# Decay studies of nuclei far from β stability line



第十七届全国核物理大会,2019年10月8-12日,华中师范大学

# 报告提纲

一、研究背景

二、研究内容

三、研究展望





<sup>11</sup>Li (4%), <sup>17</sup>B (11%), <sup>17</sup>C (7%), <sup>30,31</sup>Na (1%), <sup>32,33</sup>Na (10%)

•  $\beta$  -delayed 3,4-neutron emission cases: <sup>11</sup>Li (3n), <sup>17</sup>B (3n, 4n)

## **Mechanism of two-proton radioactivity**





## > 2p, 4p decay of <sup>18</sup>Mg



### **MICHIGAN STATE** UNIVERSITY

May 3, 2018

#### Dear Hui Hua:

The Program Advisory Committee (PAC 42) of the National Superconducting Cyclotron Laboratory met on May 2-3, 2018 to consider proposals for beam time at the Coupled Cyclotron Facility and ReA3. Your proposal, "Spectroscopy of 18Mg by 14O + 4p correlations" (No. 18015), was included in this review.

In spite of the large amount of time requested (7060 hours) compared to the limited running time available, the Committee recommended your full time request of 205 hours (178 hours on-target time and the rest beam development and delivery time) for your experiment. I concur with the Committee's recommendation.

For PAC 42 experiments, the approval duration is 24 months from the beginning of the running period on October 1, 2018.



Please notify your colleagues on the proposal of this decision. As spokesperson, you are responsible for communicating information regarding your experiment to your collaborators. For information on the scheduling process, you may find it useful to browse

the user section of the NSCL website, especially the User Guide Bradley Sherrill (http://nscl.msu.edu/users/guide.html). All collaborators should be made aware of the

NSCL Director Experimenter Responsibilities

(http://nscl.msu.edu/users/ExperimenterResponsibilities1.pdf). National

#### Superconducting

We wish you the best of luck on your experiment. Cvclotron

Laboratory Sincerely,

East Lansing, MI Buly M therill

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Bradley M. Sherrill NSCL Director

48824





- <u>反应</u>:  ${}^{20}Mg$  敲出两个中子得到目标核 ${}^{18}Mg \rightarrow {}^{18}Mg$  衰变为 ${}^{14}O+4p$
- <u>探测</u>: [轻粒子] DSSD+CsI进行粒子鉴别,得到能量和动量方向

[剩余核] 闪烁光纤测量出射角度 → S800磁谱仪进行粒子鉴别, 给出粒子能量







正面







## **Proton and Residue**



## **Some collaborators**

TIL

## Mg19 invariant mass

 $^{20}Mg \rightarrow {}^{19}Mg + 1n$ 



State	mean	sigma	Excitation energy	
g.s	0.727	0.17	0	
1Ex	1.74	0.35	1.013	2540
2Ex	3.11	0.68	2.383	
3Ex	4.74	1.03	4.013	



Mukha et al. PRC 85, 044325 (2012)

<sup>19</sup>N level scheme (from NNDC)

-0.0 271 MS

## Ne16 invariant mass



Previous experimental value: Q2p (g.s.) = 1.466 MeV Q2p (2+) = 3.156 MeV



FIG. 1 (color online). Experimental spectrum of <sup>16</sup>Ne decay energy  $E_T$  reconstructed from detected <sup>14</sup>O + p + p events. The dashed histogram indicates the contamination from <sup>15</sup>O + p + pevents. The smooth curves are predictions (without detector resolution) for the indicated <sup>16</sup>Ne states. The inset compares the contamination-subtracted data to the simulation of the g.s. peak for  $\Gamma = 0$ ,  $f_{tar} = 0.95$ , where the dotted line is the fitted background. Brown et al. PRL 113, 232501 (2014)

