

Hadron2019, Guilin, August 16-21, 2019

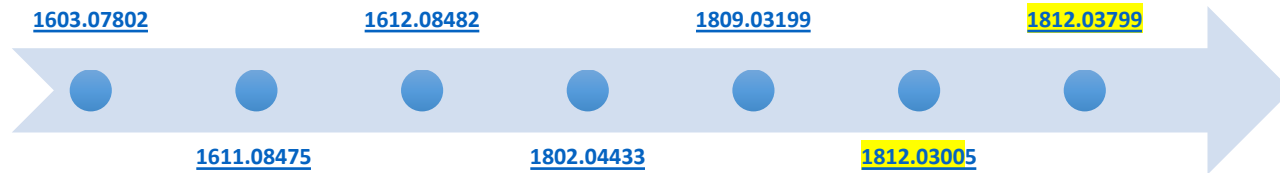


北京航空航天大学
BEIHANG UNIVERSITY



Recent **progress** in the construction of a **covariant chiral** nuclear force

Lisheng Geng (耿立升) @ Beihang U.

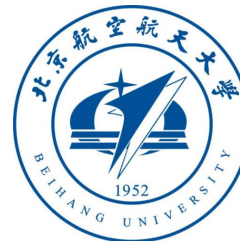


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Beihang: Yang Xiao, Jun-Xu Lu, Kai-Wen Li, Jie Song

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SCU: Bing-Wei Long

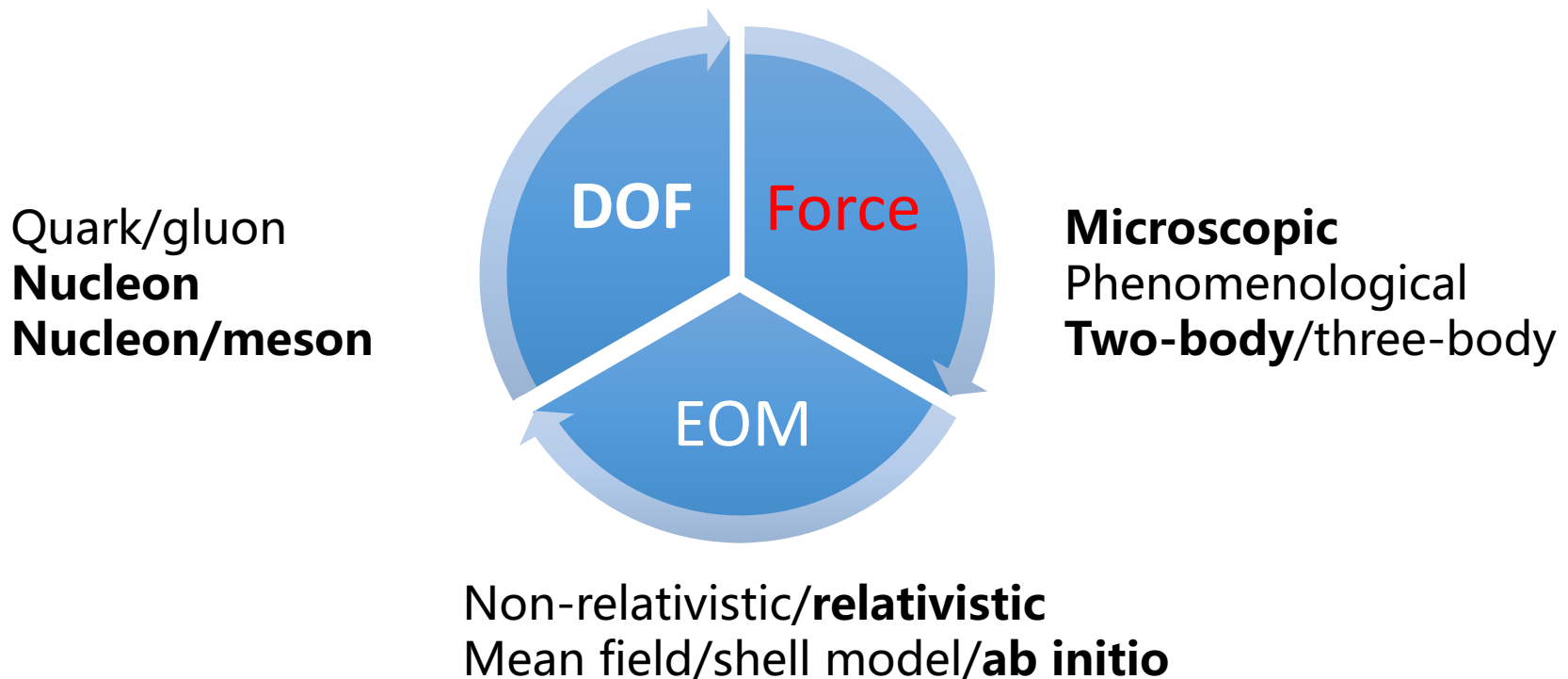


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- ① **Why relativistic/covariant chiral nuclear forces**
- ② **Our purpose and some exploratory studies**
- ③ **Covariant contact NN Lagrangians up to $O(q^4)$**
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- ⑤ ***Renormalization group invariance at LO***
- ⑥ **Summary and outlook**

Why (bare) nuclear forces

Understanding static and dynamic properties of nuclear systems in terms of bare nucleon-nucleon (NN, NNN) forces is one of the ultimate goals of nuclear theory



One of the most difficult problems

in the history of mankind

SCIENTIFIC AMERICAN, September 1953

What Holds the Nucleus Together?

by Hans A. Bethe

In the past quarter century physicists have devoted a huge amount of experimentation and mental labor to this problem – probably more man-hours than have been given to any other scientific question in the history of mankind.

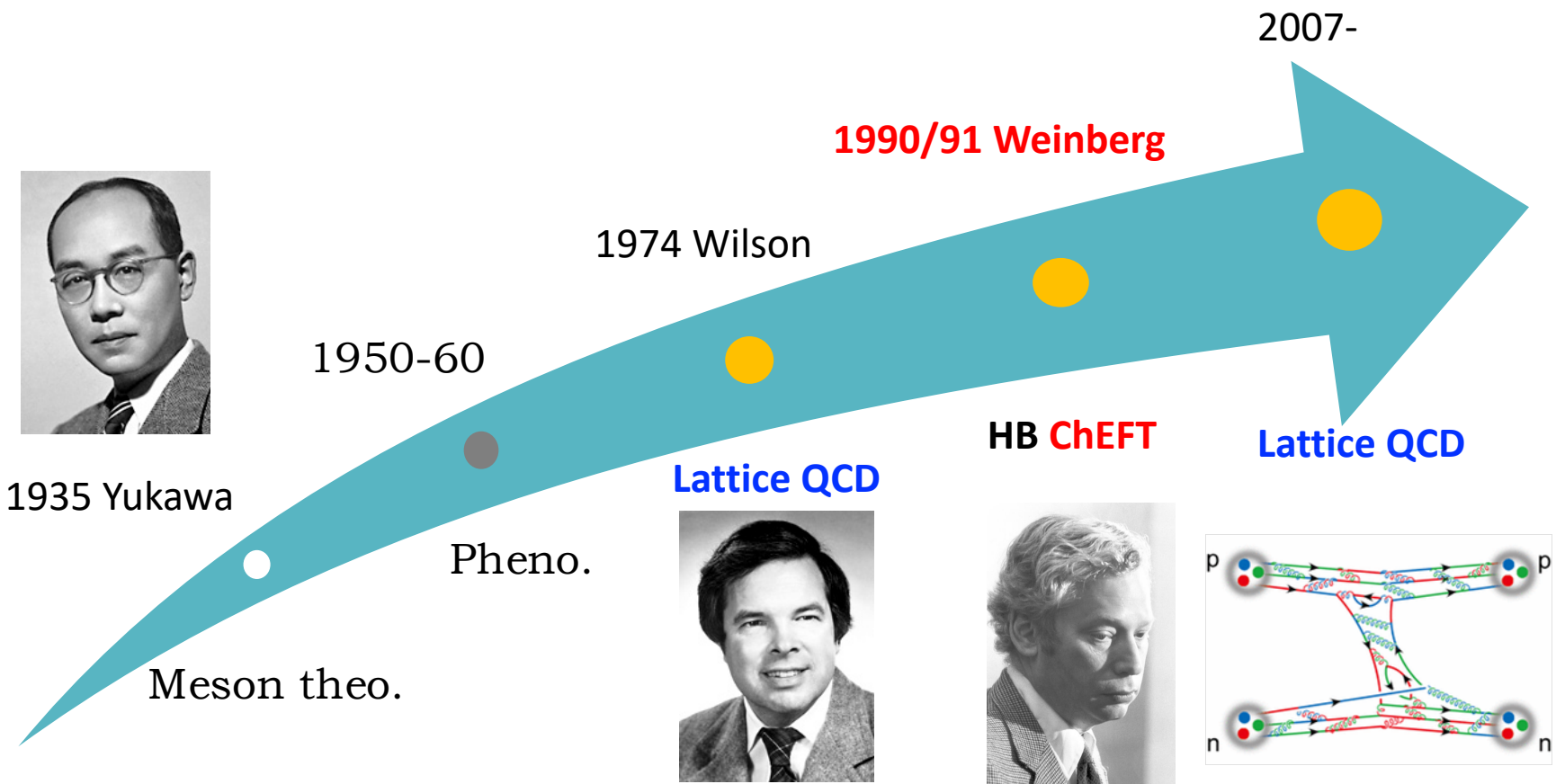


Hans Bethe
Nobel Prize in Physics
1967

A brief account of the long history

Advantages of ChEFT

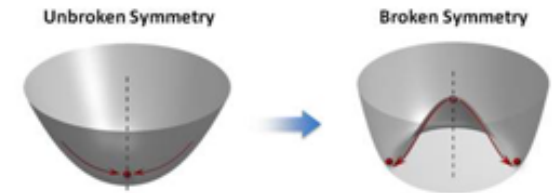
- ① Closer link with QCD
- ② Systematic/order-by-order improvements possible
- ③ Consistent descriptions of two/three/four body interactions on the same footing



Why chiral (effective field theory)

□ Chiral perturbation theory—low energy EFT of QCD

- Because of **quark confinement and asymptotic freedom**, low energy QCD can not be solved perturbatively
- Maps quark (u, d, s) dof's to those of the asymptotic states, hadrons
- Perturbative formulation of **low energy QCD** in powers of the external momenta and the light quark masses, by utilizing chiral symmetry and its breaking pattern (**the third feature of QCD**)



