



DIRECT DETECTION IN GAUGE PORTAL

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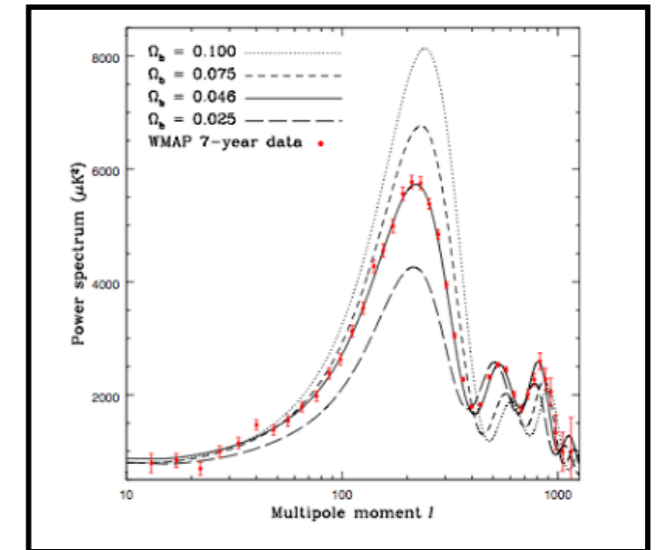
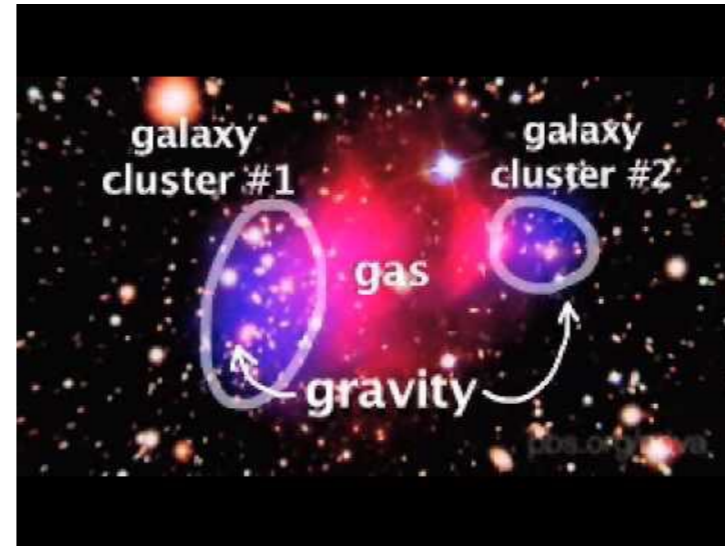
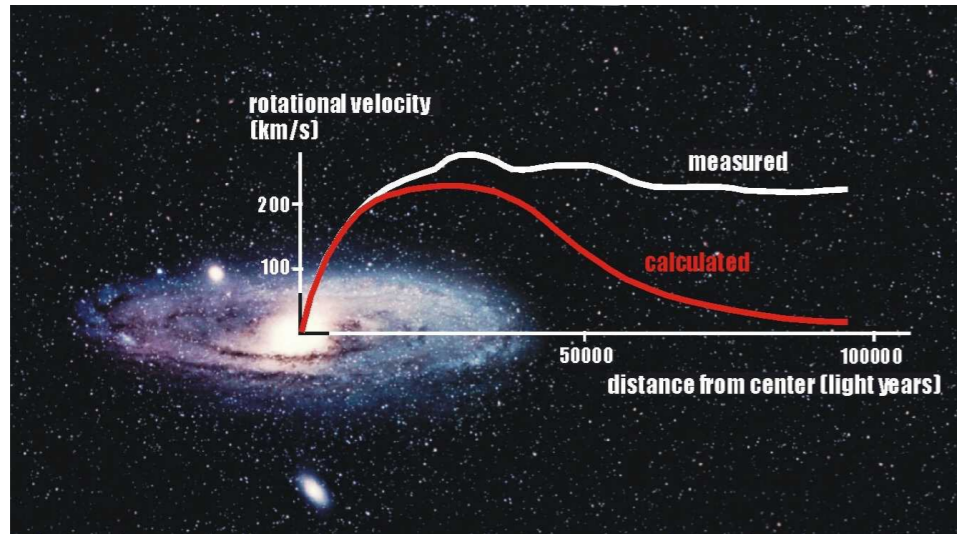
2019.05.18@重庆大学_23rd LHC mini workshop

要点预览

	Main points of DDs in gauge portal
1	Velocity suppressed cross section is reachable by the current DD technique!
2	Effects of twist-2 operators are crucial for direct detections in some parameter space.
3	Spin-dependent cross section may play important role in rolling out WIMPs
4	DD cross section of scalar-type electroweak multiplet DM are first analytically calculated by us!

W. Chao, G.J. Ding, X.G. He and M.J. Ramsey-Musolf, 1812.07829
W. Chao, 1904.09785

Evidence of DM



大尺度结构



暗物质雨

70kg 人体、60GeV 暗物质：
每秒10亿暗物质穿体而过
每年约有10个暗物质与人体发生碰撞
PLB 717 (2002)25

What is dark matter

We do not exactly know!

Mass
✗

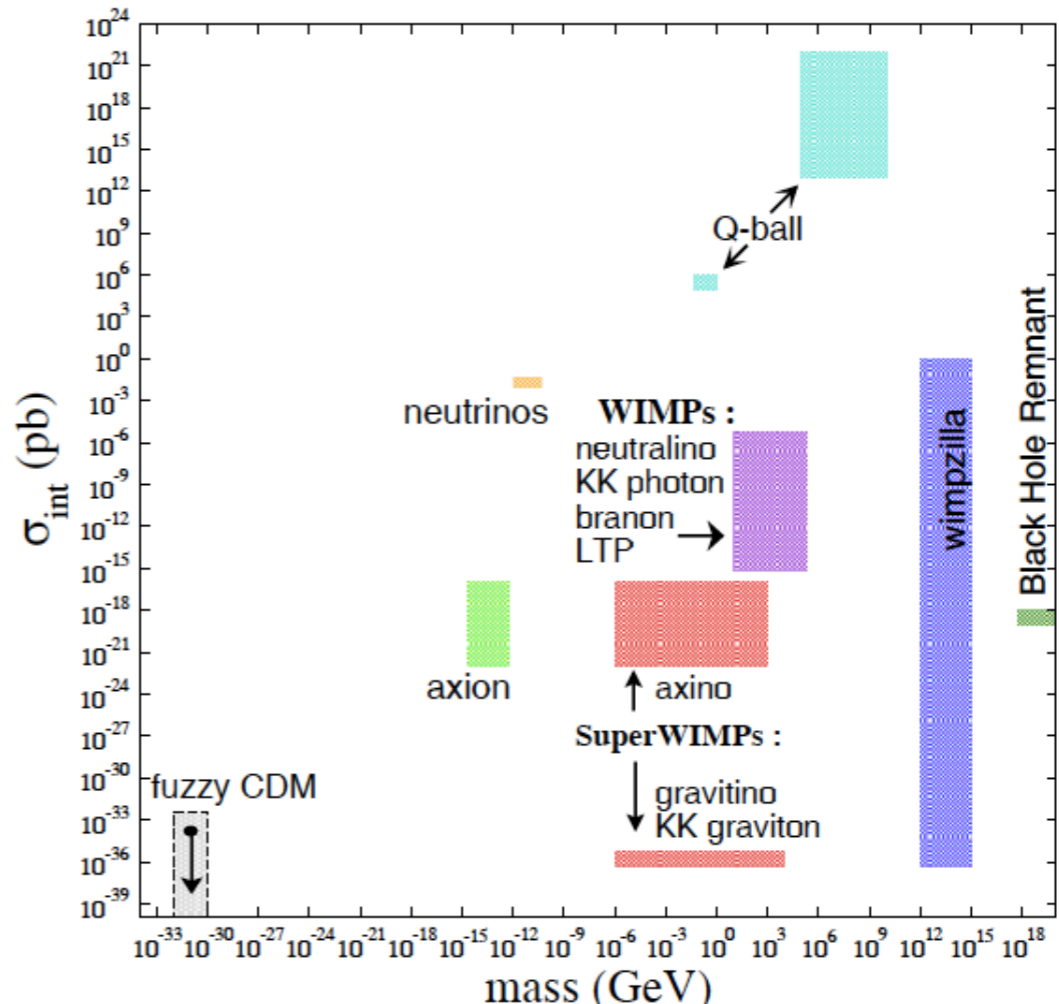
Spin
✗

Interactions
✗

Neutral, non-baryonic, weakly interacting particle!

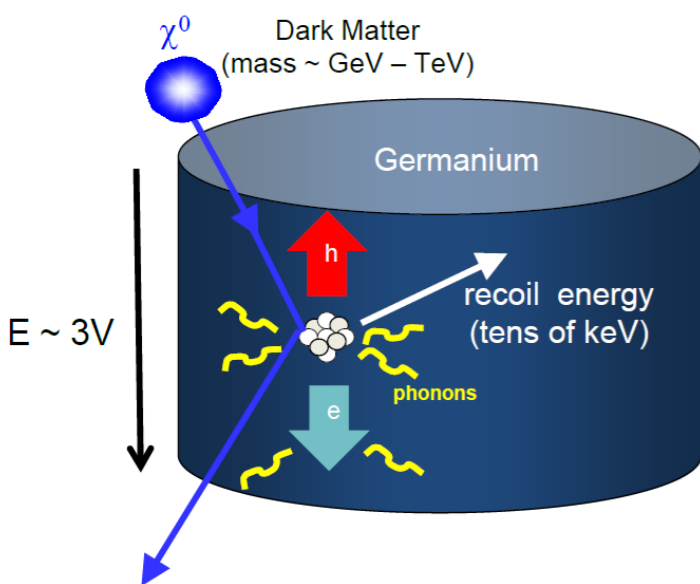
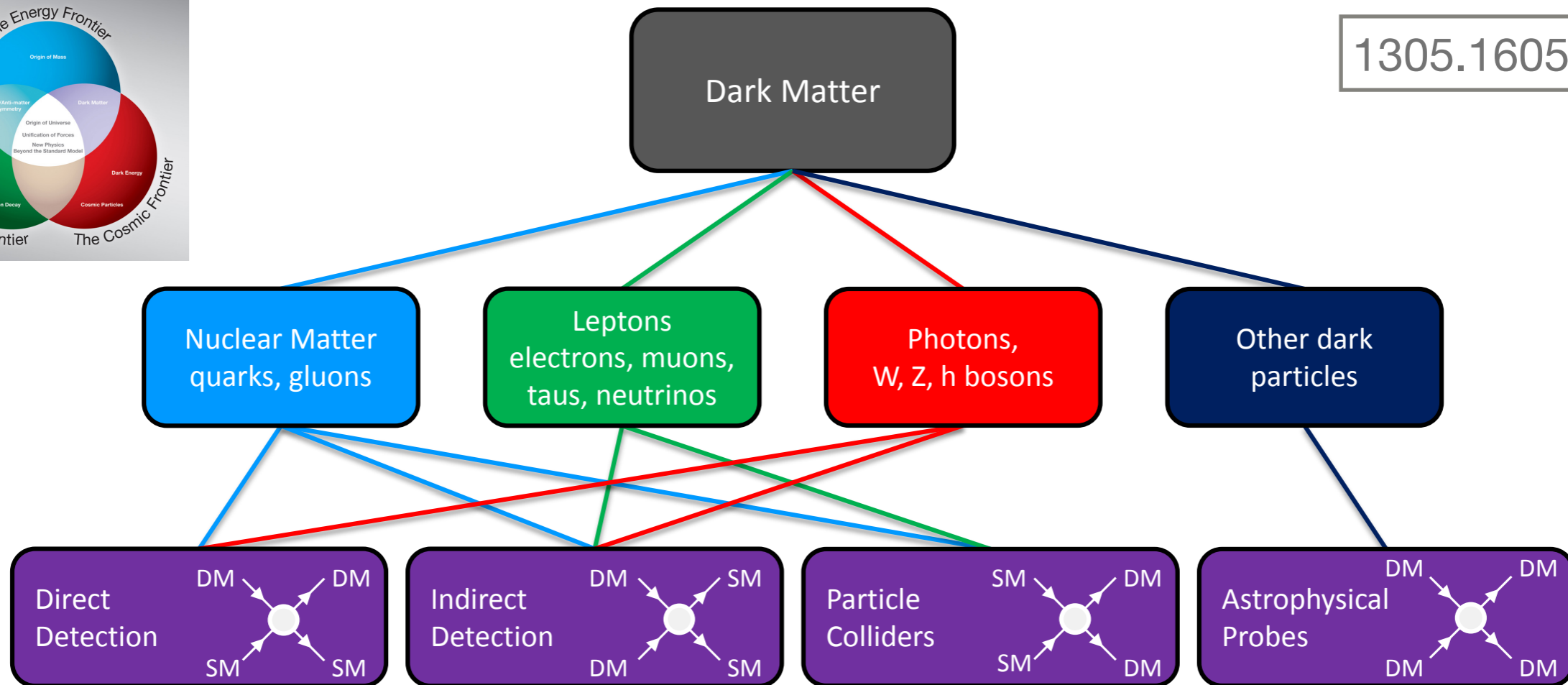
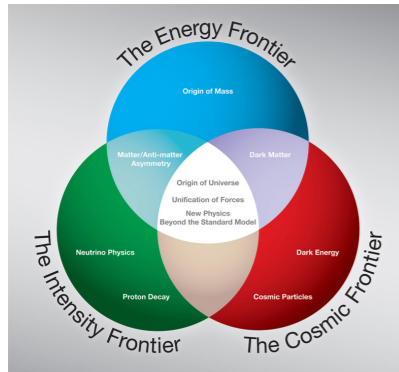
Particle Zoo

mass →	$\approx 2.3 \text{ MeV}/c^2$	$\approx 1.275 \text{ GeV}/c^2$	$\approx 173.07 \text{ GeV}/c^2$	0	$\approx 126 \text{ GeV}/c^2$
charge →	2/3	2/3	2/3	0	0
spin →	1/2	1/2	1/2	1	0
	u up	c charm	t top	g gluon	H Higgs boson
QUARKS	d down	s strange	b bottom	γ photon	DM
	$\approx 4.8 \text{ MeV}/c^2$	$\approx 95 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$	0	
	-1/3	-1/3	-1/3	0	
	1/2	1/2	1/2	1	
	e electron	μ muon	τ tau	Z Z boson	
LEPTONS	$0.511 \text{ MeV}/c^2$	$105.7 \text{ MeV}/c^2$	$1.777 \text{ GeV}/c^2$	$91.2 \text{ GeV}/c^2$	
	-1	-1	-1	0	
	1/2	1/2	1/2	1	
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	
	$< 2.2 \text{ eV}/c^2$	$< 0.17 \text{ MeV}/c^2$	$< 15.5 \text{ MeV}/c^2$	$80.4 \text{ GeV}/c^2$	
	0	0	0	± 1	
	1/2	1/2	1/2	1	
					GAUGE BOSONS



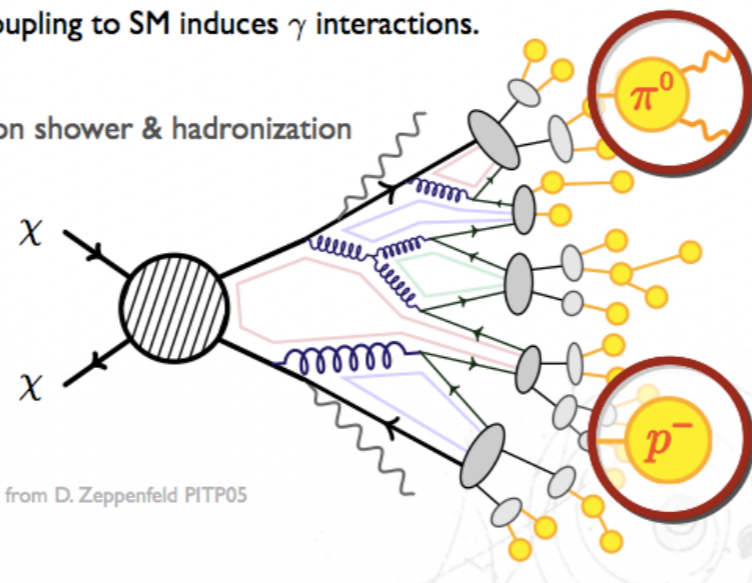
Ways of probing DM

1305.1605

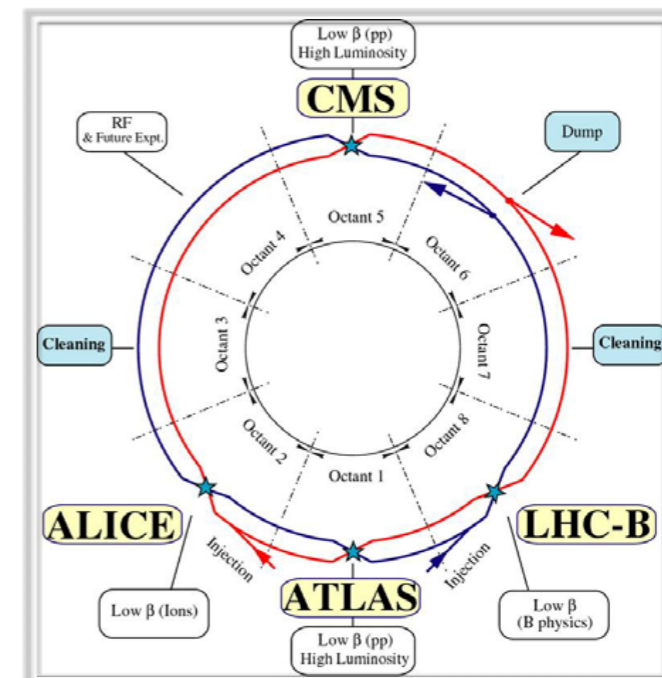


DM coupling to SM induces γ interactions.

