

The 13th Workshop on QCD Phase Transition and Relativistic Heavy-Ion
Physics (QPT 2019)

Vorticity and Λ spin polarization in heavy ion collisions

Hui Li

Fudan University, Shanghai

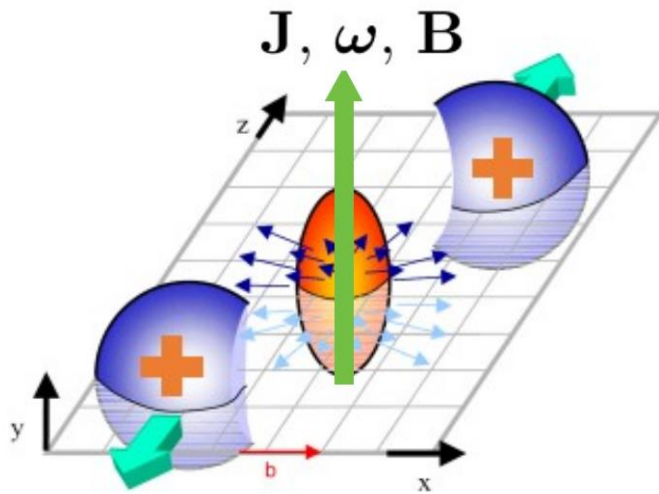
2019.08.19



X. L. Xia, H. Li, X. G. Huang, H. Z. Huang,
Phys. Rev. C 100, 014913

- Introduction
- Status and puzzles on Λ spin polarization
- Feed-down effect on Λ spin polarization
- Summary

Introduction



$$J \sim 10^6 \hbar \quad B \sim 10^{18} \text{G}$$

- Huge angular momentum and magnetic field are generated in heavy ion collisions. ω and B can induce chiral anomaly effect, such as CME, CVE, CMW, etc.
- Magnetic and vorticity field can lead to the spin polarization of final hadrons.
- From the polarization measurement, information of ω and B can be extracted.

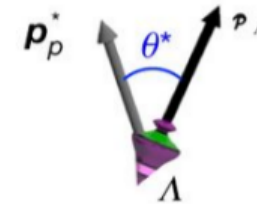
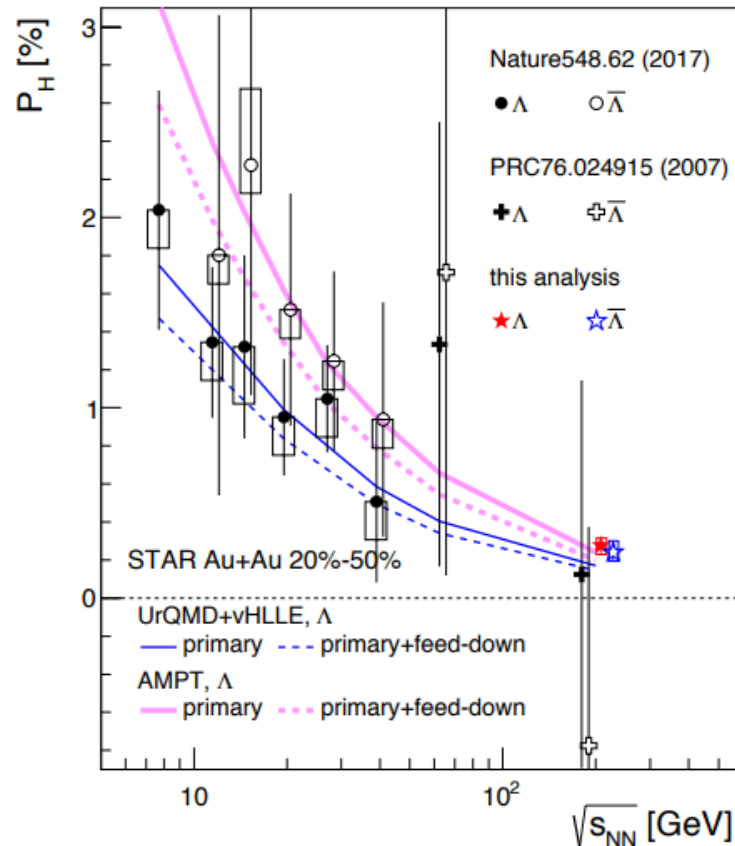
See talks by

**J. F. Liao, X. G. Huang, S. Lin, J. H. Gao,
Q. Y. Shou, S. Q. Feng, Y. C. Liu, L. Yin ...**

Global Λ polarization

Angular momentum or vorticity can lead to global polarization

Liang, Wang, PRL 94, 102301 (2005);
Voloshin nucl-th/0410089



$$\frac{dN}{d\cos\theta^*} = \frac{1}{2} \left(1 + \alpha_H |\vec{P}_H| \cos\theta^* \right).$$

Measurements:

STAR, Nature 548, 62 2017;
STAR, Phys. Rev. C 98, 014910 (2018).

