

A polarization along the beam direction in Au+Au collisions at $\sqrt{s_{NN}} = 200 \text{ GeV}$

Takafumi Niida for the STAR Collaboration



WAYNE STATE UNIVERSITY

The 5th Workshop on Chirality, Vorticity, and Magnetic Field in Heavy Ion Collisions @Tsinghua University, Beijing







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Star Important features in non-central heavy-ion collisions





Strong magnetic field

 $B \sim 10^{13} \text{ T}$ $(eB \sim \text{MeV}^2 \ (\tau = 0.2 \text{ fm}))$

D. Kharzeev, L. McLerran, and H. Warringa, Nucl.Phys.A803, 227 (2008) McLerran and Skokov, Nucl. Phys. A929, 184 (2014)



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 \rightarrow Chiral magnetic effect Chiral magnetic wave particle polarization

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Star Important features in non-central heavy-ion collisions

 \rightarrow Chiral vortical effect particle polarization











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- Z.-T. Liang and X.-N. Wang, PRL94, 102301 (2005)
- S. Voloshin, nucl-th/0410089 (2004)
- ^aNon-zero angular momentum transfers to the spin degrees of freedom (polarization)
 - Particles' and anti-particles' spins are aligned with angular momentum, L
- - ^aMagnetic field align particle's spin • Particles' and antiparticles' spins are aligned oppositely along **B** due to the opposite sign of
 - magnetic moment







Parity-violating decay of hyperons

Daughter baryon is preferentially emitted in the direction of hyperon's spin (opposite for anti-particle)

$$\frac{dN}{d\Omega^*} = \frac{1}{4\pi} (1 + \alpha_{\rm H} \mathbf{P}_{\rm H} \cdot \mathbf{p}_{\mathbf{p}}^*)$$

 P_{H} : Λ polarization p_p^* : proton momentum in the Λ rest frame $\alpha_{\rm H}$: Λ decay parameter $(\alpha_{\wedge} = -\alpha_{\bar{\wedge}} = 0.642 \pm 0.013)$



C. Patrignani et al. (PDG), Chin. Phys. C 40, 100001 (2016)

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How to measure the global polarization?

Projection onto the transverse plane

Angular momentum direction can be determined by spectator deflection (spectators deflect outwards) - S. Voloshin and TN, PRC94.021901(R)(2016)



 Ψ_1 : azimuthal angle of b ϕ_{p}^{*} : ϕ of daughter proton in Λ rest frame STAR, PRC76, 024915 (2007)









First observation of fluid vortices in HIC



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The Fastest Fluid by Sylvia Morrow

Superhot material spins at an incredible rate.

- Positive polarization signal at lower energies!
- polarization looks to increase in lower energies
- anti- Λ looks larger than Λ , possible effect of B-field?





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Becattini, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)



 μ_{Λ} : Λ magnetic moment T: temperature at thermal equilibrium

$$\omega = (P_{\Lambda} + P_{\bar{\Lambda}})k_B T/\hbar$$

~ 0.02-0.09 fm⁻¹
~ 0.6-2.7 × 10²²s⁻¹
(T=160 MeV)

The most vortical fluid ever observed!







Star Positive signal at VSNN = 200 GeV



AMPT: H. Li et al., Phys. Rev. C 96, 054908 (2017)

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Average P_H for 20-50%:

 $P_H(\Lambda) \ [\%] = 0.277 \pm 0.040 (\text{stat}) \pm {}^{0.039}_{0.049} (\text{sys})$ $P_H(\bar{\Lambda})$ [%] = 0.240 ± 0.045(stat) ±^{0.061}_{0.045} (sys)

- Having new results for 200 GeV, P_H decreases in higher energy - partly due to stronger shear flow structure in lower $\sqrt{s_{NN}}$ because of baryon transparency

Both hydrodynamic and AMPT models describe the data

- 15%-20% smearing effect in the data due to feed-down

F. Becattini, I. Karpenko, M. Lisa, I. Upsal, and S. Voloshin, PRC95.054902 (2017)

Larger polarization in in-plane than in out-of-plane

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Larger polarization in in-plane than in out-of-plane

Opposite to the hydrodynamic expectation (larger in out-of-plane)

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0.012 0.009 0.006 0.003 0.000 -0.003 -0.006 -0.009-0.012

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YT and T. Hirano, Nucl. Phys. A904-905 2013 (2013) 1023c-1026c Y. Tachibana and T. Hirano, NPA904-905 (2013) 1023