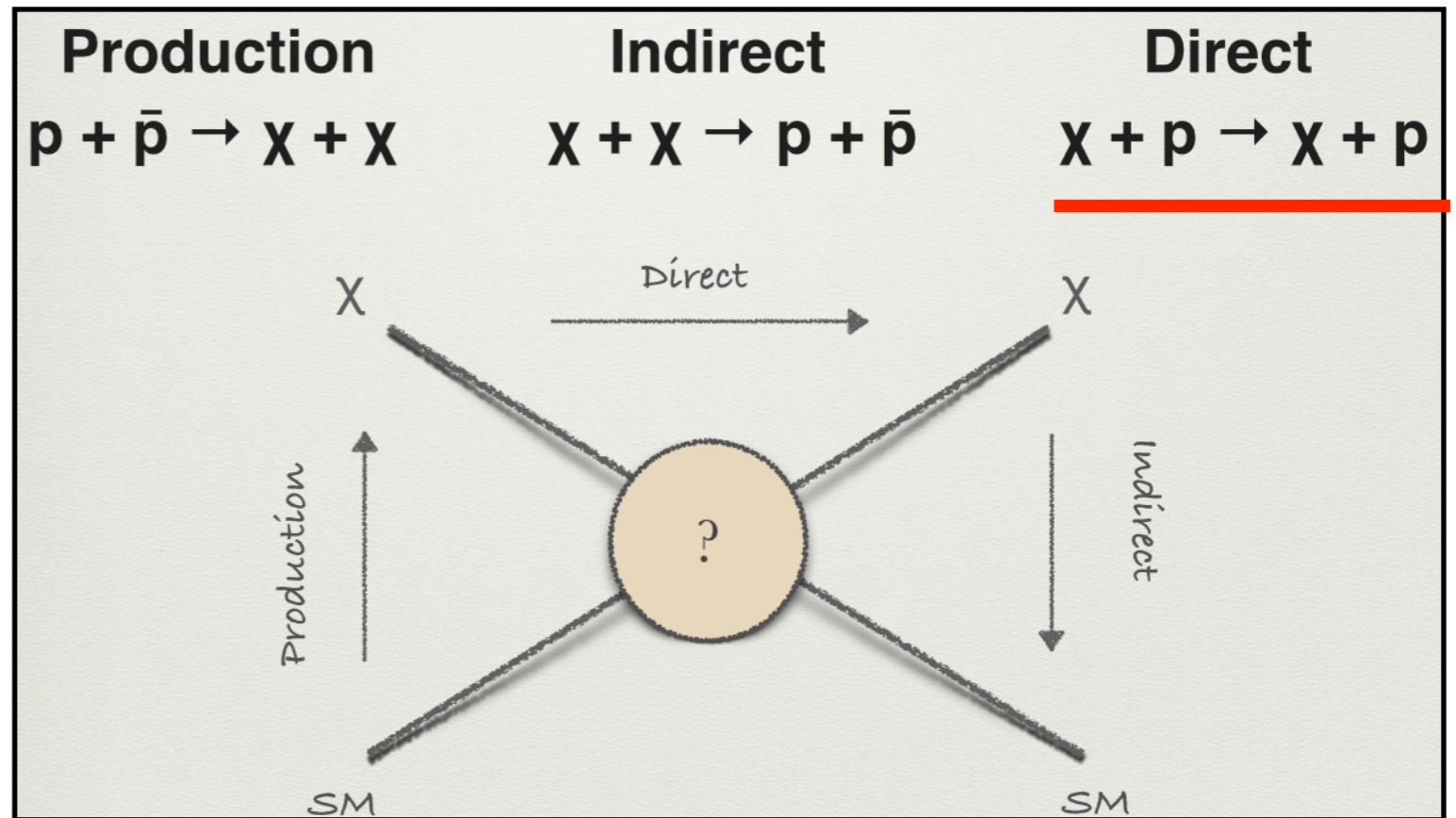
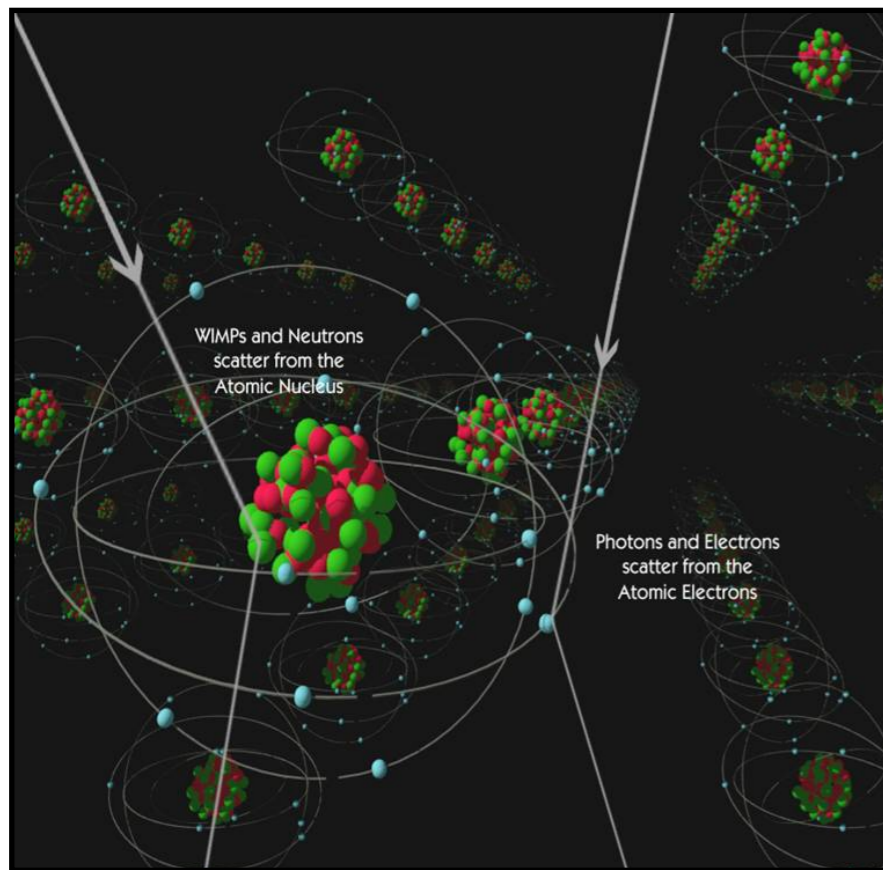
Parton shower & hadronization



Adapted from D. Zeppenfeld PITP05



DM Direct detections



Event rate per unit time

$$R \equiv N_A \frac{\rho_0}{m_\chi} \frac{m_A \sigma_0}{2\mu_A^2} \underbrace{\int F^2(E_R) dE_R}_{\text{Nuclear physics}} \int_{v_{\min}}^{v_{\max}} d^3v \frac{f(v)}{v}$$