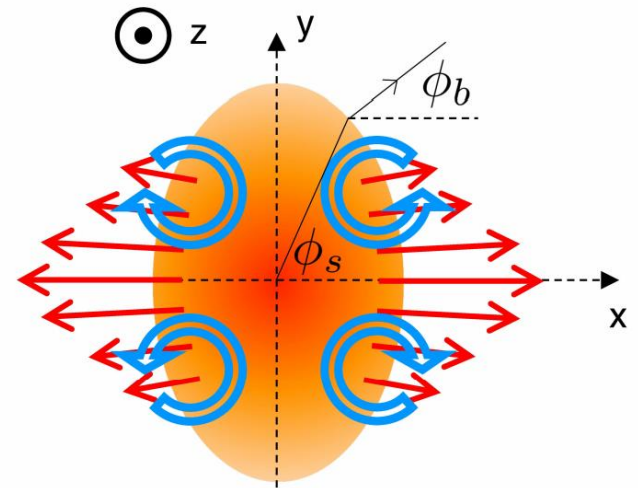
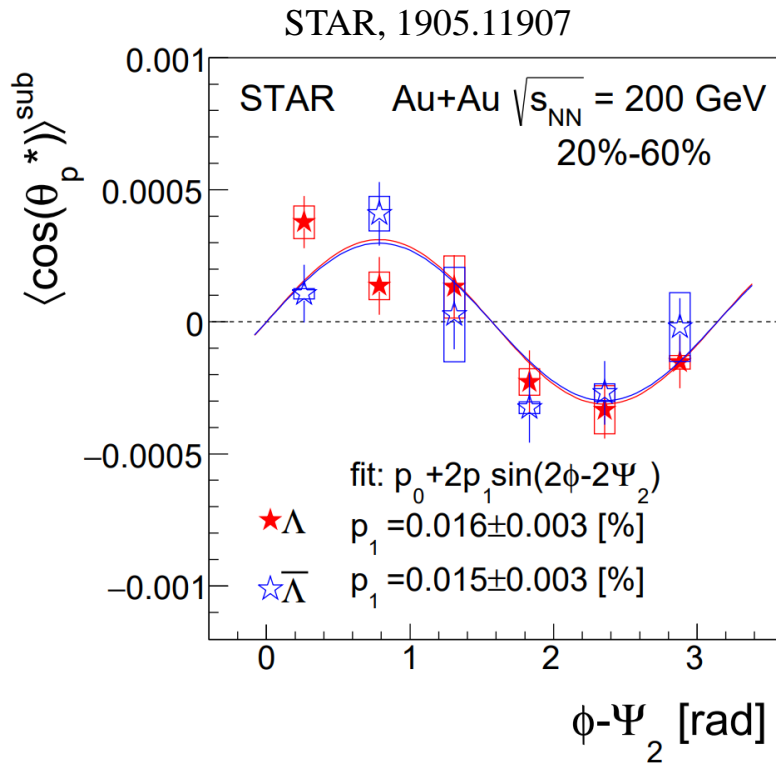
Theoretical calculation:

Karpenko-Becattini 2017;
Xie-Wang-Csernai 2017;
Li-Pang-Wang-Xia 2017;
Sun-Ko 2017;
Shi-Li-Liao 2019;
Wei-Deng-Huang 2019; ...

The model calculations agree with the data

Local Λ polarization

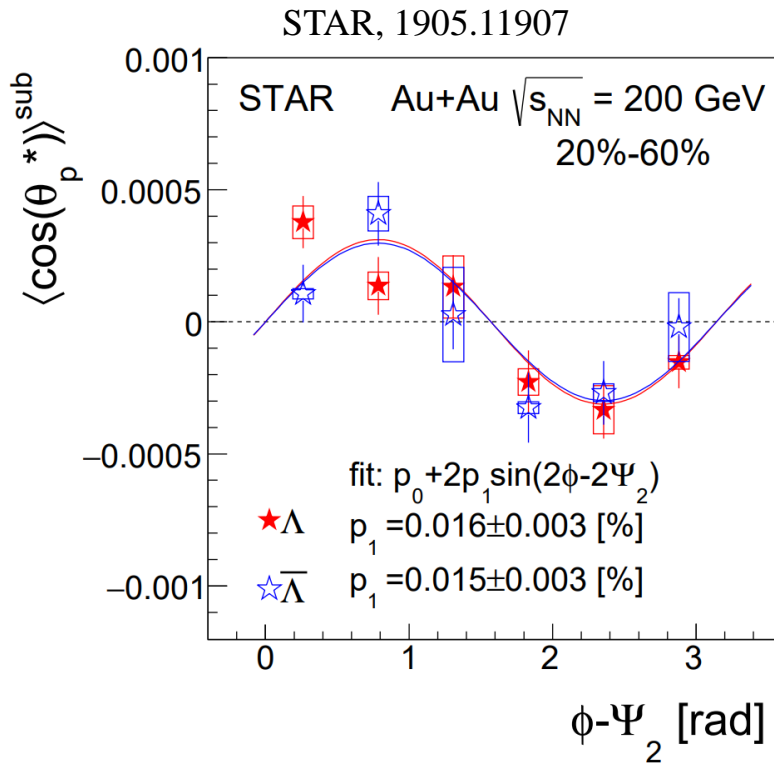
- Velocity gradients due to elliptic flow can produce vorticity along the beam direction.