## > 1n&2n decay of <sup>34</sup>Al



## <sup>35</sup>Al $\beta$ -delayed neutron



C. Timis, et. al., J. Phys. G<sup>\*</sup>. Nucl.Part.Phys., 31, S1965 (2005).

# <sup>34</sup>Al $\beta$ -delayed neutron



S. Nummela et al., Phys. Rev. C. 044316(2001)

## <sup>34</sup>Al $\beta$ -delayed neutron



## S. Nummela et al., Phys. Rev. C. 044316(2001)

## • Some collaborators



## ➤ Shell evolution near N = 82



## <sup>132</sup>Sn附近壳演化研究现状



A. Jungclaus *et al.*, PRL 99, 132501(2007). H. Watanabe *et al.*, PRL 111, 152501(2013)







## 实验中得到的粒子鉴别图

Z vs. A/Q identification plot





## Red: 新能级和跃迁 Black: 前人工作

 $E_{\gamma}$  [keV]



### PHYSICAL REVIEW LETTERS 122, 212502 (2019)

### Proton Shell Evolution below <sup>132</sup>Sn: First Measurement of Low-Lying $\beta$ -Emitting Isomers in <sup>123,125</sup>Ag

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•β -delayed 2-neutron emission cases: <sup>11</sup>Li (4%), <sup>17</sup>B (11%), <sup>17</sup>C (7%), <sup>30,31</sup>Na (1%), <sup>32,33</sup>Na (10%)

## **Toward more neutron-rich Ag isotope**

	<sup>123</sup> Sn	<sup>124</sup> Sn	<sup>125</sup> Sn	<sup>126</sup> Sn	<sup>127</sup> Sn	<sup>128</sup> Sn	<sup>129</sup> Sn	<sup>130</sup> Sn	<sup>131</sup> Sn	<sup>132</sup> Sn	<sup>133</sup> Sn	134Sn	<sup>135</sup> Sn
	<sup>122</sup> ln	<sup>123</sup> ln	<sup>124</sup> ln	<sup>125</sup> ln	<sup>126</sup> ln	<sup>127</sup> ln	<sup>128</sup> ln	<sup>129</sup> ln	<sup>130</sup> ln	<sup>131</sup> ln	<sup>132</sup> ln	<sup>1°3</sup> In	<sup>134</sup> ln
	<sup>121</sup> Cd	<sup>122</sup> Cd	<sup>123</sup> Cd	<sup>124</sup> Cd	<sup>125</sup> Cd	<sup>126</sup> Cd	<sup>127</sup> Cd	<sup>128</sup> Cd	<sup>129</sup> Cd	<sup>30</sup> Cd	<sup>131</sup> Cd	<sup>132</sup> Cd	<sup>133</sup> Cd
89/2	<sup>120</sup> Ag	<sup>121</sup> Ag	<sup>122</sup> Ag	<sup>123</sup> Ag	<sup>124</sup> Ag	<sup>125</sup> Ag	<sup>126</sup> Ag	<sup>127</sup> Ag	<sup>128</sup> Ag	29Ag	<sup>130</sup> Ag	<sup>131</sup> Ag	<sup>132</sup> Ag
	<sup>119</sup> Pd	<sup>120</sup> Pd	<sup>121</sup> Pd	<sup>122</sup> Pd	<sup>123</sup> Pd	<sup>124</sup> Pd	<sup>125</sup> Pd	<sup>126</sup> Pd	<sup>127</sup> Pd	2 <sup>8</sup> Pc	<sup>129</sup> Pd		
	<sup>118</sup> Rh	<sup>119</sup> Rh	<sup>120</sup> Rh	<sup>121</sup> Rh	<sup>122</sup> Rh	<sup>123</sup> Rh	<sup>124</sup> Rh	<sup>125</sup> Rh	<sup>126</sup> Rh	<sup>127</sup> Rh		-	
	<sup>117</sup> Ru	<sup>118</sup> Ru	<sup>119</sup> Ru	<sup>120</sup> Ru	<sup>121</sup> Ru	<sup>122</sup> Ru	<sup>123</sup> Ru	<sup>124</sup> Ru		7	-		

 $--- h_{11/2} ---- \Rightarrow$ 



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