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# Heavy Flavor Evolution with EM fields and Hot Medium

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# Outline

### **1. Vector meson photoproduction from EM fields**

2. Stochastic Schrodinger equation : (hot medium effect) stochastic potential (open quantum system) charm wave function evolution with SSE, Diffusion coefficient Ds

### 3. Schrodinger equation for charmonium:

transitions between 1S and 2S states

### Different $P_T$ physics



# Photoproduction from EM fields

## **Equivalent Photon Approximation**



charges moves at nearly speed of light → produce E-B fields

#### Strong Lorentz-contracted Electromagnetic field



## $p_T$ dependence

Compare the p<sub>T</sub> dependence of coherent photoproduction and hadroproduction



#### > 2.76 TeV forward rapidity 2.5<y<4,



## $p_T$ dependence

Coherent photoproduction:Photons interact with entire nucleus, $p_T \sim 1/R_A \sim 0.03 \text{ GeV/c}$ Exp.  $< p_T >= 0.055 \text{ GeV/c}$ 

PRL 116, 222301 (2016)

b=10.2 fm	Hadroproduction $2.5 < y < 4$	photoproduction			
$0 < p_T < 0.04$ GeV/c	0.47×10 <sup>-5</sup>	$5.54 \times 10^{-5}$			
$0 < p_T < 0.1$	2.4×10 <sup>-5</sup>	15.7×10 <sup>-5</sup>			
$0 < p_T < 0.5$	50×10 <sup>-5</sup>	~16×10 <sup>-5</sup>			
$0 < p_T < 1$	<b>179×10</b> <sup>-5</sup>		4		
$0 < p_T < 3$	772×10 <sup>-5</sup>				
		m <sup>-2</sup> )	2		
rom impact parameter b~10 fm to more central					
ollisions, hadroproduction increase significantly.					
$N_{AA}^{J/\psi} = \sigma_{pp}^{J/\psi} \int d^2 x_T T_A(x_T) T_B(x_T - b)$					
At $h \sim R_{\star}$ , they are at the same order in nT<0.1 GeV/c					



### hadro- and photo- production



#### Heavy quarks (quarkonium) + light partons (QGP) Produced in the *overlap area*.



 $\mathbf{b} < 2R_A \text{ or } \mathbf{b} \geq 2R_A$ 

Produced in the entire nucleus surface

 $\gamma A \rightarrow I/\psi A$ 



## $J/\psi$ from EM field

 $N_{\psi}^{\gamma A} \propto \int dw rac{dN_{\gamma}}{dw} \sigma_{\gamma A o J/\psi A} \Gamma_{QGP}^{decay}$ Mainly ingredients: From transport model

• Photon density  $\frac{dN_{\gamma}}{dw}$  emitted by one nucleus 

$$\frac{dN_{\gamma}}{dw} = n(w) = \frac{1}{\pi w} \int d\vec{x}_T |\vec{E}_T(\vec{r}, w)|^2$$
Photon density
$$= \frac{(Ze)^2}{\pi w} \int_0^\infty \frac{d^2 \vec{k}_T}{(2\pi)^2} [\frac{F((\frac{w}{v\gamma})^2 + k_T^2)}{(\frac{w}{v\gamma})^2 + k_T^2}]^2 \frac{k_T^2}{v^2}$$

#### Photon-nucleus cross section $\sigma_{\gamma A \rightarrow I/\psi A}$ Widely studied in UPC

$$\sigma(\gamma A \to J/\psi A) = \frac{d\sigma(\gamma A \to J/\psi A)}{dt}|_{t=0} \int_{-t_{min}}^{\infty} |F(t)|_{t=0} \int_{-t_{min}}^{\infty}$$

**Nuclear charge form** factor is the Fourier transform of Woods-Saxon distribution. For point particle, it's 1

S.R.Klein, J. Nystrand, PRC, 1999  $|^{2}dt$  Physics Roports, G.Baur, et al, 2002

