

MODELING THE NUCLEAR DUST EMISSION REVERBERATION RESPONSE IN AGN

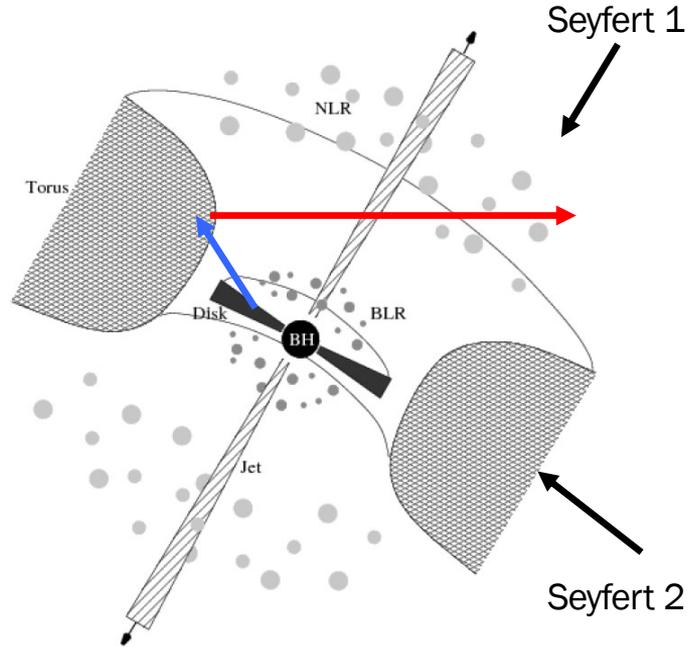
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Collaborators:

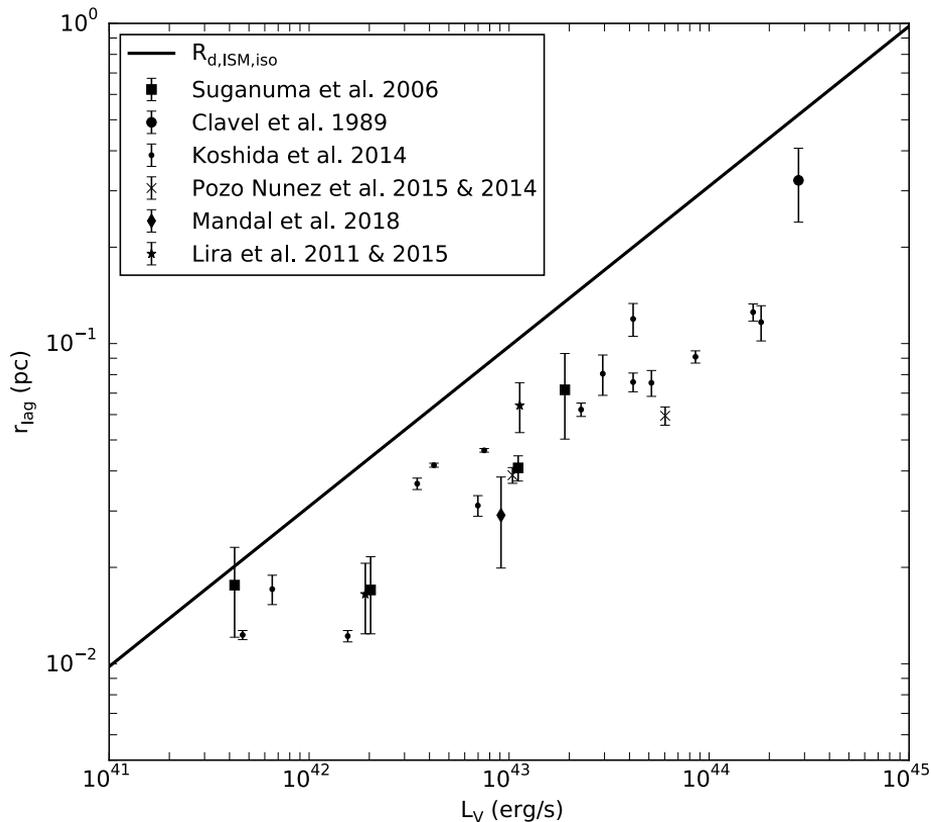
Sebastian Hönig, Andrew Robinson, Robert Nikutta, Nadiya
Ikonnikova, Bryanne McDonough, Michael Richmond, Jack Gallimore

