

# Mesonic Superfluidity in Isospin Matter under Rotation

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Based on: **HZ**, D. Hou and J. Liao, arXiv:1812.11787

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Introduction

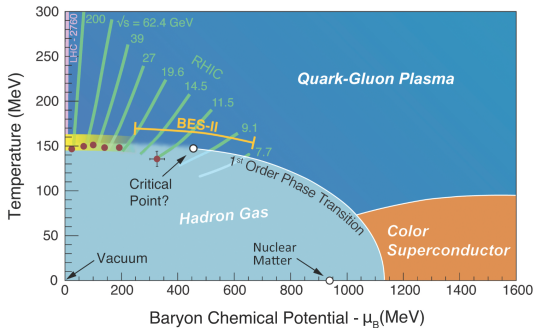
Framework

Numerical results

Summary and Outlook

# Introduction

## Typical QCD Phase Diagram



New dimension  $\Rightarrow$  change the phase diagram & CEP  
 $\mu_I$ , Magnetic field, Rotation, etc.



# Introduction

## QCD with large angular momentum: HIC

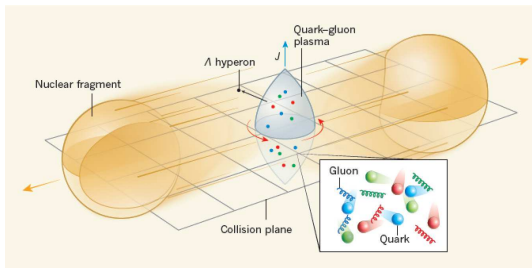


Figure: Off-central HIC

Rotational polarization effect  $\rightarrow$  Anomalous effects: Chiral vortical effect, Chiral vortical wave. . .



# Introduction

## QCD with large angular momentum: Neutron Stars

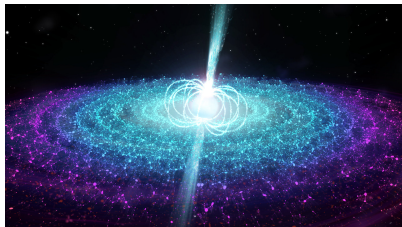


Figure: Spin Neutron Star.

Rotational Suppression of fermion pairing in  $J=0$  (PRL117, no.19(2016)192302) This study (first): isospin matter  
Isospin chemical potential: imbalance between the u-flavor and d-flavor of quarks



# Model

$$\begin{aligned}\mathcal{L} = \bar{\psi}(i\gamma_{\mu}\partial^{\mu} - m_0 + \frac{\mu_I}{2}\gamma_0\tau_3)\psi &+ G_s \left[ (\bar{\psi}\psi)^2 + (\bar{\psi}i\gamma_5\boldsymbol{\tau}\psi)^2 \right] \\ &- G_v (\bar{\psi}\gamma_{\mu}\boldsymbol{\tau}\psi)^2 .\end{aligned}\quad (1)$$

$$\mathcal{L}_R = \psi^{\dagger} \left[ (\vec{\omega} \times \vec{x}) \cdot (-i\vec{\partial}) + \vec{\omega} \cdot \vec{S}_{4 \times 4} \right] \psi \quad (2)$$

MF approximation:

$$\sigma = \langle \bar{\psi}\psi \rangle, \quad \pi = \langle \bar{\psi}i\gamma_5\boldsymbol{\tau}\psi \rangle, \quad \rho = \langle \bar{\psi}i\gamma_0\tau_3\psi \rangle. \quad (3)$$



$$\Omega = G_s(\sigma^2 + \pi^2) - G_v \rho^2 - \frac{N_c N_f}{16\pi^2} \sum_n \int dk_t^2 \int dk_z [J_{n+1}(k_t r)^2 + J_n(k_t r)^2] T \times$$

$$\left[ \ln \left( 1 + \exp \left( -\frac{\omega^+ - (n + \frac{1}{2})\omega}{T} \right) \right) + \ln \left( 1 + \exp \left( \frac{\omega^+ - (n + \frac{1}{2})\omega}{T} \right) \right) \right.$$

$$\left. + \ln \left( 1 + \exp \left( -\frac{\omega^- - (n + \frac{1}{2})\omega}{T} \right) \right) + \ln \left( 1 + \exp \left( \frac{\omega^- - (n + \frac{1}{2})\omega}{T} \right) \right) \right]$$