$$P_z = F_z \sin(2\phi)$$

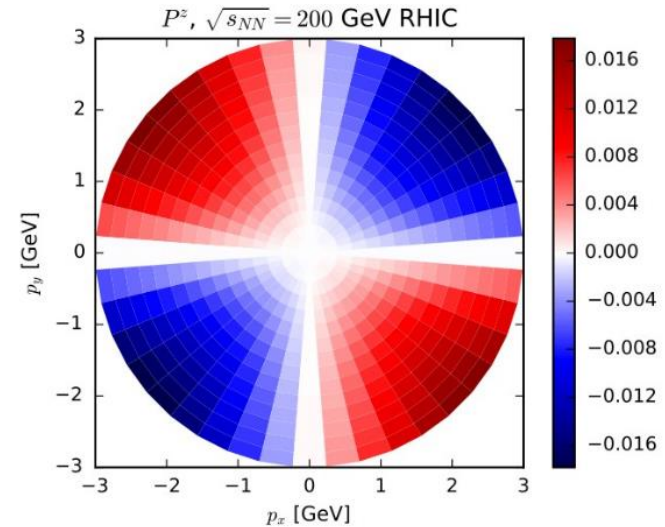
Local Λ polarization

- Local velocity gradients due to elliptic flow may produce vorticity along beam direction
- $\sin(2\phi)$ signal has been observed, but the sign is opposite between STAR data and hydro and transport models

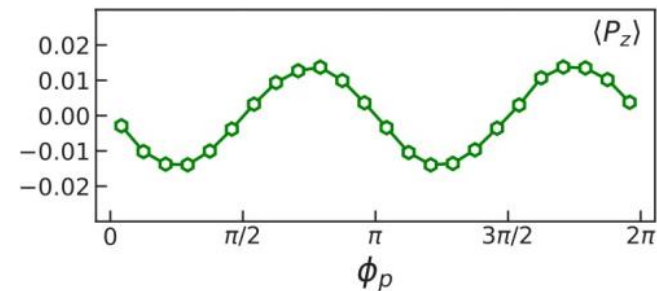


$$P_z = F_z \sin(2\phi)$$

Becattini-Karpenko, 2018



Xia-Li-Tang-Wang, 2018

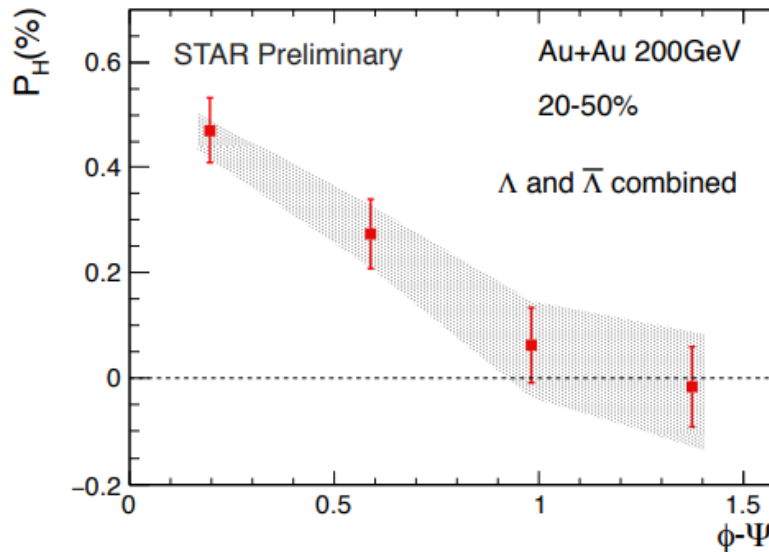


In-plane to out-of-plane difference

Global polarization also has an azimuthal-angle dependence.

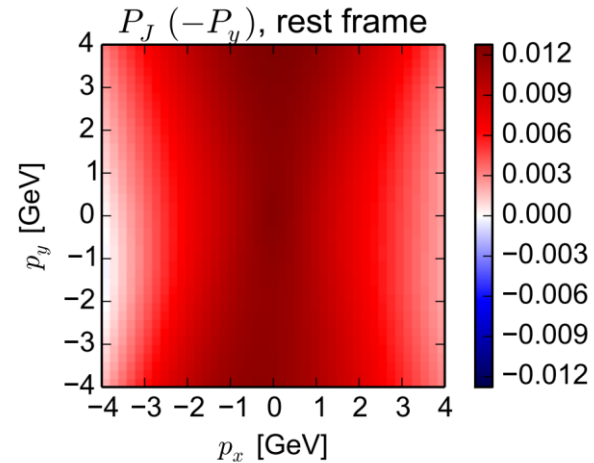
- **STAR data:** the spin polarization is stronger in the reaction plane
- **Model calculation:** the spin polarization is stronger out of plane

