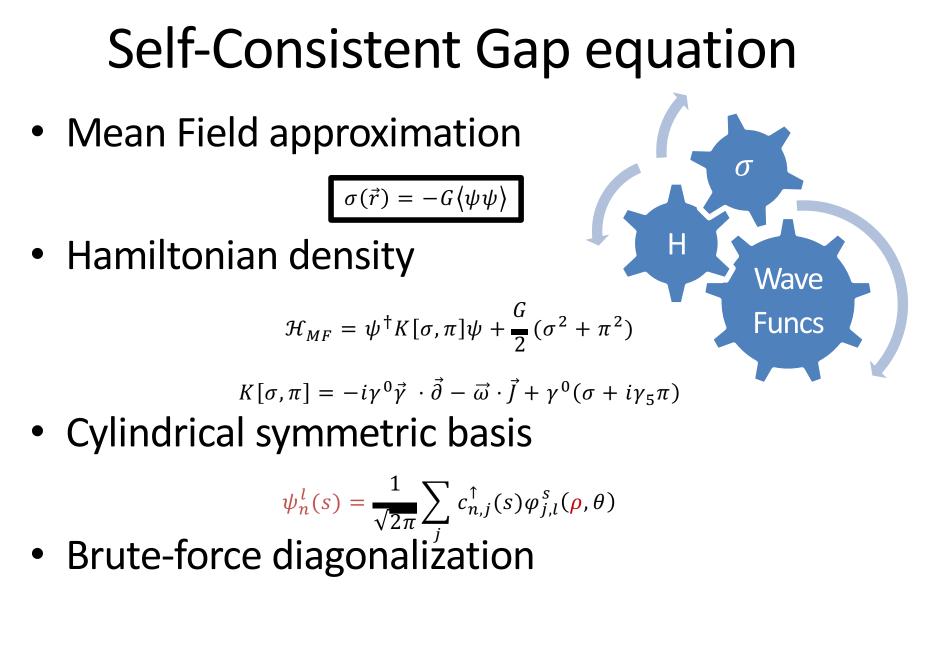
• Considering a rotating 2+1 D system at angular

velocity ω . In its rest frame

$$H = H_0 - \vec{\omega} \cdot \vec{J}$$

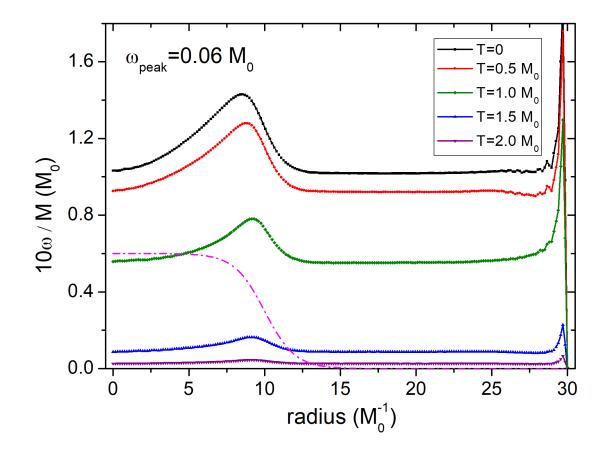


- Renormalizable to avoid subtle regularization.
- Eigen states are unavailable for a general $\omega(\rho)$.



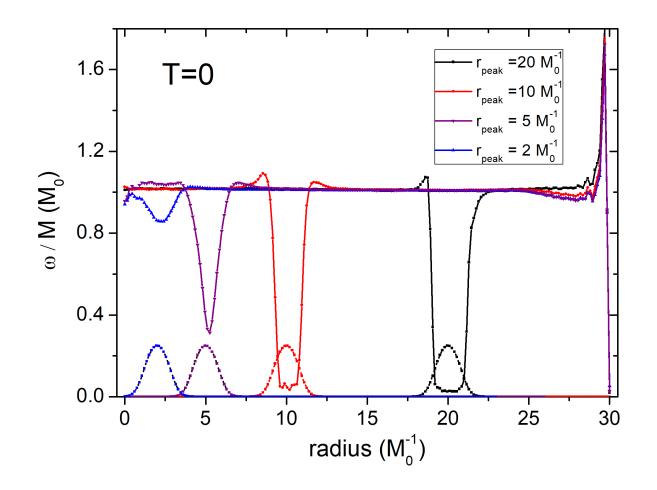
Centrifugal Effects

Condensate bump @ speed slope

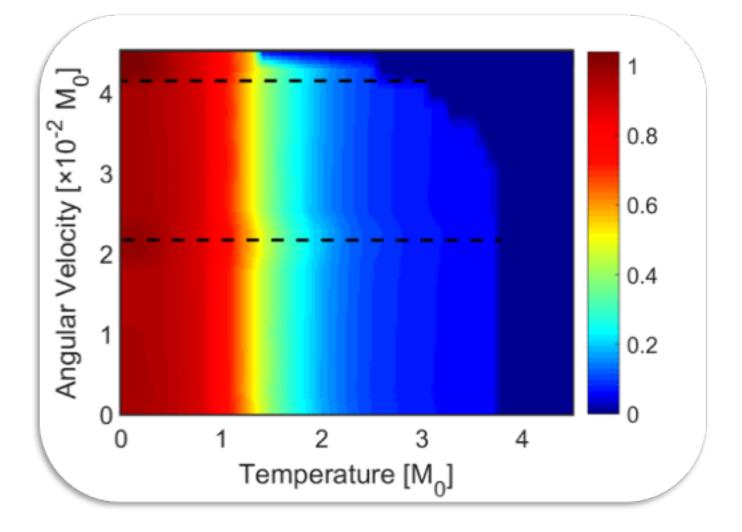


Inhomogeneous rotating vacuum

Condensate suppression @ large speed



Phase diagram in Finite systems



Conclusions & Outlook

- Fermion pairing studied with 2+1D NJL model in Self-consistent BdG framework.
- Centrifugal effects and Chiral condensate suppression are observed in rotating systems.
- T- ω phase diagram is roughly similar, but NOT equivalent to T- μ diagram in finite systems.
- More novel states(vortex states) are worth to exploring.

Thank you for your attention!

Cookbook of PT study

- Choose the model
- Write down the interaction (Lagrangian density)
- Order parameter(σ) & Approximation(Mean field) <
- Calculate the free energy $Z = \int D\phi \ e^{-S[\phi,\sigma]} = e^{-\Omega[\sigma]V}$
- Minimize the free energy
- Gap equation for order parameter $f(\sigma, T, \mu, ...) \stackrel{\checkmark}{=} 0$

Model choice

- Aiming at the chiral condensate.
- A relativistic model is necessary for quark matter.
- A fermionic model is better than a bosonic one.
- ✓ Nambu—Jona-Lasinio model is a suitable one.
- Mean field approximation as the 1st step

\mathcal{L}

$$=\psi(i\gamma^{\mu}\partial_{\mu}-m_{0}+2G\langle\psi\psi\rangle_{\vec{r}})\psi+\mu\,\psi\gamma^{0}\psi-G\langle\psi\psi\rangle^{2}$$

+ $\mathcal{L}_{rotation}$