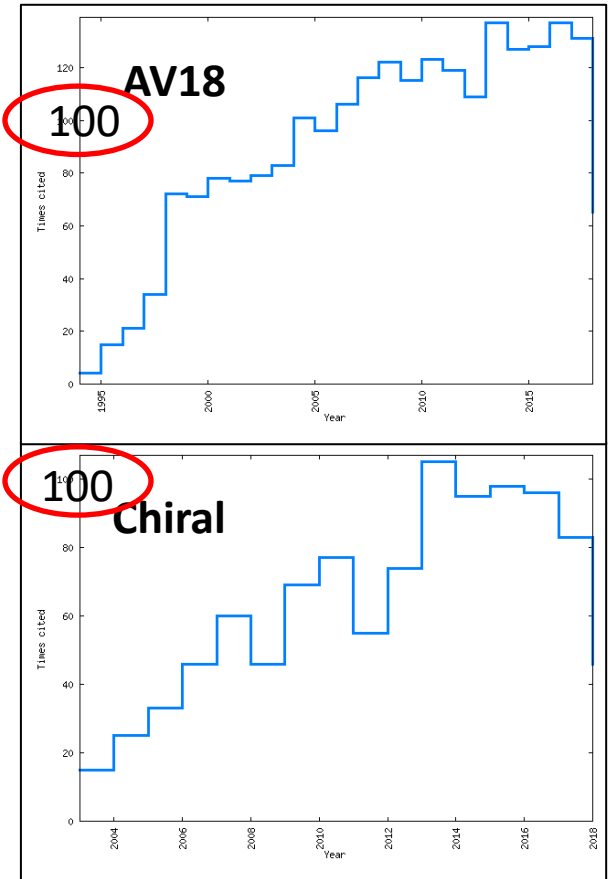
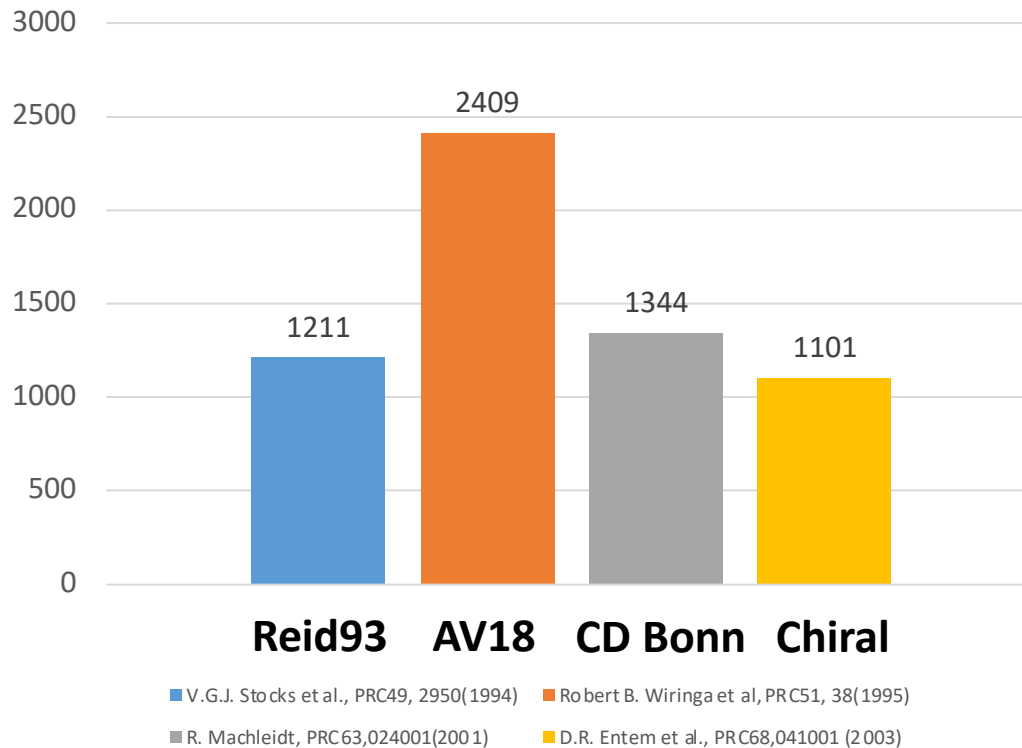
□ Development—Trilogy

- 1979, pion-pion, Weinberg
- 1989, to the one-baryon sector, Gasser, Sainio, Svarc
- 1990/91, to NN, Weinberg—**very successful**



Steven Weinberg
Nobel Prize in Physics in 1979

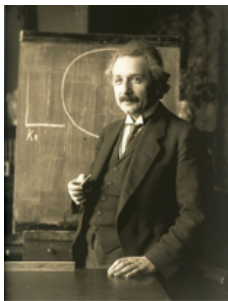
Why (bare) nuclear forces



Yearly citation about 100 times

Why covariant/relativistic

- Lorentz invariance is one of **the most important symmetries of nature**.
- Include kinematical and dynamical relativistic corrections self-consistently and simultaneously
- Relativistic approaches successful in explaining **fine structures**
 - Atomic and molecular systems, why gold is yellow
 - Nuclear system : **spin-orbit splitting, pseudospin symmetry**
 - One-baryon sector : magnetic moments, masses, sigma terms



Einstein



Dirac



Mayer



Jensen



Arima

What it is not

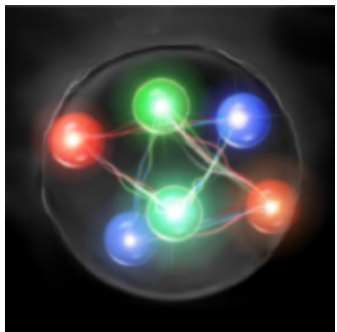
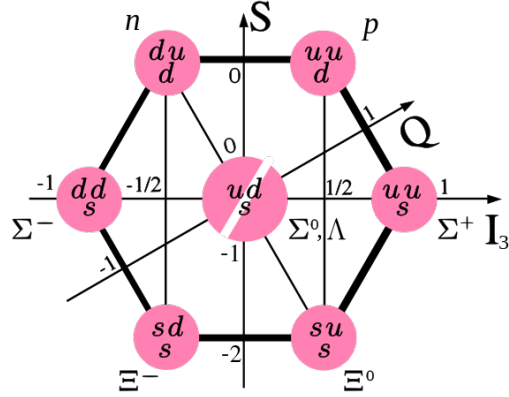
- ***Not*** to challenge or replace the successful Weinberg chiral nuclear force
- ***Not*** to compete with phenomenological forces in terms of agreement with NN phaseshifts

Our purpose 1: **Exploration**

- Provide a high precision covariant chiral nuclear force that can be used as inputs for ab initio nuclear structure and reaction studies **in a covariant setting**
- Explore the consequence of such a covariant framework in solving some long-standing open questions, such as RG invariance, A_y puzzle, ...

Our purpose 2: from NN to BB

- Nucleon-nucleon
- Hyperon-nucleon
- Hyperon-hyperon



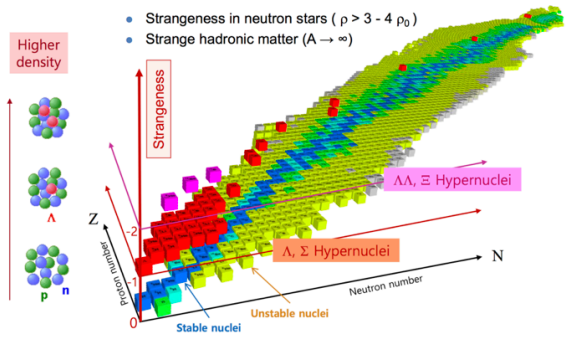
A bound H-dibaryon?

Inoue PRL 106 (2011) 162002



Neutron star

Lonardoni PRL 114 (2015) 092301

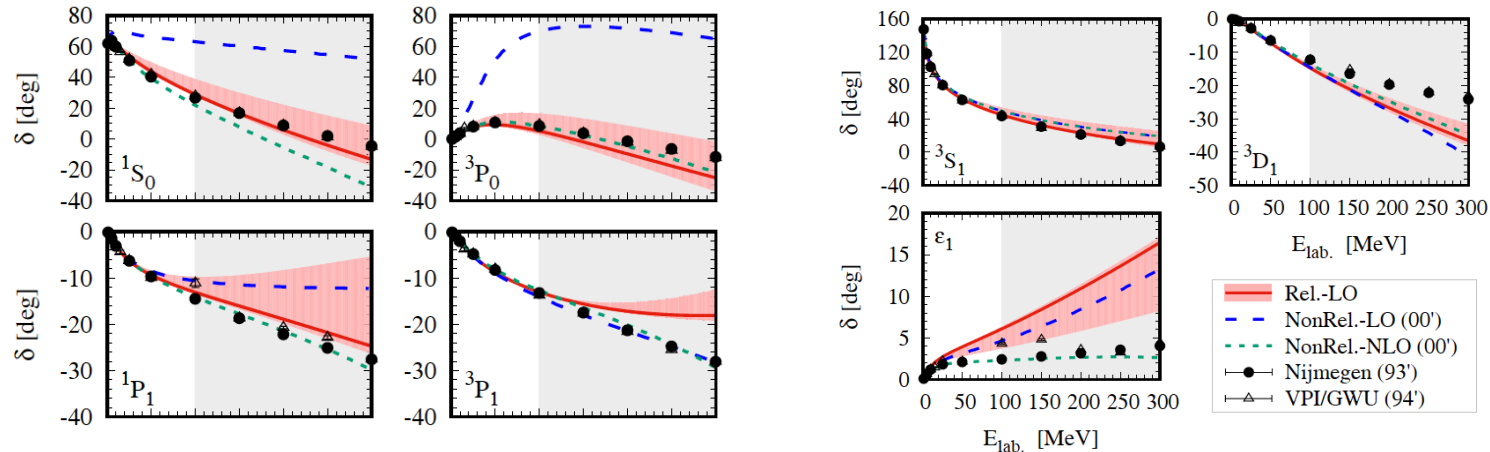


Three D nuclear chart

Kaneta M, Tohoku University, Japan)

Exploratory/feasibility studies

① Leading order relativistic chiral nucleon-nucleon interaction, 1611.08475



② Leading order relativistic hyperon-nucleon interactions in chiral effective field theory, 1612.08482