AGN Dust Reverberation Mapping

Dust absorbs UV-visible radiation and re-emits in the IR



Radius-Luminosity Relationship



- Observed radii are ~2 time smaller than theoretical dust sublimation radius (e.g., Kishimoto et al. 2007, Vazquez et al. 2015)
 - *Hot graphite dust component* (e.g., Mor et al. 2009, Mor & Netzer 2012)
 - *Larger grain size* (Kishimoto et al. 2007)
 - *Anisotropic illumination of torus* (Kawaguchi & Mori 2010, 2011)

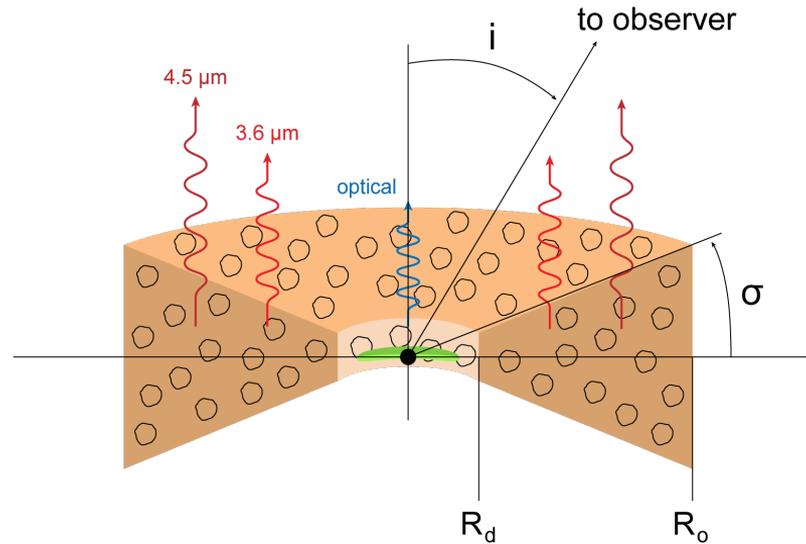
TORMAC: Torus Reverberation Mapping Code

■ Model Features:

- Inner Radius set to Dust Sublimation Radius
- Sharp or “fuzzy” boundary
- Isotropic or anisotropic illumination, s
- ISM dust composition

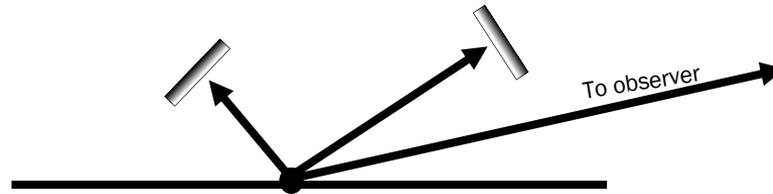
■ Free Parameters:

- Spherical or disk, $\sigma=0-90^\circ$
- Inclination, $i=0-90^\circ$
- Radial distribution of clouds, $\propto r^p$
- Radial depth, $Y=R_o/R_d$
- Optical depth, τ_V
- Volume filling factor, Φ