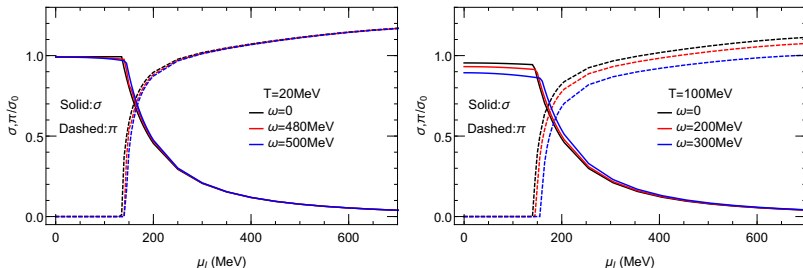
$$\omega^\pm = \sqrt{4G_s^2\pi^2 + (\sqrt{(m_0 - 2G_s\sigma)^2 + k_t^2 + k_z^2} \pm \tilde{\mu}_l)^2}, \quad \tilde{\mu}_l = \frac{\mu_l}{2} + G_v \rho$$

Gap equation:  $\frac{\partial \Omega}{\partial \sigma} = \frac{\partial \Omega}{\partial \pi} = \frac{\partial \Omega}{\partial \rho} = 0$



# Results— $\sigma, \pi$ Channel

## Rotational Suppression on Pion Superfluidity



**Figure:** Suppression effect is consistent with (PRL117, no.19 (2016) 192302). inverse catalysis effect





# Results— $\sigma, \pi$ Channel

## Rotational Suppression on Pion Superfluidity

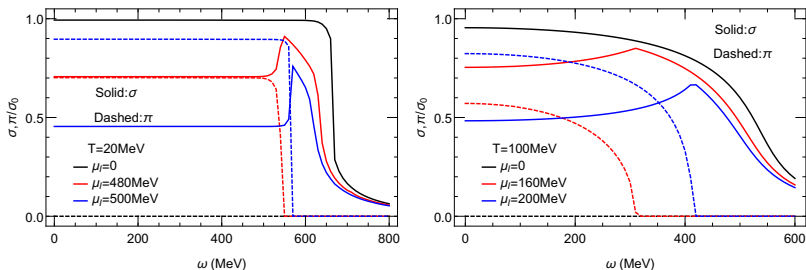


Figure: Prefer  $\sigma$  than  $\pi$

$\sigma : s = 1, L = 1, J = 0;$

$\pi : s = 0, L = 0, J = 0$

Rotation weaken spin 0 condensate: inverse catalysis effect



# Pion superfluidity phase diagram in $T - \mu_I$

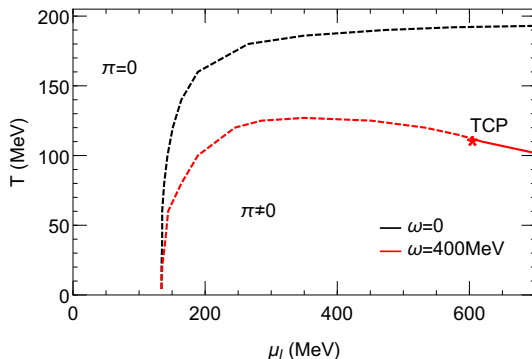


Figure: Dashed line stands for the second-order phase transition, while solid for the first-order. The star denotes a tri-critical point (TCP).