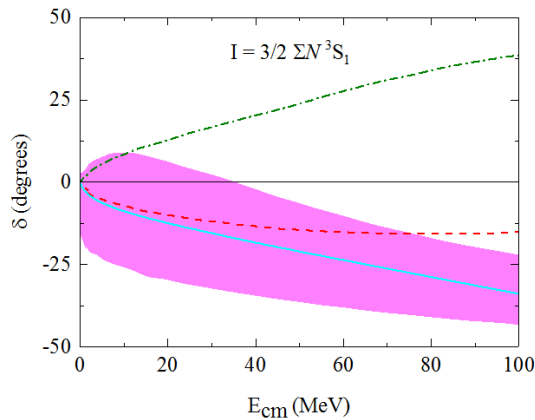
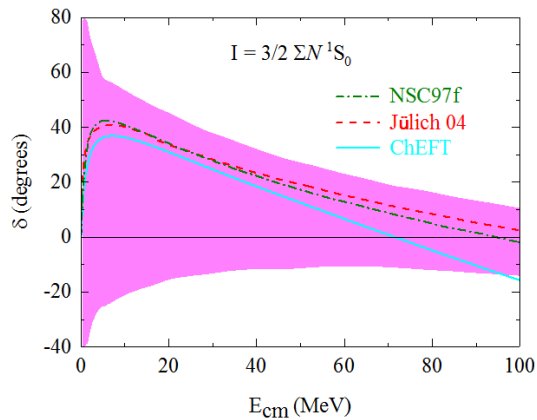
36 YN data	Weinberg's approach	Covariant ChEFT	NSC97f [§]
No. of LECs	5 (LO*)	23 (NLO [#])	29
χ^2	28.3	16.2	16.7

*Polinder NPA799 (2006) 244

[#]Haidenbauer NPA915 (2013) 24

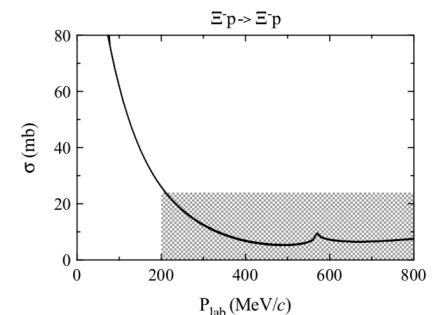
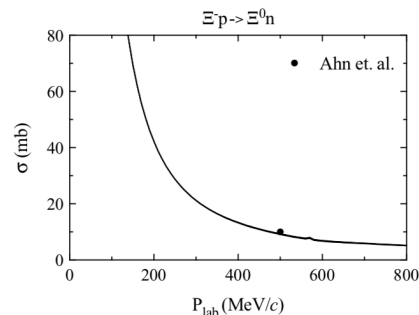
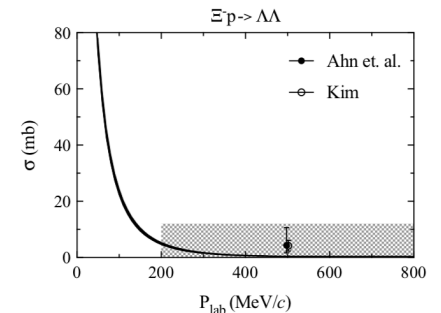
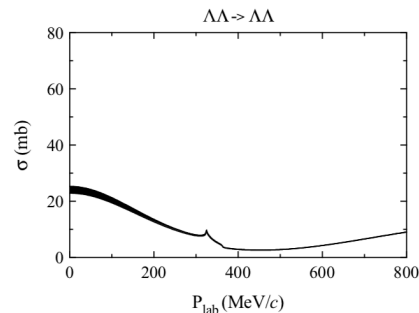
[§]Rijken PRC 59 (1999) 21

Exploratory/feasibility studies



③ Strangeness $S=-1$ hyperon-nucleon interactions: Chiral effective field theory versus lattice QCD, [1802.04433](#)

④ Strangeness $S=-2$ baryon-baryon interactions in relativistic chiral effective field theory, [1809.03199](#)



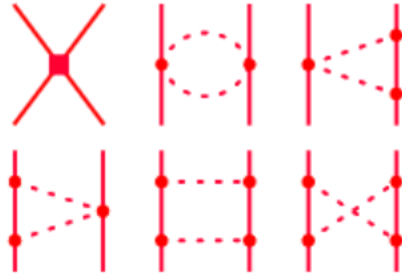
Going to higher orders—challenging

2N Force

LO



NNLO



N3LO



N4LO



Three key inputs

- 1) Four nucleon (baryon) vertices, [1812.03005](#)
- 2) Meson-baryon vertices, [1812.03799](#)
- 3) *Two-pion exchanges, to come soon*

3: How to construct non-relativistic Lagrangians

- **Basic requirement:** A scalar.
- **Building blocks:** $N, N^\dagger, \sigma, \vec{\nabla}, \overleftarrow{\nabla}$.
- **Symmetry constraint:** Even ∇ operators (Parity).
- **Power counting rules:**
 - $\vec{\nabla} \sim O(p^1), \overleftarrow{\nabla} \sim O(p^1), \sigma \sim O(p^0), N \sim O(p^0), N^\dagger \sim O(p^0)$.
- **Linear relations (center-of-mass frame):**
 1. $a \cdot b = b \cdot a, a \times b = -b \times a,$
 2. $a \cdot b \times c = -b \cdot a \times c = -c \cdot b \times a,$
 3. $a \times (b \times c) = b(a \cdot c) - c(a \cdot b),$
 4. $(a \times b) \cdot (c \times d) = (a \cdot c)(b \cdot d) - (a \cdot d)(b \cdot c),$
 5. $(N^\dagger N)(N^\dagger \vec{\nabla} N) \equiv -(N^\dagger \vec{\nabla} N)(N^\dagger N),$
 6. $(N^\dagger N)(N^\dagger \overleftarrow{\nabla} N) \equiv -(N^\dagger \overleftarrow{\nabla} N)(N^\dagger N).$

$$a, b, c, d \in \{\sigma, \vec{\nabla}, \overleftarrow{\nabla}\}$$

3: Non-relativistic Lagrangians (NNLO)

O_S	$(N^\dagger N)(N^\dagger N)$	O_{11}	$(N^\dagger \vec{\nabla} \cdot \overleftarrow{\nabla} N)(N^\dagger \vec{\nabla} \cdot \overleftarrow{\nabla} N)$
O_T	$(N^\dagger \boldsymbol{\sigma} N) \cdot (N^\dagger \boldsymbol{\sigma} N)$	O_{12}	$i(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} \times \overleftarrow{\nabla} N)(N^\dagger \vec{\nabla}^2 N) + \text{h.c.}$
O_1	$(N^\dagger N)(N^\dagger \vec{\nabla}^2 N) + \text{h.c.}$	O_{13}	$i(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} \times \overleftarrow{\nabla} N)(N^\dagger \vec{\nabla} \cdot \overleftarrow{\nabla} N)$
O_2	$(N^\dagger N)(N^\dagger \vec{\nabla} \cdot \overleftarrow{\nabla} N)$	O_{14}	$(N^\dagger \sigma^j N)(N^\dagger \sigma^j \vec{\nabla}^4 N) + \text{h.c.}$
O_3	$i(N^\dagger \boldsymbol{\sigma} N) \cdot (N^\dagger \vec{\nabla} \times \overleftarrow{\nabla} N)$	O_{15}	$(N^\dagger \sigma^j \vec{\nabla} \cdot \overleftarrow{\nabla} N)(N^\dagger \sigma^j \vec{\nabla}^2 N) + \text{h.c.}$
O_4	$(N^\dagger \sigma^j N)(N^\dagger \sigma^j \vec{\nabla}^2 N) + \text{h.c.}$	O_{16}	$(N^\dagger \sigma^j \vec{\nabla}^2 N)(N^\dagger \sigma^j \overleftarrow{\nabla}^2 N)$
O_5	$(N^\dagger \sigma^j N)(N^\dagger \sigma^j \vec{\nabla} \cdot \overleftarrow{\nabla} N)$	O_{17}	$(N^\dagger \sigma^j \vec{\nabla} \cdot \overleftarrow{\nabla} N)(N^\dagger \sigma^j \vec{\nabla} \cdot \overleftarrow{\nabla} N)$
O_6	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N) + \text{h.c.}$	O_{18}	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} \vec{\nabla}^2 N) + \text{h.c.}$
O_7	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \overleftarrow{\nabla} N)$	O_{19}	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \overleftarrow{\nabla} \vec{\nabla}^2 N) + \text{h.c.}$
O_8	$(N^\dagger N)(N^\dagger \vec{\nabla}^4 N) + \text{h.c.}$	O_{20}	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} \vec{\nabla} \cdot \overleftarrow{\nabla} N) + \text{h.c.}$
O_9	$(N^\dagger \vec{\nabla}^2 N)(N^\dagger \vec{\nabla} \cdot \overleftarrow{\nabla} N) + \text{h.c.}$	O_{21}	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} \overleftarrow{\nabla}^2 N) + \text{h.c.}$
O_{10}	$(N^\dagger \vec{\nabla}^2 N)(N^\dagger \overleftarrow{\nabla}^2 N)$	O_{22}	$(N^\dagger \boldsymbol{\sigma} \cdot \vec{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \overleftarrow{\nabla} \vec{\nabla} \cdot \overleftarrow{\nabla} N)$

2+7+15 terms up to NNLO, consistent with the known Weinberg chiral nuclear force

3: How to construct **cov.** Lagrangians

□ Symmetry Constraints

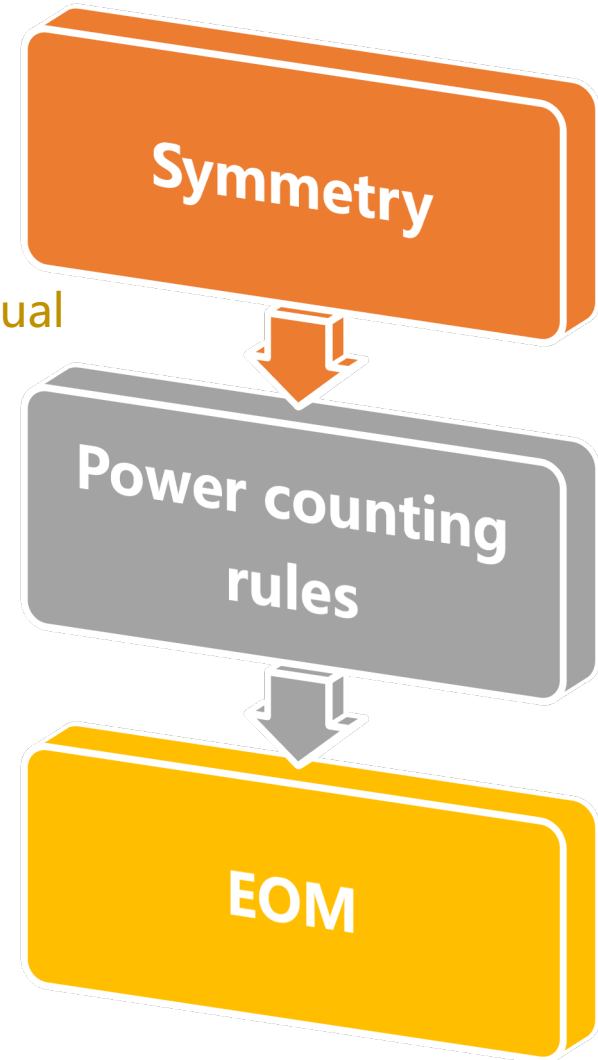
- ✓ Lorentz invariance: $\alpha, \beta, \gamma \dots$
- ✓ Chiral symmetry: matter field $\psi \rightarrow K\psi K^\dagger$, NGB as usual
- ✓ Hermitian conjugation: add an appropriate "i" .
- ✓ Parity and Charge conjugation symmetries:
- ✓ Time inverse symmetry: CPT theory.