Radiative Transfer Effects

■ Cloud orientation

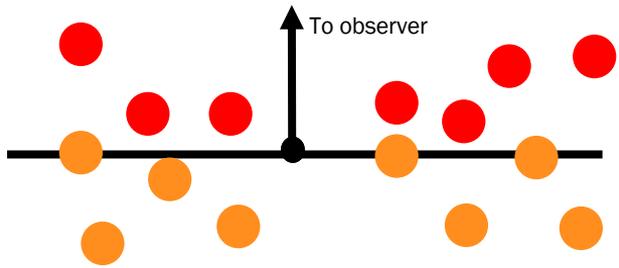


■ Cloud shadowing



● Shadowed clouds
● Non-shadowed clouds

■ Cloud occultation

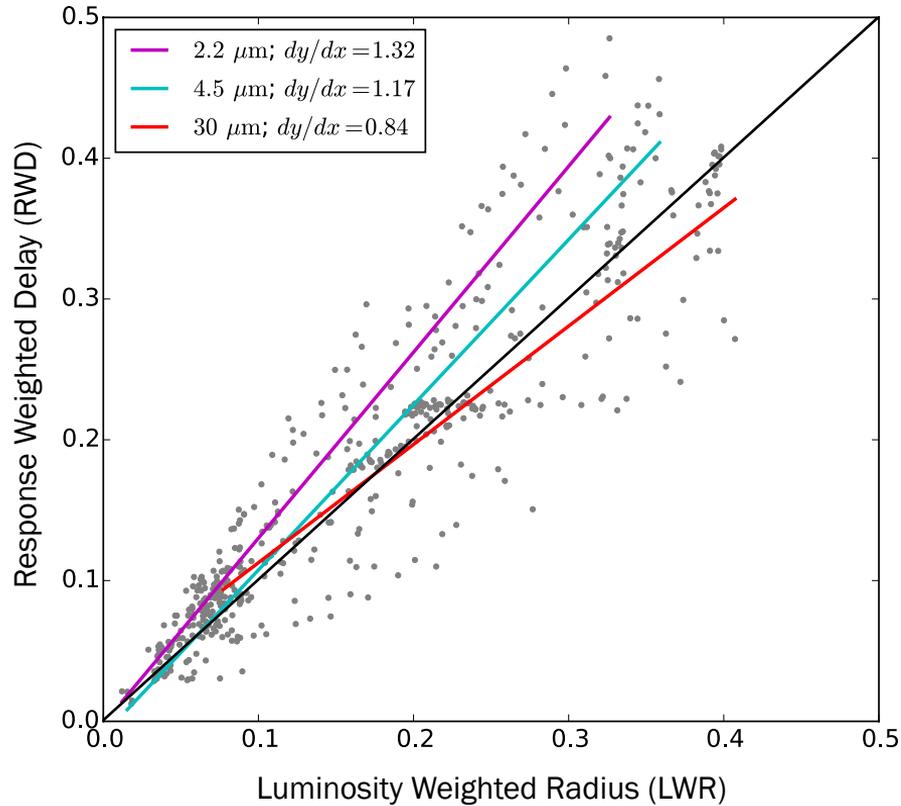


● Occulted clouds
● Non-occulted clouds

Descriptive Parameters

- Response Weighted Delay (RWD)
 - *Characteristic lag of transfer function*
- Torus Luminosity Weighted Radius (LWR)
 - *Effective radius of the torus*
- When torus is composed of isotropically emitting clouds, $RWD=LWR$
 - True for both sphere and disk