STAR, 1808.10482

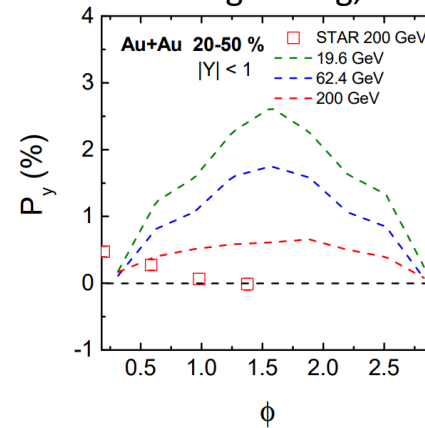


$$P_y = F_0 + F_2 \cos(2\phi)$$

Karpenko-Becattini, 2017

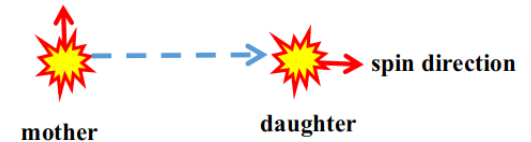
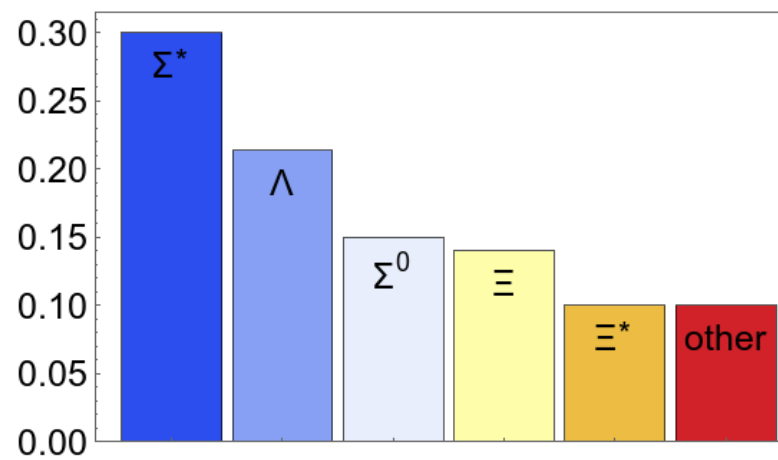


Wei-Deng-Huang, 2019



Motivations

- The spin polarization calculated in the above models is the polarization of primordial Λ s, which are thermally produced.



- In experimental measurement, more than 75% Λ s are produced by decay of heavier states.
- The spin transfer in decay should be considered!

Motivations

- For the global polarization, i.e. the mean value of spin polarization.

$$\langle \mathbf{P}_D \rangle = C \mathbf{P}_P$$

	spin and parity	$\langle \mathbf{P}_D \rangle / \mathbf{P}_P$
strong decay	$1/2^+ \rightarrow 1/2^+ 0^-$	-1/3
strong decay	$1/2^- \rightarrow 1/2^+ 0^-$	1
strong decay	$3/2^+ \rightarrow 1/2^+ 0^-$	1
strong decay	$3/2^- \rightarrow 1/2^+ 0^-$	-3/5
weak decay	$1/2 \rightarrow 1/2 \ 0$	$(2\gamma+1)/3$
EM decay	$1/2^+ \rightarrow 1/2^+ 1^-$	-1/3

- The global polarization was estimated to be suppressed by a factor of 80%-95%, compared to the primordial polarization.

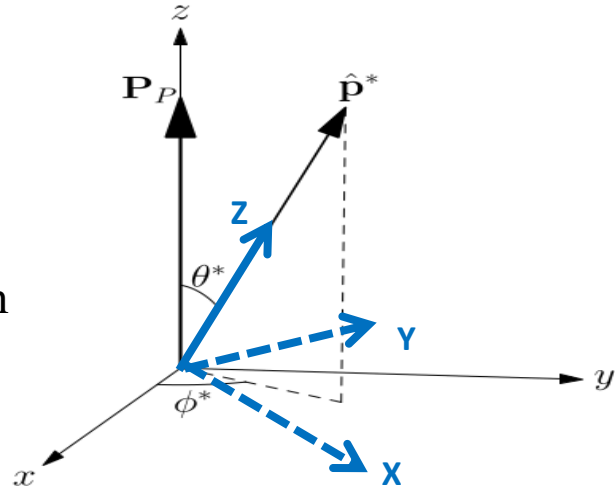
However, to study the azimuthal-angle dependence, linear equation does not apply.

Two-body decay

Decay process $P \rightarrow D + X$

With given spin state of P, we want:

- (1) angular distribution of D,
- (2) polarization of D as a function of D's momentum direction.



Resonance Decay

In the helicity basis, the spin density matrix of D is related to that of P by:

$$\rho_{\lambda_D \lambda_X; \lambda'_D \lambda'_X}^f(\theta^*, \phi^*) = \sum_{M_P, M'_P} H_{\lambda_D \lambda_X; M_P} \rho_{M_P; M'_P}^i H_{M'_P; \lambda'_D \lambda'_X}^\dagger$$

$$H_{\lambda_D \lambda_X; M_P} = \sqrt{\frac{2S_P + 1}{4\pi}} D_{M_P; \lambda_D - \lambda_X}^{S_P*}(\phi^*, \theta^*, 0) A_{\lambda_D; \lambda_X}$$