### **Total** $J/\psi$ from EM field + QGP



# **Total** $J/\psi$ from EM field + QGP



Also significant enhancement at N<sub>p</sub>  $\approx 100$ , where  $T_0^{QGP} = 2T_c$ , similar with T at RHIC 200 GeV Au-Au (most central)

When N<sub>part</sub> → 0 (b > 2R<sub>A</sub>),
hadroproduction → 0, photoproduction → nonzero,  $R_{AA}$ → infinity
Baoyi Chen HENPIC 2018.01.18

## Photoproduced 2S/1S



# SSE and Stochastic potential

### Thermal medium effect

Quantum approach (open quantum system):

- -Lindblad equation[Lindblad et al, 1976, ...]-Solid theoretical foundation for open quantum system,<br/>-Evolving density matrices, computationally-intensive,<br/>approx. needed-Schrödinger-Langevin equation<br/>-Originates from Heisenberg-Langevin equation<br/>-Nonlinear in  $|\phi\rangle$  Generalized Langevin approachJ. P. Blaizot, et al
  - pNRQCD

J. P. Blaizot, et al X. Yao, B. Muller

- Stochastic Schrodinger equation (our approach)

Can be employed in Spin thermal theory, Bose-Einstein conden. Thermalization process, et al.

### Framework of SSE



### Heavy quark-Medium interaction

Like QED

• We model the medium with gluon field:

$$H_I = \int d^3 x g \bar{\psi} \gamma^{\mu} \psi A_{\mu}$$

• Take non-relativistic approximation and keep only zeroth component:

$$H_I = \int d^3 x g \bar{\psi} \gamma^0 \psi A_0$$

•  $A_0$  in momentum space

$$A_0(\vec{x}) = \int \frac{d^3 p_g}{(2\pi)^3} \sqrt{\frac{2}{E_{p_g}}} a(\vec{p}) e^{-i\vec{p}\cdot\vec{x}}$$

• Based on Boltzmann distribution:  $|a(p_g)|^2 = \rho(p_g) = d_g V e^{-\beta E_g}$ , we conjecture

$$a(\vec{p}) = \sqrt{\frac{d_g V}{e^{-\beta E_g}}} e^{-i\theta_{p_x}^r} e^{-i\theta_{p_y}^r} e^{-i\theta_{p_z}^r}$$

where  $\theta_{pp'} \equiv \theta_p - \theta_p$ , is a random phase with  $\theta_{pp'} = -\theta_{p'p}$  (Hermiticity) as  $a(-\vec{p}) = a^*(\vec{p})$ ,  $\theta_{-p}^r = -\theta_p^r$ 

Numerical results show that  $A_0(x)^2 s'$  mean value are intensive, not extensive

SSE in momentum space and interaction picture

$$i\frac{d}{dt}\psi(p,t) = \sum_{p'} \sqrt{\frac{d_g V}{e^{\beta E_g}}} e^{i(E_p - E_{p'})t/2} e^{i\theta_{pp'}} e^{\beta(E_{p'} - E_p)/2} \psi(p',t)$$
Square root of
distribution of
gluon field
$$\begin{cases}
\text{"damping" factor: from ensemble average;} \\
\text{weighting transition matrix elements with} \\
\text{number of microscopic states of the medium}
\end{cases}$$

 $\theta_{pp}$ , is time-dependent, updating period is a parameter, currently is taken at 1/T.

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$$E_{p'} - E_p$$
 lead to non-Hermitian Hamiltonian  $E_g \sim |E_p - E_{p'}|$ 

### Wave function in momentum space

 $\tau_c = 0.03 \text{ fm}/c, p_c = 0.3 \text{ GeV}, L = 64 \text{fm}/c$ 



### Wave function in spatial space





### compare

#### **Our SSE**

## (1) Equation (interaction picture) $i \frac{d}{dt} |\psi\rangle = V_I^{stochastic} |\psi\rangle$ (2) ingredients

**External field**  $A_0$  for QGP (with random phase  $e^{i\theta}$ )

damping factor  $e^{-E_p/T}$ 

(from the information of environment)

**Detailed balance** 

#### (3) properties of random phase term

Introduce  $\tau_c$ ,  $p_c$  for its correlation in momentum and time.