□ How to raise chiral order ?

→ **Power counting rules**

□ How to deal with redundant terms ?

→ **Equation of motion (EOM)**



3: Symmetry requirements

□ Building blocks (Dirac matrices & partial derivatives)

	$\mathbb{1}$	γ_5	γ_μ	$\gamma_5\gamma_\mu$	$\sigma_{\mu\nu}$	$\epsilon_{\mu\nu\rho\sigma}$	$\overleftrightarrow{\partial}_\mu$	∂_μ
\mathcal{P}	+	-	+	-	+	-	+	+
\mathcal{C}	+	+	-	+	-	+	-	+
h.c.	+	-	+	+	+	+	-	+
\mathcal{O}	0	1	0	0	0	-	0	1

□ General form of a Lagrangian term

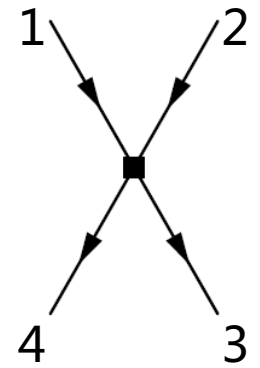
$$\frac{1}{(2m)^{N_d}} \left(\bar{\psi} i \overleftrightarrow{\partial}^\alpha i \overleftrightarrow{\partial}^\beta \dots \Gamma_A \psi \right) \partial^\lambda \partial^\mu \dots \left(\bar{\psi} i \overleftrightarrow{\partial}^\sigma i \overleftrightarrow{\partial}^\tau \dots \Gamma_B \psi \right),$$

Note $\overleftrightarrow{\partial}^\alpha = \bar{\psi}(\overrightarrow{\partial}^\alpha - \overleftarrow{\partial}^\alpha)\psi$ vs. $\partial^\alpha = \partial^\alpha(\bar{\psi}\Gamma\psi)$

3: Power counting rules

$$\frac{1}{(2m)^{N_d}} \left(\bar{\psi} i \overleftrightarrow{\partial}^\alpha i \overleftrightarrow{\partial}^\beta \dots \Gamma_A \psi \right) \partial^\lambda \partial^\mu \dots \left(\bar{\psi} i \overleftrightarrow{\partial}^\sigma i \overleftrightarrow{\partial}^\tau \dots \Gamma_B \psi \right) \quad N_d \text{ is the number of } \overleftrightarrow{\partial} = \vec{\partial} - \partial$$

- Nucleon field: $\psi = \binom{p}{n} \sim O(p^0)$, Nucleon mass: $m \sim O(p^0)$,
- Dirac matrices: $\Gamma \in \{1, \gamma_\mu, \gamma_5 \gamma_\mu, \sigma_{\mu\nu} \sim O(p^0), \gamma_5 \sim O(p^1)\}$
- **Covariant derivative:** $\partial(\bar{\psi} \Gamma \psi) \sim O(p^1)$, $(\bar{\psi} \overleftrightarrow{\partial} \psi) \sim O(p^0)$, **except**
 $(\bar{\psi} \sigma_{\mu\nu} \psi)(\bar{\psi} \overleftrightarrow{\partial}^\mu \Gamma \psi) \sim O(p^1)$, $(\bar{\psi} \gamma_5 \gamma_\mu \psi)(\bar{\psi} \overleftrightarrow{\partial}^\mu \Gamma \psi) \sim O(p^1)$
- **Treatment for covariant derivative:**



$$\tilde{O}_{\Gamma_A \Gamma_B}^{(n)} = \frac{1}{(2m)^{2n}} \left(\bar{\psi} i \overleftrightarrow{\partial}^{\mu_1} i \overleftrightarrow{\partial}^{\mu_2} \dots i \overleftrightarrow{\partial}^{\mu_n} \Gamma_A^\alpha \psi \right) \left(\bar{\psi} i \overleftrightarrow{\partial}_{\mu_1} i \overleftrightarrow{\partial}_{\mu_2} \dots i \overleftrightarrow{\partial}_{\mu_n} \Gamma_B^\alpha \psi \right)$$

-Expansion of such structure:

-up to $O(q^2)$: $n = 0, 1$;

-up to $O(q^4)$: $n = 0, 1, 2$.

$$\boxed{\frac{[(p_1 + p_3) \cdot (p_2 + p_4)]^n}{(2m)^{2n}}} \longleftrightarrow \boxed{\left[1 + \frac{(s - 4m^2) - u}{4m^2} \right]^n}$$

3: Reduction using equation of motion (EOM)

□ **Equation of motion** : $\not{D}B = \gamma^\mu D_\mu B = -iM_0 B + \mathcal{O}(q)$

□ Beyond the obvious replacements one can bring terms that do not containing $\not{D}B$ into a form where they do. *Annals Phys., 283:273, (2000)*

$$-2im(\bar{\psi}\Gamma\psi) \approx 2(\bar{\psi}\Gamma \times \gamma_\lambda \partial^\lambda \psi) = (\bar{\psi}\Gamma'_\lambda \overleftrightarrow{\partial}^\lambda \psi) + \partial^\lambda (\bar{\psi}\Gamma''_\lambda \psi),$$

TABLE II. Decomposition of the Dirac matrix products $\Gamma \times \gamma_\lambda$ into charge conjugation even (Γ'_λ) and charge conjugation odd (Γ''_λ) parts [43].

Γ	Γ'_λ	Γ''_λ
$\mathbb{1}$	γ_λ	0
γ_μ	$g_{\mu\lambda}\mathbb{1}$	$-i\sigma_{\mu\lambda}$
γ_5	0	$\gamma_5\gamma_\lambda$
$\gamma_5\gamma_\mu$	$\frac{1}{2}\epsilon_{\mu\lambda\rho\tau}\sigma^{\rho\tau}$	$g_{\mu\lambda}\gamma_5$
$\sigma_{\mu\nu}$	$\epsilon_{\mu\nu\lambda\tau}\gamma_5\gamma^\tau$	$-i(g_{\mu\lambda}\gamma_\nu - g_{\nu\lambda}\gamma_\mu)$
$\epsilon_{\mu\nu\rho\tau}\gamma^\tau$	$\epsilon_{\mu\nu\rho\lambda}\mathbb{1}$	$g_{\mu\lambda}\gamma_5\sigma_{\nu\rho} + g_{\rho\lambda}\gamma_5\sigma_{\mu\nu} + g_{\nu\lambda}\gamma_5\sigma_{\rho\mu}$
$\epsilon_{\mu\nu\rho\tau}\gamma_5\gamma^\tau$	$g_{\mu\lambda}\sigma_{\nu\rho} + g_{\rho\lambda}\sigma_{\mu\nu} + g_{\nu\lambda}\sigma_{\rho\mu}$	$\epsilon_{\mu\nu\rho\lambda}\gamma_5$
$\epsilon_{\mu\nu\rho\alpha}\sigma_\tau^\alpha$	$\gamma_5\gamma_\rho(g_{\lambda\nu}g_{\mu\tau} - g_{\lambda\mu}g_{\nu\tau}) + \gamma_5\gamma_\nu(g_{\lambda\mu}g_{\rho\tau} - g_{\lambda\rho}g_{\mu\tau}) + \gamma_5\gamma_\mu(g_{\lambda\rho}g_{\nu\tau} - g_{\lambda\nu}g_{\rho\tau})$	$i g_{\lambda\tau}\epsilon_{\mu\nu\rho\alpha}\gamma^\alpha - i\epsilon_{\mu\nu\rho\lambda}\gamma_\tau$
$\frac{i}{2}\epsilon_{\mu\nu\rho\tau}\sigma^{\rho\tau} = \gamma_5\sigma_{\mu\nu}$	$\frac{1}{i}(g_{\mu\lambda}\gamma_5\gamma_\nu - g_{\nu\lambda}\gamma_5\gamma_\mu)$	$\epsilon_{\mu\nu\lambda\rho}\gamma^\rho$

3: Covariant NN contact Lagrangians (N2LO)

\tilde{O}_1	$(\bar{\psi}\psi)(\bar{\psi}\psi)$
\tilde{O}_2	$(\bar{\psi}\gamma^\mu\psi)(\bar{\psi}\gamma_\mu\psi)$
\tilde{O}_3	$(\bar{\psi}\gamma_5\gamma^\mu\psi)(\bar{\psi}\gamma_5\gamma_\mu\psi)$
\tilde{O}_4	$(\bar{\psi}\sigma^{\mu\nu}\psi)(\bar{\psi}\sigma_{\mu\nu}\psi)$
\tilde{O}_5	$(\bar{\psi}\gamma_5\psi)(\bar{\psi}\gamma_5\psi)$
\tilde{O}_6	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\alpha i\overleftrightarrow{\partial}_\mu\psi)$
\tilde{O}_7	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu} i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\alpha} i\overleftrightarrow{\partial}_\nu\psi)$
\tilde{O}_8	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\mu\psi)\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
\tilde{O}_9	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\alpha}\psi)\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
\tilde{O}_{10}	$\frac{1}{4m^2}(\bar{\psi}\psi)\partial^2(\bar{\psi}\psi)$
\tilde{O}_{11}	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_\mu\psi)$
\tilde{O}_{12}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\mu\psi)$
\tilde{O}_{13}	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}\psi)\partial^2(\bar{\psi}\sigma_{\mu\nu}\psi)$
\tilde{O}_{14}	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_1$
\tilde{O}_{15}	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_2$
\tilde{O}_{16}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_3$
\tilde{O}_{17}	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu} i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\nu} i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_4$

O_S	$(N^\dagger N)(N^\dagger N)$
O_T	$(N^\dagger \boldsymbol{\sigma} N) \cdot (N^\dagger \boldsymbol{\sigma} N)$
O_1	$(N^\dagger N)(N^\dagger \overleftrightarrow{\nabla}^2 N) + \text{h.c.}$
O_2	$(N^\dagger N)(N^\dagger \overleftrightarrow{\nabla} \cdot \overleftrightarrow{\nabla} N)$
O_3	$i(N^\dagger \boldsymbol{\sigma} N) \cdot (N^\dagger \overleftrightarrow{\nabla} \times \overleftrightarrow{\nabla} N)$
O_4	$(N^\dagger \sigma^j N)(N^\dagger \sigma^j \overleftrightarrow{\nabla}^2 N) + \text{h.c.}$
O_5	$(N^\dagger \sigma^j N)(N^\dagger \sigma^j \overleftrightarrow{\nabla} \cdot \overleftrightarrow{\nabla} N)$
O_6	$(N^\dagger \boldsymbol{\sigma} \cdot \overleftrightarrow{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \overleftrightarrow{\nabla} N) + \text{h.c.}$
O_7	$(N^\dagger \boldsymbol{\sigma} \cdot \overleftrightarrow{\nabla} N)(N^\dagger \boldsymbol{\sigma} \cdot \overleftrightarrow{\nabla} N)$