⇒ Angular distribution

$$Tr(\rho^f) = \frac{1}{N} \frac{dN}{d\Omega^*}$$

Polarization vector

$$\mathbf{P}_D = tr_D(\hat{\mathbf{P}} \rho^D) / tr_D(\rho^D)$$

$$\rho^D = tr_X(\rho^f)$$

Ex1: Polarization vector and angular distribution

Strong decay $1/2^\pm \rightarrow 1/2^+ 0^-$

Parent density matrix:

$$\rho_{M_P;M'_P}^i = \text{diag} \left(\frac{1+P_P}{2}, \frac{1-P_P}{2} \right)$$

Daughter density matrix:

$$\rho_{\lambda_D;\lambda'_D}^D = \frac{1}{4\pi} \begin{pmatrix} |A_{1/2}|^2 (1+P_P \cos \theta^*) & -A_{1/2} A_{-1/2}^* P_P \sin \theta^* \\ -A_{1/2}^* A_{-1/2} P_P \sin \theta^* & |A_{-1/2}|^2 (1-P_P \cos \theta^*) \end{pmatrix}$$

Since parity is conserved, amplitude is constrained by

$$A_{\lambda_D;\lambda_X} = \pi_P \pi_D \pi_X (-1)^{S_P - S_D - S_X} A_{-\lambda_D; -\lambda_X}$$

$$\rho_{\lambda_D;\lambda'_D}^D = \frac{1}{8\pi} \begin{pmatrix} 1+P_P \cos \theta^* & \pm P_P \sin \theta^* \\ \pm P_P \sin \theta^* & 1-P_P \cos \theta^* \end{pmatrix} \Longrightarrow \frac{1}{N} \frac{dN}{d\Omega^*} = \frac{1}{4\pi}$$

$$P_X(\theta^*, \phi^*) = \pm P_P \sin \theta^*,$$

$$P_Y(\theta^*, \phi^*) = 0,$$

$$P_Z(\theta^*, \phi^*) = P_P \cos \theta^*.$$



$$\mathbf{P}_D = 2(\mathbf{P}_P \cdot \hat{\mathbf{p}}^*) \hat{\mathbf{p}}^* - \mathbf{P}_P$$

$$\mathbf{P}_D = \mathbf{P}_P$$

Ex2: Polarization vector and angular distribution

Weak decay $1/2 \rightarrow 1/2 0$



Parent density matrix:

$$\rho_{M_P; M'_P}^i = \text{diag} \left(\frac{1+P_P}{2}, \frac{1-P_P}{2} \right)$$

Daughter density matrix:

$$\rho_{\lambda_D; \lambda'_D}^D = \frac{1}{4\pi} \begin{pmatrix} |A_{1/2}|^2 (1 + P_P \cos \theta^*) & -A_{1/2} A_{-1/2}^* P_P \sin \theta^* \\ -A_{1/2}^* A_{-1/2} P_P \sin \theta^* & |A_{-1/2}|^2 (1 - P_P \cos \theta^*) \end{pmatrix}$$

parity is not conserved, amplitude can be decomposed into parity-odd and parity-even terms:

$$A_{\pm 1/2} = \frac{A_s \pm A_p}{\sqrt{2(|A_s|^2 + |A_p|^2)}}$$

$$\mathbf{P}_D = \frac{(\alpha + \mathbf{P}_P \cdot \hat{\mathbf{p}}^*) \hat{\mathbf{p}}^* + \beta (\mathbf{P}_P \times \hat{\mathbf{p}}^*) + \gamma \hat{\mathbf{p}}^* \times (\mathbf{P}_P \times \hat{\mathbf{p}}^*)}{1 + \alpha \mathbf{P}_P \cdot \hat{\mathbf{p}}^*}$$

weak decay parameter

$$\frac{1}{N} \frac{dN}{d\Omega^*} = \frac{1}{4\pi} (1 + \alpha P_P \cos \theta^*)$$

T.D. Lee and C.N. Yang 1957

$$\bullet \alpha = 0, \beta = 0, \gamma = 1 \implies \mathbf{P}_D = 2(\mathbf{P}_P \cdot \hat{\mathbf{p}}^*) \hat{\mathbf{p}}^* - \mathbf{P}_P$$

$$\bullet \alpha = 0, \beta = 0, \gamma = -1 \implies \mathbf{P}_D = \mathbf{P}_P$$

$$\left\{ \begin{array}{l} \alpha = \frac{2\text{Re}(A_s^* A_p)}{|A_s|^2 + |A_p|^2}, \\ \beta = \frac{2\text{Im}(A_s^* A_p)}{|A_s|^2 + |A_p|^2}, \\ \gamma = \frac{|A_s|^2 - |A_p|^2}{|A_s|^2 + |A_p|^2}. \end{array} \right.$$

Polarization vector and angular distribution

For all decay channels, the polarization vector are obtained.

TABLE I. Daughter angular distribution and polarization vector \mathbf{P}_D in different decay channels.