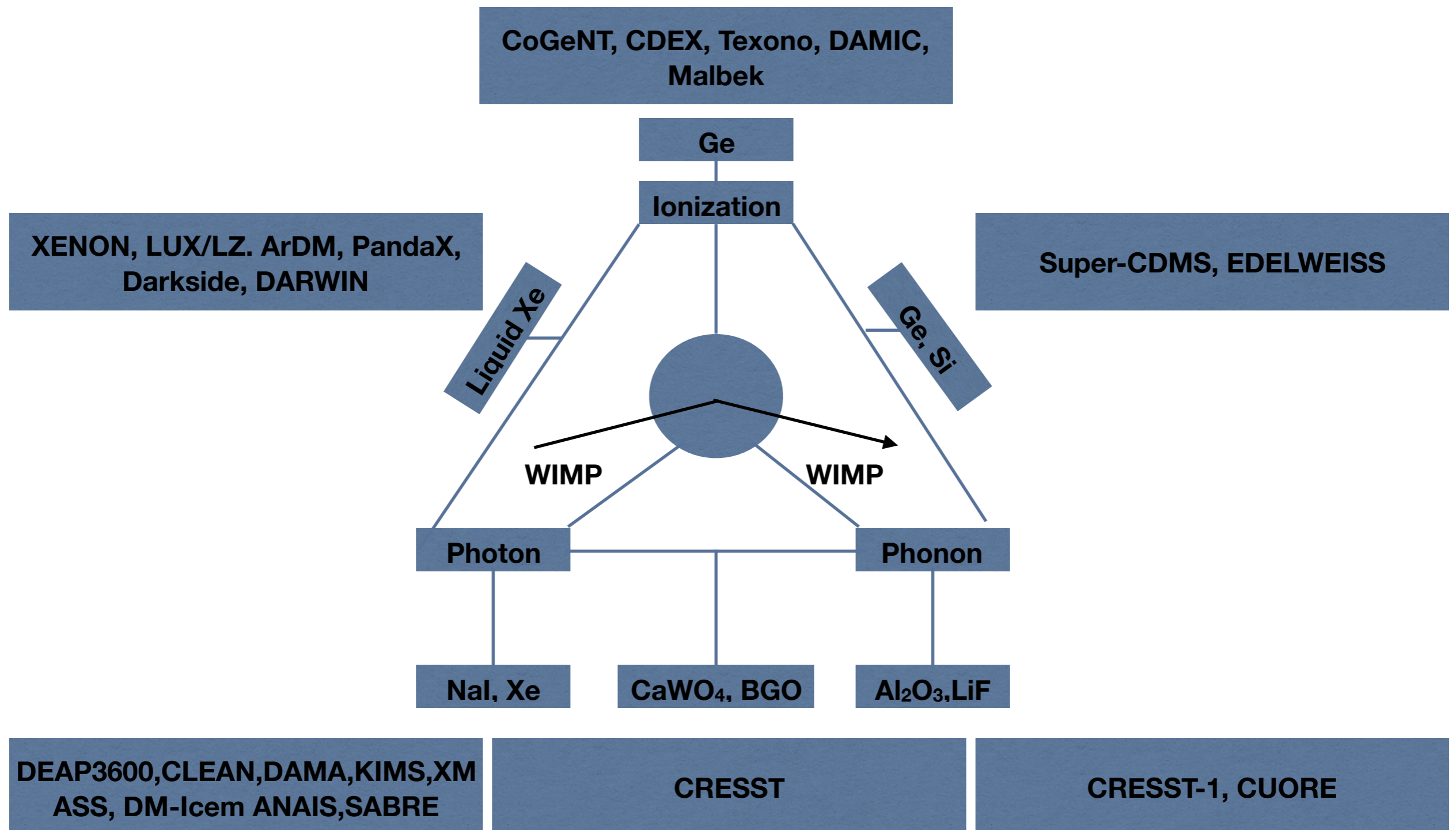
DM number density around the earth

Nuclear physics

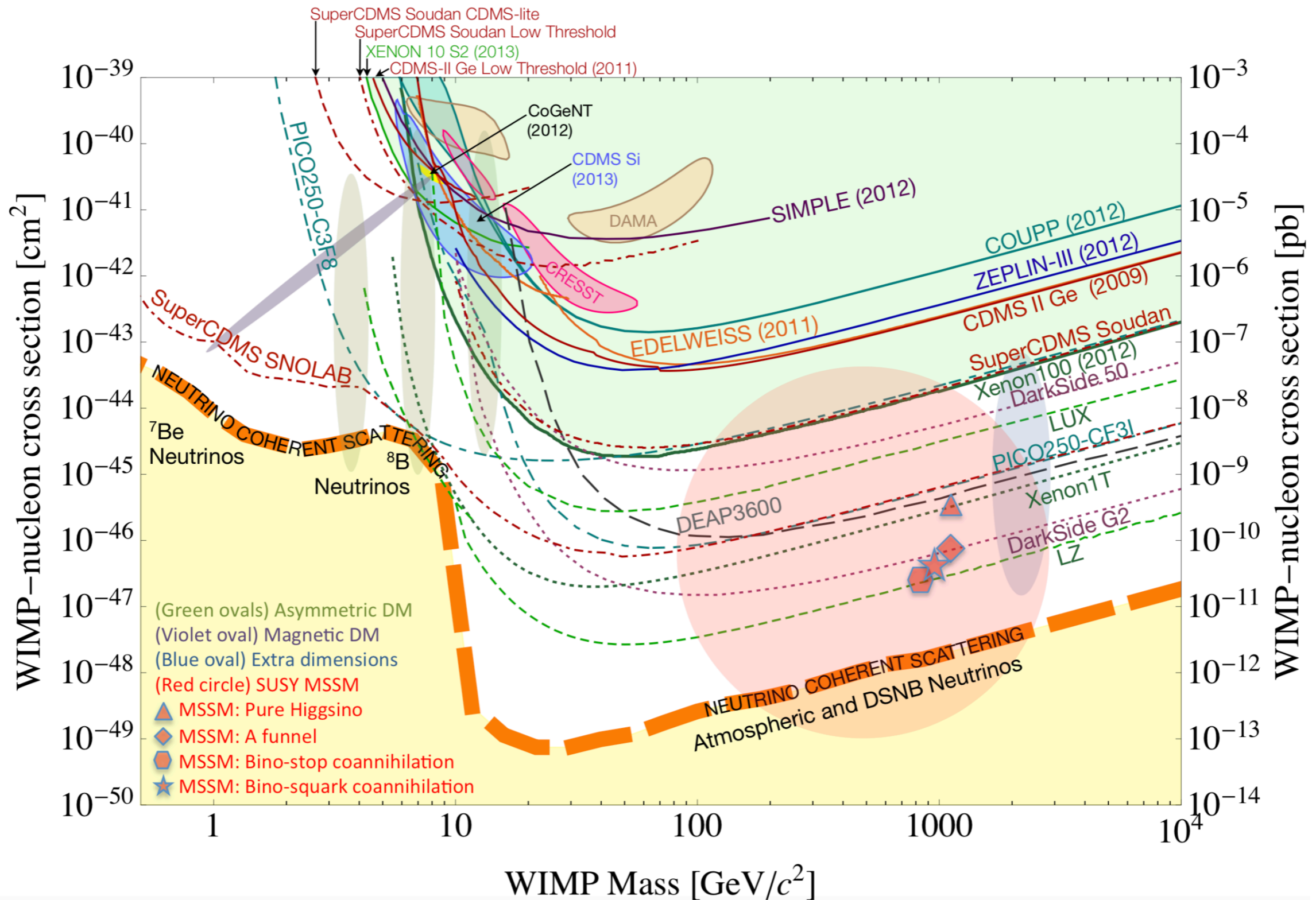
Particle physics

Maxwell-Boltzmann velocity distribution

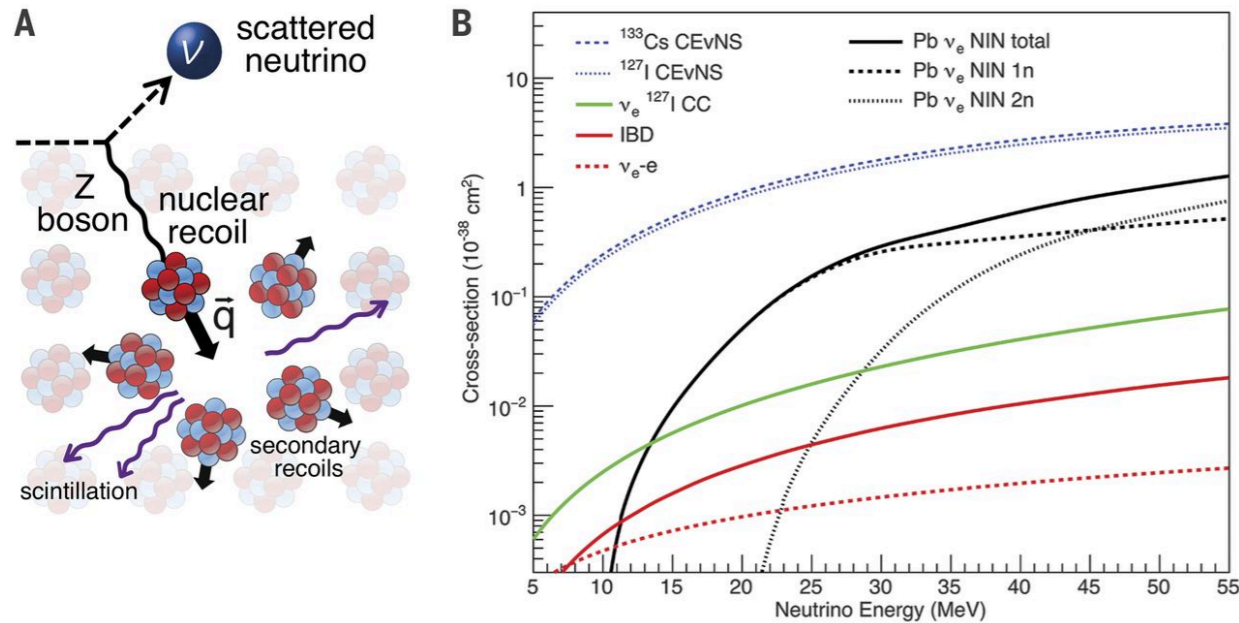
Detecting technologies



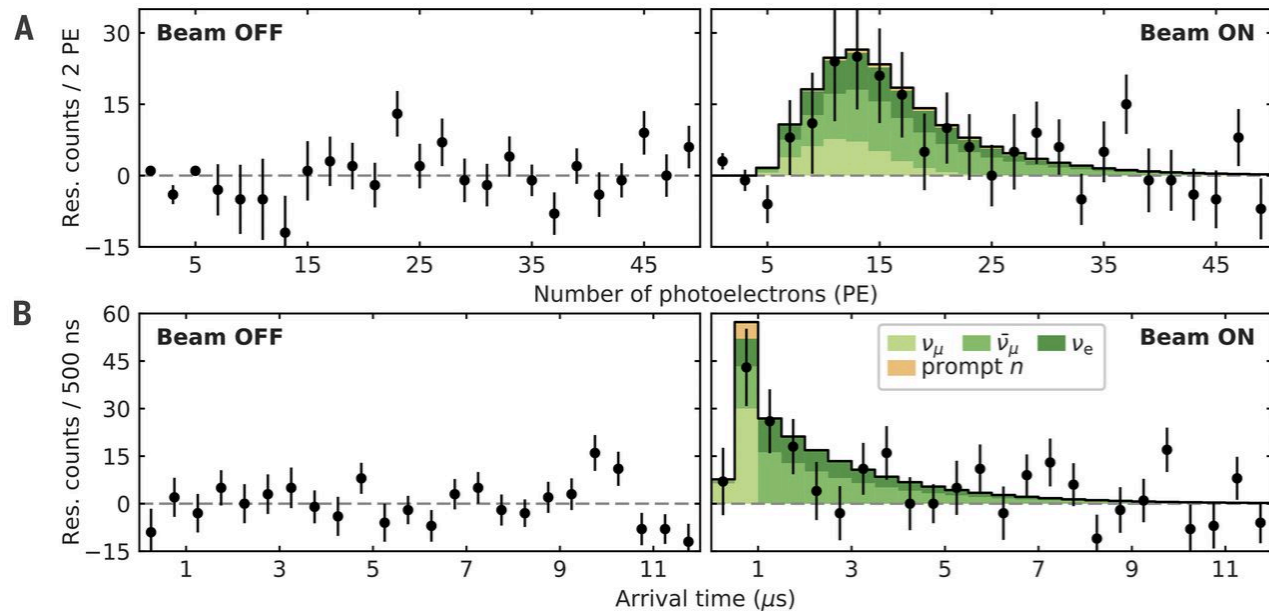
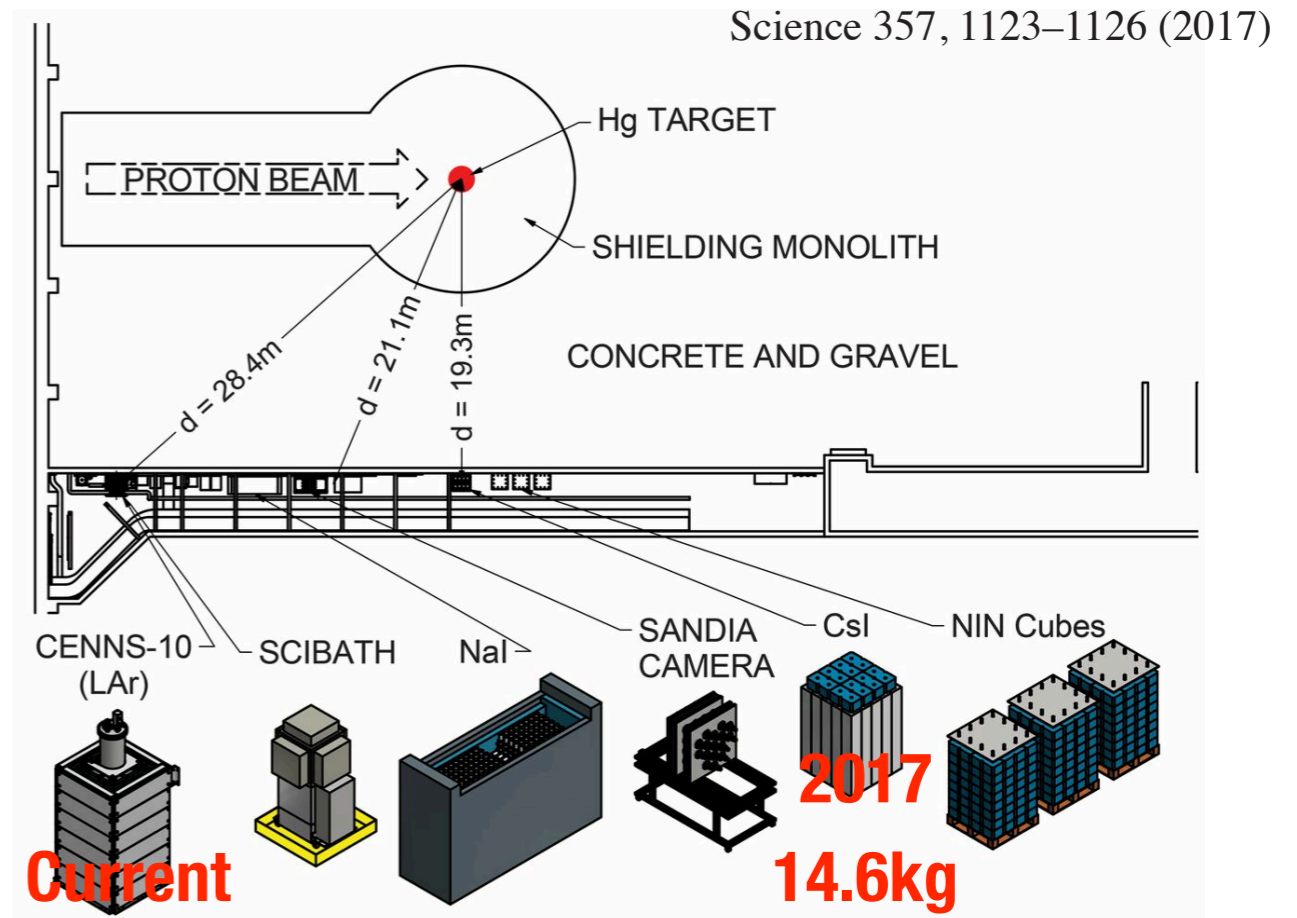
Where to go for Direct detections



CONHERENT



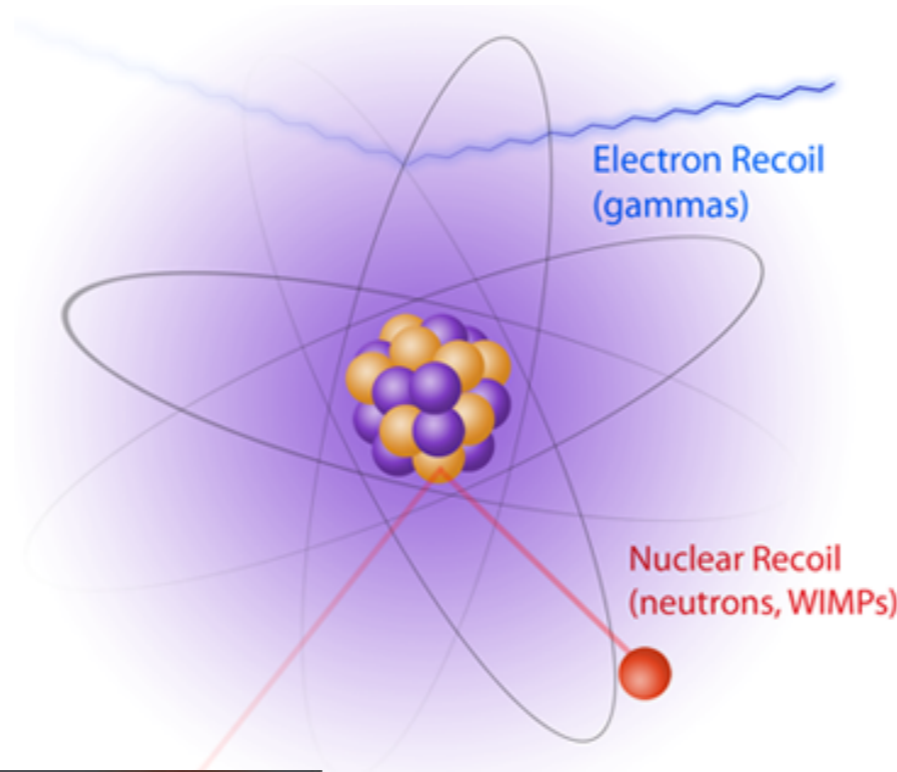
GOAL: Measure N^2 dependence of CEvNS process



Confirm CEvNS at 6.7 sigma

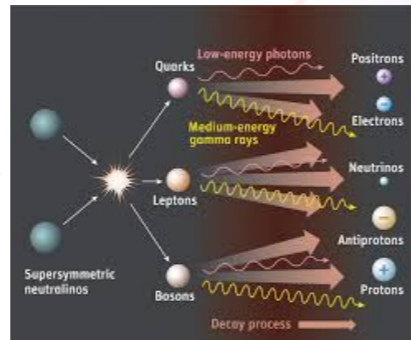
Beam ON coincidence window	547 counts
Anticoincidence window	405 counts
Beam-on bg: prompt beam neutrons	7.0 ± 1.7
Beam-on bg: NINs (neglected)	4.0 ± 1.3
Signal counts, single-bin counting	136 ± 31
Signal counts, 2D likelihood fit	134 ± 22
Predicted SM signal counts	173 ± 48

Be serious about the DD



SD cross section

SI cross section



LO effects

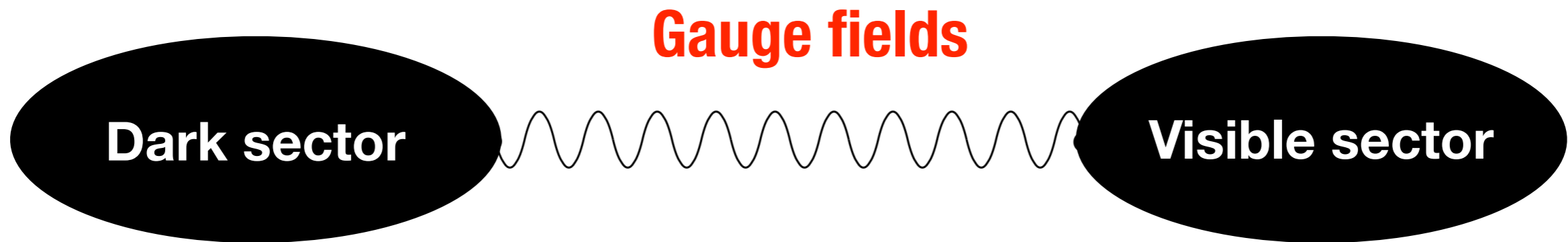
ν -suppressed

NLO effects



SERIOUS

DDs in gauge portal



Type of the gauge field:

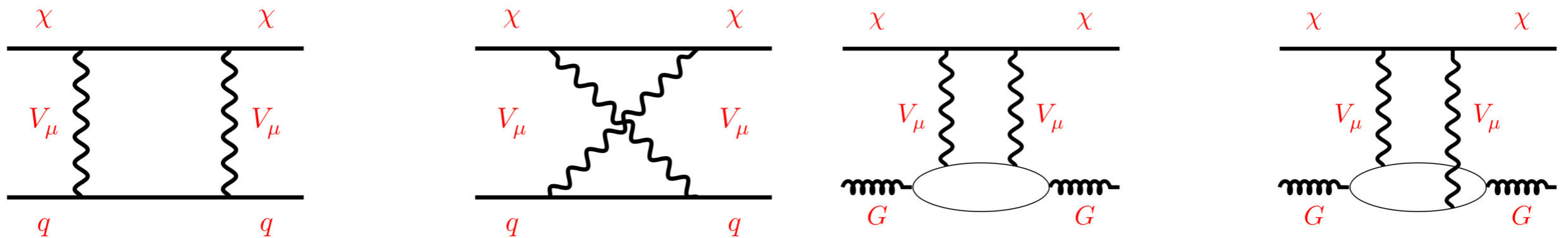
- Z'** **U(1) extension of SM**
- W/Z** **Minimal dark matter**

Lagrangian	Scenario	DD cross section
$\mathcal{L} \sim \frac{1}{2} g_V \bar{\chi} \gamma^\mu \gamma^5 V_\mu \chi - \frac{1}{2} m_\chi \bar{\chi} \chi + \zeta g_V \bar{q} \gamma^\mu q V_\mu$	A	SI but v suppressed
$\mathcal{L} \sim \frac{1}{2} g_V \bar{\chi} \gamma^\mu \gamma^5 V_\mu \chi - \frac{1}{2} m_\chi \bar{\chi} \chi + \zeta g_V \bar{q} \gamma^\mu \gamma^5 q V_\mu$	B	SD