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L.-G. Pang, H. Peterson, Q. Wang, and X.-N. Wang PRL117, 192301 (2016)

F. Becattini and I. Karpenko, PRL120.012302 (2018) S. Voloshin, EPJ Web Conf. 171, 07002 (2018)

YT and T. Hirano, Nucl. Phys. A904-905 2013 (2013) 1023c-1026c Y. Tachibana and T. Hirano, NPA904-905 (2013) 1023

Vorticity (polarization) along the beam direction due to the elliptic flow

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local vorticity induced by collective flow

L.-G. Pang, H. Peterson, Q. Wang, and X.-N. Wang PRL117, 192301 (2016)

F. Becattini and I. Karpenko, PRL120.012302 (2018) S. Voloshin, EPJ Web Conf. 171, 07002 (2018)

- S. Voloshin, SQM2017
- F. Becattini and I. Karpenko, PRL120.012302 (2018)

Stronger flow in in-plane than in out-of-plane could make local polarization along beam axis!

arXiv:1501.04468v3 [nucl-th] 17 Aug 2015

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- Effect of Ψ_2 resolution is not corrected here

^D Sine structure as expected from the elliptic flow!

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- Effect of Ψ_2 resolution is not corrected here

^a Sine structure as expected from the elliptic flow!

^D Opposite sign to the hydrodynamic model and transport model (AMPT)

- Hydro model: F. Becattini and I. Karpenko, PRL.120.012302 (2018)

- AMPT model: X. Xia, H. Li, Z. Tang, Q. Wang, arXiv:1803.0086

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^aStrong centrality dependence as in v₂ ^aSimilar magnitude to the global polarization ^a~5 times smaller magnitude than the hydro and AMPT with the opposite sign!

Opposite sign to hydrodynamic model and AMPT model

- F. Becattini and I. Karpenko, PRL.120.012302 (2018)
 3D viscous hydrodynamic model with UrQMD initial condition assuming a local thermal equilibrium
- AMPT: X. Xia, H. Li, Z. Tang, Q. Wang, PRC98.024905 (2018)

Same sign as chiral kinetic approach

- Y. Sun and C.-M. Ko, PRC99, 011903(R) (2019)
- Assuming non-equilibrium of spin degree of freedom
- Smaller quark scattering cross section changes the sign

Suggest incomplete thermal equilibrium of spin degree of freedom as it may develop later in time unlike the global polarization?

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I. Karpenko, QM2018

Longitudinal quadrupole f_2 :

 P_z dominated by temperature gradient and relativistic term, but not by kinematic vorticity based on the hydro model.

Can we get such a small kinetic vorticity in the blast-wave

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Quadrupole or sine structure of ω_z is expected with the factor [b_n-a_n]. The sign could be negative depending on the relation of flow and spatial anisotropy.

blast-wave model

S. Voloshin, SQM2017 EPJ Web Conf.171, 07002 (2018)

 $r_{max} = R[1 - a\cos(2\phi_s)],$ $\rho_t = \rho_{t,max}[r/r_{max}(\phi_s)][1 + b\cos(2\phi_s)] \approx \rho_{t,max}(r/R)[1 + (a+b)\cos(2\phi_s)].$

Approximation of the kinetic vorticity in the blast-wave model:

 $\omega_z = 1/2(\nabla \times \mathbf{v})_z \approx (\rho_{t,nmax}/R) \sin(n\phi_s)[b_n - a_n].$

an: spatial anisotropy R: reference source radius b_n: flow anisotropy ρ_t: transverse flow velocity

- Hydro-inspired model parameterized with freeze-out condition assuming the longitudinal boost invariance
 - Freeze-out temperature T_f
 - Radial flow rapidity ρ_0 and its modulation ρ_2 -
 - Source size R_x and R_y

$$\rho(r,\phi_s) = \tilde{r}[\rho_0 + \rho_2 \cos(2\phi_b)]$$
$$\tilde{r}(r,\phi_s) = \sqrt{(r\cos\phi_s)^2/R_x^2 + (r\sin\phi_s)^2}$$

Calculate vorticity at the freeze-out using the parameters • extracted from spectra, v₂, and HBT fit

$$\begin{split} \langle \omega_z \sin(2\phi) \rangle &= \frac{\int d\phi_s \int r dr \, I_2(\alpha_t) K_1(\beta_t) \omega_z \sin(2\phi_b)}{\int d\phi_s \int r dr \, I_0(\alpha_t) K_1(\beta_t)} \\ \omega_z &= \frac{1}{2} \left(\frac{\partial u_y}{\partial x} - \frac{\partial u_x}{\partial y} \right), \end{split}$$

u: local flow velocity, In, Kn: modified Bessel functions

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 $(s)^2/R_u^2$

FIG. 2. Schematic illustration of an elliptical subshell of the source. Here, the source is extended out of the reaction plane $(R_v > R_x)$. Arrows represent the direction and magnitude of the flow boost. In this example, $\rho_2 > 0$ [see Eq. (4)].

