Recent advances and perspectives of cluster physics

M. Kimura (Hokkaido Univ.)

- 1. A new theoretical approach to nuclear structure & reaction problems
- 2. A new approach to stellar fusion reactions + α
- 3. Nuclear data evaluation by Machine Learning

Hoyle stateの記述

Introduction Cluster structure and α gas-like states



Energy [MeV]

Recent studies for gas-like stats.

⁸Be(2α), ¹²C(3α), ¹⁶O(4α) are studied in detail

A. Tohsaki et.al PRL, **87**.192501 (2001) Y. Funaki et.al PRC, **82**.024312 (2010) Y. Kanada-En'yo PRC, **89**.024304 (2014) $\geq {}^{20}$ Ne(5 α), 24 Mg(6 α) and heavier systems are also expected to have gas-like state.

 \bigcirc How many α particles can form dilute-gas like state?

Introduction Candidates of 9α and 14α gas-like state

From ⁴He(³⁶Ar, $N\alpha$) ⁴He(⁵⁶Ni, $N\alpha$)experiment, signature of the 9α and 14α gas-like state.



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T. Yamada and P. Schuck, PRC 69, 024309 (2004)

 \odot Gas-like state possibility exist up to $\sim 10 \alpha$ sysytem .



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The prediction should be verified by the microscopic cluster model without Boson approximation.

 \Rightarrow We need microscopic model that can describe gas-like states of many α particles

Introduction microscopic description of gas-like state

Gas-like state is described by the superposition of "many" Slater determinants



 \succ Number of Slater determinants becomes huge as the number of α increases

 \Rightarrow We need effective way to generate basis wave functions.

Various methods to genarate Nα wave functions
➢ Random generation T. Ichikawa et.al PRC, 83.061301 (2011)
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Framework Time evolution of wave packets



Framework New generator coordinate method

The time evolution generates all important configurations



Superposition of the wave functions obtained by the time evolution should give a good description for the gas-like state

In other words, we perform GCM calculation using the real time t as the generator coordinate

Step1: Time evolution to generate basis wave function

$$\hbar \sum_{j,\sigma} C_{i\rho j\sigma} \frac{\partial \vec{Z}_{j\sigma}}{\partial t} = \frac{\partial H}{\partial \vec{Z}_{i\rho}^*}$$



Step2: Superpose basis wave function and solve GCM equation

$$\psi(\vec{r}_1, \vec{r}_2, \cdots, \vec{r}_{4N}) = \int_0^\infty dt \underline{F_{MK}^{J\pi}(t)} P_{MK}^{J\pi} \phi(\vec{Z}_1(t), \vec{Z}_2(t), \cdots, \vec{Z}_N(t))$$

Determined by diagonalization of Hamiltonian

 $P_{MK}^{J\pi}$; Projection operator of parity and angler momentum

Results & Discussions for 3α and 4α systems

Result for 3α system (¹²C)

M.Kamimura NPA.**351,**456 (1981)

Exp.

Y.Funaki PRC.92 021302(2015)

 $E_x - E_{3\alpha}$ [MeV]

This work

8					<u>л</u> +			
	This work		THSR		RGM			
J^{π}	E_x [MeV]	$\sqrt{\langle r^2 \rangle}$ [fm]	E_x [MeV]	$\sqrt{\langle r^2 \rangle} [{ m fm}]$	E_x [MeV]	$\sqrt{\langle r^2 \rangle} [\mathrm{fm}]$	•	
0_{1}^{+}	-7.63	2.39	-7.48	2.4	-7.56	2.40		
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							_	

THSR

RGM

Discussion Check of the convergence of calc.

Our method is based on the ergodic nature

- \Rightarrow Following two pointes should be checked
- 1. When the time evolution is long enough, the result should be converged.
- 2. The result should be independent of the initial configuration of α particles at t = 0.

Result for 4α system (¹⁶O)



Summary

- For ⁸ Be(2 α), ¹²C(3 α), ¹⁶O(4 α), gas state is well known but those of ²⁰Ne(5 α), ²⁴Mg(6 α) \cdots are not known.
- > To solve the difficulty for generating wave function increasing number of α particles, we introduced real time evolution setting time *t* as a generator coordinate.
- For ¹²C(3α), ¹⁶O(4α) nuclei, this method yielded the results consistent with or better than the previous researches.

Future work

- Separate from continuum state.
- Introduce three-body interaction.

Benchmark calculations for ${}^{12}C$ (3 α cluster system)





○ All results are consistent with or better than THSR wave function

 \bigcirc Computational cost is greatly reduced



A new theoretical approach to nuclear structure and reactions

Coherent stateの時間発展に関 する問題を追加しておく

Motivation

O Most of nuclear structure models describes a nuclear state by a superposition of model wave functions



Newly developed our method "Real-time evolution method" is a very natural and effective way to choose Φ_n and N_{max} .

It realizes very accurate description of nuclei

Model wave function (time-dependent wave packets)

 \bigcirc Slater determinant of nucleon wave packets

 $\Phi(t) = \mathcal{A}\left\{\phi(\mathbf{Z}_1(t)), ..., \phi(\mathbf{Z}_A(t))\right\}$

 $\phi(\mathbf{Z}_i(t)) = \exp\left\{-\nu(\mathbf{r} - \mathbf{Z}_i(t))^2\right\} (\alpha_i(t) \downarrow \downarrow)$

 \bigcirc Dynamical variables of the model (* *e*-dependent p meters) $\mathbf{Z}_i(t)$: Centroids of wave packets (pc for and momentum) $\alpha_i(t) \ \beta_i(t)$: Spin directions

O Hamiltonian
$$H = \sum_{i=1}^{A} t(i) - t_{cm} + \sum_{i < j} q_{ij}$$

 \bigcirc Microscopic Hamiltonian with effective/bare NN interactions

Time-dependent variational principle

$$\delta \int dt \, \frac{\langle \Phi(t) | i\hbar d/dt - H | \Phi(t) \rangle}{\langle \Phi(t) | \Phi(t) \rangle} = 0$$

Equation of Motion for nucleon wave packets

⁶He (6 nucleons)



© Each wave function is not accurate (Gaussian approx.)

but the ensemble of the wave functions has nice properties

J. Schnack and H. Feldmeier, NPA601, 181 (1996). A. Ono and H. Horiuchi, PRC53, 845 (1996), PRC53, 2341 (1996).



O Actually, this ensemble properties (ergodicity & quantum statistcs) have been utilized to discuss Liquid-Gas phase transition of nuclear matter.



♥ We superpose time dependent wave function and diagonalize the Hamiltonian

$$\Psi^{J\pi} = f_1 + f_2 + f_3 + f_4 + f$$

Benchmark calculations for few-body systems



Nuclear response functions

O It is also possible to study the nuclear responses

The ensemble of wave function contains the information of the excited states



Application to the Hoyle state

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Collaborators

 \odot It is noted that this work was initially motivated by Prof. Yabana

O Structure and Response of light nuclei
 Y. Taniguchi(Kagawa), W. Horiuchi(Hokkaido)

O Extension to DFTJ. A. Maruhn(Frankfurt)

○ Exotic clusters in neutron-rich nuclei
 Zhou Bo (Hokkaido), T. Baba(Kitami), H. Masui(Kitami)

Clusters & Astrophysics
 K. Ogata(RCNP), Y. Kanada-En'yo(Kyoto), Y. Taniguchi(Kagawa)
 Y. Chiba(Osaka-city), K. Yoshida(JAEA)