# Enhanced $\rho$ Superfluidity under Rotation— $\sigma, \pi, \rho$ Channel

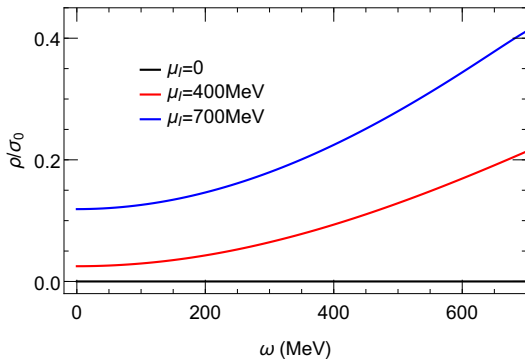


Figure: Rotation weaken spin 0 condensate, Rotation enhance spin 1 condensate



# Enhanced $\rho$ Superfluidity under Rotation— $\sigma, \pi, \rho$ Channel

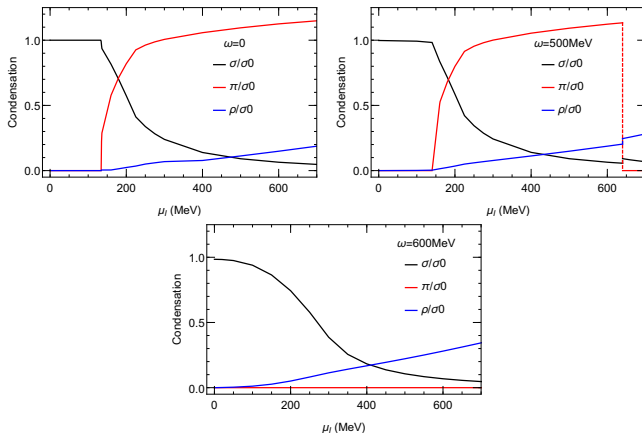


Figure:  $\sigma, \pi, \rho$  dominated phase



# Phase diagram in $\omega - \mu_I$

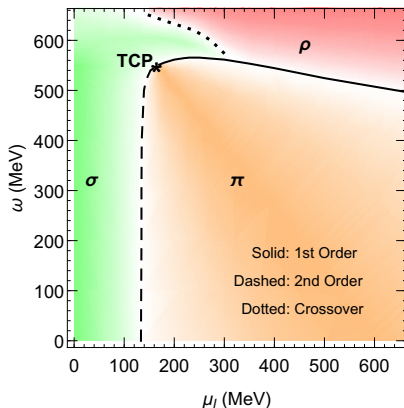


Figure: New phase diagram, New Tri-Critical End Point



# Summary

- ▶  $\pi$  and  $\rho$  meson superfluidity under rotation in NJL model.
- ▶ inverse catalysis effect on the  $\pi$  superfluidity (spin-0 channel).
- ▶ Rotation weaken spin 0 condensate (1606.03808). And enhance nonzero ones (this work).
- ▶  $\rho$  condensate at  $T = \mu = 0$  with none zero isospin chemical potential under rotation.
- ▶ A new type phase diagram in the  $\omega - \mu_I$  plane and a new TCP  $\sim (\mu_I^c = 165, \omega^c = 548)$  MeV.



## Rotational Isospin Matter

- ▶ Boundary effect
- ▶ Introduce Confinement
- ▶ possible splitting of deconfinement and chiral transitions
- ▶ Isospin matter is relevant for understanding the properties of neutron stars.

**Thank you for your attention!**



# Backup

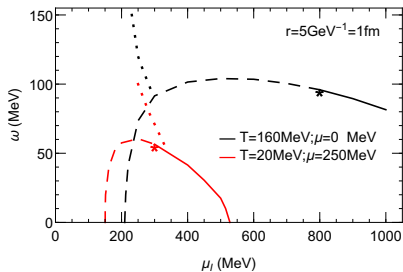


Figure: More realistic case.