	Spin and parity	$(1/N)dN/d\Omega^*$	\mathbf{P}_D	$\langle \mathbf{P}_D \rangle / \mathbf{P}_P$
Strong decay	$1/2^+ \rightarrow 1/2^+ 0^-$	$1/(4\pi)$	$2(\mathbf{P}_P \cdot \hat{\mathbf{p}}^*)\hat{\mathbf{p}}^* - \mathbf{P}_P$	$-1/3$
Strong decay	$1/2^- \rightarrow 1/2^+ 0^-$	$1/(4\pi)$	\mathbf{P}_P	1
Strong decay	$3/2^+ \rightarrow 1/2^+ 0^-$	$3[1 - 2\Delta/3 - (1 - 2\Delta)\cos^2\theta^*]/(8\pi)$	Eq. (42)	1
Strong decay	$3/2^- \rightarrow 1/2^+ 0^-$	$3[1 - 2\Delta/3 - (1 - 2\Delta)\cos^2\theta^*]/(8\pi)$	Eq. (43)	$-3/5$
Weak decay	$1/2 \rightarrow 1/2 0$	$(1 + \alpha P_P \cos\theta^*)/(4\pi)$	Eq. (29)	$(2\gamma + 1)/3$
EM decay	$1/2^+ \rightarrow 1/2^+ 1^-$	$1/(4\pi)$	$-(\mathbf{P}_P \cdot \hat{\mathbf{p}}^*)\hat{\mathbf{p}}^*$	$-1/3$

$$\mathbf{P}_D = \frac{-4\delta(\mathbf{P}_P \cdot \hat{\mathbf{p}}^*)\hat{\mathbf{p}}^* + [1 - 2\delta - (1 - 10\delta)(\hat{\mathbf{P}}_P \cdot \hat{\mathbf{p}}^*)^2]\mathbf{P}_P}{1 - 2\Delta/3 - (1 - 2\Delta)(\hat{\mathbf{P}}_P \cdot \hat{\mathbf{p}}^*)^2}, \quad (42)$$

$$\mathbf{P}_D = \frac{2[1 - 4\delta - (1 - 10\delta)(\hat{\mathbf{P}}_P \cdot \hat{\mathbf{p}}^*)^2](\mathbf{P}_P \cdot \hat{\mathbf{p}}^*)\hat{\mathbf{p}}^* - [1 - 2\delta - (1 - 10\delta)(\hat{\mathbf{P}}_P \cdot \hat{\mathbf{p}}^*)^2]\mathbf{P}_P}{1 - 2\Delta/3 - (1 - 2\Delta)(\hat{\mathbf{P}}_P \cdot \hat{\mathbf{p}}^*)^2}. \quad (43)$$

$$\mathbf{P}_D = \frac{(\alpha + \mathbf{P}_P \cdot \hat{\mathbf{p}}^*)\hat{\mathbf{p}}^* + \beta(\mathbf{P}_P \times \hat{\mathbf{p}}^*) + \gamma\hat{\mathbf{p}}^* \times (\mathbf{P}_P \times \hat{\mathbf{p}}^*)}{1 + \alpha\mathbf{P}_P \cdot \hat{\mathbf{p}}^*} \quad (29)$$

Use Monte Carlo simulation to estimate the feed down effect on the polarization.

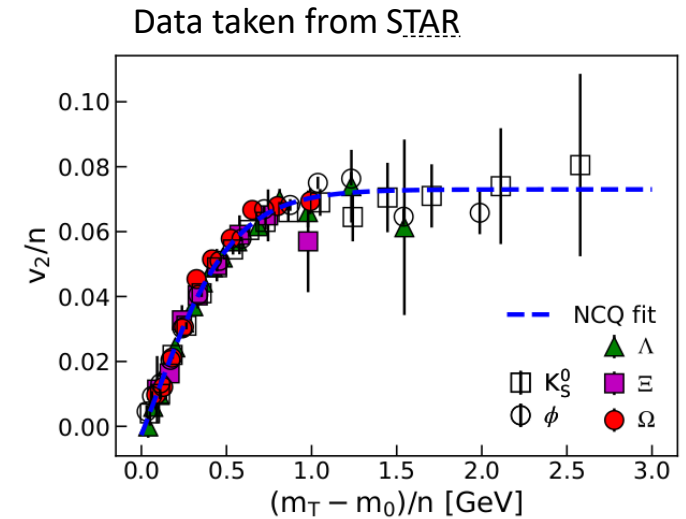
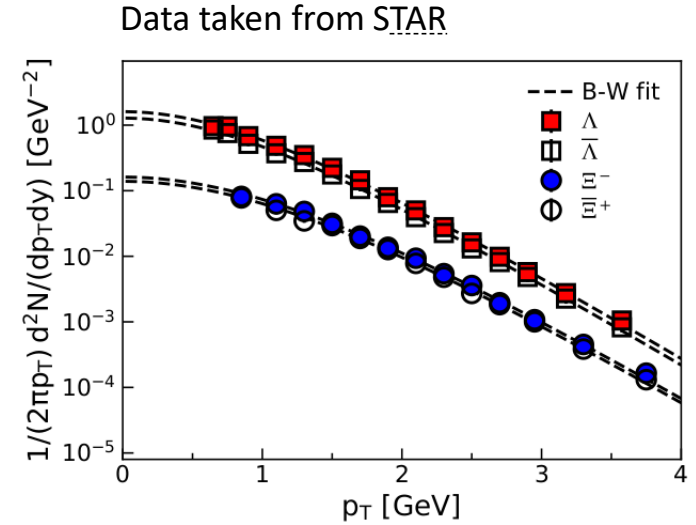
See also: F. Becattini, G. Cao, E. Speranza, arXiv: 1905.03123

Transverse momentum spectrum and elliptic flow

- Primordial yield ratio are generated from THERMUS model
- p_T spectrum and v_2 are fitted to data, and then, are used to sample parent particles.