Relativistic: 17
vs.
Non-relativistic: 9

3: Comparison with other works (N2LO)

	Terms	Procedure	Advantage	Disadvantage
L.Girlanda [1]	36	① $n=0,1$ $\left(\frac{1}{(2m)^{2n}} (\bar{\psi} \partial^{\mu_1} \dots \partial^{\mu_n} \psi) (\bar{\psi} \partial_{\mu_1} \dots \partial_{\mu_n} \psi)\right)$	A complete set of NN contact Lagrangians	Not minimal
Stefan Petschauer [2]	25 (NN case)	① $n=0,1,2;$ ② Apply EOM; ③ Ignore Lagrangians with $\partial^\mu (\bar{\psi} \sigma_{\mu\nu} \psi)$ cause they claim it contribute to higher order $O(p^1)$ and can be subsumed in higher order Lagrangians	Contains less terms compared with [1]	Not complete
Our work	17	① $n=0,1;$ ② Apply EOM; ③ Include Lagrangians with $\partial^\mu (\bar{\psi} \sigma_{\mu\nu} \psi)$ cause it contains unique Lorentz structure	A complete and minimal set of NN contact Lagrangians	

[1] PRC81 (2010) 034005 [2] NPA916 (2013) 1

3: Covariant NN contact Lagrangians (N4LO)

\bar{O}_1	$(\bar{\psi}\psi)(\bar{\psi}\psi)$	\bar{O}_{21}	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\mu\psi)\partial^2\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
\bar{O}_2	$(\bar{\psi}\gamma^\mu\psi)(\bar{\psi}\gamma_\mu\psi)$	\bar{O}_{22}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\alpha}\psi)\partial^2\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
\bar{O}_3	$(\bar{\psi}\gamma_5\gamma^\mu\psi)(\bar{\psi}\gamma_5\gamma_\mu\psi)$	\bar{O}_{23}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^\beta\partial_\nu(\bar{\psi}\sigma_{\alpha\beta}i\overleftrightarrow{\partial}_\mu\psi)$
\bar{O}_4	$(\bar{\psi}\sigma^{\mu\nu}\psi)(\bar{\psi}\sigma_{\mu\nu}\psi)$	\bar{O}_{24}	$\frac{1}{16m^4}(\bar{\psi}\psi)\partial^4(\bar{\psi}\psi)$
\bar{O}_5	$(\bar{\psi}\gamma_5\psi)(\bar{\psi}\gamma_5\psi)$	\bar{O}_{25}	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu\psi)\partial^4(\bar{\psi}\gamma_\mu\psi)$
\bar{O}_6	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\alpha i\overleftrightarrow{\partial}_\mu\psi)$	\bar{O}_{26}	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu\psi)\partial^4(\bar{\psi}\gamma_5\gamma_\mu\psi)$
\bar{O}_7	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu\psi)$	\bar{O}_{27}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}\psi)\partial^4(\bar{\psi}\sigma_{\mu\nu}\psi)$
\bar{O}_8	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\mu\psi)\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$	\bar{O}_{28}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5 i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5 i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_5$
\bar{O}_9	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\alpha}\psi)\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$	\bar{O}_{29}	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_5\gamma_\alpha i\overleftrightarrow{\partial}_\mu i\overleftrightarrow{\partial}_\beta\psi) - \bar{O}_6$
\bar{O}_{10}	$\frac{1}{4m^2}(\bar{\psi}\psi)\partial^2(\bar{\psi}\psi)$	\bar{O}_{30}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu i\overleftrightarrow{\partial}_\beta\psi) - \bar{O}_7$
\bar{O}_{11}	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_\mu\psi)$	\bar{O}_{31}	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\mu i\overleftrightarrow{\partial}^\beta\psi)\partial^\alpha(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\beta\psi) - \bar{O}_8$
\bar{O}_{12}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\mu\psi)$	\bar{O}_{32}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\alpha}i\overleftrightarrow{\partial}^\beta\psi)\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\beta\psi) - \bar{O}_9$
\bar{O}_{13}	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}\psi)\partial^2(\bar{\psi}\sigma_{\mu\nu}\psi)$	\bar{O}_{33}	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_{10}$
\bar{O}_{14}	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_1$	\bar{O}_{34}	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_{11}$
\bar{O}_{15}	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_2$	\bar{O}_{35}	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_{12}$
\bar{O}_{16}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_3$	\bar{O}_{36}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_{13}$
\bar{O}_{17}	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha\psi) - \bar{O}_4$	\bar{O}_{37}	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\bar{O}_{14} - \bar{O}_1$
\bar{O}_{18}	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\psi)\partial^2(\bar{\psi}\gamma_5\psi)$	\bar{O}_{38}	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\bar{O}_{15} - \bar{O}_2$
\bar{O}_{19}	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\nu\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\nu i\overleftrightarrow{\partial}_\mu\psi)$	\bar{O}_{39}	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\bar{O}_{16} - \bar{O}_3$
\bar{O}_{20}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu\psi)$	\bar{O}_{40}	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\bar{O}_{17} - \bar{O}_4$

Relativistic: 40
vs.
Non-relativistic: 24

3: Non-relativistic reduction

□ **Why need non-relativistic reduction: self consistency check**

□ **Non-relativistic expansion:** $\psi \rightarrow N$, expand Lagrangians in terms of $1/m$

- **Relativistic nucleon field operator:** $\psi(x) = \int \frac{d\mathbf{p}}{(2\pi)^3} \frac{m}{E_p} \tilde{b}_s(\mathbf{p}) u^{(s)}(\mathbf{p}) e^{-ip \cdot x}$,
- **Non-relativistic nucleon field operator:** $N(x) = \int \frac{d\mathbf{p}}{(2\pi)^3} b_s(\mathbf{p}) \chi_s e^{-ip \cdot x}$,
- **Expansion of field operator**

$$\psi(x) = \left[\begin{pmatrix} 1 \\ 0 \end{pmatrix} - \frac{i}{2m} \begin{pmatrix} 0 \\ \boldsymbol{\sigma} \cdot \boldsymbol{\nabla} \end{pmatrix} + \frac{1}{8m^2} \begin{pmatrix} \nabla^2 \\ 0 \end{pmatrix} - \frac{3i}{16m^3} \begin{pmatrix} 0 \\ \boldsymbol{\sigma} \cdot \boldsymbol{\nabla} \nabla^2 \end{pmatrix} + \frac{11}{128m^4} \begin{pmatrix} \nabla^4 \\ 0 \end{pmatrix} \right] N(x) + \mathcal{O}(Q^5).$$

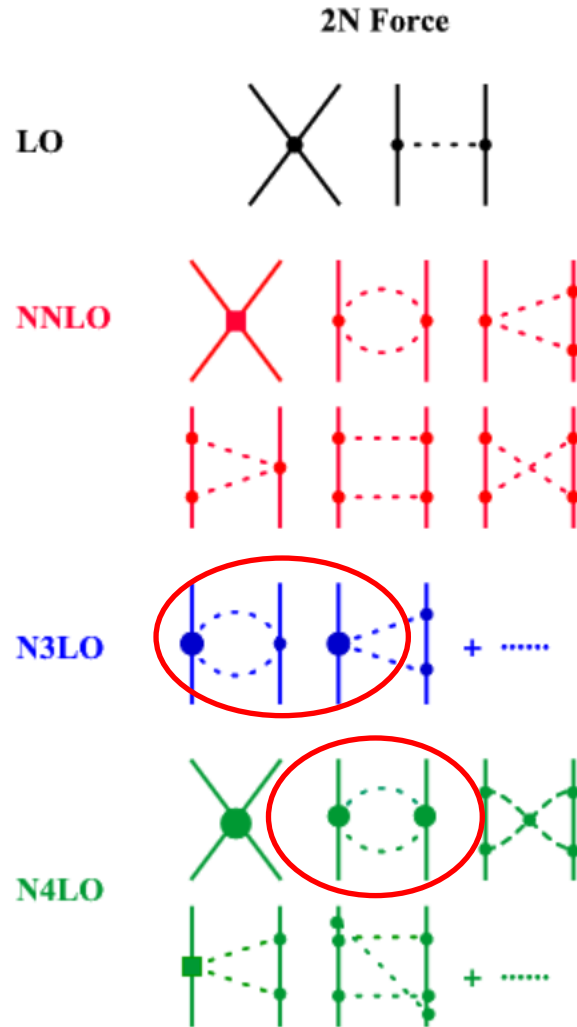
- **Dirac matrices expressed in term of pauli matrices**

$$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \gamma^5 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \vec{\gamma} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix}, \sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu].$$

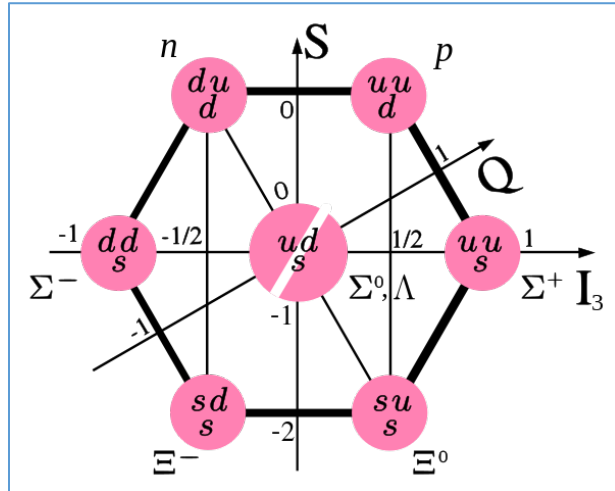
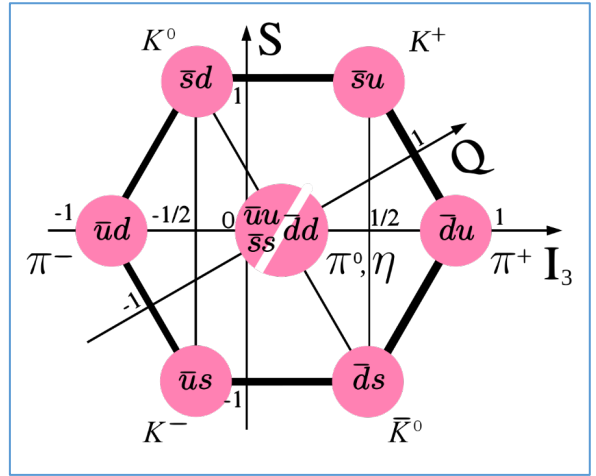
3: Non-relativistic reduction