DD cross sections

Scenario	DD cross section at leading order
A	$\sigma_{\text{SI}}^{\text{LO}} \approx \frac{36\pi\zeta^2\alpha_V^2\mu^2v^2}{m_V^4}$
B	$\sigma_{\text{SD}}^{\text{LO}} = 64\pi\zeta^2\alpha_V^2\frac{\mu^2}{m_V^4}\left(\sum_q \Delta_q^N\right)^2 J_N(J_N + 1)$

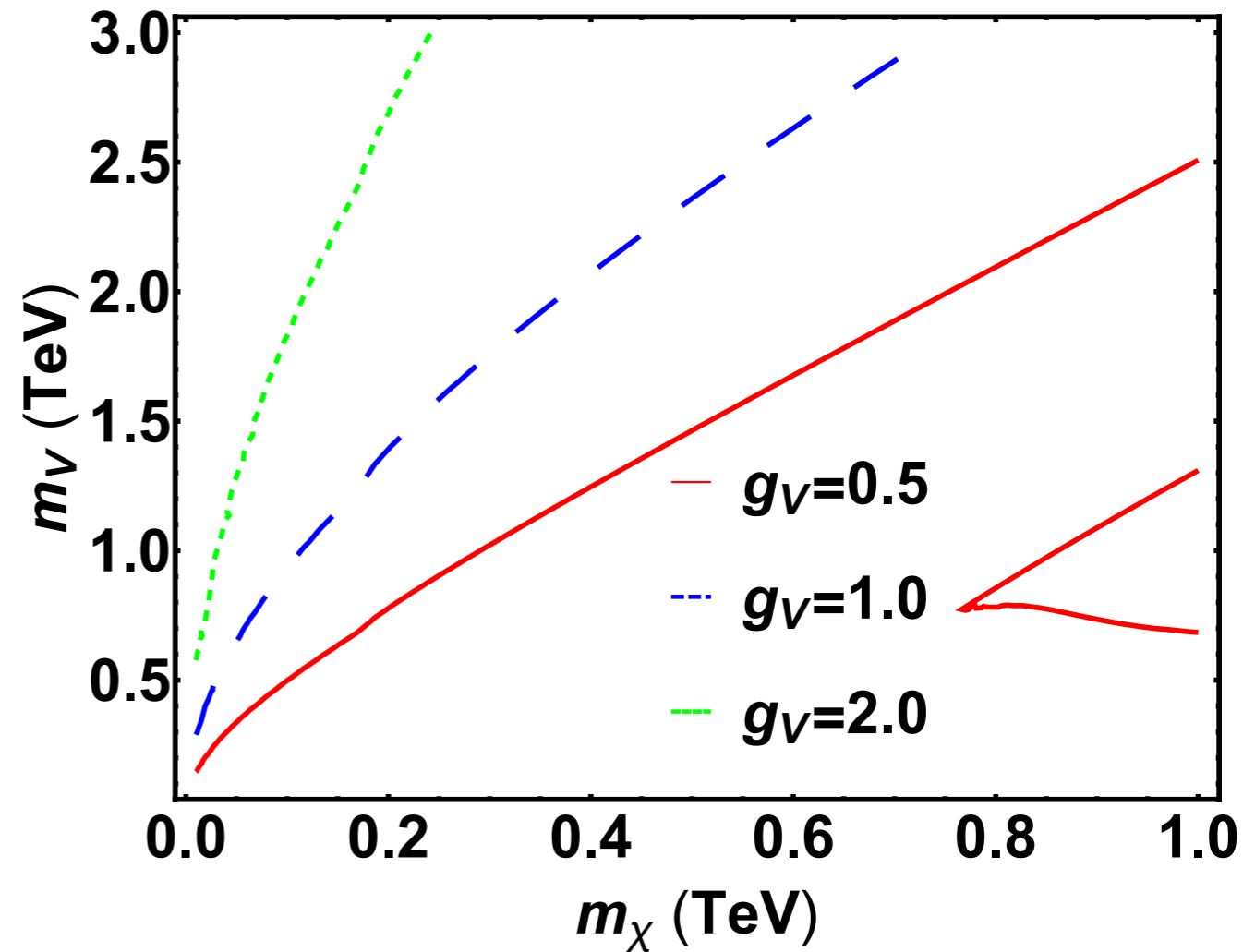
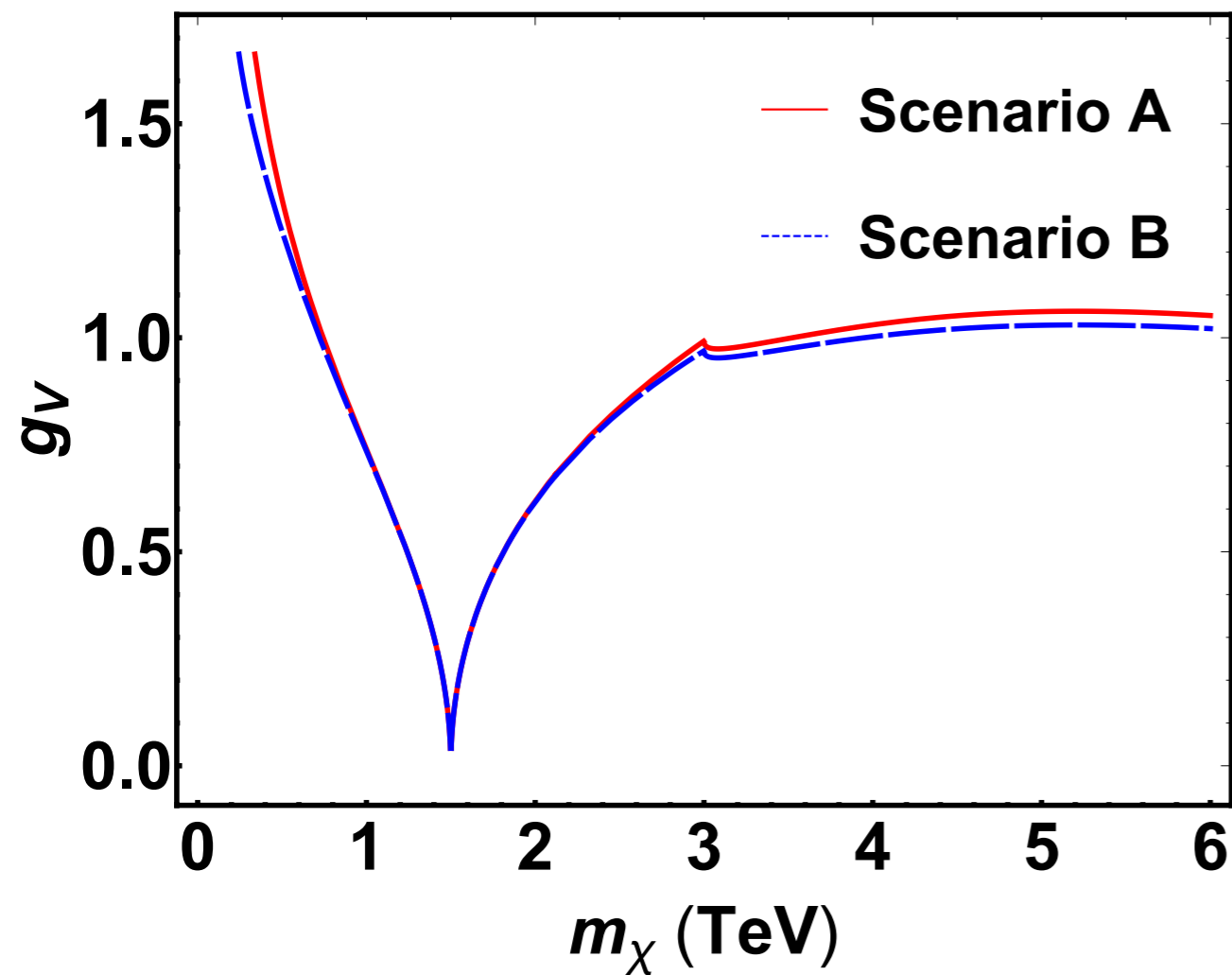
At the next-to-leading order:



$$\mathcal{L}_{\text{eff}} = \sum_{p=q,g} \frac{1}{2} \kappa_{1p} \bar{\chi} \chi \mathcal{O}_s^p + \frac{1}{2} \kappa_{2q} \bar{\chi} i \partial^\mu \gamma^\nu \chi \mathcal{O}_{\mu\nu}^q + \frac{1}{2} \kappa_{3q} \bar{\chi} i \partial^\mu i \partial^\nu \chi \mathcal{O}_{\mu\nu}^q,$$

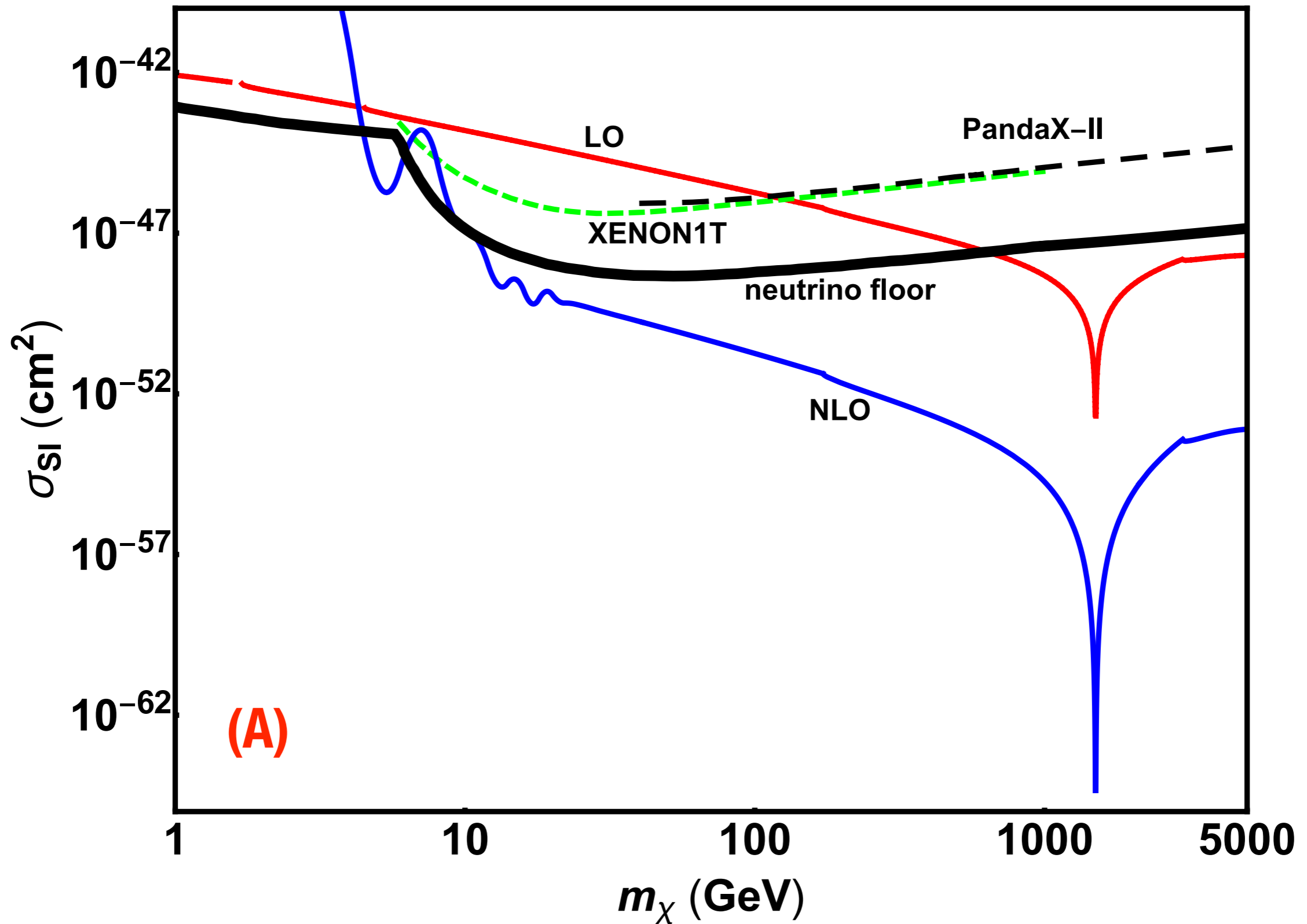
$$\mathcal{O}_{\mu\nu}^q = \frac{1}{2} \bar{q} \left(\partial_\mu \gamma_\nu + \partial_\nu \gamma_\mu - \frac{1}{2} g_{\mu\nu} \gamma \cdot \partial \right) q$$