TABLE II. The primordial yield ratio N_i/N_Λ , spin, parity, and decay channels of strange particles.

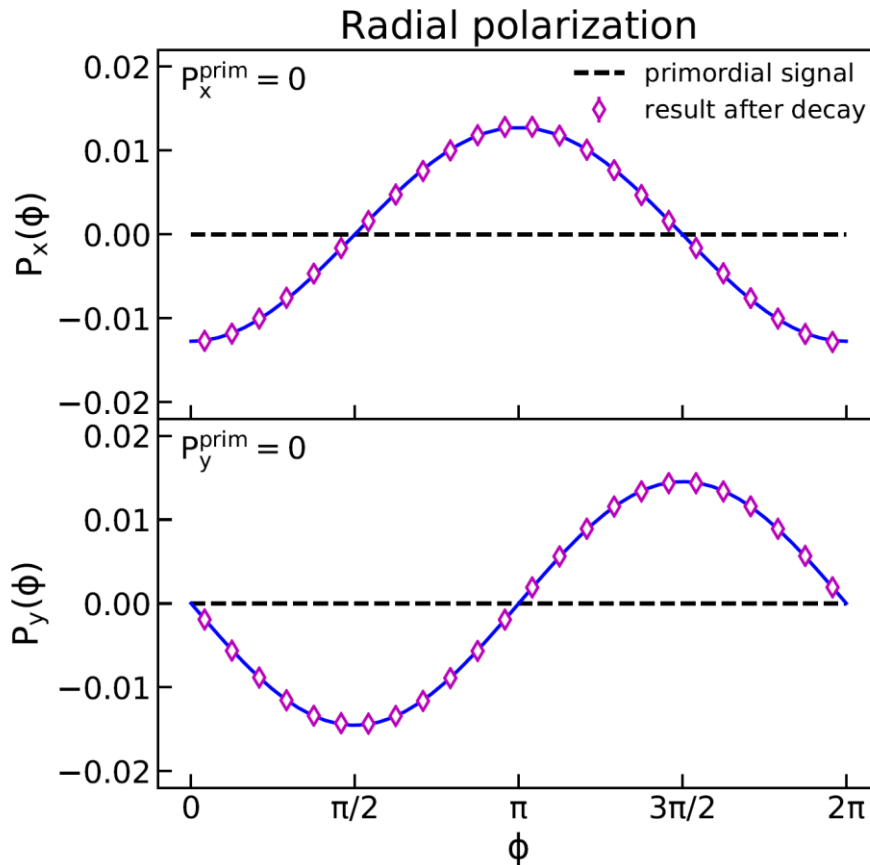
	N_i/N_Λ	Spin and parity	Decay channel
Λ	1	$1/2^+$	
$\Lambda(1405)$	0.236	$1/2^-$	$\Sigma^0\pi$
$\Lambda(1520)$	0.265	$3/2^-$	$\Sigma^0\pi$
$\Lambda(1600)$	0.098	$1/2^+$	$\Sigma^0\pi$
$\Lambda(1670)$	0.061	$1/2^-$	$\Sigma^0\pi, \Lambda\eta$
$\Lambda(1690)$	0.112	$3/2^-$	$\Sigma^0\pi$
Σ^0	0.686	$1/2^+$	$\Lambda\gamma$
Σ^{*0}	0.533	$3/2^+$	$\Lambda\pi$
Σ^{*+}	0.535	$3/2^+$	$\Lambda\pi, \Sigma^0\pi$
Σ^{*-}	0.524	$3/2^+$	$\Lambda\pi, \Sigma^0\pi$
$\Sigma(1660)$	0.068	$1/2^+$	$\Lambda\pi, \Sigma^0\pi$
$\Sigma(1670)$	0.125	$3/2^-$	$\Lambda\pi, \Sigma^0\pi$
Ξ^0	0.343	$1/2^+$	$\Lambda\pi$
Ξ^-	0.332	$1/2^+$	$\Lambda\pi$
Ξ^{*0}	0.228	$3/2^+$	$\Xi\pi$
Ξ^{*-}	0.224	$3/2^+$	$\Xi\pi$



Contribution from Ξ

- We first input zero polarization to parent particles

Weak decay can contribute a radial polarization.