RWD vs LWR

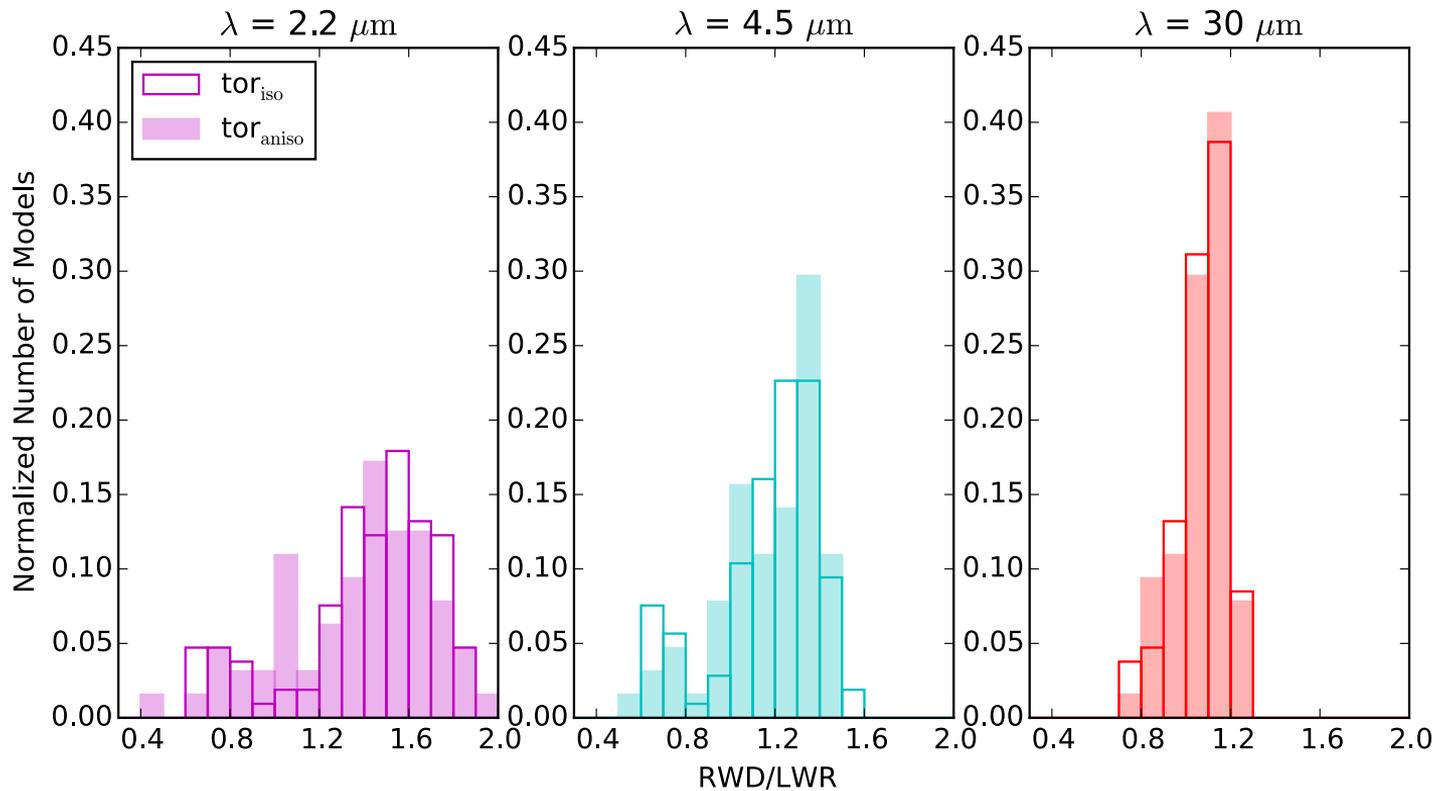


Model Parameters:

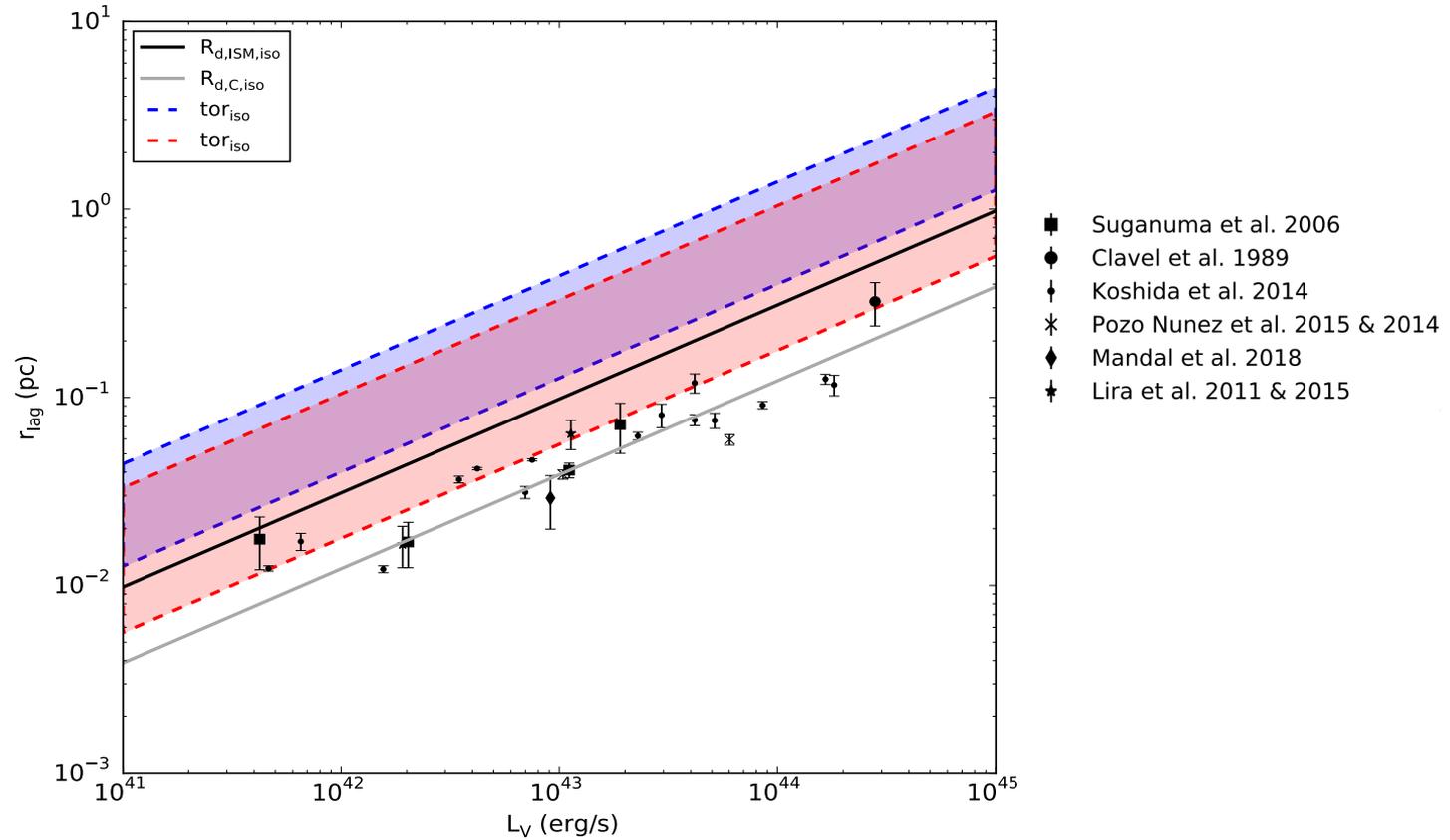
- $\propto r^p$; $p=-2-4$
- $Y=R_o/R_d=2-50$
- $i=0-90^\circ$
- $\sigma=15-60^\circ$
- $s=0.01-1$
- $\tau_V=5-100$
- $\Phi=0.0001-0.1$
- Sharp and fuzzy