 ϕ_s : azimuthal angle of the source element ϕ_b : boost angle perpendicular to the elliptical subshell

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e.g. Blast-wave fit to spectra and v_2

PHENIX, PRC93.051902(R) (2016)

Calculated vorticity ω_z shows the sine modulation. Assuming a local thermal equilibrium, z-component of polarization is estimated as follows: $P_z \approx \omega_z / (2T)$

BW parameters obtained with HBT: STAR, PRC71.044906 (2005)

a AMPT model

• opposite sign and 5 times larger in magnitude X. Xia, H. Li, Z. Tang, Q. Wang, PRC98.024905 (2018)

Blast-wave model

- simple estimate for kinematic vorticity
- similar magnitude to the data
- inclusion of HBT in the fit affects the sign in \bullet peripheral collisions

T. Niida, S. Voloshin, A. Dobrin, and R. Bertens, in preparation

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- Observation of Λ global polarization at $\sqrt{s_{NN}} = 7.7-200$ GeV
 - Polarization decreases at higher energies → Quantitatively consistent with hydrodynamic and AMPT models
 - Larger signal in in-plane than in out-of-plane
 - → Disagree with hydrodynamic and AMPT model
 - Charge-asymmetry dependence with different slopes between Λ and anti- Λ (~2 σ level) \rightarrow A possible relation to the axial current induced by B-field?
- \Box First study of Λ polarization along the beam direction at $\sqrt{s_{NN}} = 200$ GeV
 - Quadrupole structure of the polarization relative to the 2nd-order event plane → Qualitatively consistent with a picture of the elliptic flow but agree/disagree among the data and theoretical calculations in the sign
 - Strong centrality dependence as in the elliptic flow
 - Sign problem among different models and data, but the blast-wave model predicts the same sign and similar magnitude to the data

Outlook

Isobar collision data (Ru+Ru, Zr+Zr) already taken in 2018! o Same mass number but different number of protons \rightarrow 10% difference in the magnetic field \rightarrow More P_H splitting between Λ and anti- Λ in Ru?

 \square New 27 GeV data taken in 2018! (x10 events with ~1.5 better EP resolution) o Possible probe of the magnetic field from Λ vs anti- Λ global polarization

Beam Energy Scan II (2019+) with STAR detector upgrade o x10 events for $\sqrt{s_{NN}} = 7.7-19.6$ GeV (collider mode) + $\sqrt{s_{NN}} = 3-7.7$ GeV (Fixed target) o How about at forward/backward rapidity? How about for multi-strangeness?

D.-X. Wei et al., arXiv:1810.00151 Au+Au 20-50% (%) 10 $\sqrt{s_{_{\rm NN}}}$ (GeV)

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In non-central collisions,

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the initial collective longitudinal flow velocity depends on x.

In non-central collisions, the initial collective longitudinal flow velocity depends on x.

$$\omega_y = \frac{1}{2} (\nabla \times v)_y \approx -\frac{1}{2} \frac{\partial v_z}{\partial x}$$

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from $\Sigma^* \rightarrow \Lambda \pi$, $\Sigma^0 \rightarrow \Lambda \gamma$, $\Xi \rightarrow \Lambda \pi$

 \Box Polarization of parent particle R is transferred to its daughter Λ

$$\begin{split} \mathbf{S}_{\Lambda}^{*} &= C \mathbf{S}_{R}^{*} \qquad \langle S_{y} \rangle \propto \frac{S(S+1)}{3} (\omega + \frac{\mu}{S} B) \\ \text{hi, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)} \qquad \begin{array}{c} C_{\Lambda R} : \text{coefficient of spin transfer from parent} \\ S_{R} : \text{parent particle's spin} \\ f_{\Lambda R} : \text{fraction of } \Lambda \text{ originating from parent } R \\ \mu_{R} : \text{magnetic moment of particle } R \\ \end{array}$$