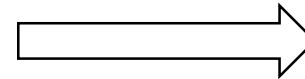


$$P_x(\phi) = K_{1x} \cos \phi,$$

$$P_y(\phi) = K_{1y} \sin \phi,$$

For $\mathbf{P}_P = 0$

$$\mathbf{P}_D = \frac{(\alpha + \mathbf{P}_P \cdot \hat{\mathbf{p}}^*) \hat{\mathbf{p}}^* + \beta (\mathbf{P}_P \times \hat{\mathbf{p}}^*) + \gamma \hat{\mathbf{p}}^* \times (\mathbf{P}_P \times \hat{\mathbf{p}}^*)}{1 + \alpha \mathbf{P}_P \cdot \hat{\mathbf{p}}^*}$$



$$\mathbf{P}_D = \alpha \hat{\mathbf{p}}^*$$

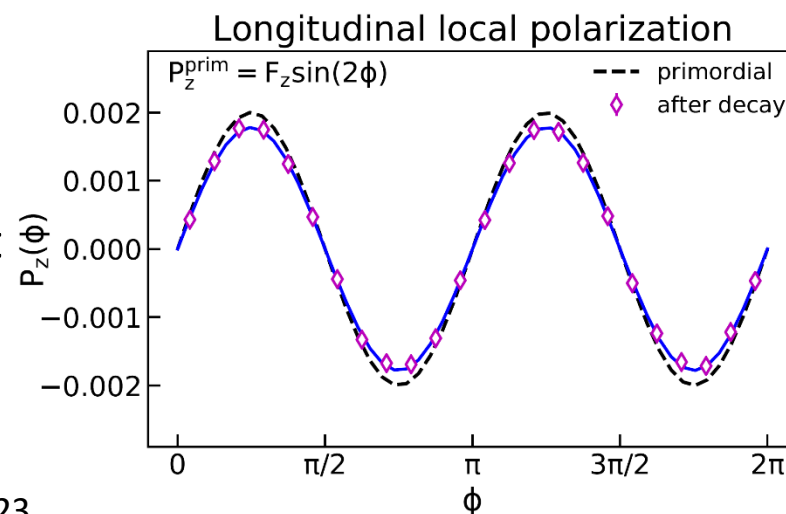
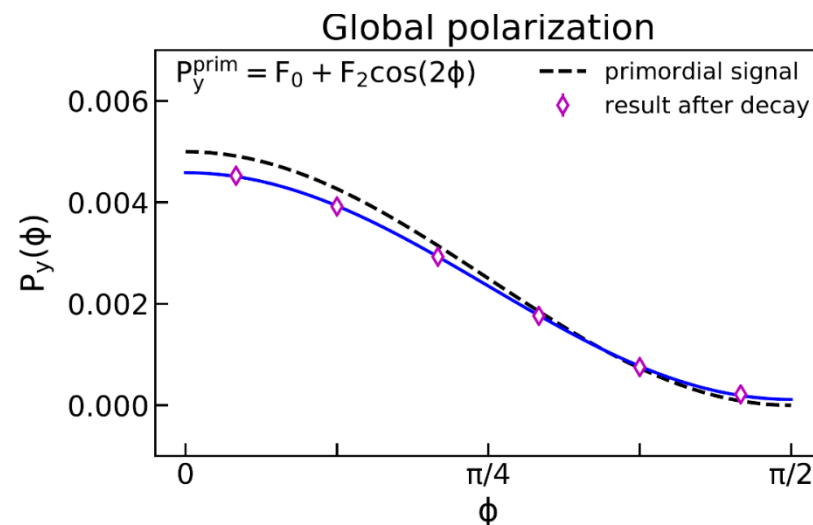
$$\alpha_{\Xi} < 0$$

Simulation on the global polarization

- We then input non-zero polarization to parent particles

- $P_y(\phi) = F_0 + F_2 \cos(2\phi)$

- $P_z(\phi) = F_z \sin(2\phi)$



- Conclusion:

Feed-down decays suppress around 10% the primordial polarization, but it does not significant change the original harmonic behaviors.

See also: F. Becattini, G. Cao, E. Speranza, arXiv: 1905.03123

Summary

Spin polarization in heavy ion collisions:

- Global polarization---initial angular momentum
- Local polarization---the inhomogeneous expansion

Some puzzles exist in the present data and theoretical calculation.

Feed down effect on the Λ spin polarization is studied

- Spin transfer in decay **can not flip the sign** of the polarization signal.
- The Ξ weak decay can contribute **a radial polarization**.

Thanks for your attention!

Ensemble average of the spin vector

$$S^\mu(x, p) = -\frac{1}{8m} (1 - n_F) \epsilon^{\mu\nu\rho\sigma} p_\nu \varpi_{\rho\sigma}(x)$$

Becattini, Piccinini, Ann. Phys.323 (2008) 2452

Becattini, Chandra, Zanna, Grossi, Annals Phys. 338, 32 (2013)

Becattini, Inghirami, Eur.Phys.J. C75 (2015) no.9, 406

Fang, Pang, Wang, Wang, PRC 94, 024904(2016)

- thermodynamic equilibrium
- Thermal vorticity: $\varpi_{\mu\nu} = -\frac{1}{2}(\partial_\mu\beta_\nu - \partial_\nu\beta_\mu)$ $\beta_\mu = u_\mu/T$.

$$S^\mu = \frac{\int \frac{d^3p}{E} \int d\Sigma_\lambda p^\lambda f(x, p) S^\mu(x, p)}{\int \frac{d^3p}{E} \int d\Sigma_\lambda p^\lambda f(x, p)}.$$