Numerical results in Z' portal

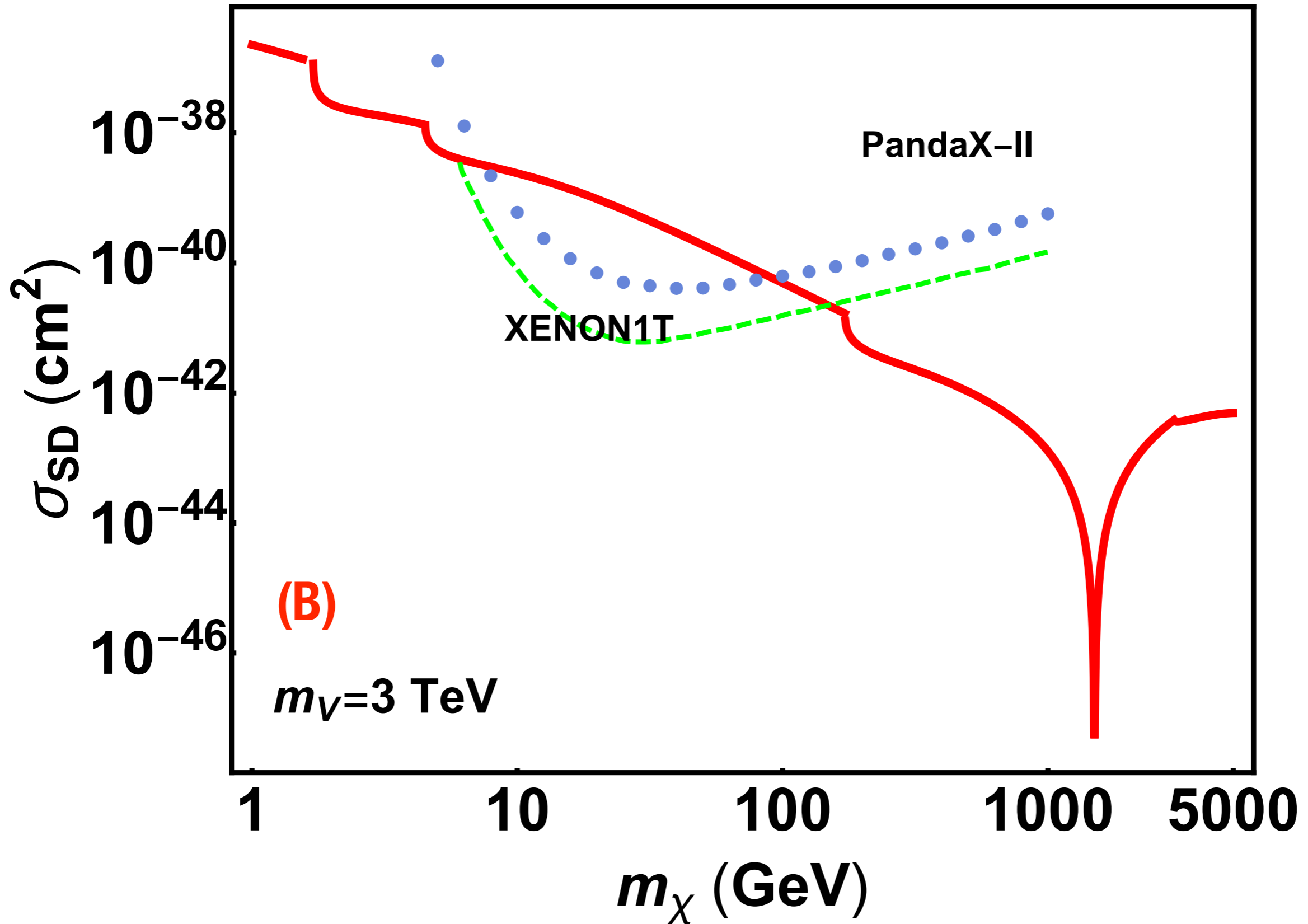


Constraints of the relic abundance

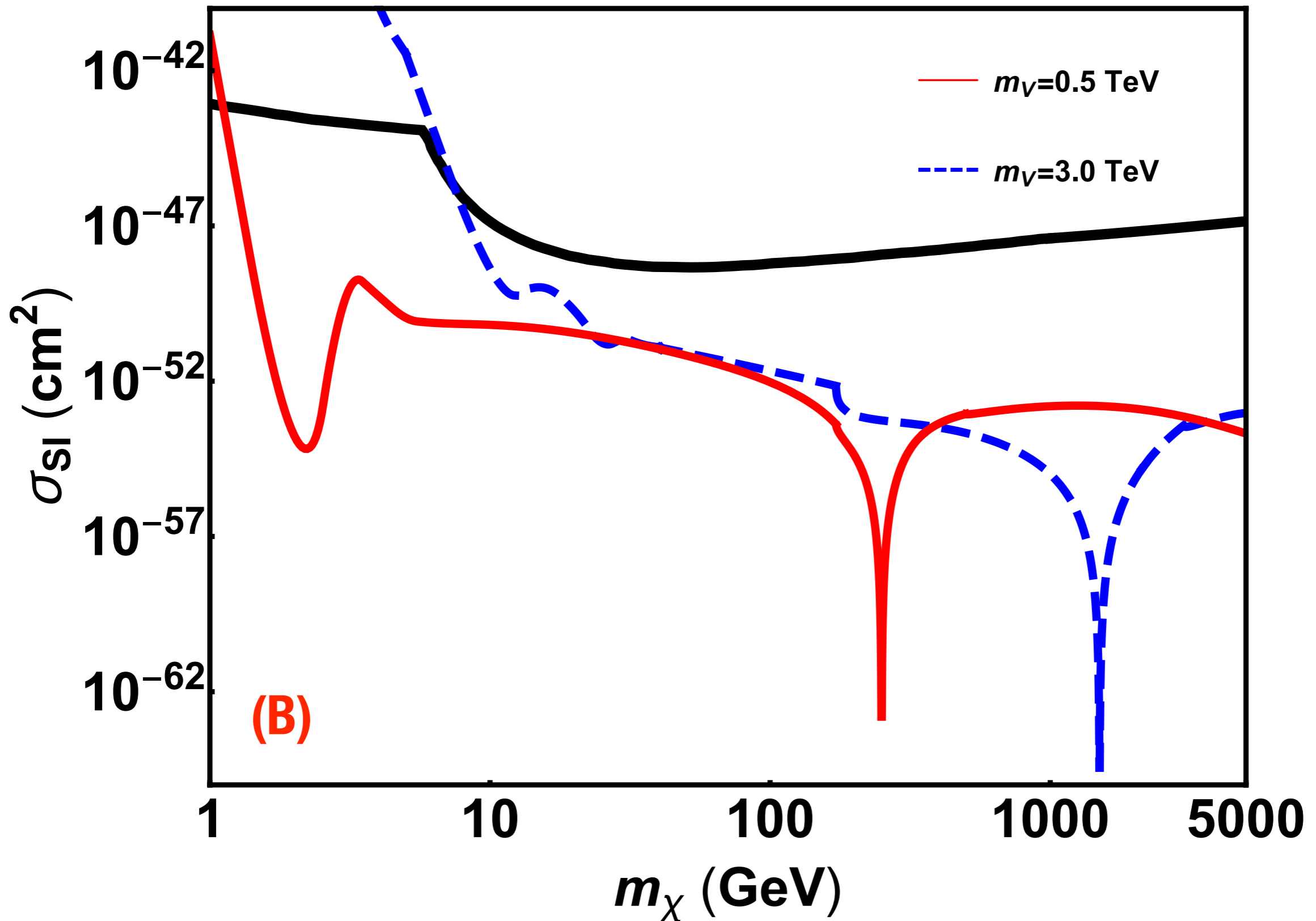
Numerical results in Z' portal



Numerical results in Z' portal



Numerical results in Z' portal



Scalar-type EW multiplet DM in W-portal

Why interesting?

Strumia, et al., NPB 2006

Quantum numbers			DM can	DM mass	$m_{\text{DM}^\pm} - m_{\text{DM}}$	Events at LHC	σ_{SI} in
$\text{SU}(2)_L$	$\text{U}(1)_Y$	Spin	decay into	in TeV	in MeV	$\int \mathcal{L} dt = 100/\text{fb}$	10^{-45} cm^2
2	1/2	0	EL	0.54 ± 0.01	350	$320 \div 510$	0.2
2	1/2	1/2	EH	1.1 ± 0.03	341	$160 \div 330$	0.2
3	0	0	HH^*	2.0 ± 0.05	166	$0.2 \div 1.0$	1.3
3	0	1/2	LH	2.4 ± 0.06	166	$0.8 \div 4.0$	1.3
3	1	0	HH, LL	1.6 ± 0.04	540	$3.0 \div 10$	1.7
3	1	1/2	LH	1.8 ± 0.05	525	$27 \div 90$	1.7
4	1/2	0	HHH^*	2.4 ± 0.06	353	$0.10 \div 0.6$	1.6
4	1/2	1/2	(LHH^*)	2.4 ± 0.06	347	$5.3 \div 25$	1.6
4	3/2	0	HHH	2.9 ± 0.07	729	$0.01 \div 0.10$	7.5
4	3/2	1/2	(LHH)	2.6 ± 0.07	712	$1.7 \div 9.5$	7.5
5	0	0	(HHH^*H^*)	5.0 ± 0.1	166	$\ll 1$	12
5	0	1/2	—	4.4 ± 0.1	166	$\ll 1$	12
7	0	0	—	8.5 ± 0.2	166	$\ll 1$	46

■ Natural DM candidate without requiring any other symmetry

■ Higgs vacuum stability

■ Avoiding constraints of DD as they are superheavy

Scalar-type EW multiplet DM in W-portal

How about the direct detection cross section

Fermion	$\mathcal{L}_{\text{eff}}^W = (n^2 - (1 \pm 2Y)^2) \frac{\pi\alpha_2^2}{16M_W} \sum_q \left[\left(\frac{1}{M_W^2} + \frac{1}{m_h^2} \right) [\bar{\mathcal{X}}\mathcal{X}] m_q [\bar{q}q] - \frac{2}{3M} [\bar{\mathcal{X}}\gamma_\mu\gamma_5\mathcal{X}] [\bar{q}\gamma_\mu\gamma_5q] \right]$
Scalar	No effective lagrangian! only some arguments.

Questions:

- No twist-2 effective operators in the Lagrangian?
- No effective gluon-DM operators in the Lagrangian?
- What is the situation when Higgs-portal interactions are included in SDM?

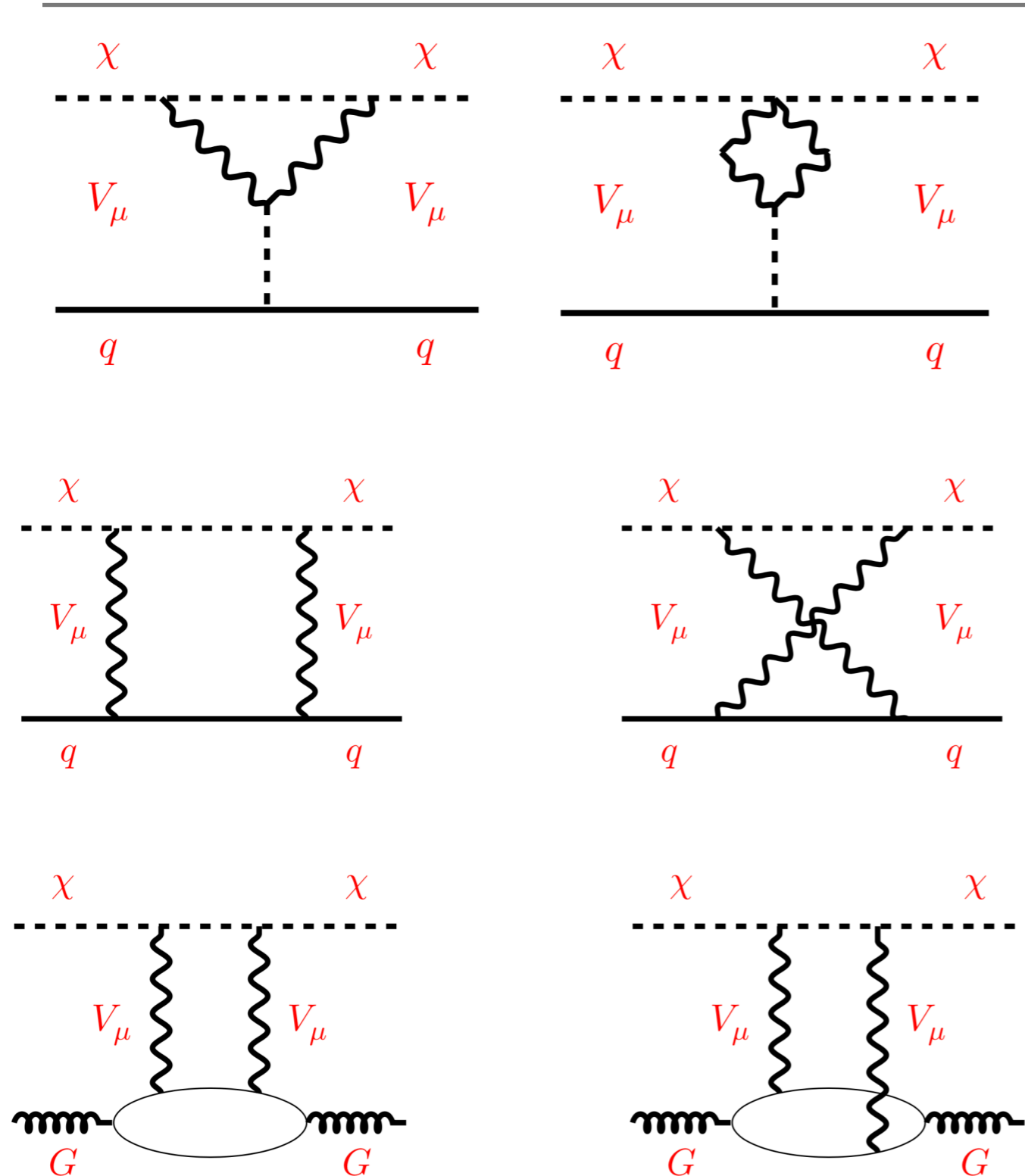


能动手就别吵吵



撸起袖子就是干

Scalar-type EW multiplet DM in W-portal



Results :

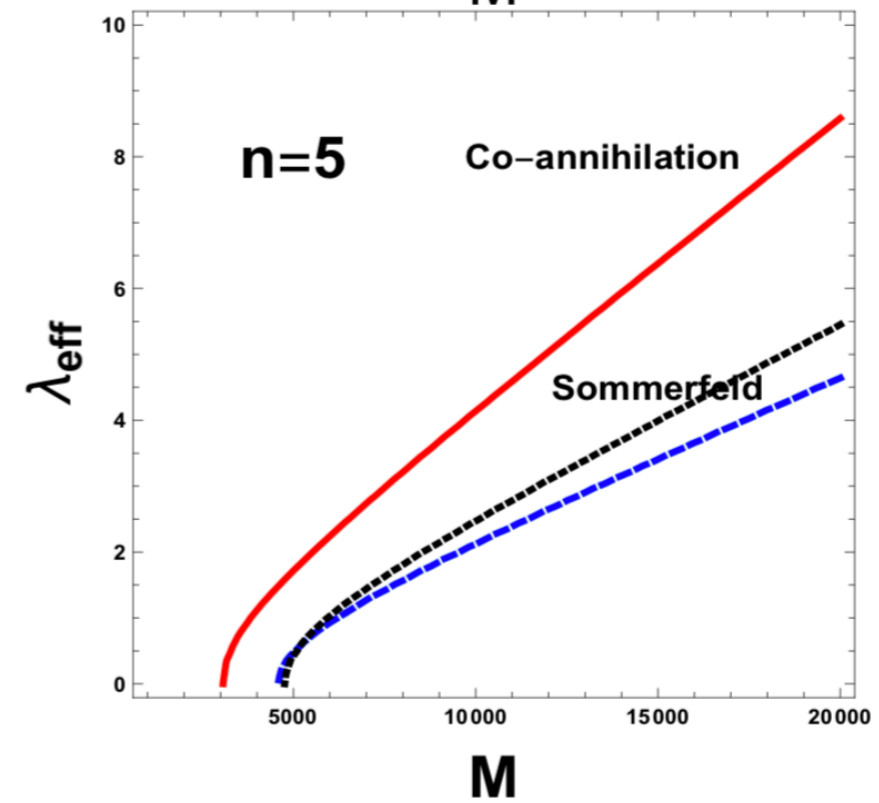
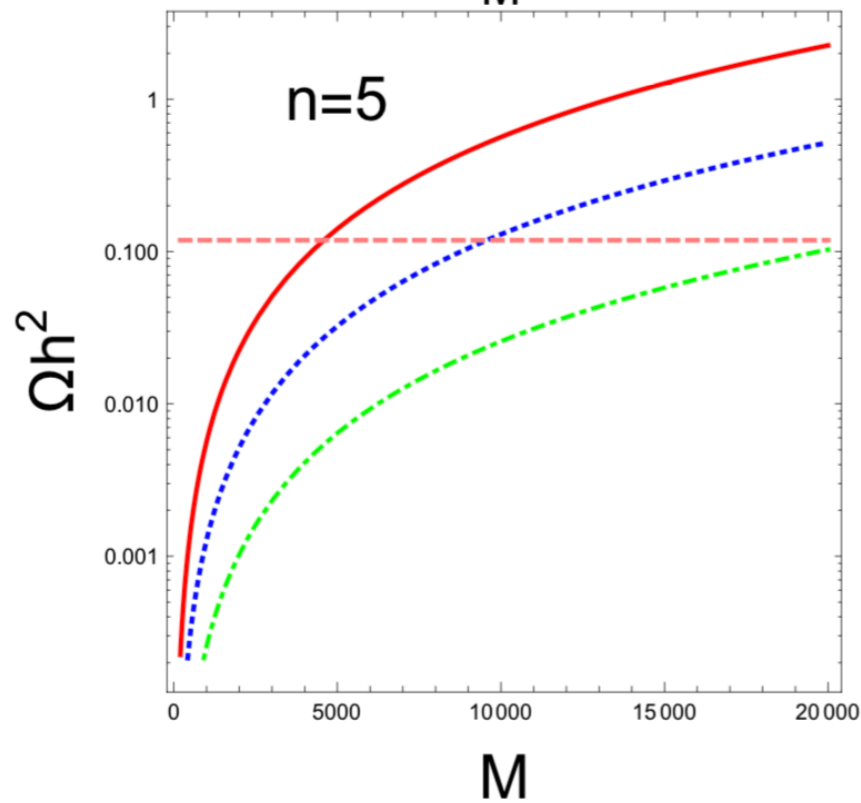
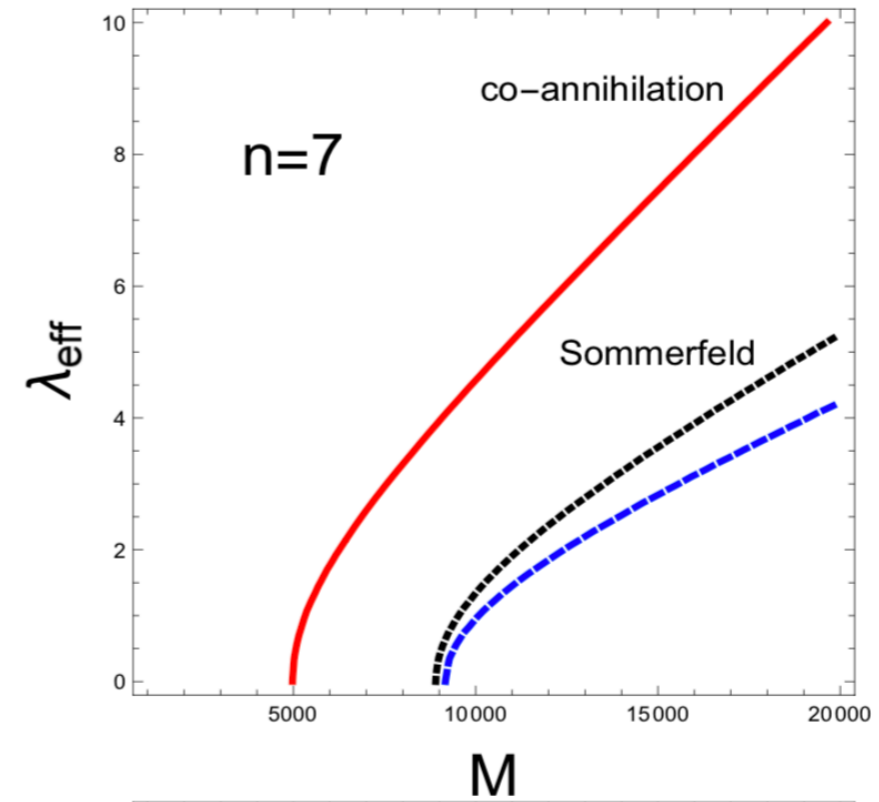
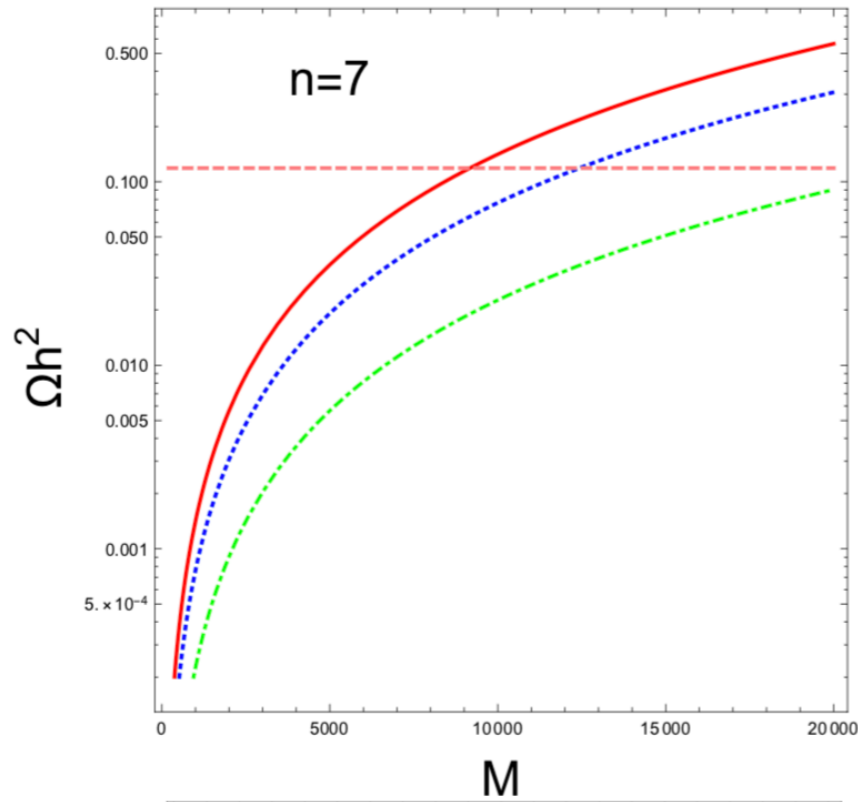
$$\mathcal{L} = \frac{(n^2 - 1)\alpha^2}{4} \left[\zeta_1 \phi^2 \bar{q} m_q q + \zeta_2 \phi i \partial^\mu i \partial^\nu \phi O_{\mu\nu}^q \right]$$