After expansion and keeping only appropriate powers of $1/mN$, we can reduce the **40** relativistic terms into the **24(2+7+15)** non-relativistic terms

4: Towards N⁴LO We will need pion-nucleon inputs



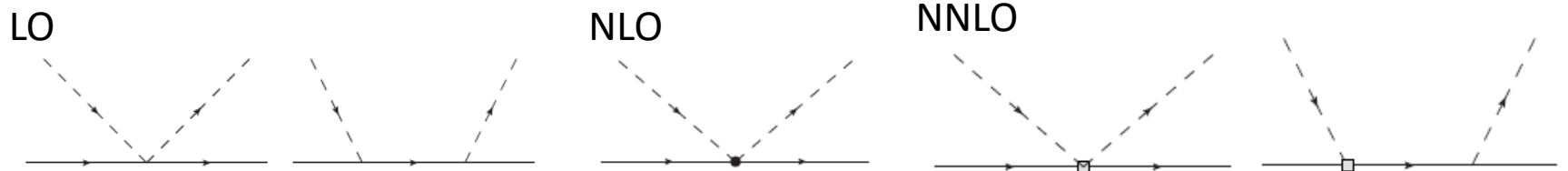
Simultaneous description of meson-baryon scattering



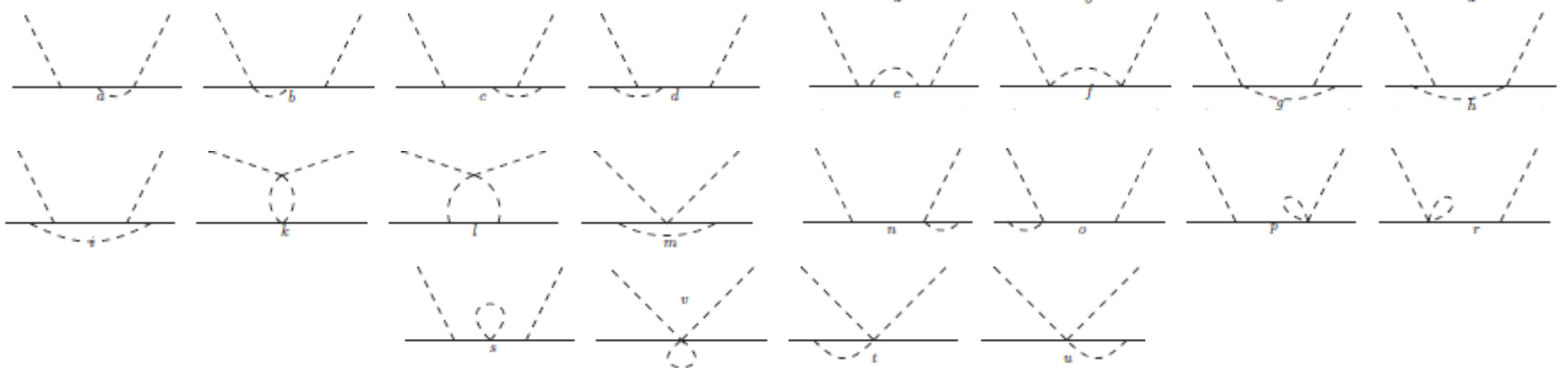
- Kaon-nucleon (s=1): some data
- Pion-nucleon (s=0): plenty of data
- Kbar-nucleon (s=-1): some data and **nonperturbative**
- Other channels (s=-2, -3): very limited lattice QCD data

4: Meson-baryon scattering upto N2LO

□ Tree



□ Loop



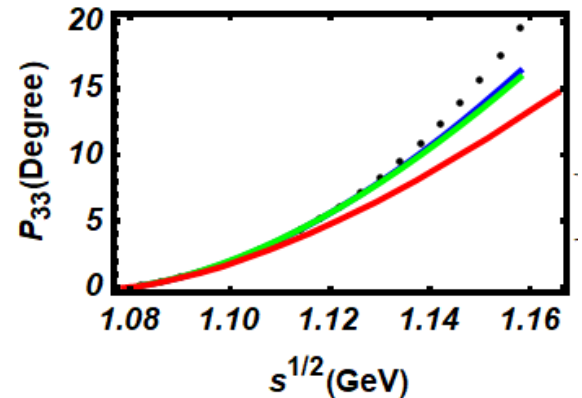
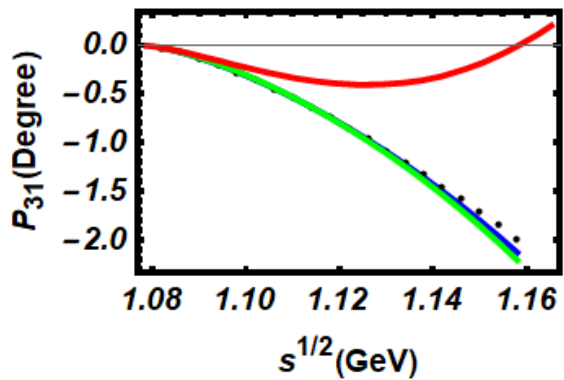
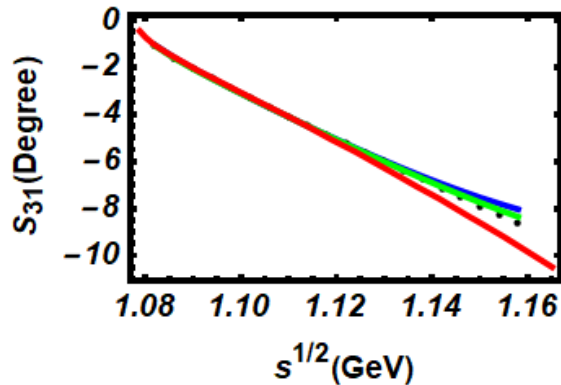
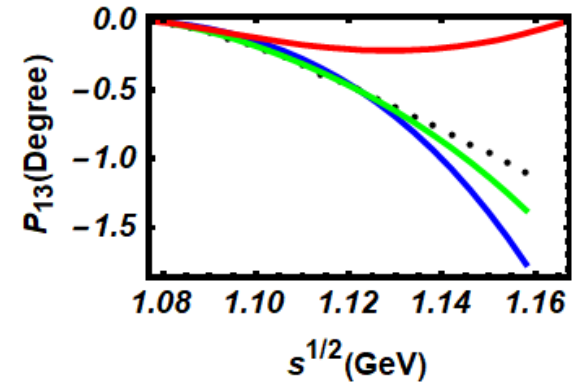
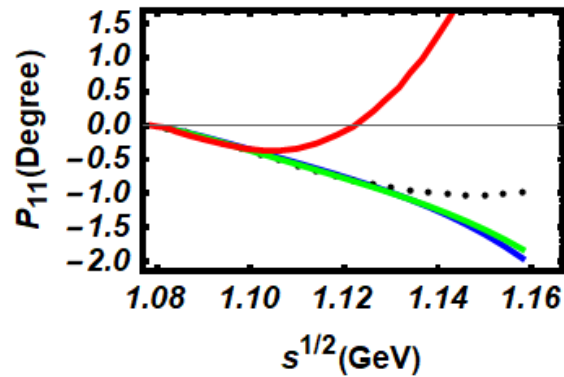
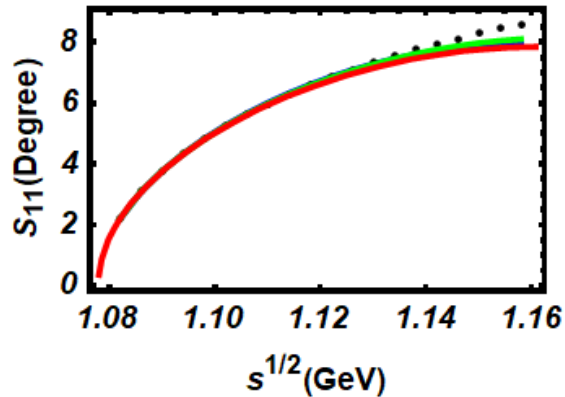
4: Meson-baryon scattering

In total 37 LECs, but decoupled in $S=0$ & $S=1$

24

πN	$KN_{I=0}$	$KN_{I=1}$
$b_1 + b_2 + b_3 + 2b_4$	$b_3 - b_4$	$b_1 + b_2 + b_4$
$b_5 + b_6 + b_7 + b_8$	$2b_6 - b_8$	$2b_5 + 2b_7 + b_8$
$c_1 + c_2$	$4c_1 + c_3$	$4c_2 + c_3$
$2b_0 + b_D + b_F$	$b_0 - b_F$	$b_0 + b_D$
d_2	$d_1 + d_2 + d_3$	$d_1 - d_2 - d_3$
d_4	$d_4 + d_5 + d_6$	$d_4 - d_5 + d_6$
$d_8 + d_{10}$	$d_7 - d_8 + d_{10}$	$d_7 + d_8 + d_{10}$
d_{49}	$d_{48} + d_{49} + d_{50}$	$d_{48} + d_{49} - d_{50}$