$$0.4LWR < RWD < 1.8LWR$$

RWD vs LWR



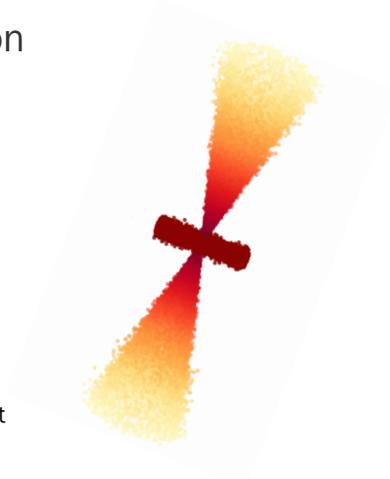
Radius-Luminosity Relationship



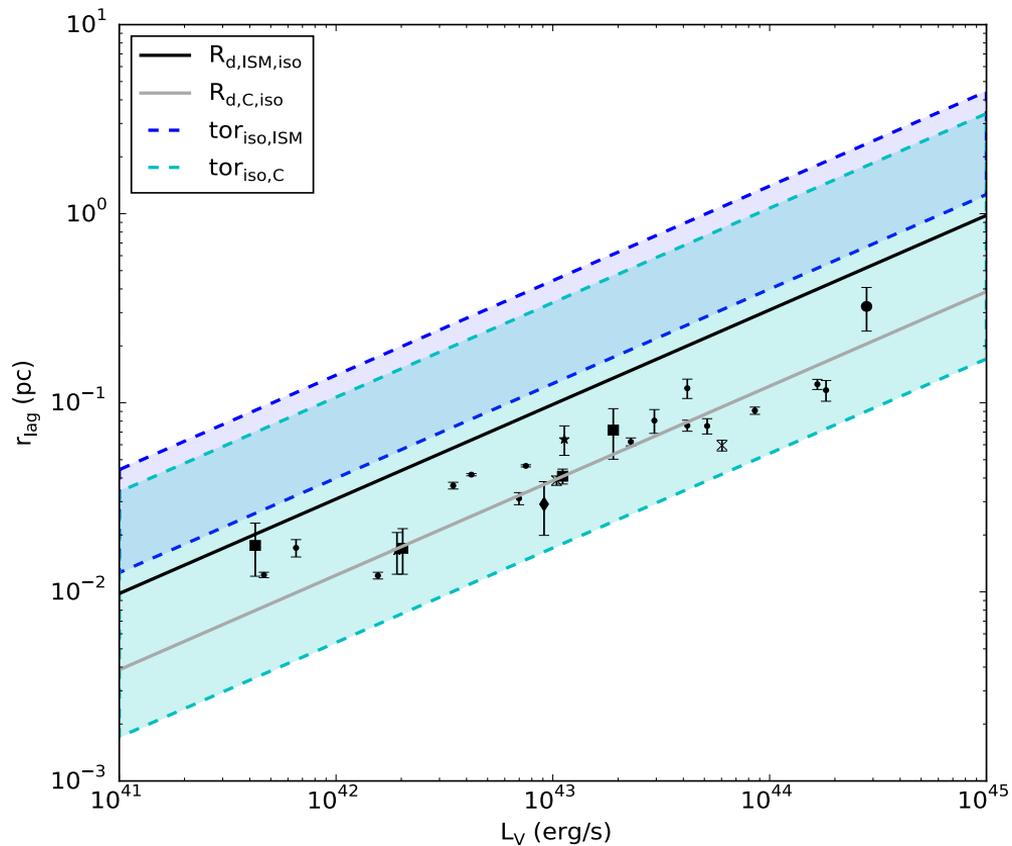
Almeida et al, submitted

Current Capabilities

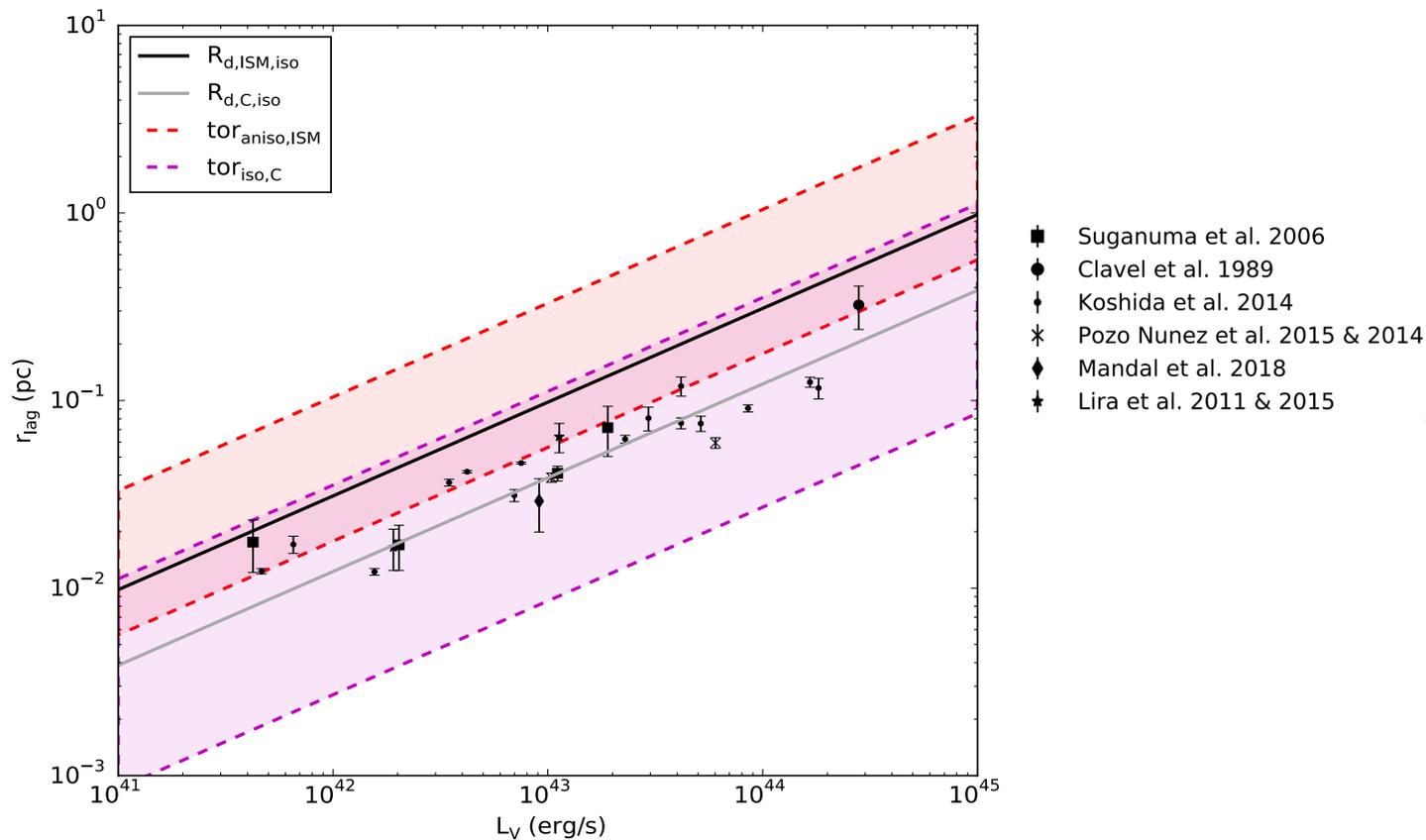
- TORMAC compatible with any radiative transfer grid/database
- “hot dust” component
 - *Gradient in composition of clouds from Carbon-dominated to full ISM composition*
 - *<http://cat3d.sungrazer.org/>
- Polar dust distribution