$$\zeta_1 = m^2(2C_0 + C_1 - 2Z_{11}) + 8Z_{00} - Y_2 - 4Z_{00}$$

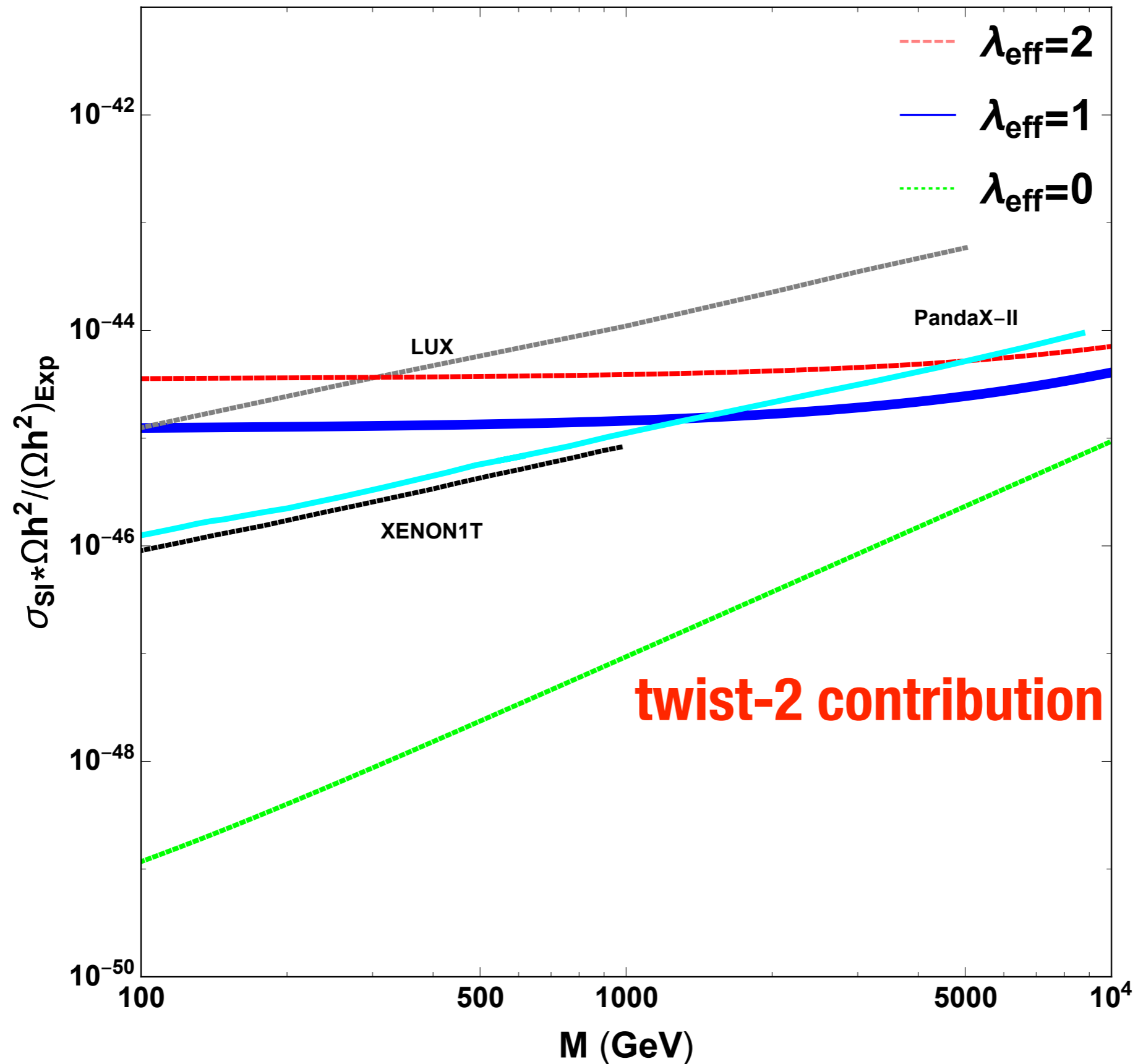
$$\zeta_2 = -8m^2 Z_{11} - 32Z_{00} - 8Y_2$$

- **C,Z,Y, are integration functions**
- **Gluon results are on the way**
- **RGE running should be included in DD**
- **When Higgs portal interactions are included, only twist-2 contribution is unique.**