4: Pion-nucleon scattering

 $L_{2I,2J}$ 

Blue: SU3 EOMS

Green: SU2 EOMS Chen et al. PRD87(2013)054019

Red: SU3 HB Huang et al. PRD96(2017)016021

Black dot: EXP

4: Pion-nucleon scattering

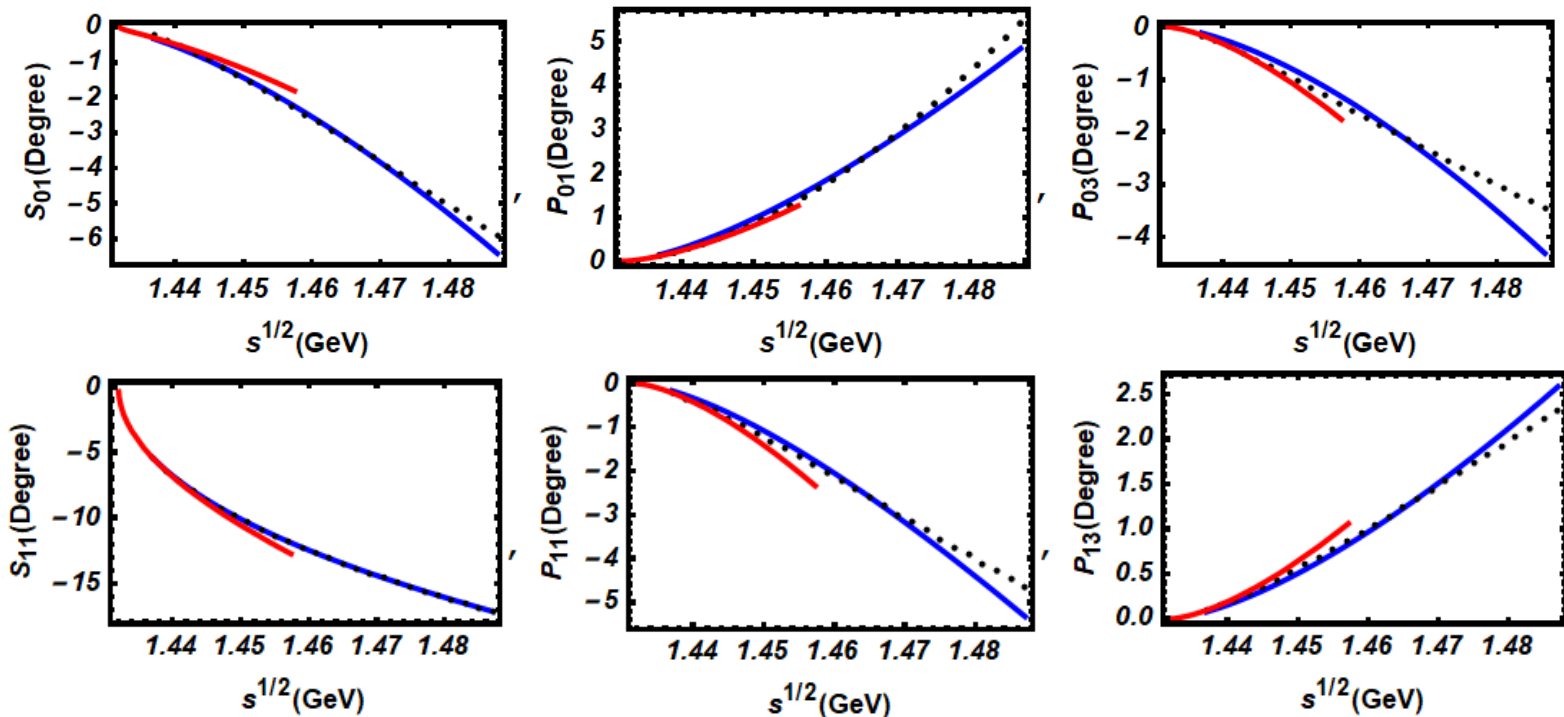
□ Pion-nucleon phase shifts only (up to $s^{1/2}=1.13$ GeV)

α_1	α_2	α_3	α_4	α_5	α_6	α_7	α_8	$\chi^2/d.o.f.$
-7.64(6)	1.42(2)	1.34(1)	-1.36(6)	0.61(2)	3.25(6)	1.45(3)	-0.32(12)	0.154

4: Kaon-Nucleon scattering

$L_{2I,2J}$

- Isospin=0, $\chi^2/\text{d.o.f} = 1.046$
- Isospin=1, $\chi^2/\text{d.o.f} = 0.507$



Blue: SU3 EOMS
Black dot: EXP

Red: SU3 HB Huang et al. PRD96(2017)016021

5.1: Renormalization Group Invariance

- **A fundamental feature of EFT** that physics at long-distance scales is **insensitive** to the details at short-distance scales



Peter Lepage

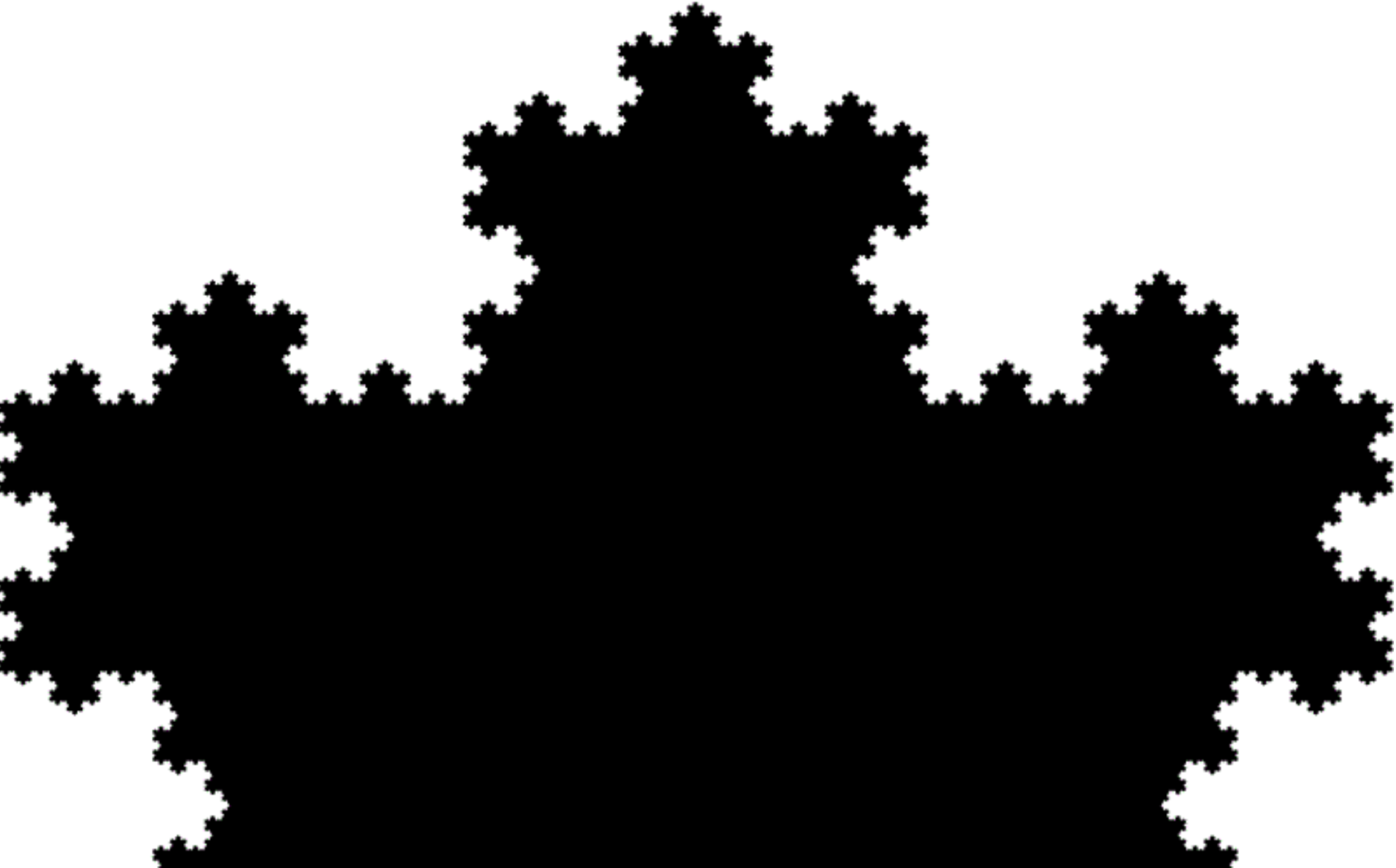
- The NN interaction can be considered properly renormalized **when the calculated observables are independent of the cutoff scale** within the range of validity of the ChEFT

- In the language of **Wilson's renormalization group**, this means that **the LECs must run with the cutoff** scale in such a way that the scattering amplitude becomes RG invariant (RGI)



Kenneth G. Wilson

5.1: RGI vs. self-similarity



5.1: RGI-a highly debated topic



Kaplan



van Kolck



Epelbaum



Machleidt

Key issues

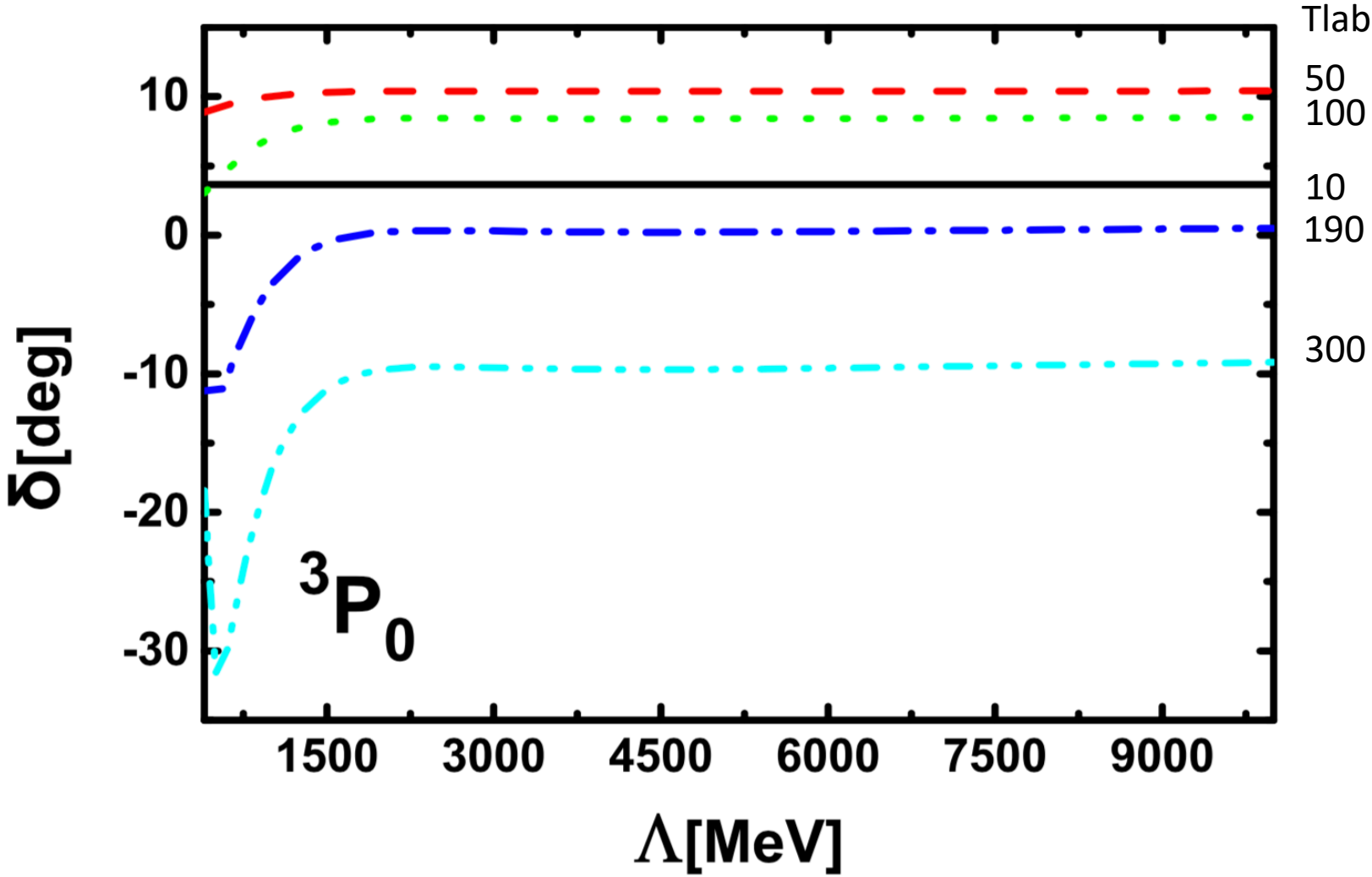
- Whether RGI is essential?
- How to implement it?