Becattir

$$\begin{pmatrix} \varpi_{c} \\ B_{c}/T \end{pmatrix} = \begin{bmatrix} \frac{2}{3} \sum_{R} \left(f_{\Lambda R} C_{\Lambda R} - \frac{1}{3} f_{\Sigma^{0} R} C_{\Sigma^{0} R} \right) S_{R}(S_{R} + 1) & \frac{2}{3} \sum_{R} \left(f_{\Lambda R} C_{\Lambda R} - \frac{1}{3} f_{\Sigma^{0} R} C_{\Sigma^{0} R} \right) (S_{R} + 1) \mu_{R} \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) & \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) (S_{\overline{R}} + 1) \mu_{\overline{R}} \end{bmatrix}^{-1} \begin{pmatrix} P_{\Lambda}^{\text{meas}} \\ P_{\overline{\Lambda}}^{\text{meas}} \\ P_{\overline{\Lambda}}^{\text{meas}} \end{pmatrix} = \begin{bmatrix} \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} C_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} + 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{1}{3} f_{\overline{\Sigma}^{0} \overline{R}} \right) S_{\overline{R}}(S_{\overline{R}} - 1) \\ \frac{2}{3} \sum_{\overline{R}} \left(f_{\overline{\Lambda R}} C_{\overline{\Lambda R}} - \frac{$$

Decay	С
Parity conserving: $1/2^+ \rightarrow 1/2^+ 0^-$	-1/3
Parity conserving: $1/2^- \rightarrow 1/2^+ 0^-$	1
Parity conserving: $3/2^+ \rightarrow 1/2^+ 0^-$	1/3
Parity-conserving: $3/2^- \rightarrow 1/2^+ 0^-$	-1/5
$\Xi^0 ightarrow \Lambda + \pi^0$	+0.900
$\Xi^- ightarrow \Lambda + \pi^-$	+0.927
$\frac{\Sigma^0 \to \Lambda + \gamma}{}$	-1/3

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$^{\Box}$ Only ~25% of measured Λ and anti- Λ are primary, while ~60% are feed-down

15%-20% dilution of primary Λ polarization (model-dependent)

nt R to Λ

Star Possible probe of magnetic field

Becattini, Karpenko, Lisa, Upsal, and Voloshin, PRC95.054902 (2017)

$$P_{\Lambda} \simeq \frac{1}{2} \frac{\omega}{T} + \frac{\mu_{\Lambda} B}{T}$$

 $P_{\overline{\Lambda}} \simeq \frac{1}{2} \frac{\omega}{T} - \frac{\mu_{\Lambda} B}{T}$
 μ_{Λ} : Λ magnetic moment
 $B = (P_{\Lambda} - P_{\overline{\Lambda}}) k_B T / \mu_N$
 $\sim 5.0 \times 10^{13}$ [Tesla]

nuclear magneton $\mu_N = -0.613 \mu_{\Lambda}$

Extracted B-field is close to our expectation. Need more data with better precision →BES-II and Isobaric collisions

In most central collision \rightarrow no initial angular momentum As expected, the polarization decreases in more central collisions

I. Karpenko and F. Becattini, EPJC(2017)77:213 W.-T. Deng and X.-G. Huang, arXiv:1609.01801

^aThe data do not show significant η dependence • Maybe due to baryon transparency at higher energy ^a Also due to event-by-event C.M. fluctuations

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- [□]No significant p_T dependence, as expected from the initial angular momentum of the system
- ^aHydrodynamic model underestimates the data. Initial conditions affect the magnitude and dependence on p_T
 - 3D viscous hydrodynamic model with two initial conditions (ICs)
 - UrQMD IC
 - Glauber with source tilt IC
 - F. Becattini and I. Karpenko, PRL120.012302, 2018

Case of 200 GeV as an example

- Event plane determination: ~22%
- Dethods to extract the polarization signal: ~21%
- ^a Possible contribution from the background: ~13%
- ^a Topological cuts: <3%
- \Box Uncertainties of the decay parameter: ~2% for Λ , ~9.6% for anti- Λ \Box Extraction of Λ yield (BG estimate): <1%
- Also, the following studies were done to check if there is no experimental effect: ^a Two different polarities of the magnetic field for TPC
- Acceptance effect
- ^D Different time period during the data taking ^D Efficiency effect