Numerical results



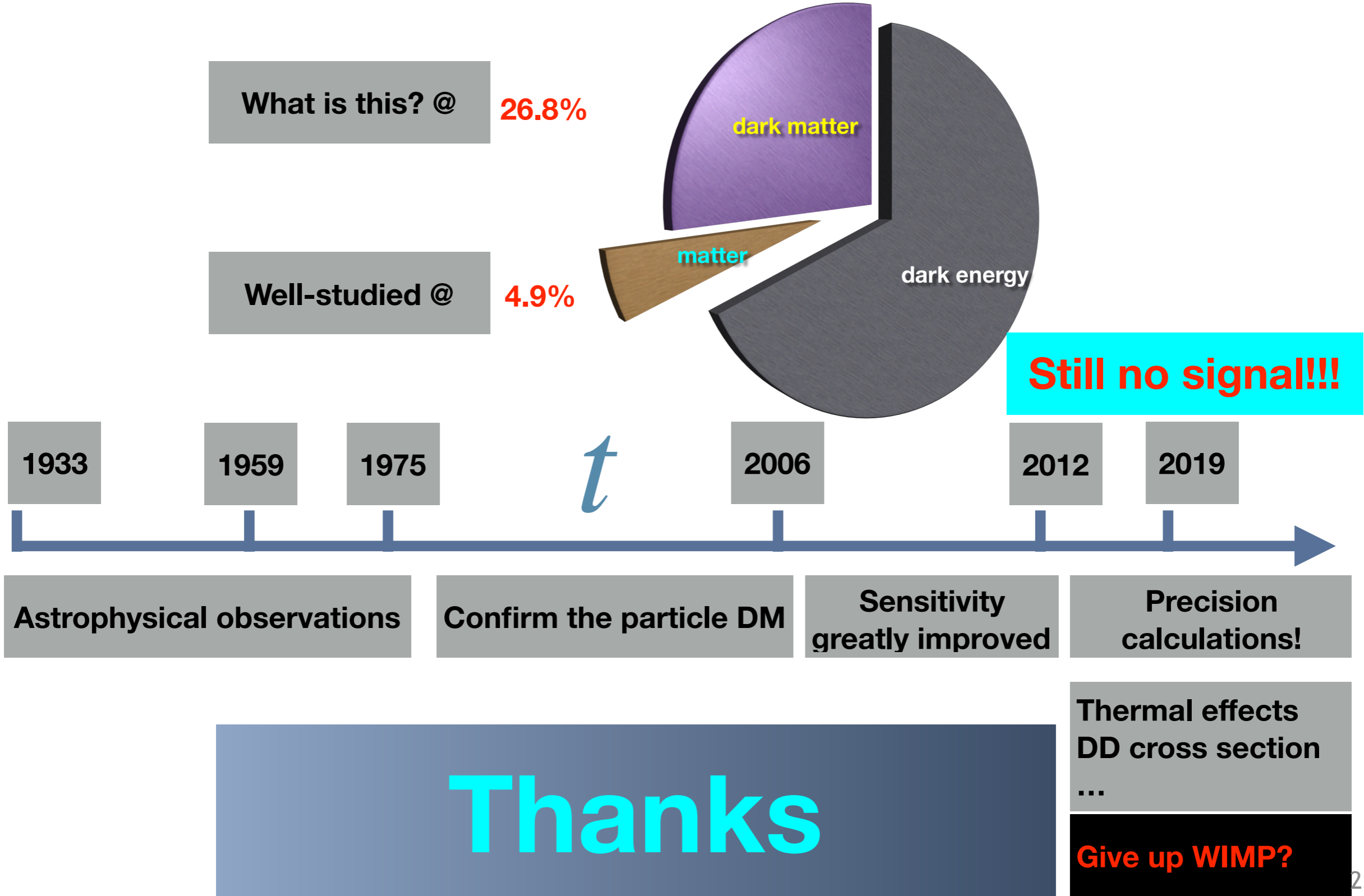
Numerical results



Conclusions

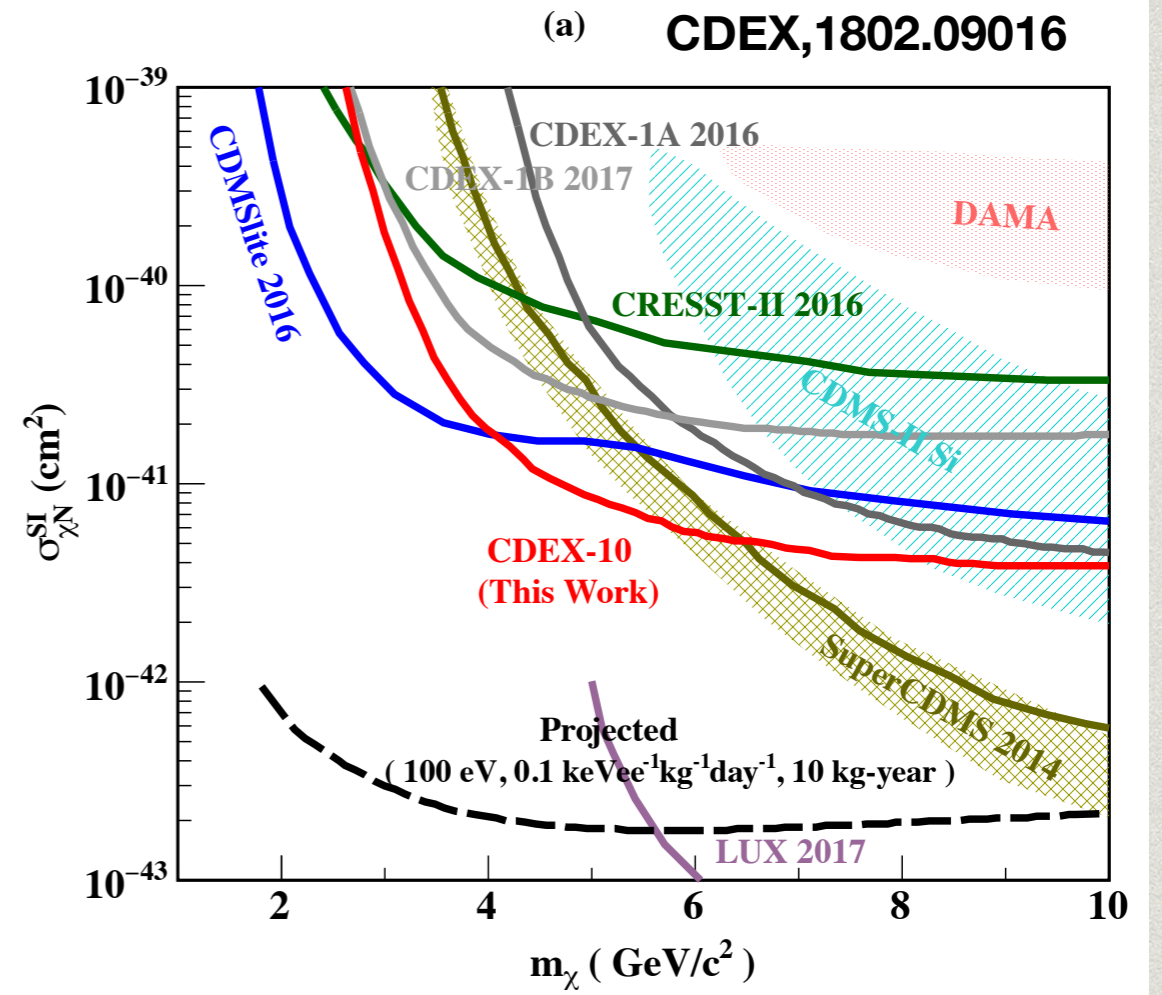
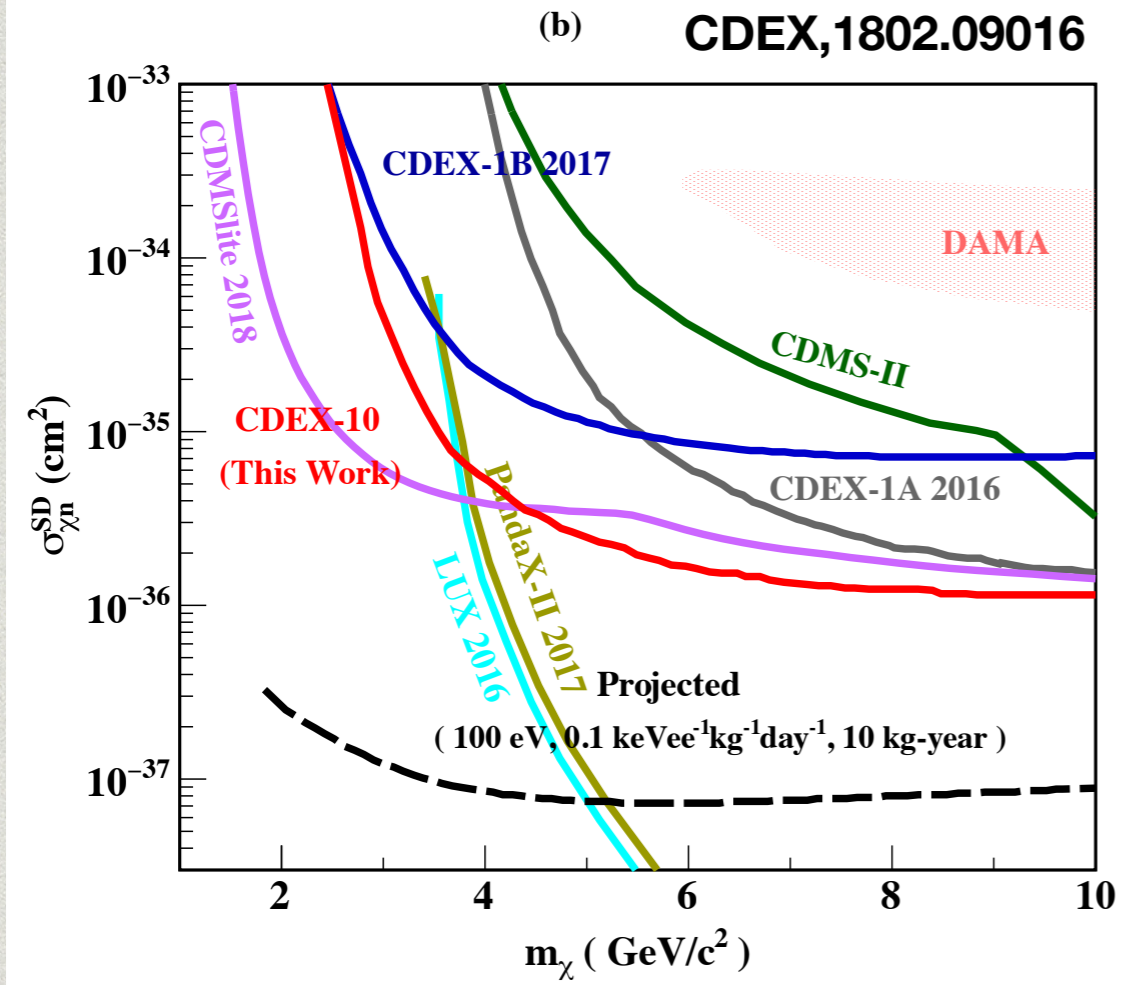
	Main points of DDs in gauge portal
1	Velocity suppressed cross section is reachable by the current DD technique!
2	Effects of twist-two operators are crucial for light DM.
3	Spin-dependent cross section may play important role in rolling out WIMPs
4	DD cross section of scalar-type electroweak multiplet DM are first analytically calculated by us!

Outlooking



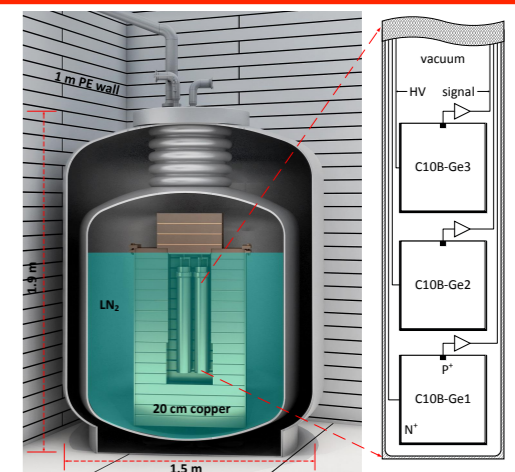
Back-up

Recent Progress-CDEX

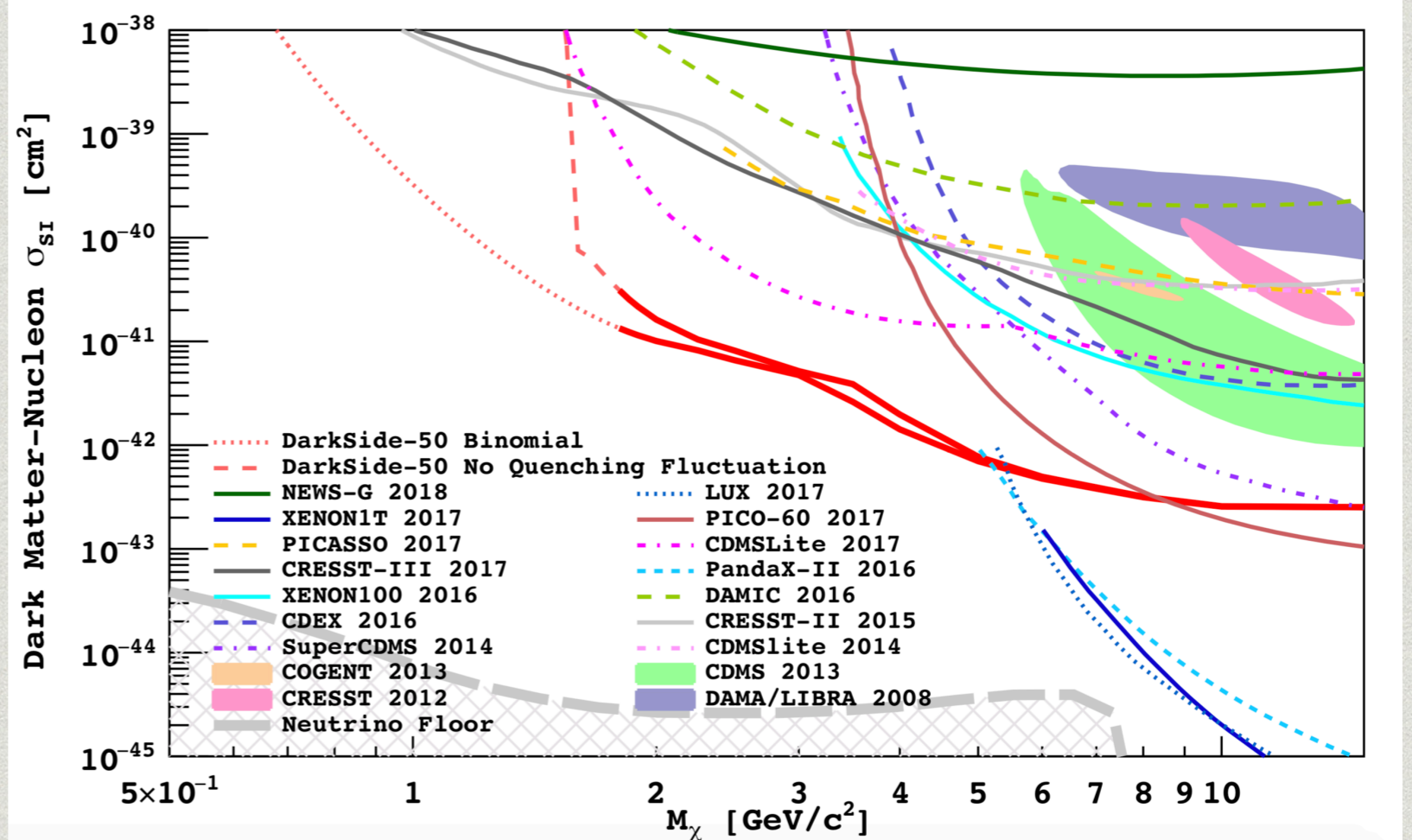
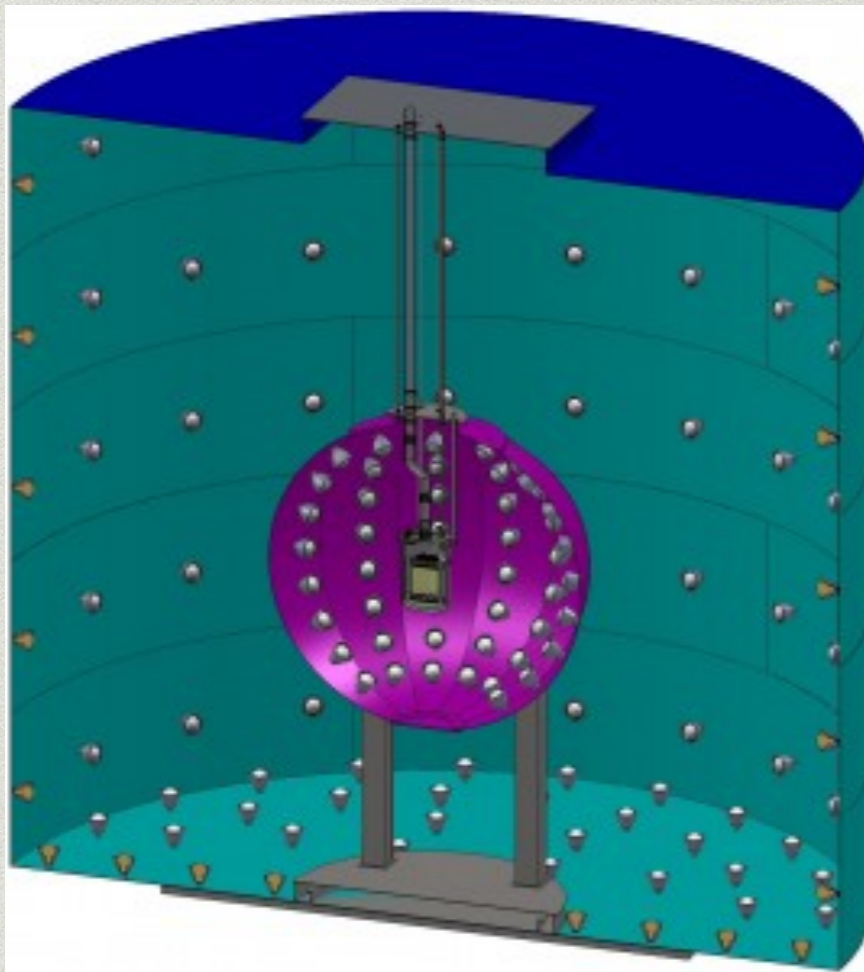


CDEX-10 ID:

- ◆ 10 Kg Ge
- ◆ Ionization
- ◆ 102.8 kg-days exposure
- ◆ Underground in CJPL



Recent Progress-DarkSide



ID of the DarkSide-50

a **50 kg** dual-phase Liquid **Argon** TPC

Using Underground Argon: **depleted in ³⁹Ar**

In a **30 ton** borated liquid scintillator **neutron veto**

In a **1000 ton Water** Cherenkov Veto

Underground in Gran Sasso National Lab, Italy

- ☼ **Dashed line: zero quench model, quenching a non-stochastic process**
- ☼ **Dotted line: binomial quench model, binomial quenching fluctuations**

The best exclusion limit in the range 1.8 GeV-6 GeV, 1802.06994

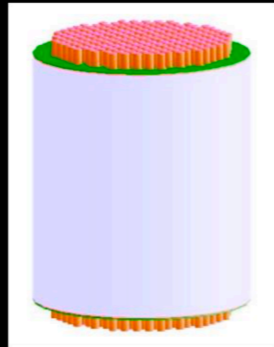
Recent Progress-PandaX



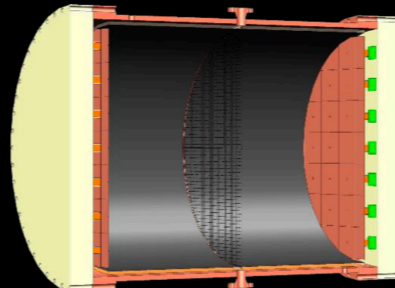
PandaX-I: 120 kg DM experiment
2009-2014