Random phase term satisfies a correlation function (not delta-function)

#### Other SSE

J.Phys.Condens.Matter, 24(2012) 273201



### Properties of noise & Ds

 $\tau_c = 0.03 \text{ fm}/c, p_c = 0.3 \text{ GeV}, L = 64 \text{fm}/c$ T= 0.1 GeV

**Based on classical Langevin equation:** 

 $\frac{dp}{dt} = -\gamma p + f(t) \qquad D_S = \frac{kT}{\gamma m}$ 



Preliminary:  $2 < D_S(2\pi T) < 3$ 

# Charmonium (two-body)



### Transitions between charmonia

V.S.





 $c\bar{c}$  evolutions in Bjorken hydro, With only **transition mechanism**.



**Transport model** With dissociations, No transitions.

Chen, Zhuang, PLB 765 (2017) 323-327



### Transitions between charmonia

#### Baoyi, Xiaojian, Carsten, Ralf, in preparation



Its suppression is mainly due to the internal evolution of wavefunction, Particle inelastic dissociation is now absent.

> P-Pb system: More suitable to study the internal evolutions of ccbar system. Such as 1S-2S transition mechanism

### If without Transitions



- Photoproduction from EM fields charmonium photo-production from strong electromagnetic fields and hadro-production
- Stochastic-Schrodinger-Equation (SSE)
   We construct the Stochastic Schrodinger Equation
   (SSE) and stochastic potential to study the wave function evolutions of heavy quark.

SSE can provide a way to include both **particle collision process** and **color screening** in Schrodinger equation for charm and  $\psi$ .

Charmonium transition within Schrodinger
 Charmonium (1S, 2S) transitions based on Schrodinger equation are also studied in small systems (p-Pb)

## More slides

## **Photoproduction contribution**



and peripheral and semi-central collisions

TABLE I: Information of QGP based on (2+1)D ideal hydrodynamics

• At Np=100, 
$$T_0^{QGP} = 2T_c$$

QGP effect important ! Photoproduction important !

Hydro in LHC $\sqrt{s_{NN}} = 2.76$ TeV Pb-Pb, $2.5 < y < 4$						
b(fm)	$N_p$	$T_0^{ m QGP}/T_c$	$\tau_{\rm f}^{\rm QGP} ~({\rm fm/c})$			
0	406	2.6	7.3			
9	124	2.1	4.2			
9.6	103	2.06	3.9			
10.2	83	1.95	3.5			
10.8	64	1.84	3.1			

## $J/\psi$ from electromagnetic field

#### • Our formula for $J/\psi$ photo-production with QGP effect





## $A_0(\vec{x})$ field is intensive



Numerical results show that  $A_0(x)^2 s'$  mean value are intensive, not extensive

### Compare with master equation



Thermalization process from SSE and master equation of density matrix are diefferent.

# **Classical Theoretical Models**

#### **Boltzmann transport models**





#### Primordial production

Transport two-component model Tsinghua Group:

> Chen, Zhuang, Phys.Lett. B726 (2013) 725-728 Chen, Zhuang, Phys.Lett. B765 (2017) 323-327

#### TAMU group:

Xingbo, Ralf, Nucl.Phys. A859 (2011) 114-125

#### Recombination of $\mathbf{c}\overline{\mathbf{c}}$ during QGP expansion $\mathbf{c} + \overline{\mathbf{c}} \leftrightarrow \mathbf{J}/\boldsymbol{\psi} + \mathbf{g}$ dominates in AA

Che-ming Ko, Ralf Rapp, R. L. Thews, P. Braun-Munzinger Jiaxing Zhao, Baoyi Chen

 $N_{J/\psi} \propto N_{N_{c\bar{c}}}^2$ 

Large uncertainties of (N<sub>cc̄</sub>)
 theoretical calculations and experimental
 data