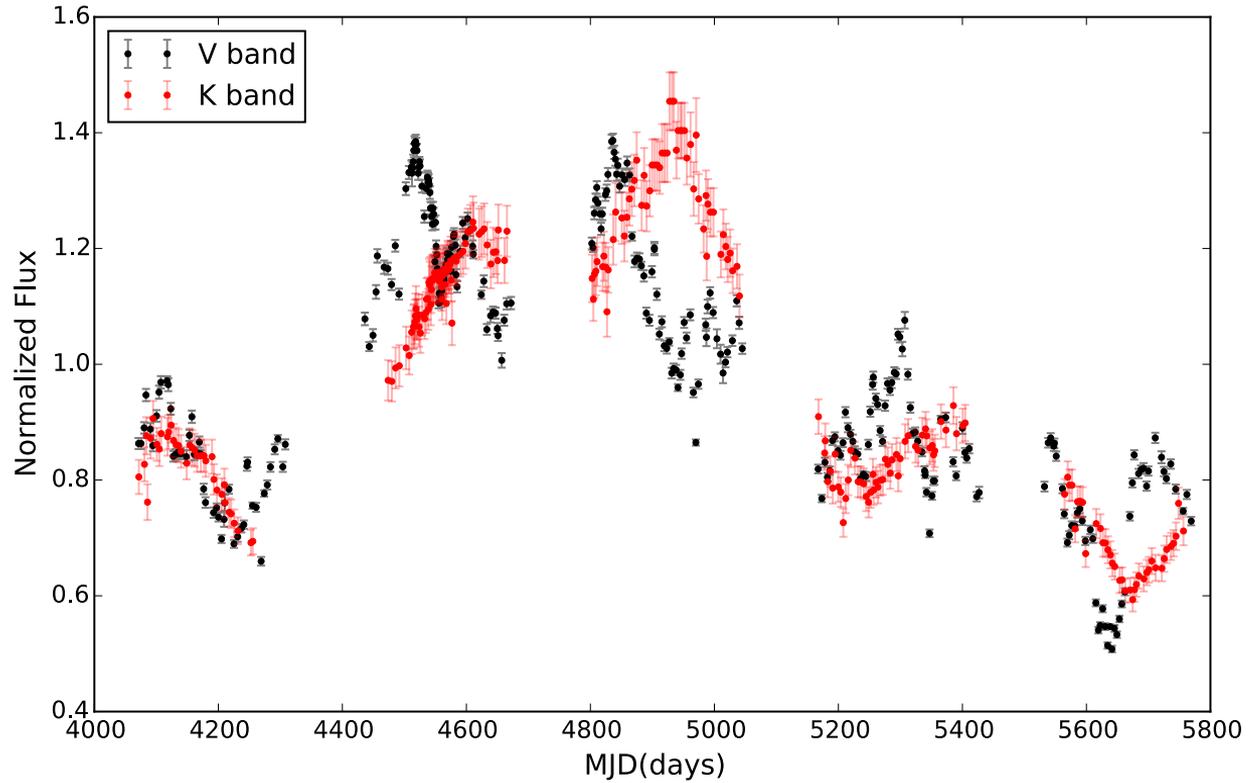
R-L Relationship with Carbon grains



R-L Relationship with Carbon grains

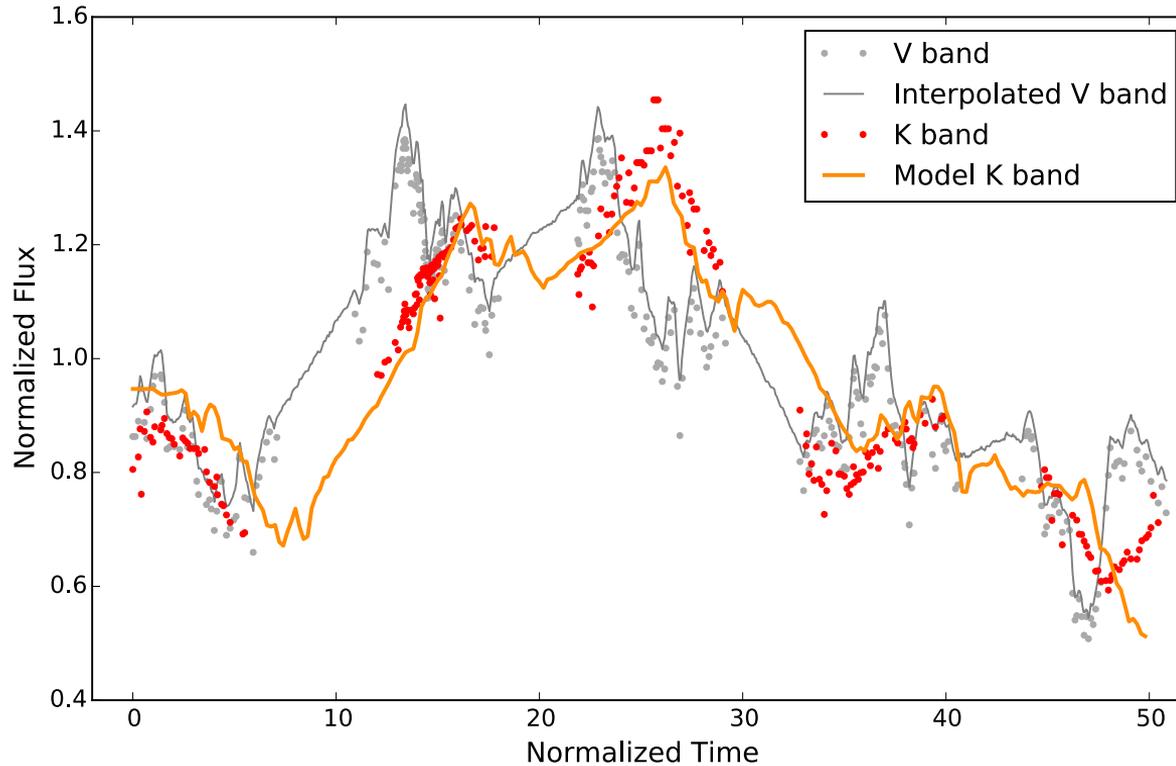


NGC 3783



Light curves courtesy of Paulina Lira

NGC 3783: Models



Summary

- Multiwavelength dust emission response for cloud ensemble using radiative transfer models
- Simulate IR response LC given any driving LC
- Incorporates anisotropic illumination, global opacity effects
- Need hot graphite dust component to match observations
- $0.4\text{LWR} < \text{RWD} < 2\text{LWR}$