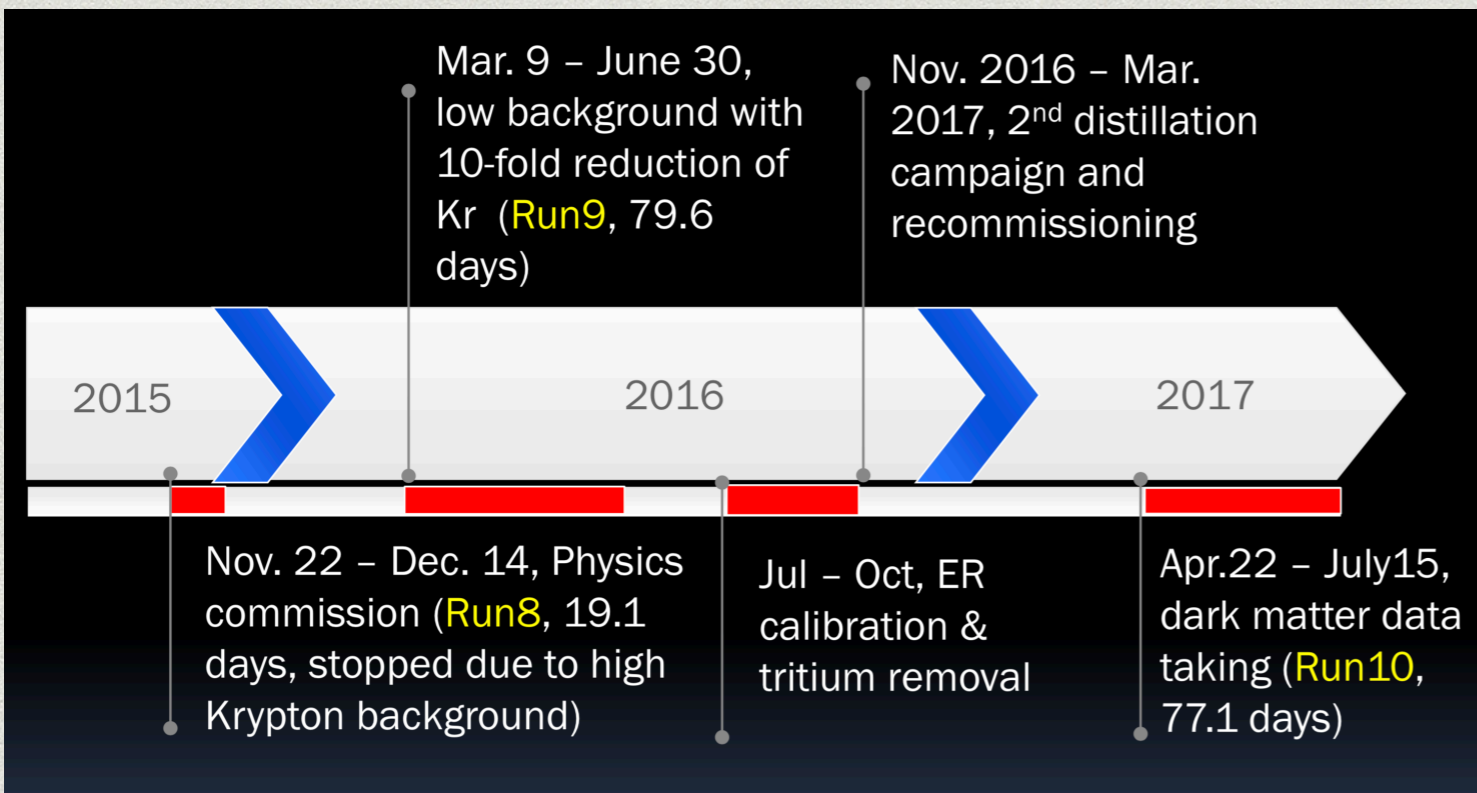
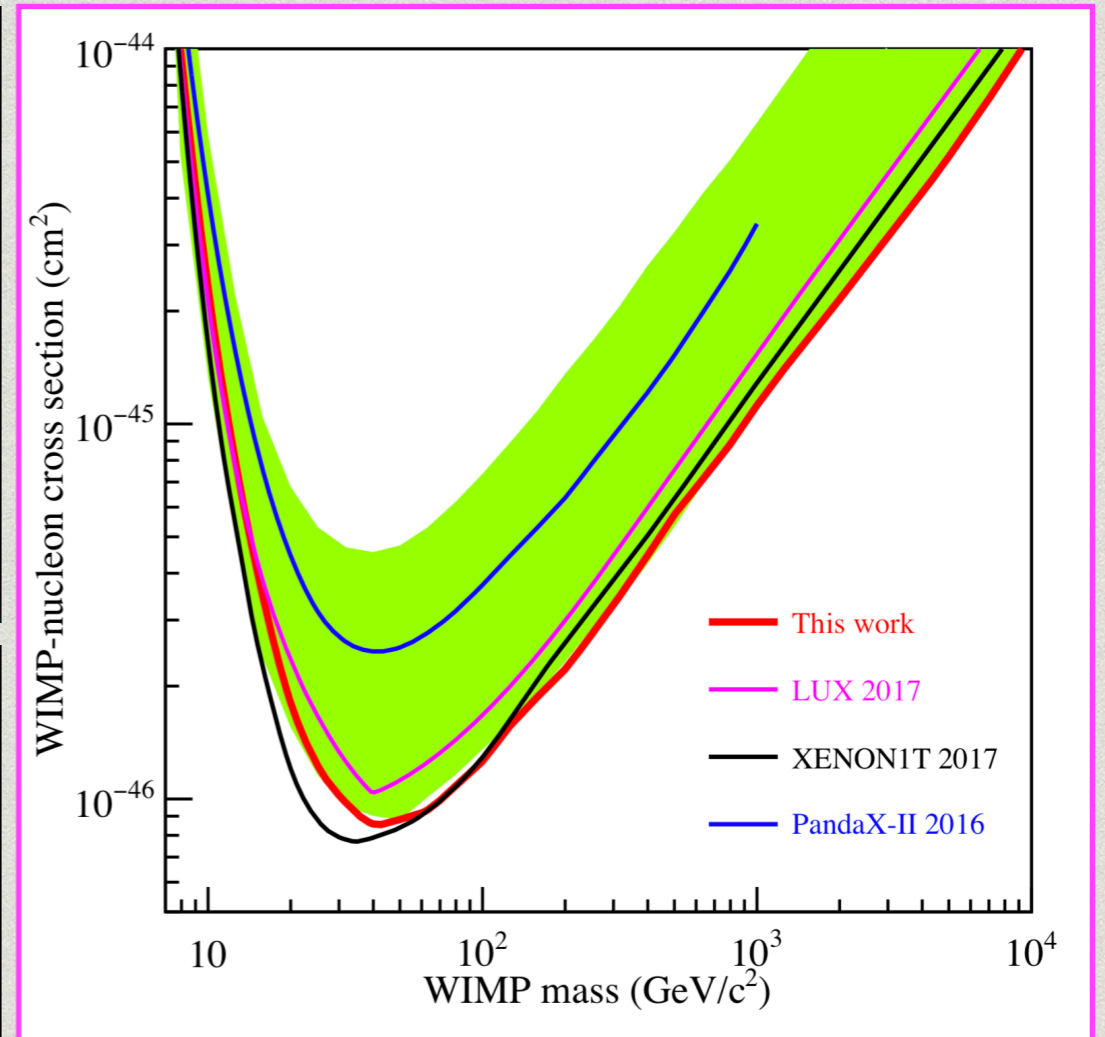
PandaX-II: 580 kg DM experiment
2014-2018



PandaX-xT: multi-ton DM experiment
Future



PandaX-III: 200 kg to 1 ton HP gas ^{136}Xe OvDBD experiment
Future



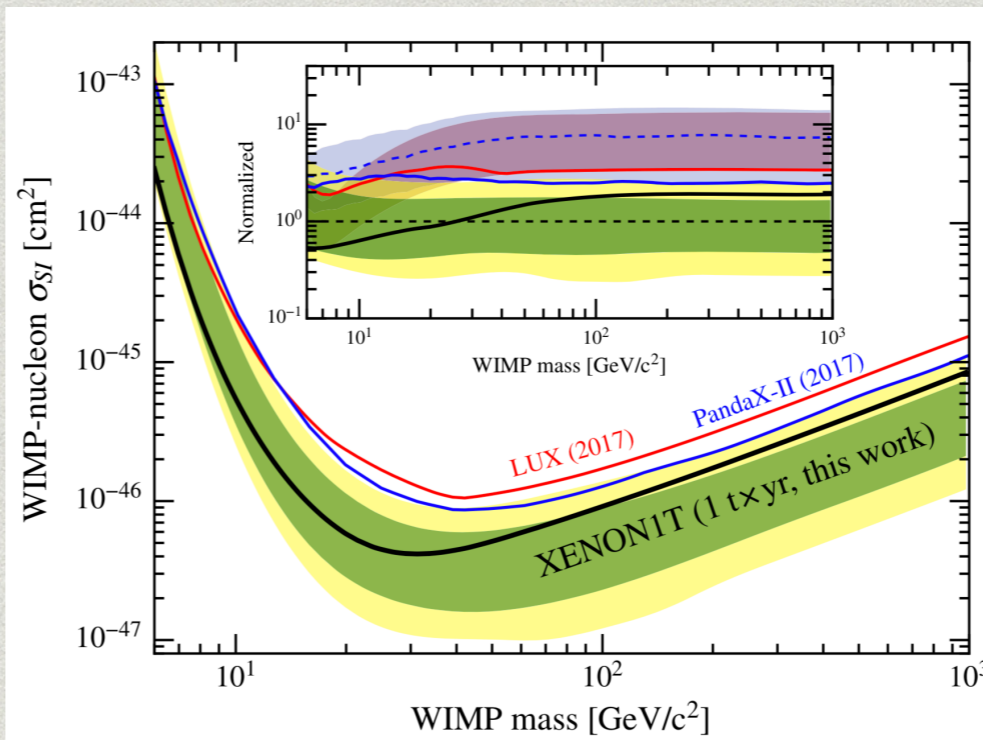
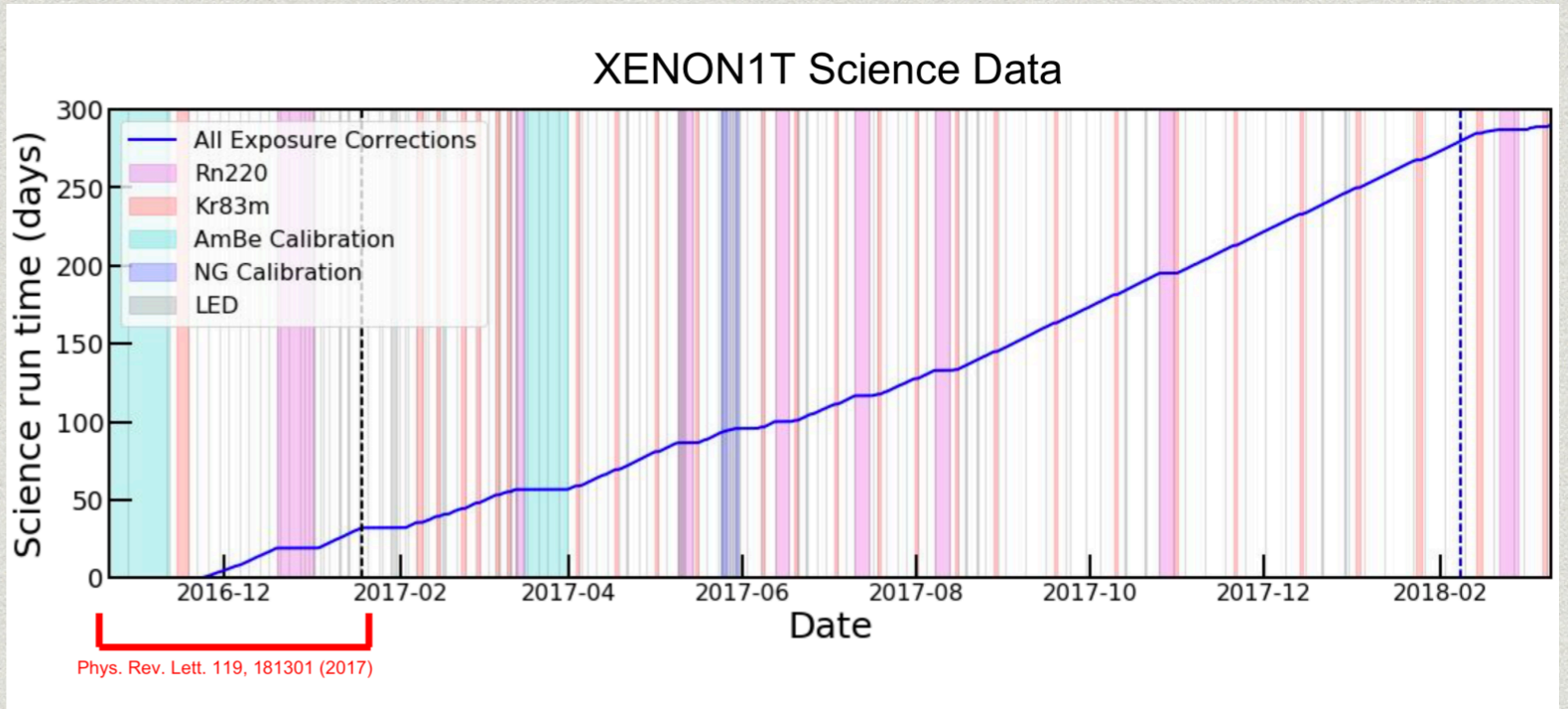
Exposure time: 54-Ton-Day

Recent Progress-XENON1T



XENON1T

Time: From 2016
Total: 3500 kg
Target: 2000 kg
Fiducial: >1000 kg
Sensitivity: $\sim 10^{-47} \text{ cm}^{-2}$



Exposure time

Updated exclusion limit

XENONnT

Time: From 2019
Total: 7500 kg
Target: 5900 kg
Fiducial: 4000 kg
Sensitivity: $\sim 10^{-48} \text{ cm}^{-2}$

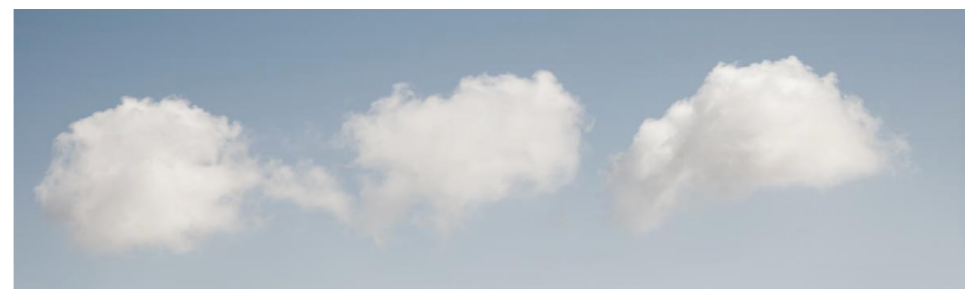
粒子物理天空中的乌云

二十世纪初物理学天空的乌云

历史总是有些相似!

- * 黑体辐射
- * 迈克逊-莫雷实验

量子力学
广义相对论



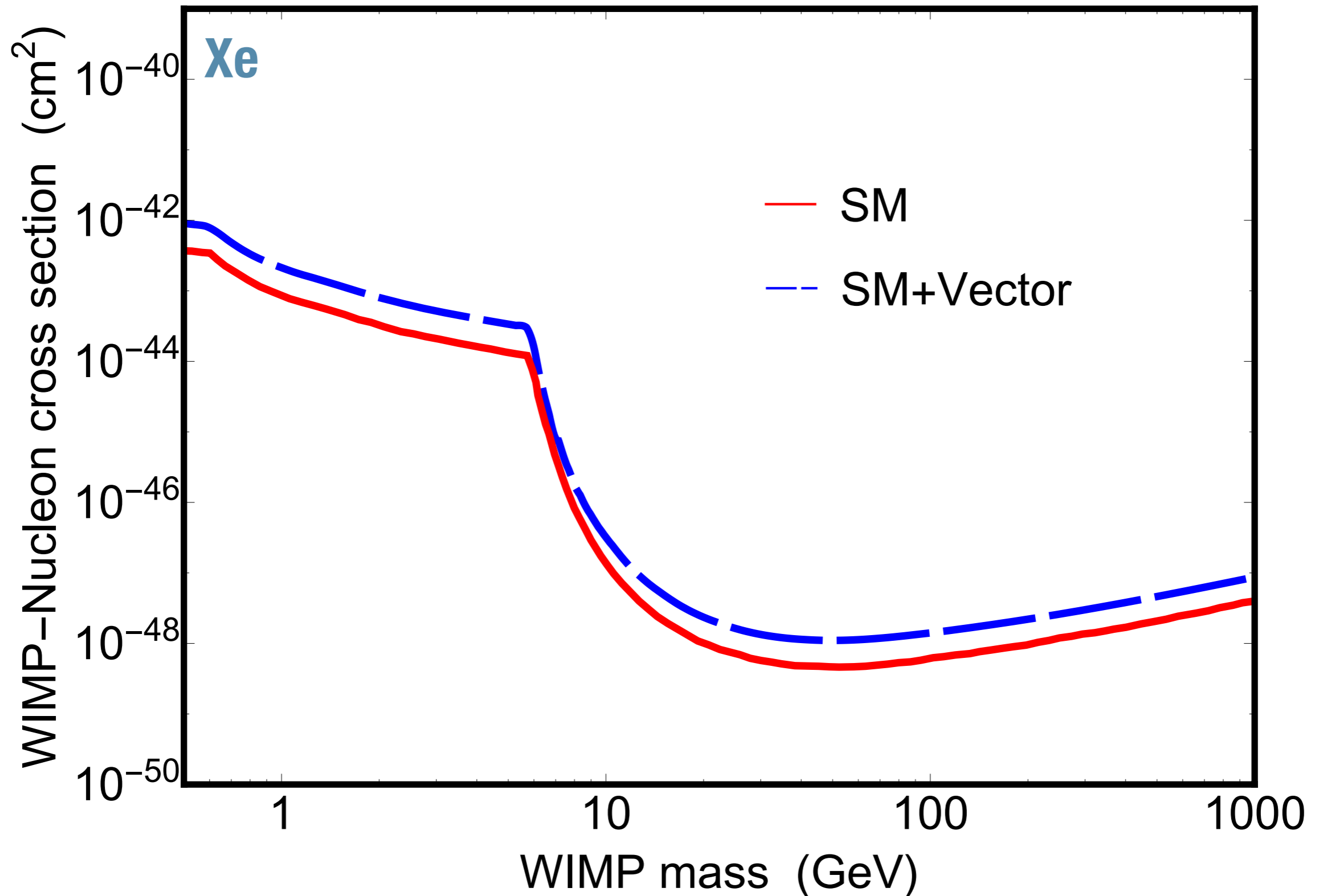
具有确凿试验证据的超出标准模型
的新物理

* 中微子物理;

* 暗物质;

* 重子数不对称;

Numerical results-1: vector interactions with **X131**



Numerical results-3: scalar interactions with **X131**