Key references

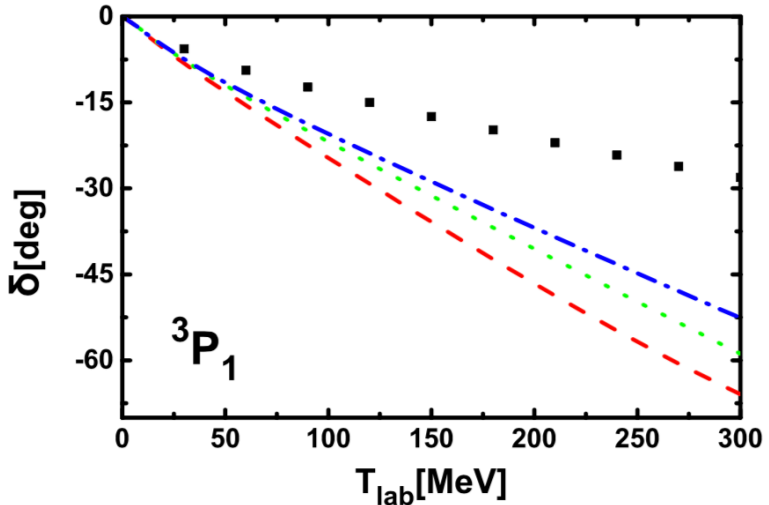
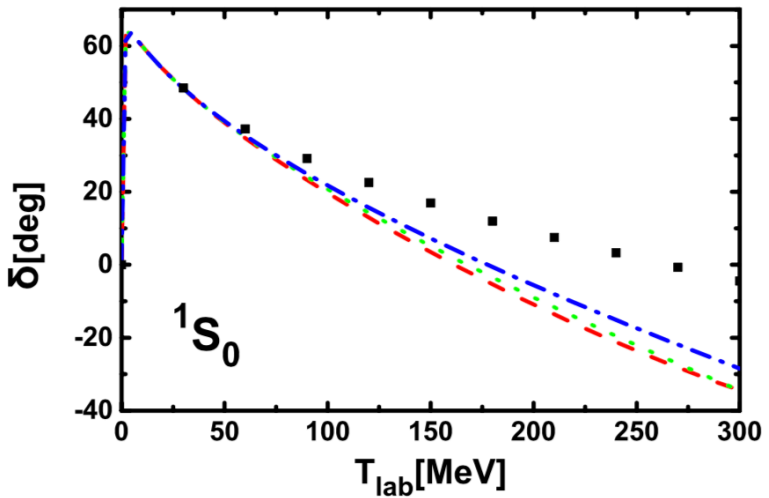
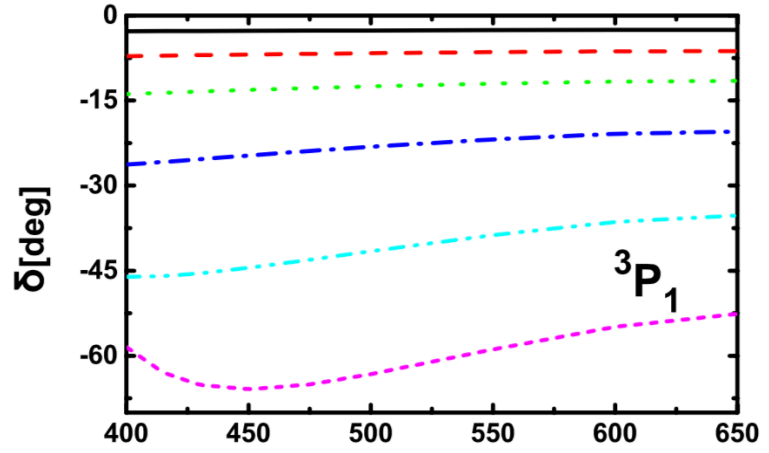
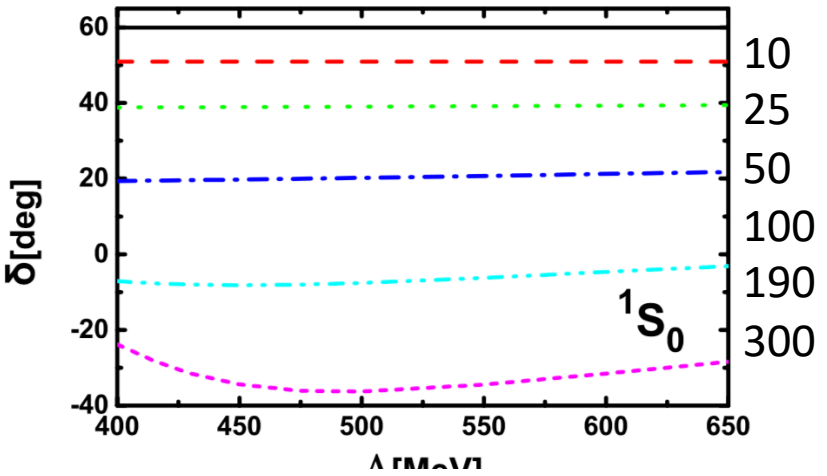
- ✓ *A New expansion for nucleon-nucleon interactions, David B. Kaplan et al., Phys.Lett. B424 (1998) 390*
- ✓ *Towards a perturbative theory of nuclear forces, S.R. Beane et al., Nucl.Phys. A700 (2002) 377*
- ✓ *How to renormalize the Schrodinger equation, P. Lepage, nucl-th/9706029*

5.1: **3P0** issue nicely solved

Chun-Xuan Wang et al.,

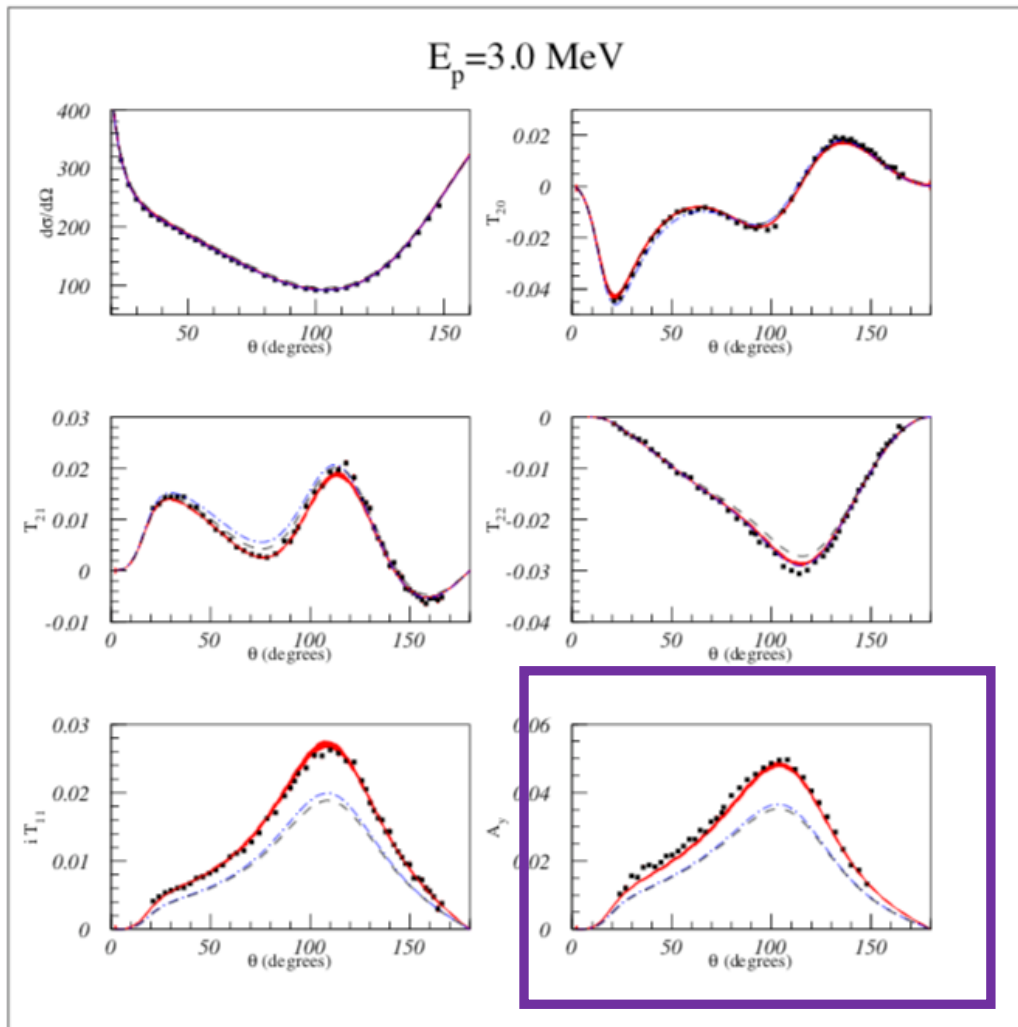


5.1: Interesting correlation **1S0** & 3P1



- Pure contact interactions cannot describe 1S0 – large cutoff limit
- NLO 3P1 in the Weinberg case loses RGI

5.2. Three-body study **promising**



1811.09398

L. Girlanda, A. Kievsky, M. Viviani and L.E. Marcucci

*“Fit results in the leading order of **the relativistic counting** to a set of cross section and polarization $p - d$ observables at 3 MeV proton energy, for $\Lambda = 200 - 500 \text{ MeV}$ (red bands) as compared to the purely two-body AV18 interaction (dashed, black lines) and to the AV18+UIX two- and three-nucleon interaction (dashed-dotted, blue lines).”*

Ay puzzle— a 32 year old persistent problem!

Summary and outlook

- Based on the consideration of symmetries and also requested by nuclear structure community, we **proposed to build** a high-precision nuclear force based on covariant baryon chiral perturbation theory
- Exploratory studies show that **the idea is feasible**
- **Steady progress is being made** and **the results are encouraging**
 - ✓ NLO & NNLO contact terms
 - ✓ NLO meson-baryon scattering
- A high precision covariant chiral nuclear force will come soon, stay tuned



北京航空航天大学
BEIHANG UNIVERSITY



Thanks for your attention !

August 20th , 2019

□ Lorentz invariance or relativistic corrections might play an important role in **speeding up** the convergence of ChEFT

T_{lab} [MeV]	1	50	100	150	200	250	300
P_{cm} [MeV/c]	21.67	153.22	216.68	265.38	306.43	342.60	375.30
P_{cm}/M_N	0.023c	0.16c	0.23c	0.28c	0.33c	0.36c	0.40c

In comparison $m_{\pi}/m_N = 138/939 \sim 0.15$